The physics case of a Low Energy Neutrino Factory

Pilar Coloma Center for Neutrino Physics

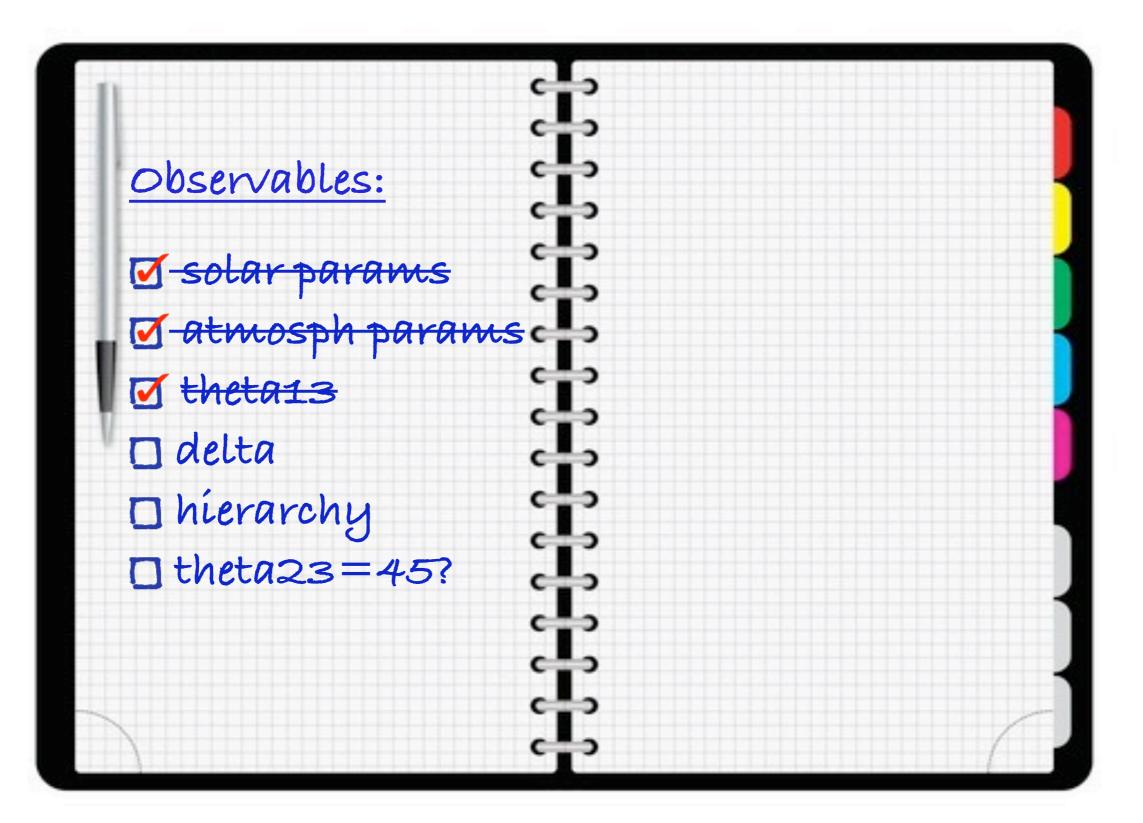
UrginiaTech

nuTURN workshop May 10th 2012 Laboratori Nazionali del Gran Sasso, Italy

Outline

- General landscape for the future
- Why a NF?
- The LENF concept
- Physics reach of a LENF from RAL to GS
- Precision measurements

The shopping list in V oscillations



General Landscape

- Super-Beams
 - Japan: T2HK
 - USA: NOvA, LBNE
 - Europe: LAGUNA-LBNO (C2P? SPL?)
- Beta-Beams
 - Low gamma $(\gamma \sim 100)$
 - High gamma (γ ~ 350 580)
- Neutrino Factories
 - High energy ($E_{\mu} = 25 50 \text{ GeV}$)
 - Low energy ($E_{\mu} = 4.5 10 \text{ GeV}$)

General landscape

BB100,BB350: hep-ph/0406132 hep-ph/0503021

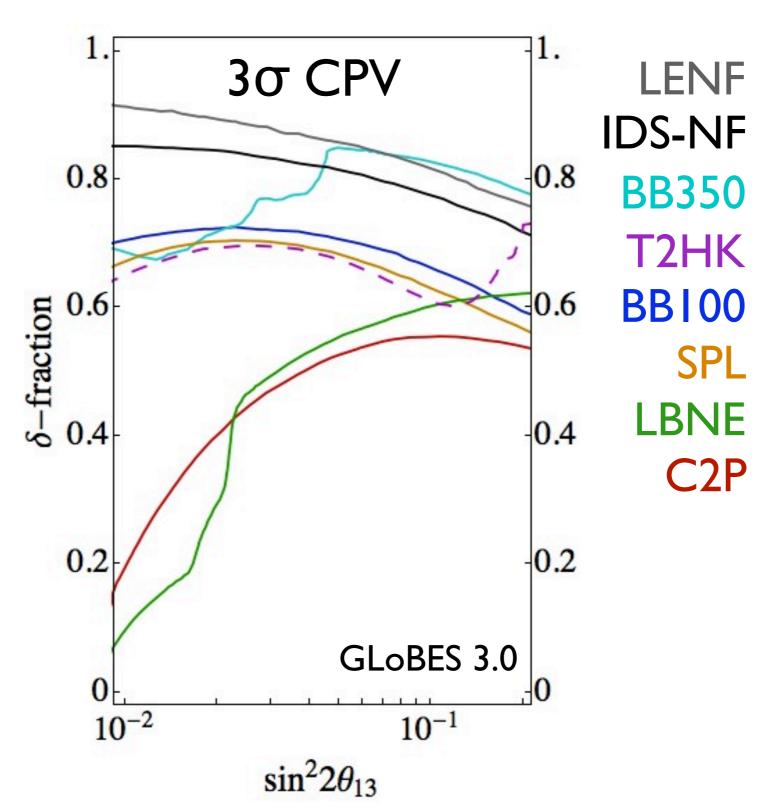
T2HK: hep-ex/0106019 1109.3262 [hep-ex]

C2P, SPL: 1001.0077 [physics.ins-det] hep-ex/0411062 1106.1096 [physics.acc-ph]

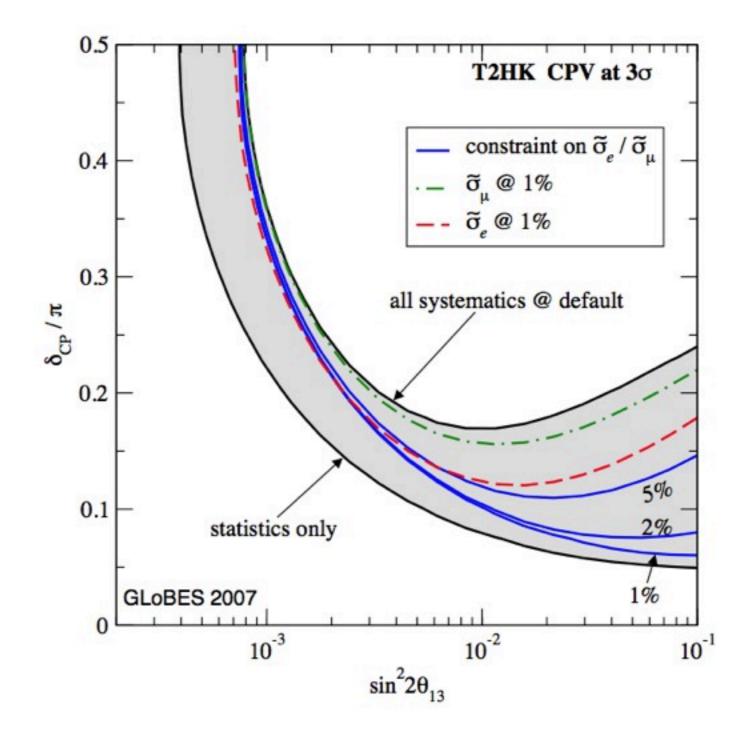
LENF: 1012.1872 [hep-ph]

LBNE: 1110.6249 [hep-ex]

IDS: 1112.2853 [hep-ex]

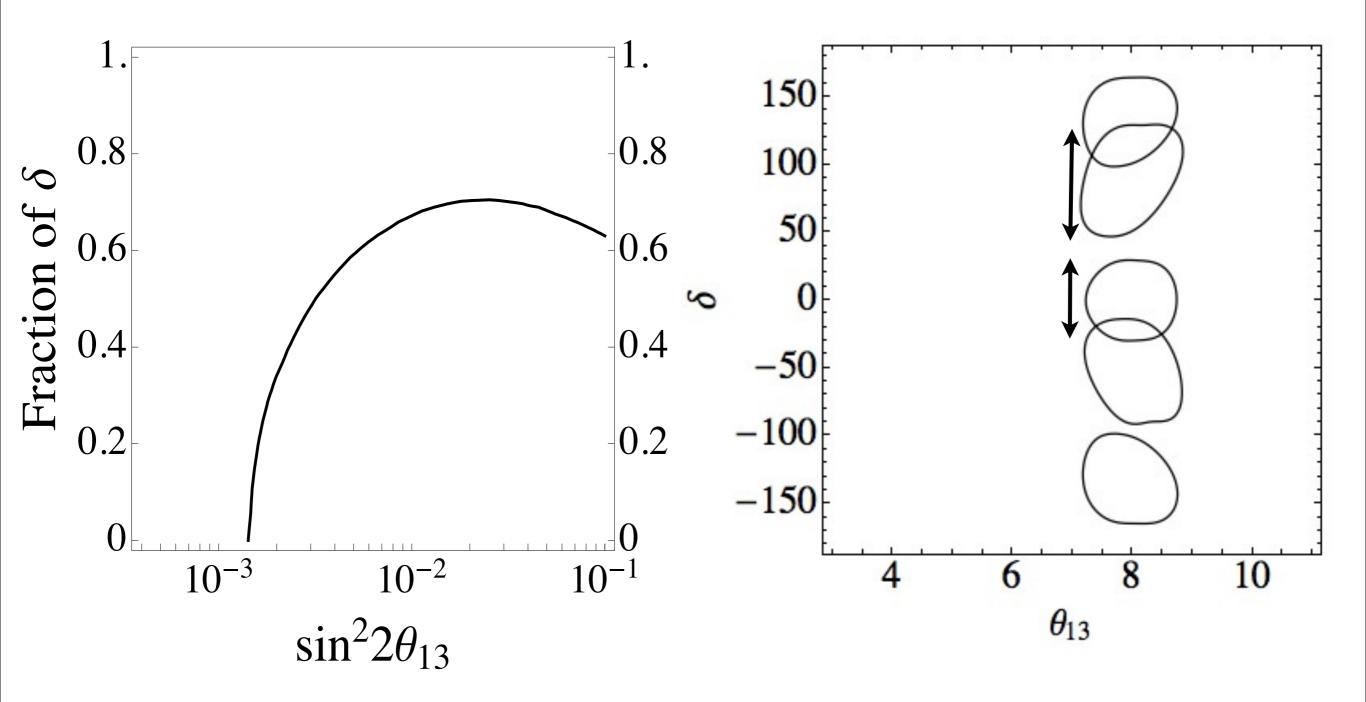


The impact of systematics



Huber, Mezzetto, Schwetz, 0711.2950 [hep-ph]

Discovery vs precision



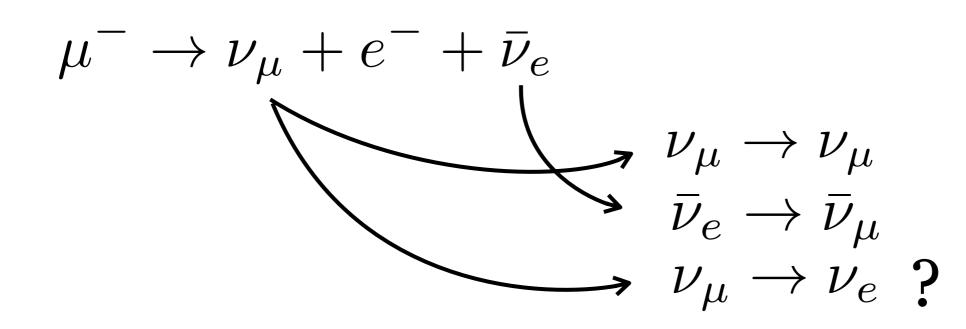
Does a large θ_{13} change anything?

Observables: Solar params 🗹 atmosph params - reduce budget? I theta13 🗆 delta hierarchy \Box theta23=45? precision

TO DO:

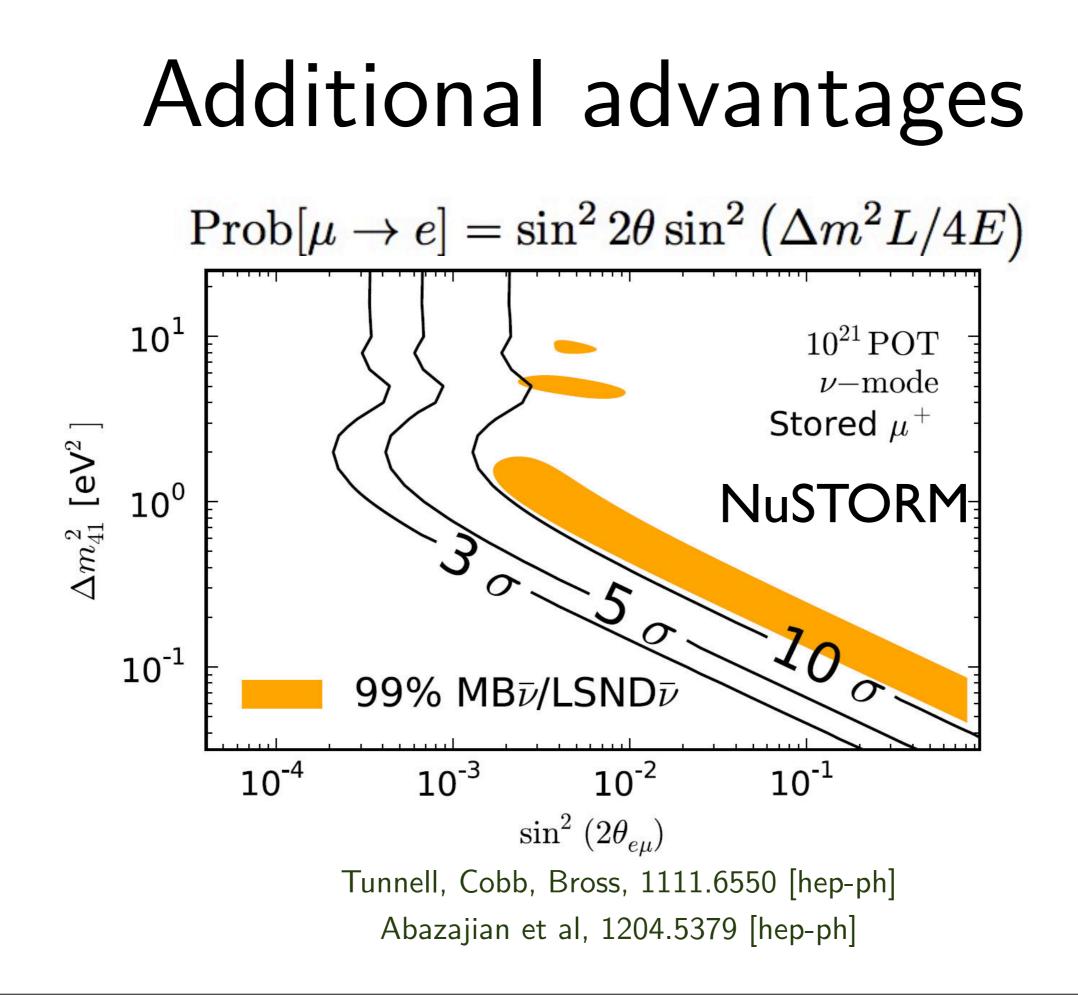
- how much can we (incremental appr.) - can we live with P larger bgs? - what is the effect of systematics?

Why a NF?

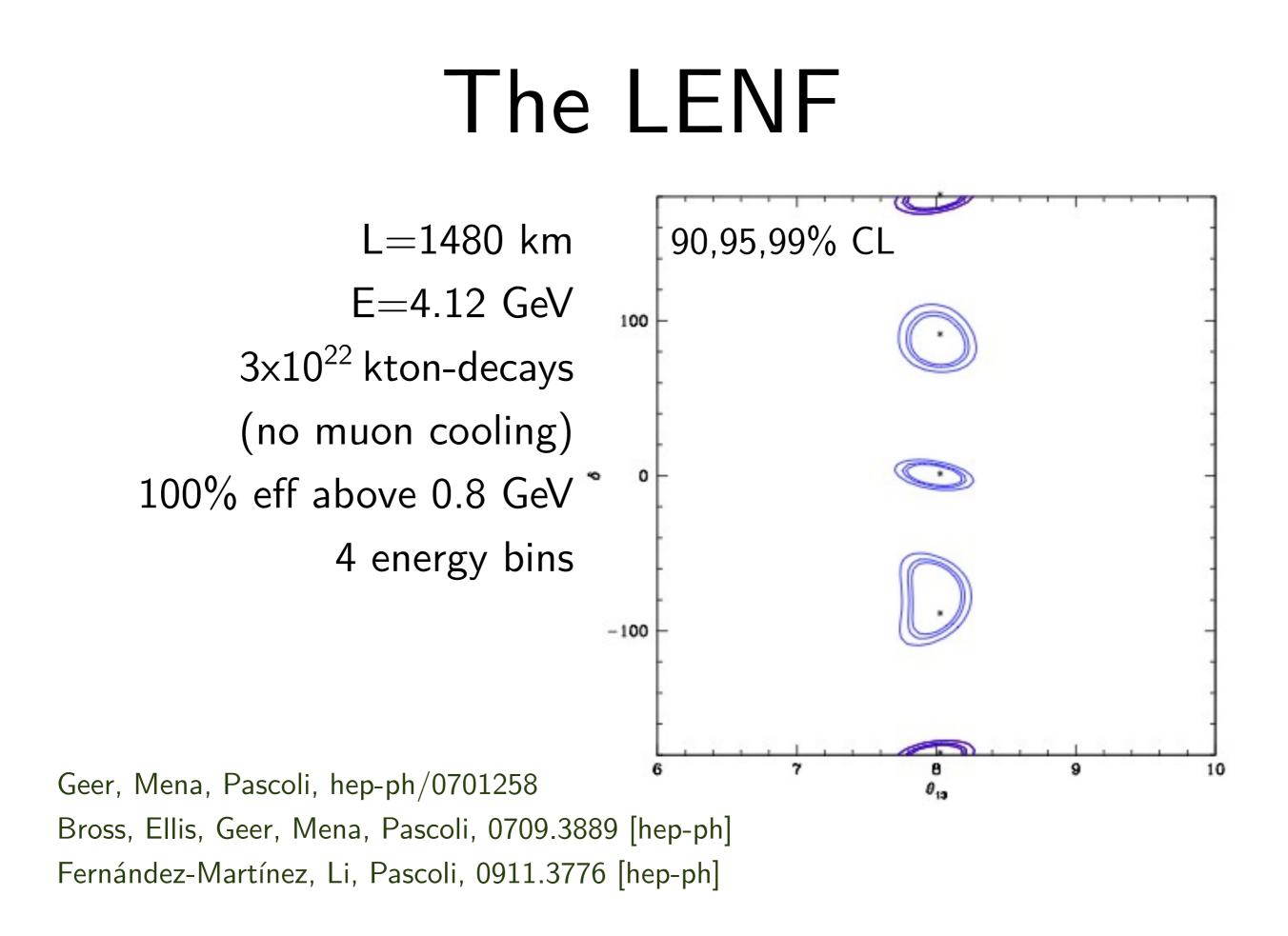


- 1) Offers multiple channels
- 2) No beam backgrounds
- 3) Low systematics
- 4) Great discovery potential and precision
- 5) Statistically less limited than other facilities

Thursdav. Mav 10.

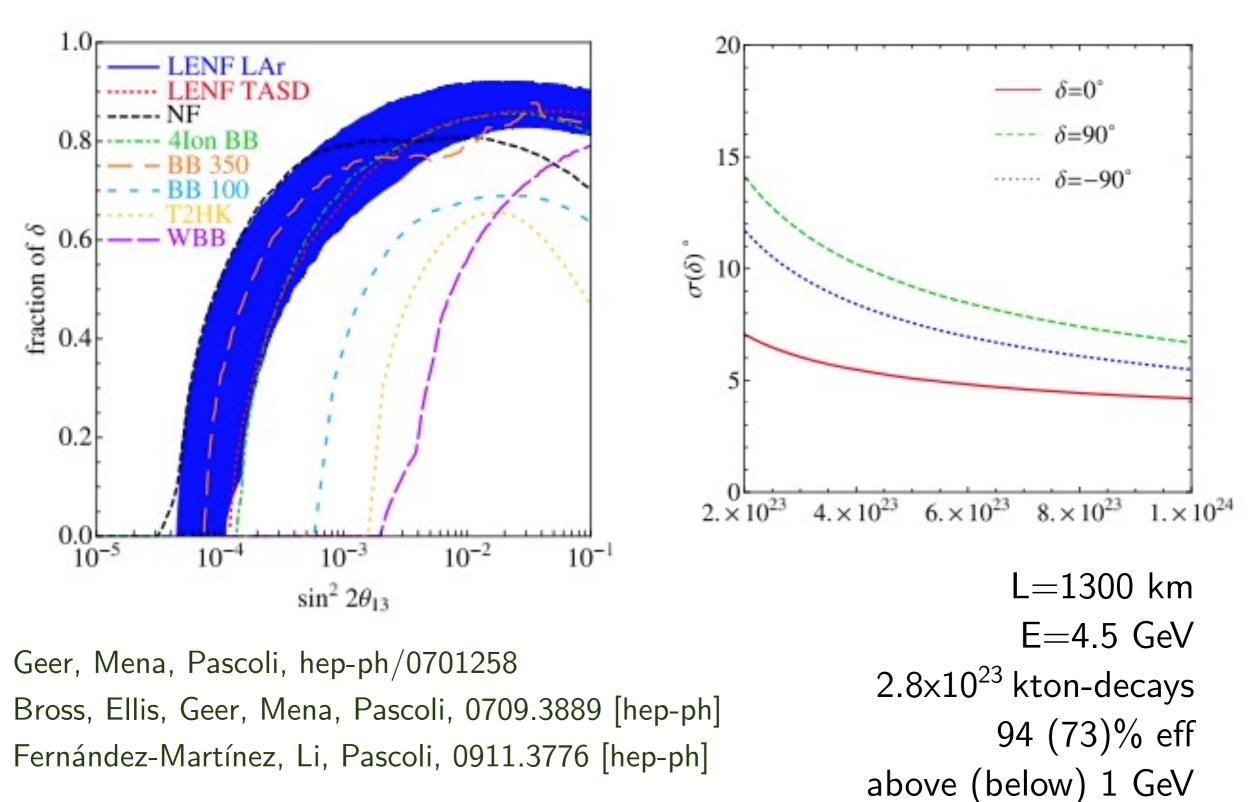


The concept of a Low Energy Neutríno Factory

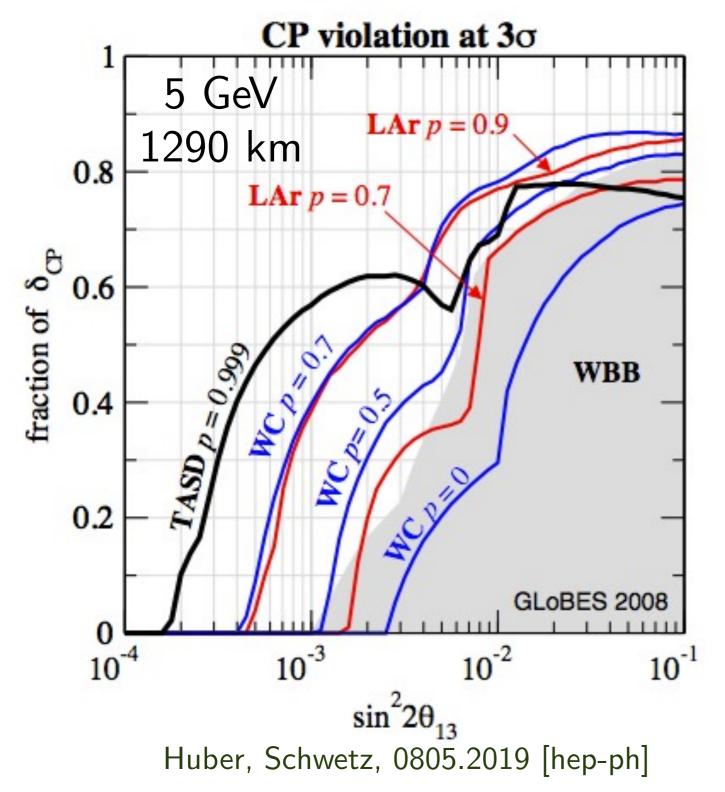


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The LENF



The LENF

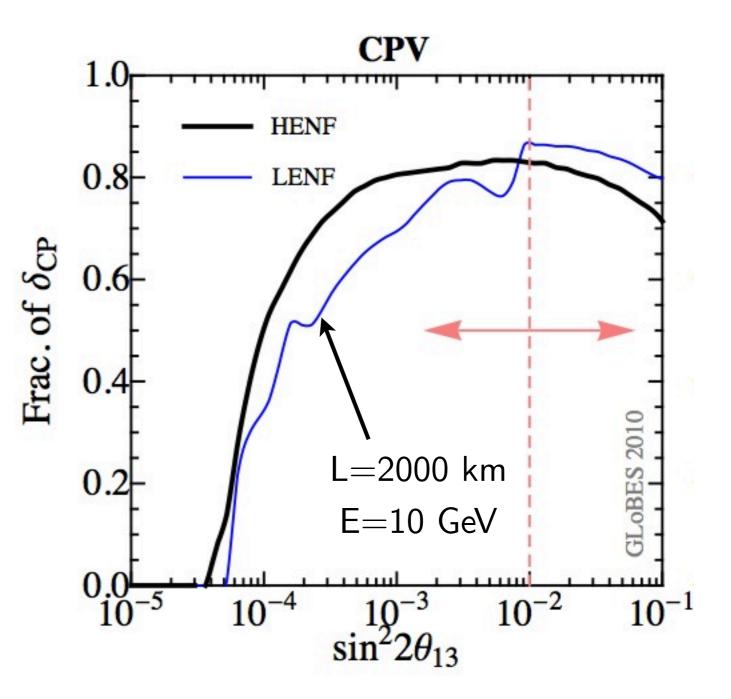


Charge signatures in nonmagnetized detectors: 1) µ capture in nuclei 2) different angles between incoming neutrino and outgoing lepton in lab frame for QE events 3) tag protons for neutrinos and neutrons for antineutrinos in QE interactions

The LENF

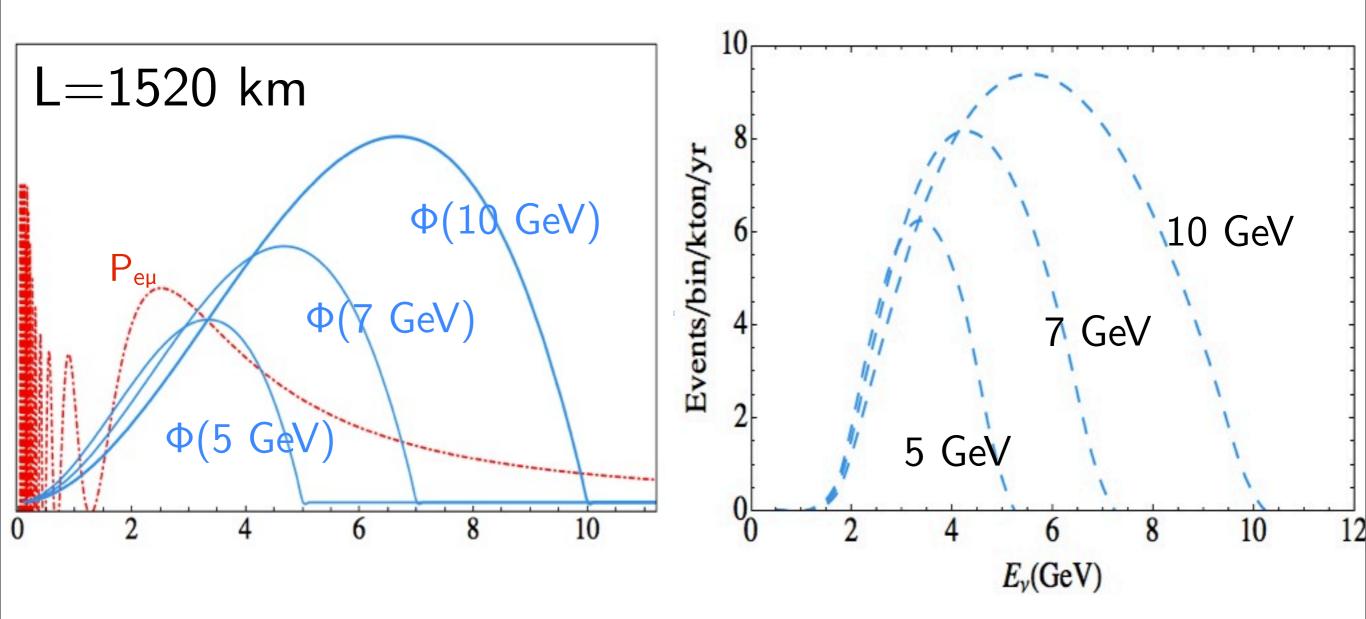
Latest MIND simulations show that a LENF could also use magnetized iron:

- well-known technology
- can be built on large
 scale and easily
 magnetized



Agarwalla, Huber, Tang, Winter, 1012.1872 [hep-ph] Cervera, Laing, Martín-Albo, Soler, 1004.0358 [hep-ex]

A setup from RAL to GS



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Incremental approach

- The NF at full luminosity (10²¹ muons/yr) relies on a high power proton driver and muon cooling...
 - Muon cooling needed? (~1.7 less muon decays)
 - 4 MW needed? what could be done with 800 kW?
 - 100 kton detector needed?
 - Huge bg rejection factors (10⁻⁴) needed?
- Exposure-thinking: how much can beam/detector requirements be reduced?

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These are equivalent to an order of magnitude in total exposure

Possible detectors

Totally-Active Detector (TAD)

- Considered magnetizable, but at most 30 kton
- High efficiency and energy resolution
- Bin width of 125 MeV
- Bg rejection of 10⁻³ assumed
- NC backgrounds migrated with matrices from LBNE
- Platinum channel can be added (10⁻² bg rejection assumed)

MIND

- Magnetized, at most 100 kton fiducial
- Migration matrices from 1004.0358 [hep-ex]
- Bin width ranging from 1 to 5
 GeV

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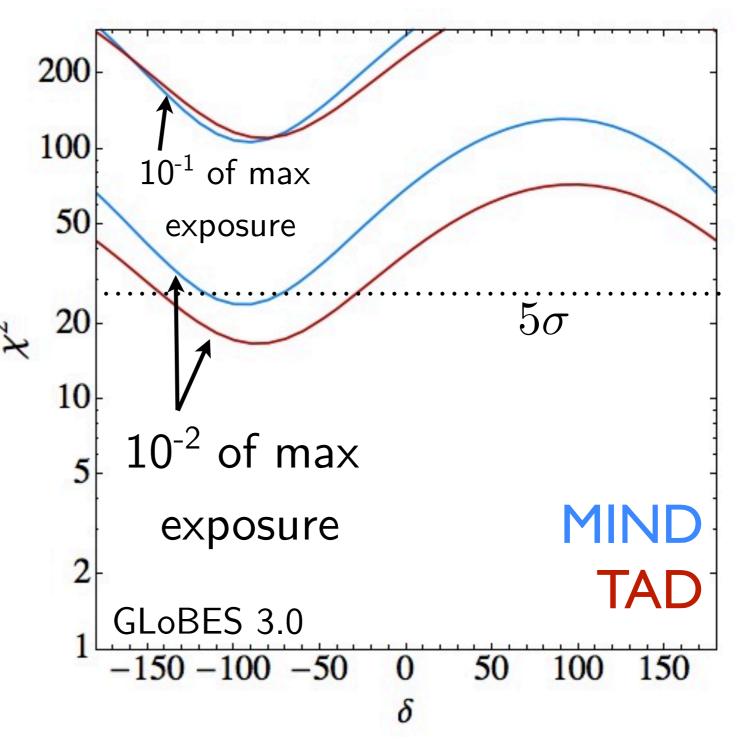
Mass hierarchy and CPV discovery potentials

Mass hierarchy

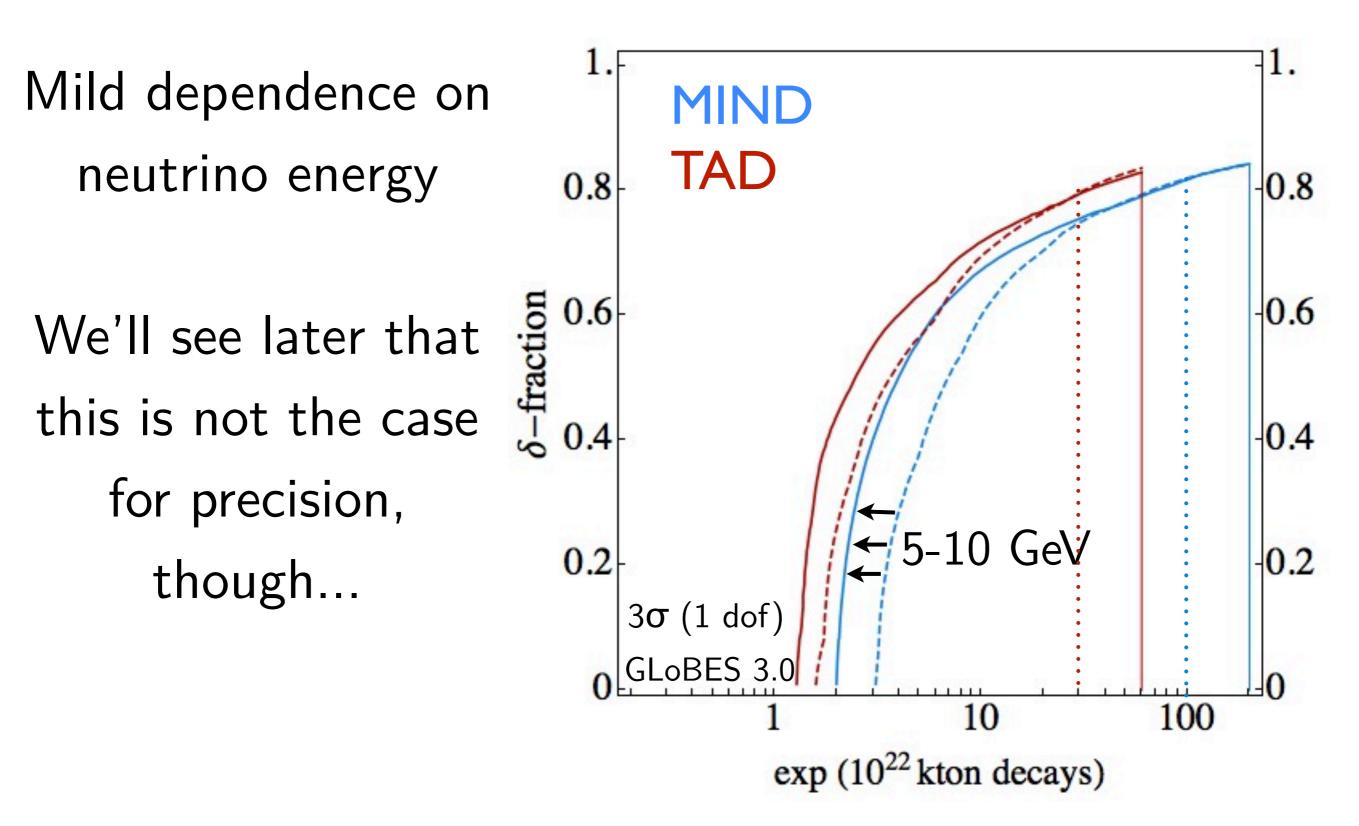
Maximum exposure:

- 5x10²⁰ muon decays
 polarity/yr
- 10 yr
- 2×10^7 useful sec/yr
- 100 kton for MIND
- 30 kton for TAD

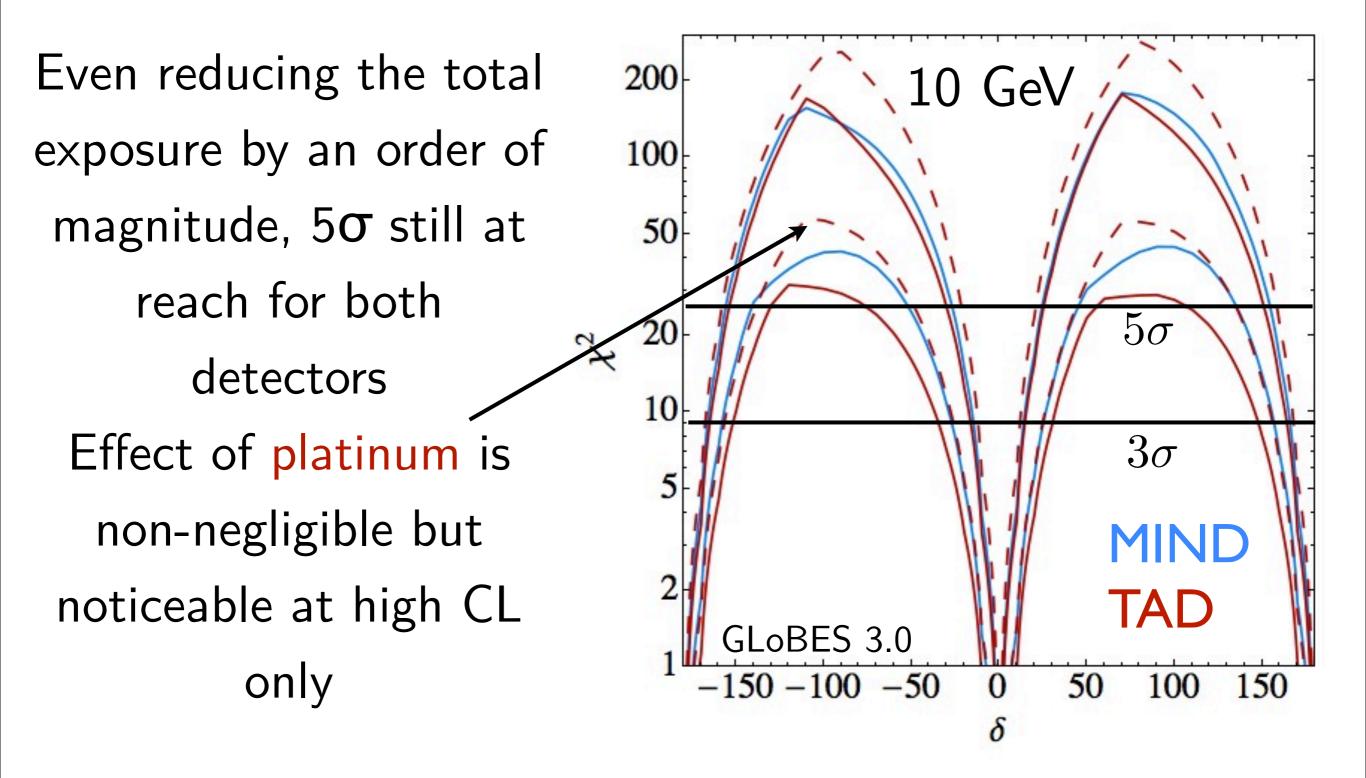
for MIND: $(1-2)\times10^{24}$ kton decays for TAD: $(3-6)\times10^{23}$ kton decays



CP violation



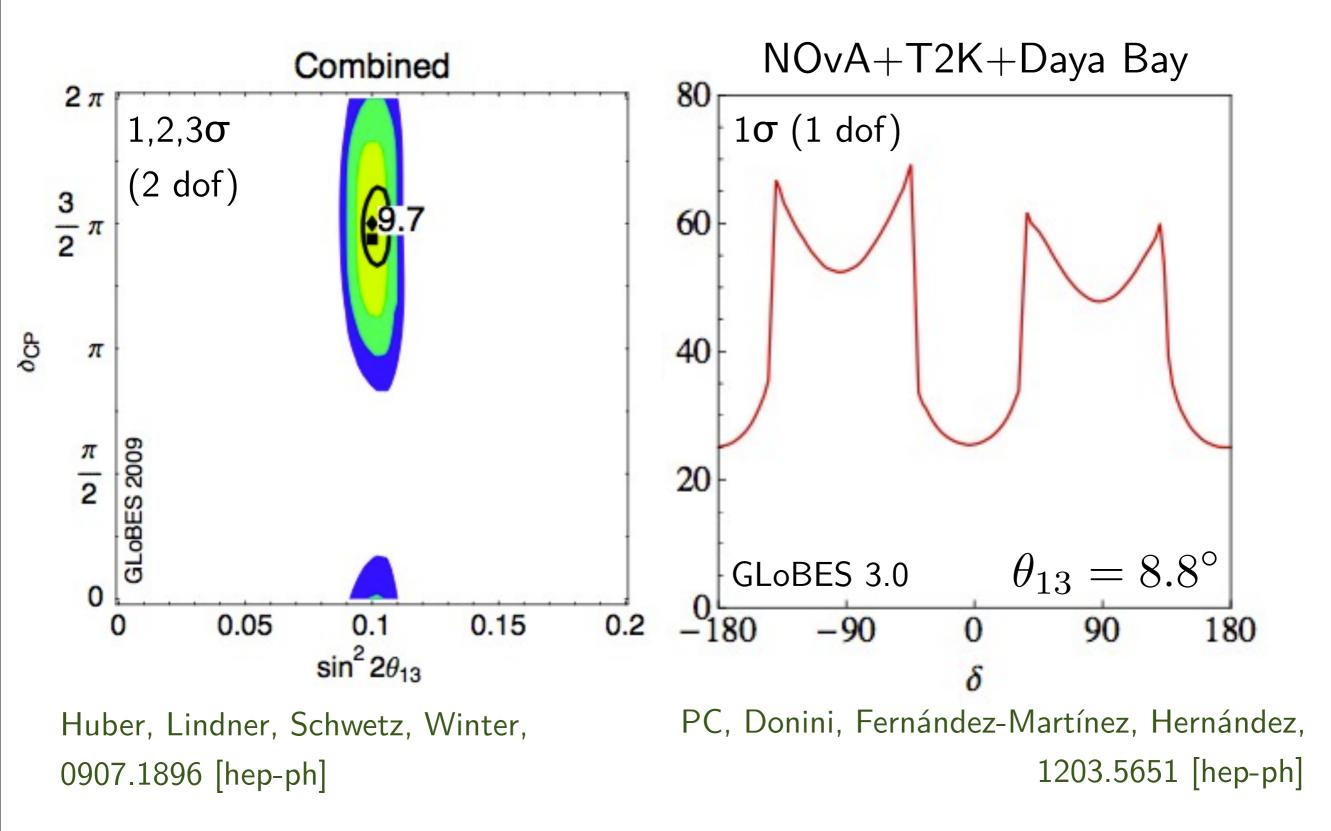
CP violation



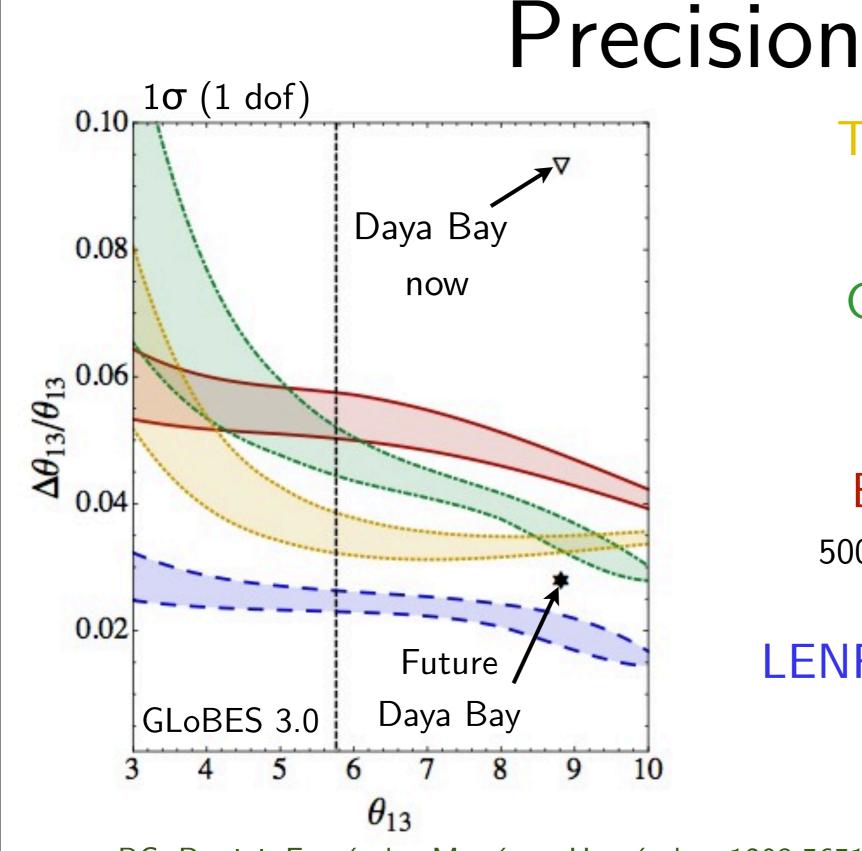
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Precision

The starting point



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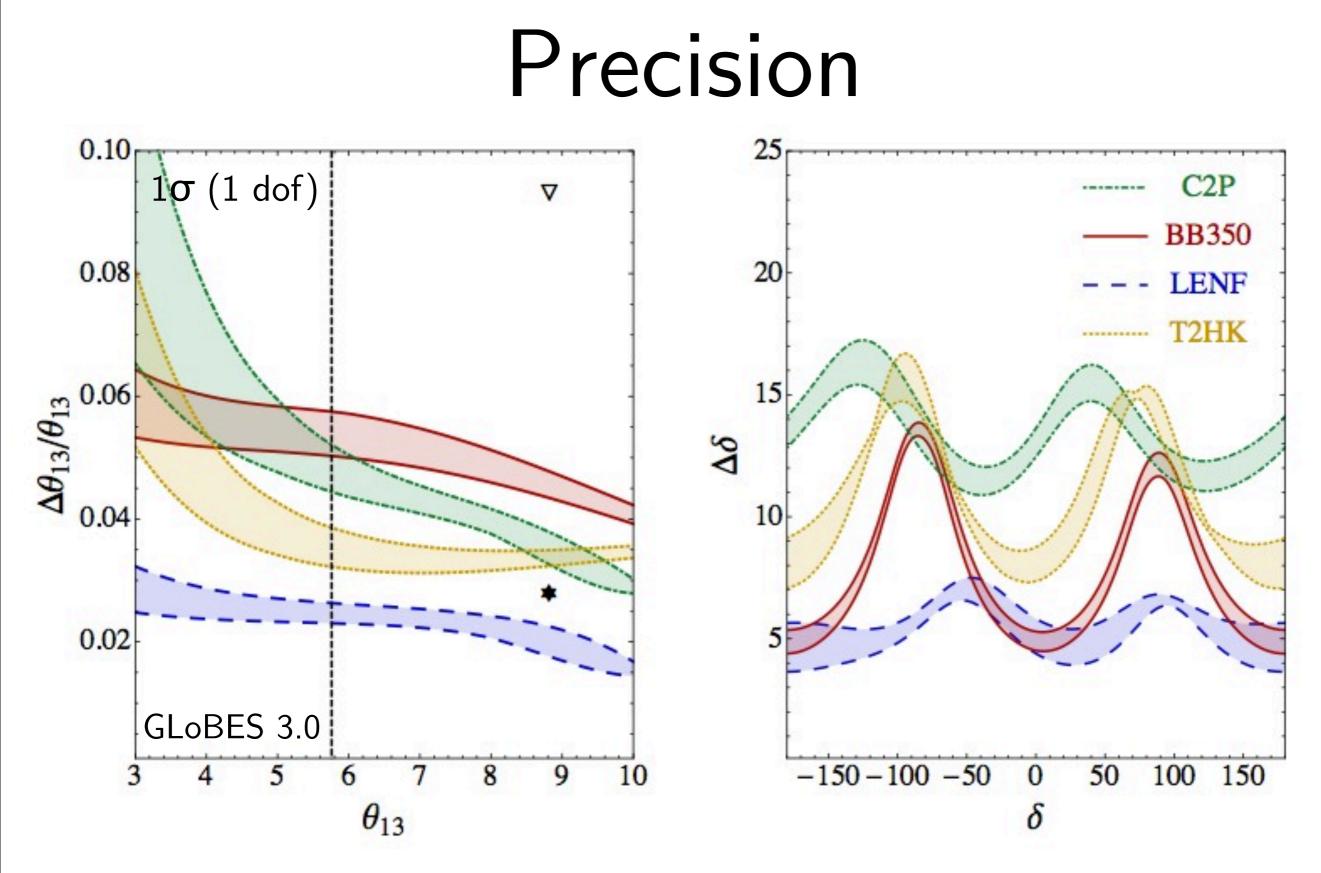
T2HK: 4 MW, 500 kton WC, 295 km, 5% sys

C2P: 800 kW, 100 kton LAr, 2300 km, 5% sys

BB350: 1.1(2.8) x10¹⁸ ions, 500 kton WC, 650 km, 2.5% sys

LENF@2000: 1.4x10²¹ µ decays 100 kton MIND, 2000 km, E=10 GeV,2.5% sys

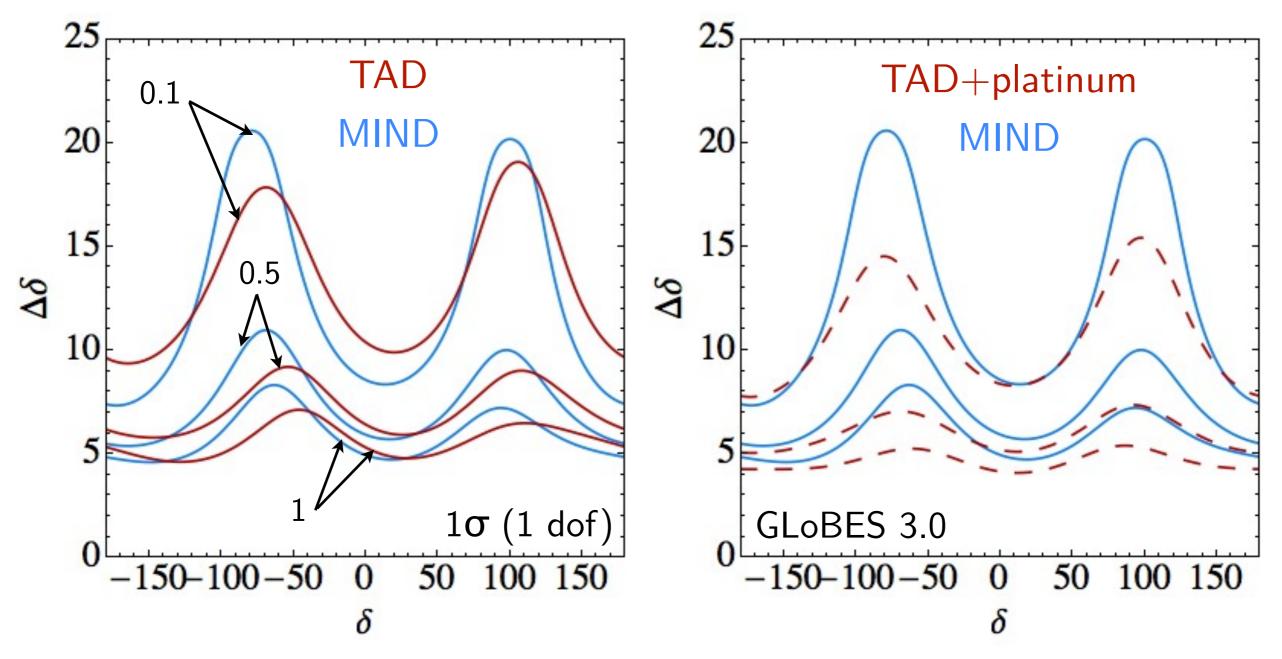
PC, Donini, Fernández-Martínez, Hernández, 1203.5651 [hep-ph]



PC, Donini, Fernández-Martínez, Hernández, 1203.5651 [hep-ph]

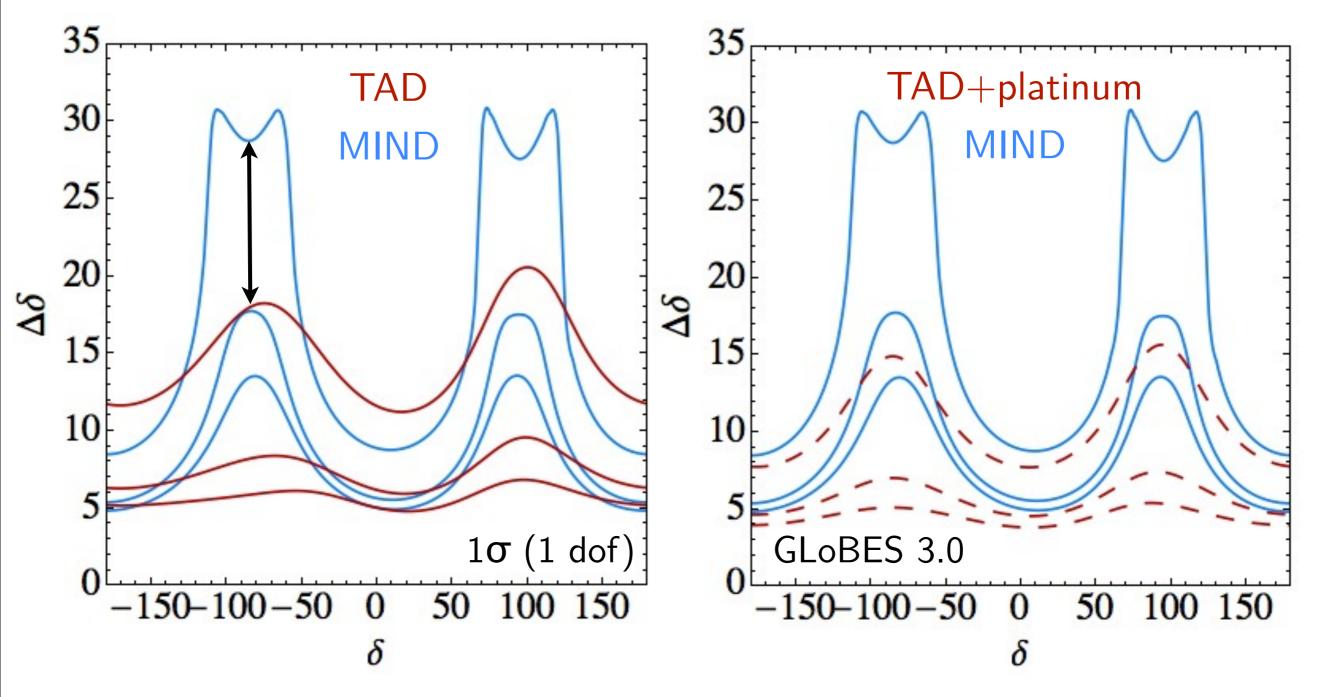
Precision at the LENF

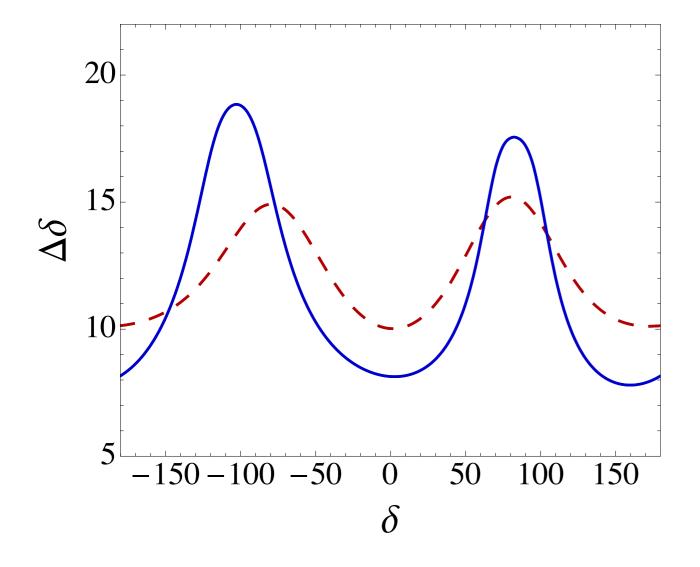
For E=10 GeV and different exposures

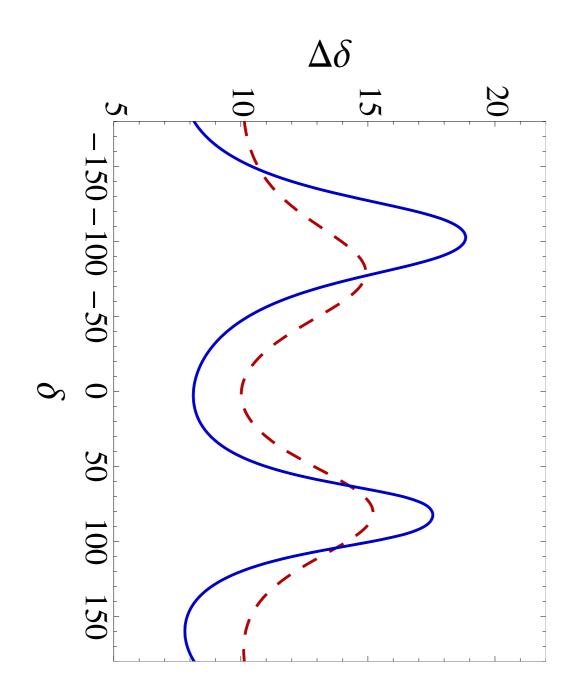


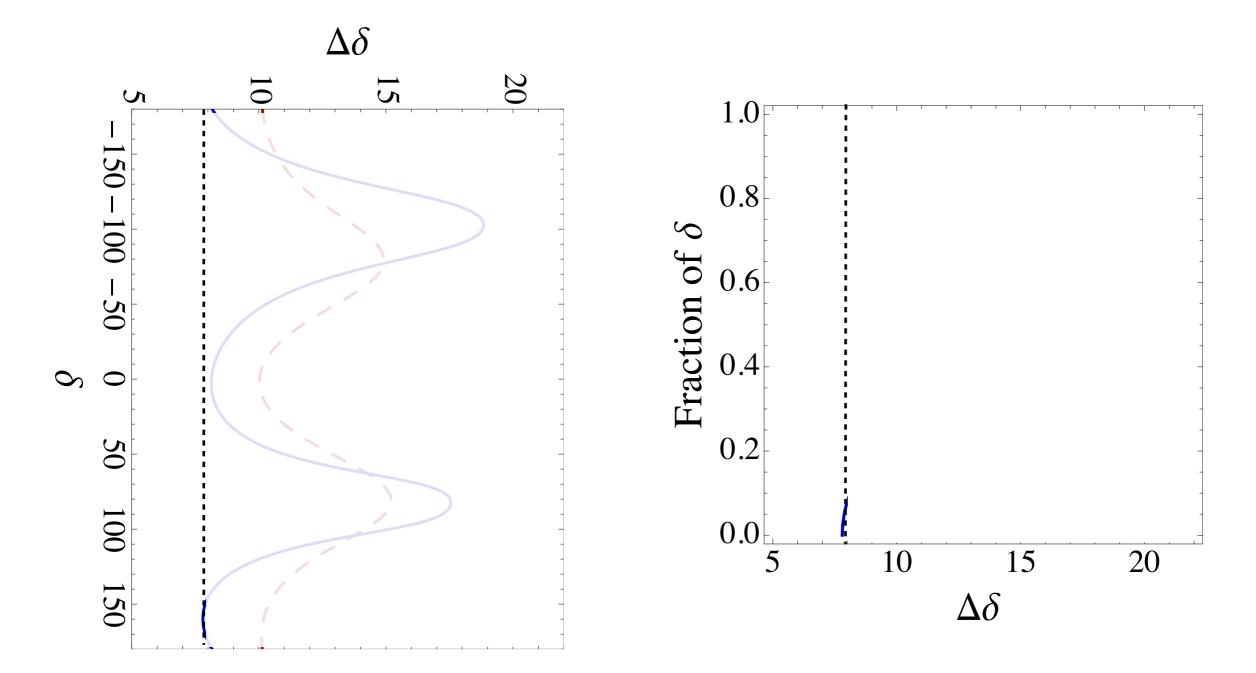
Precision at the LENF

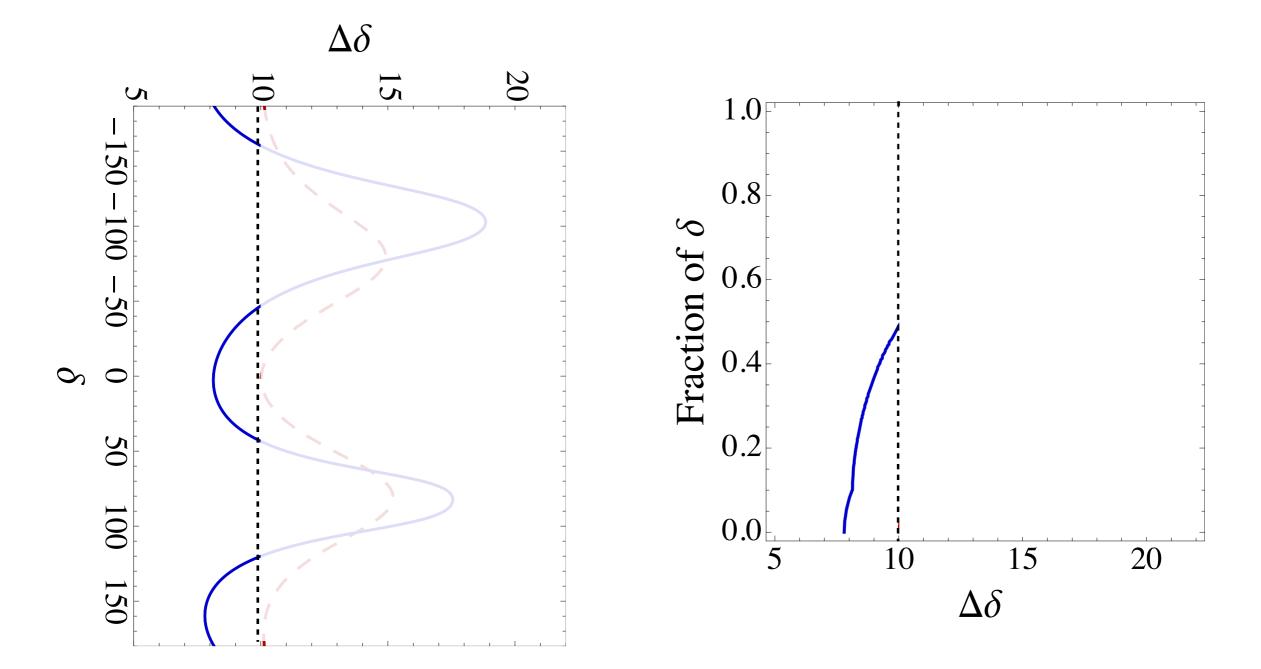
For E=5 GeV and different exposures

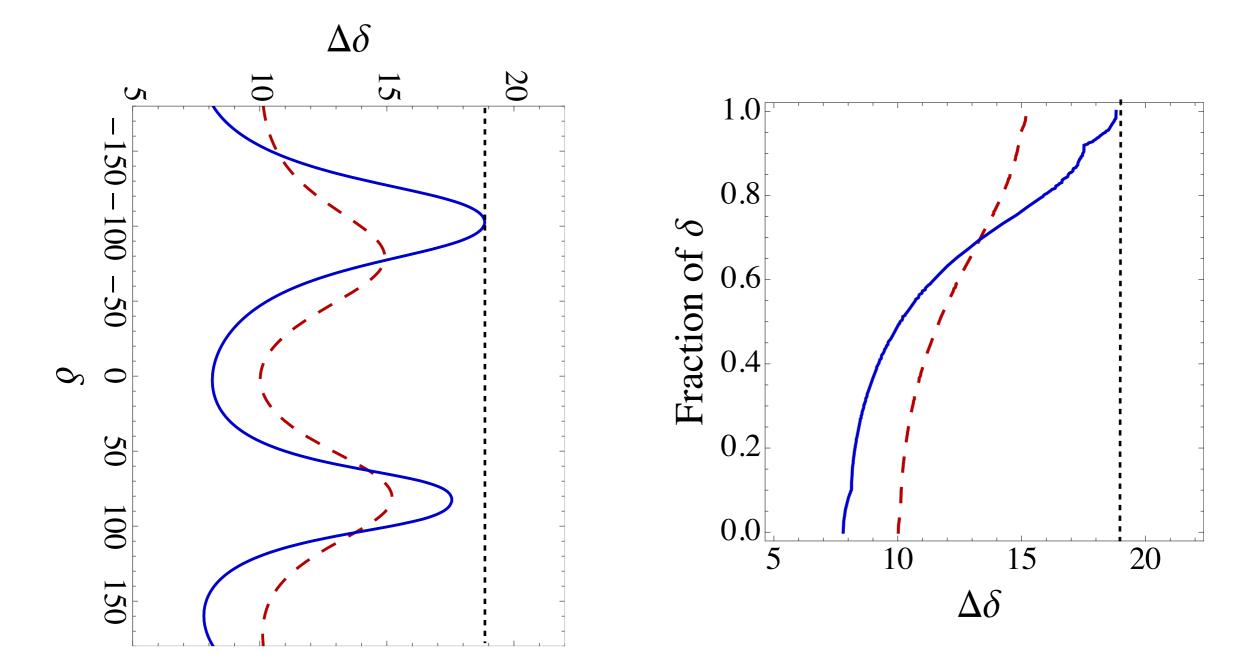




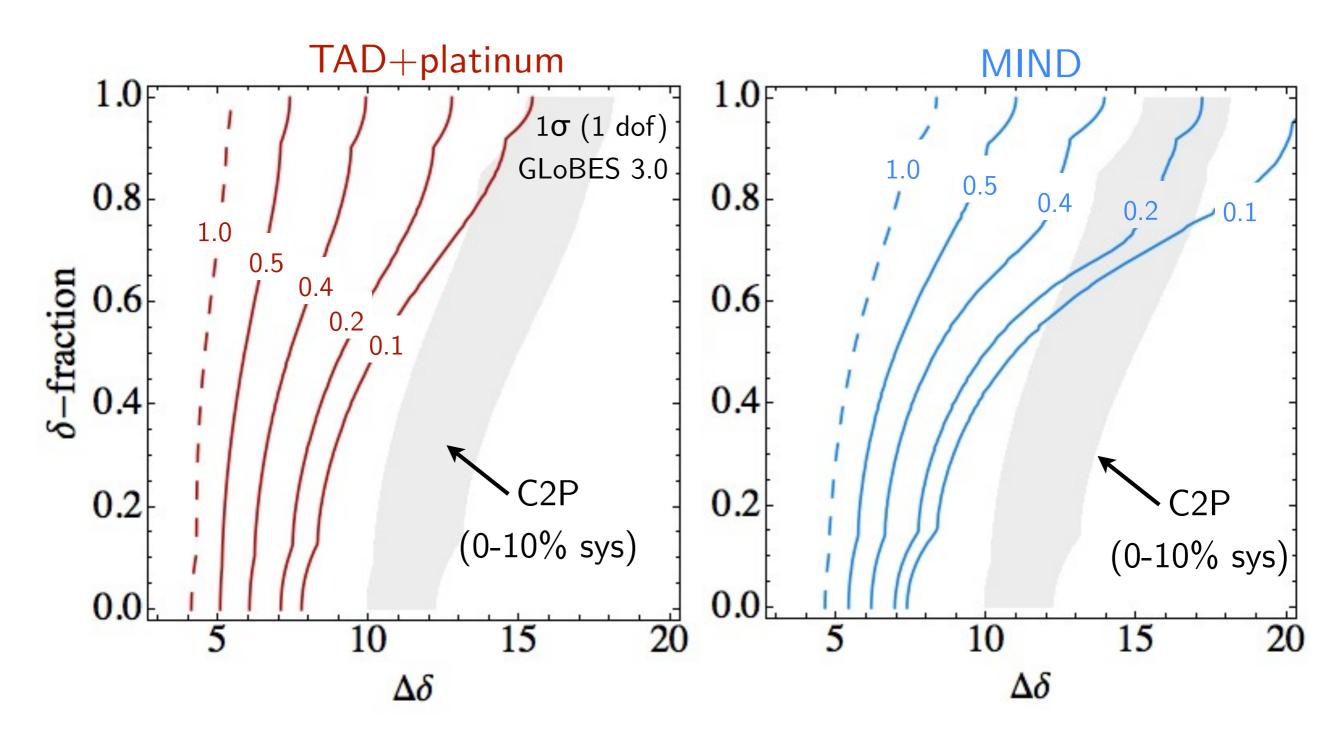




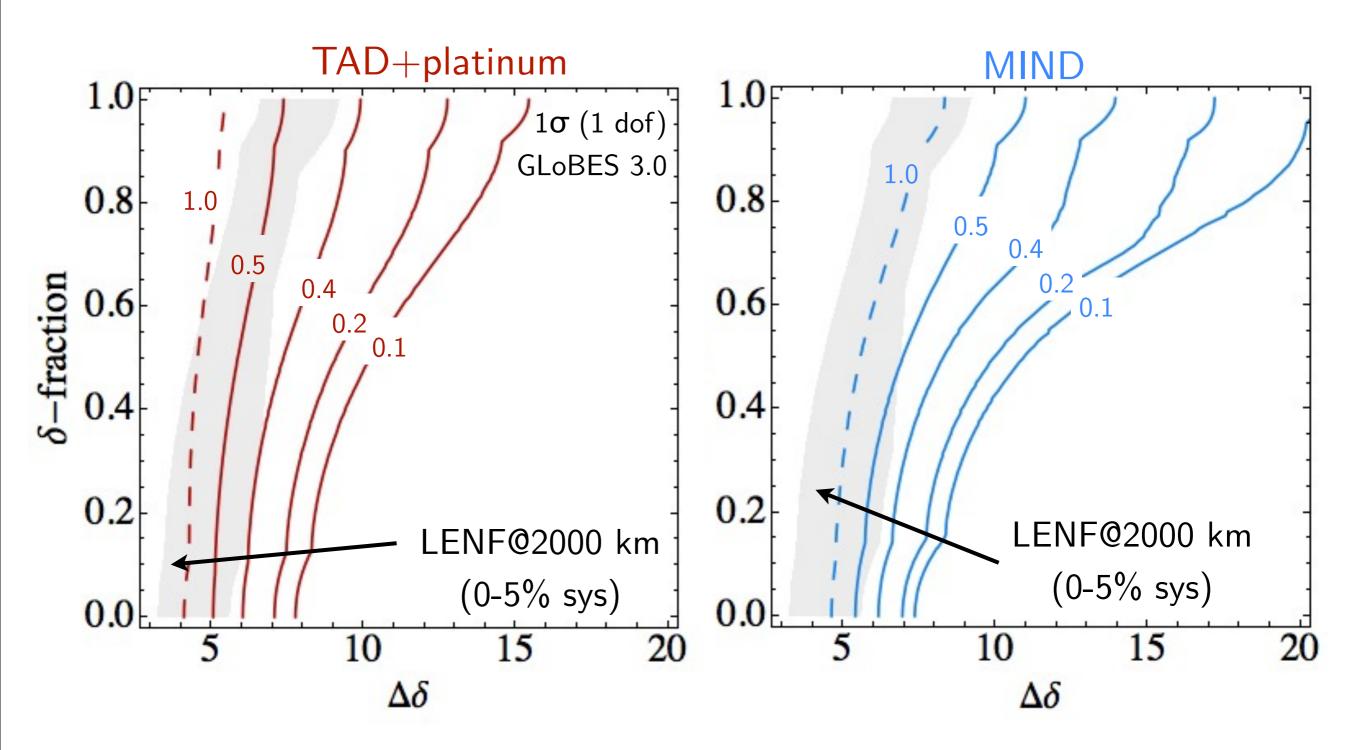




Precision at the LENF



Precision at the LENF



Conclusions

- A LENF from RAL to GS offers an excellent opportunity to study neutrino oscillations
- 10 GeV proves to be much better than 5 GeV even if slightly offpeak, independently of detector technology
- TAD offers better global performance
- A 800 kW NF w/o cooling is significantly better than a long baseline SB with the same power
- If the maximum luminosity can be reached, a NF from RAL to GS is totally comparable to the LENF@2000 (new IDS setup)
- Can we live with larger backgrounds? is magnetization necessary? can we use MIND for low energies?

Backup

Thursdav. Mav 10.