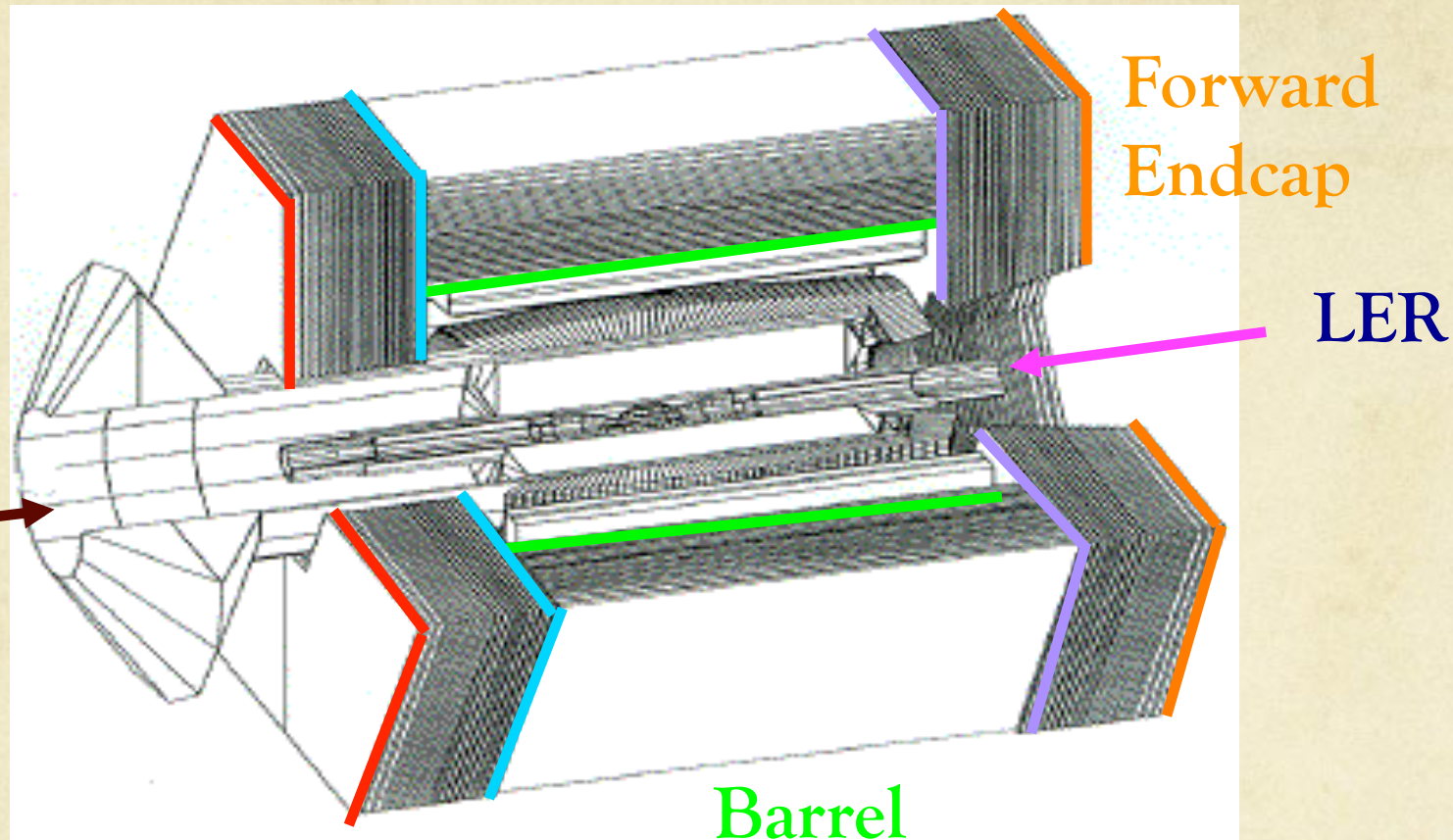




IFR Background Report

Valentina Santoro
INFN Ferrara

Hot regions



Barrel: innermost layers, mostly neutrons

FWD encaps (hottest region) : inner layers and outer layers (BEAM halo), electron and photons

BWD encaps: inner layers and small radii

What's new from the Frascati CM Meeting

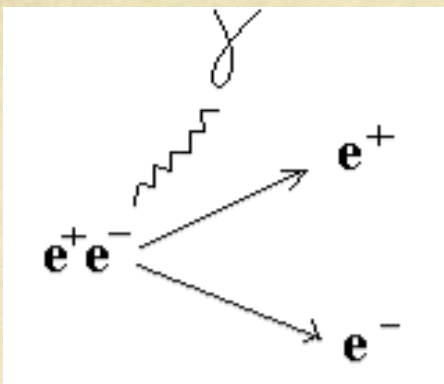


Frascati Dec 2011

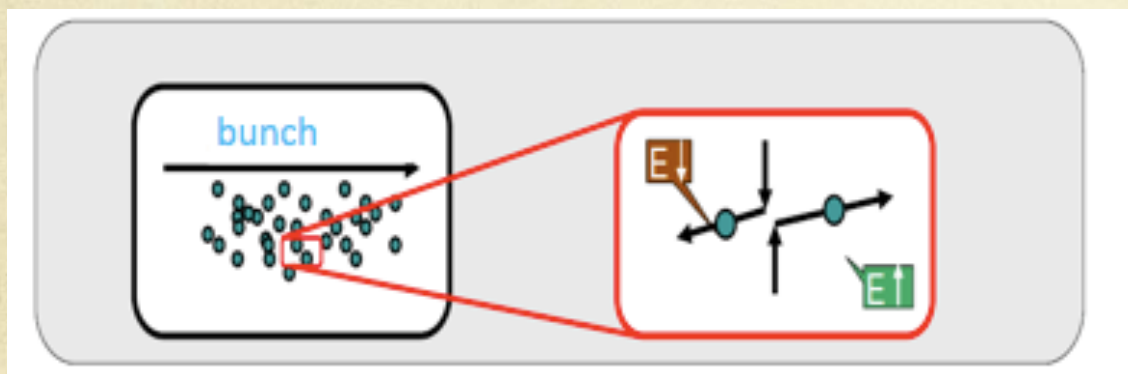
- ✓ Beam Composition for the IFR background
- ✓ Radiative Bhabha Background Studies (neutrons, photons and electron)
- ✓ Touschek background (neutrons, photons and electron)
- ✓ Pair background (neutrons, photons and electron)
- ✓ Background Studies and Absorbed dose on our FEEs

New

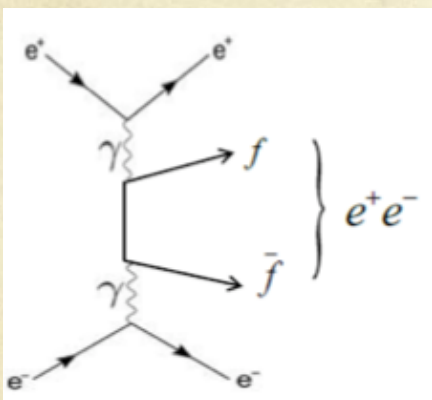
- ✓ Proton study
- ✓ FEEs studies improved
- ✓ Beam-Gas Background
- ✓ Neutron Background Shielding



Radiative Bhabha



Touschek scattering : results from a Coulomb collision of two relativistic electrons in a particle beam, producing an instantaneous change in particle energy



Pair production

$$e^+e^- \rightarrow e^+e^- \gamma\gamma \rightarrow e^+e^- e^+e^-$$

New Beam-Gas Scattering

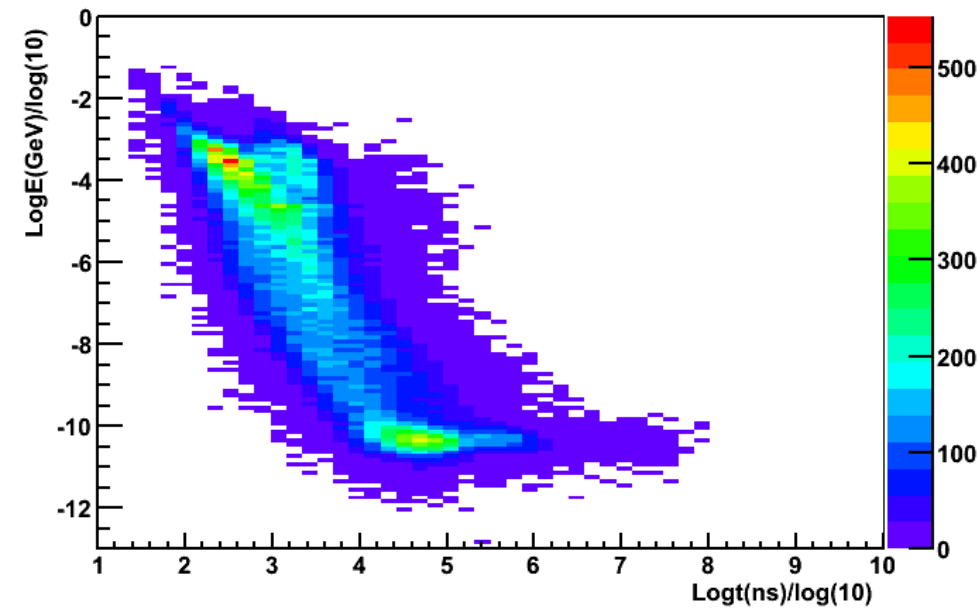
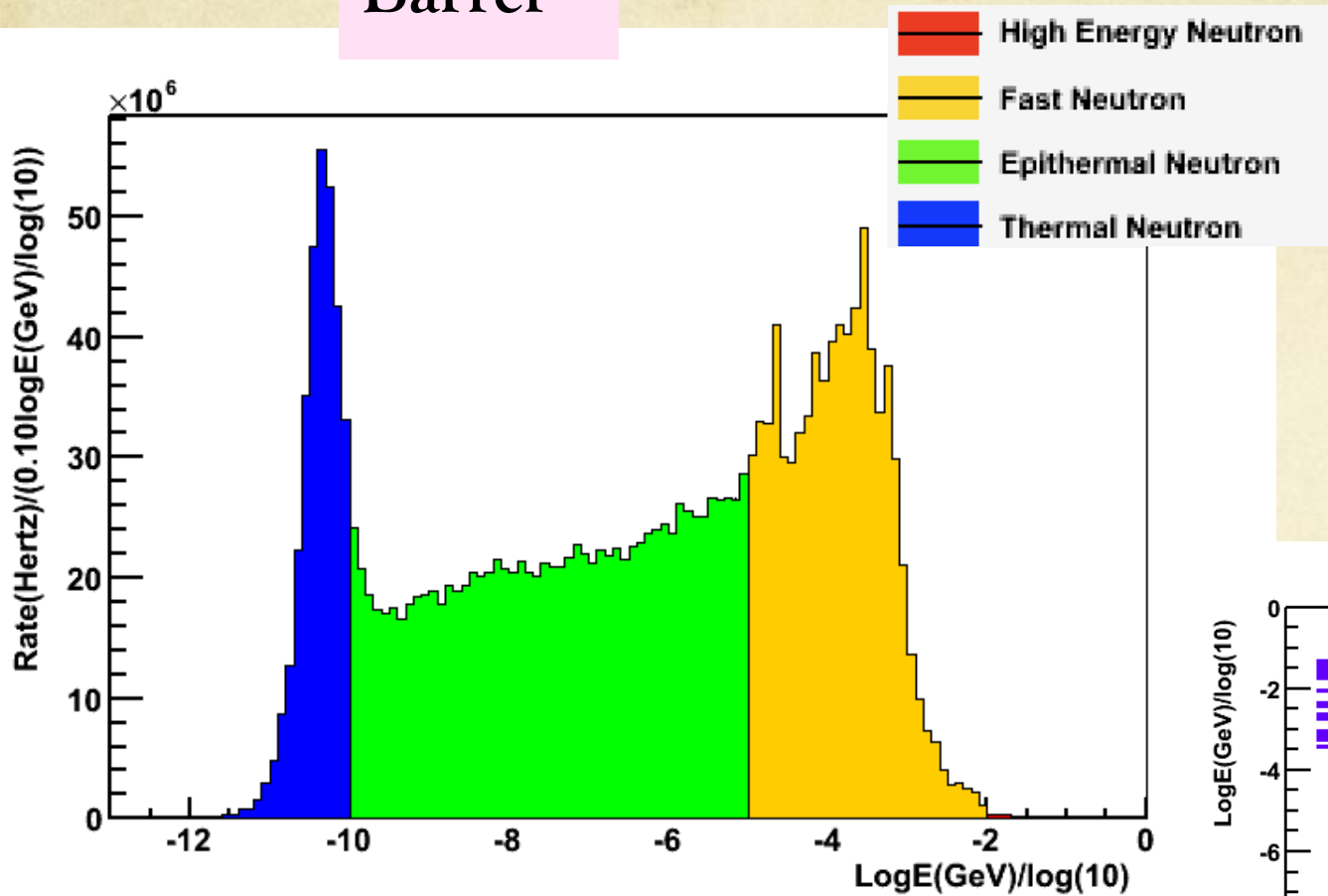
Neutron Energy Distributions for Radiative Bhabha events



Barrel

Barrel

Energy vs time distributions



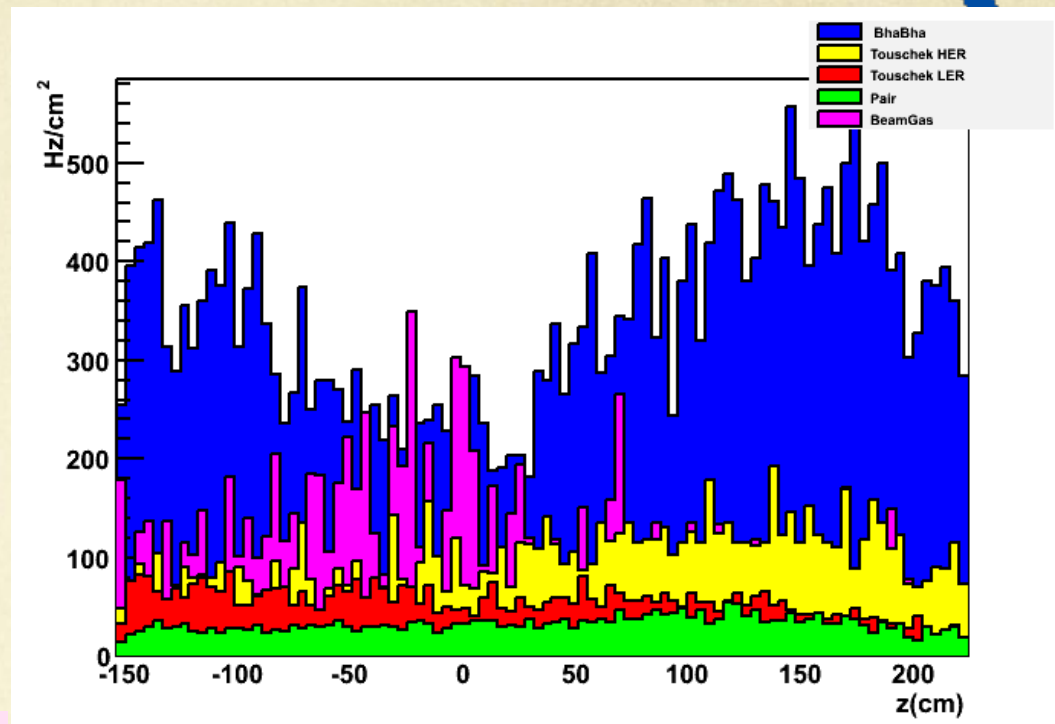
The Energy distribution for FWD and BWD Endcap are similar

Neutron Rates (for different background sources)

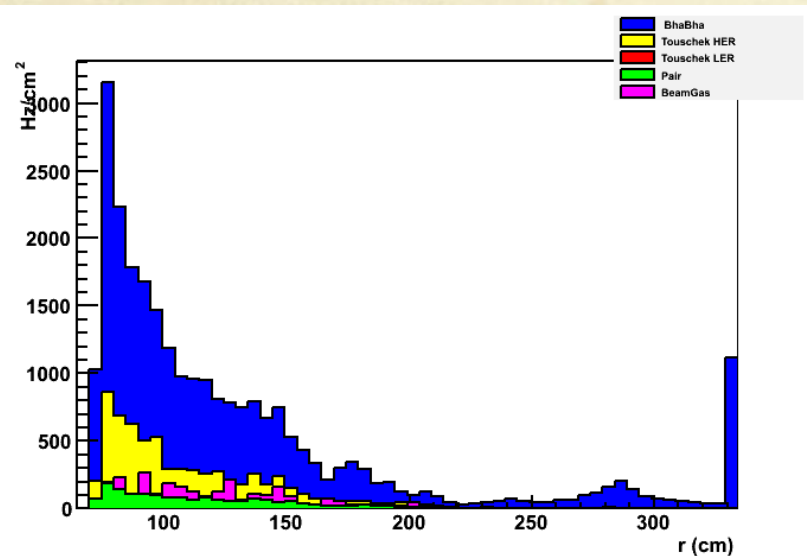


Rate vs Z-coordinate for Barrel

Rate of 450 Hz/cm^2 - \rightarrow about 3×10^9 neutrons/cm² for a year



Rate vs radius for FWD Endcap

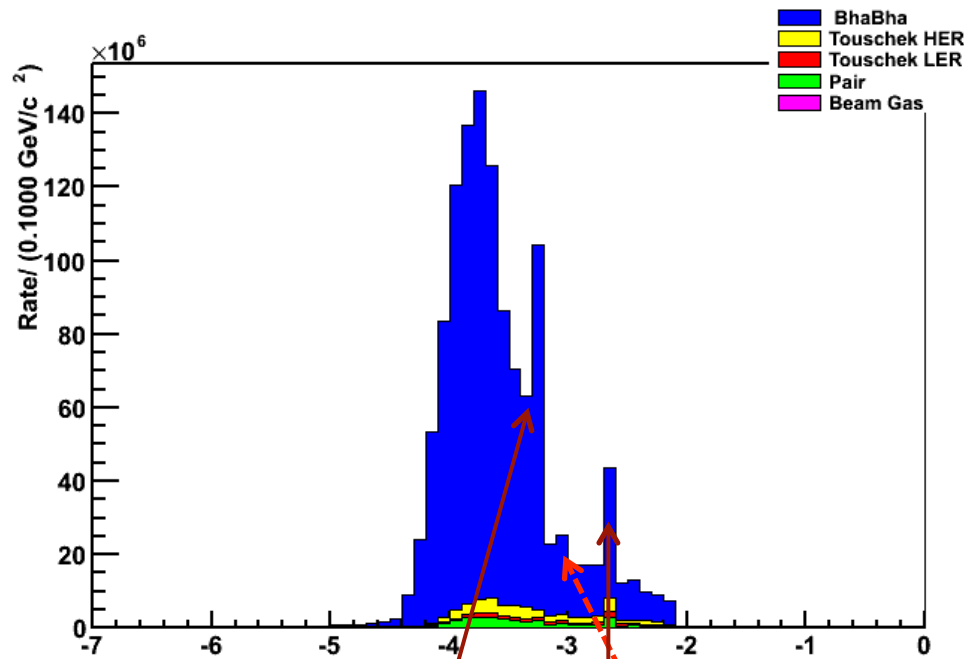


All the rate are normalized to 1MeV energy

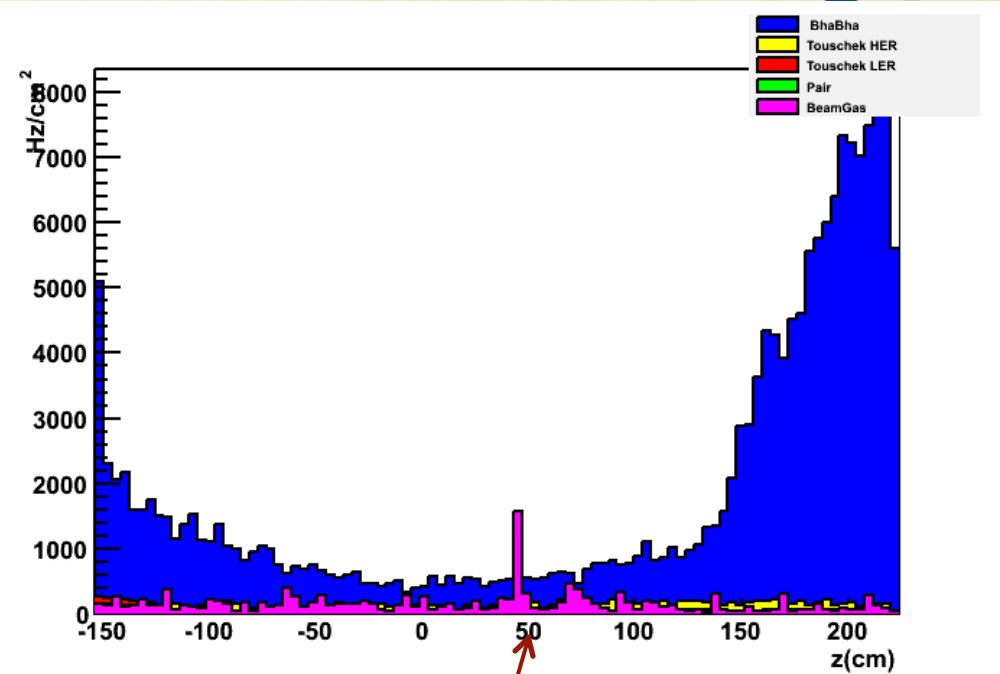
Photons for different background sources



Barrel: Photon Energy Distribution



Rate vs Z-coordinate for Barrel



The Energy distribution for FWD and BWD Endcap are similar

Photons of energy ~ 0.512 MeV are from annihilation radiation

Photons of energy ~ 0.847 MeV are due from neutron inelastic scattering on Fe^{56}

Hot spots
Located in sextant 1

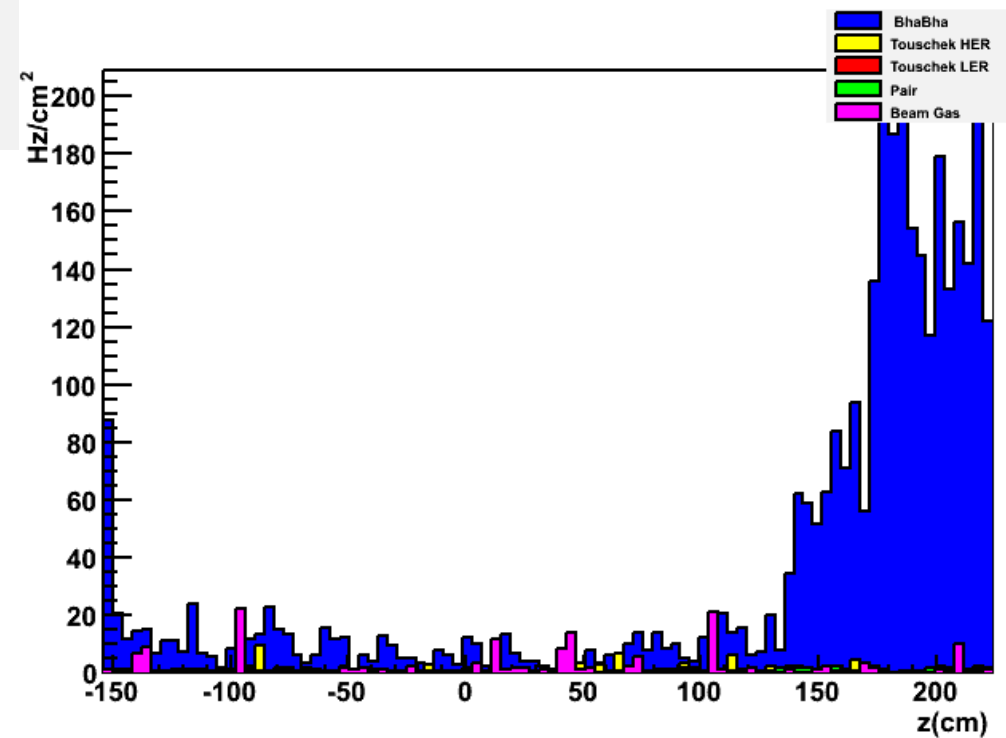
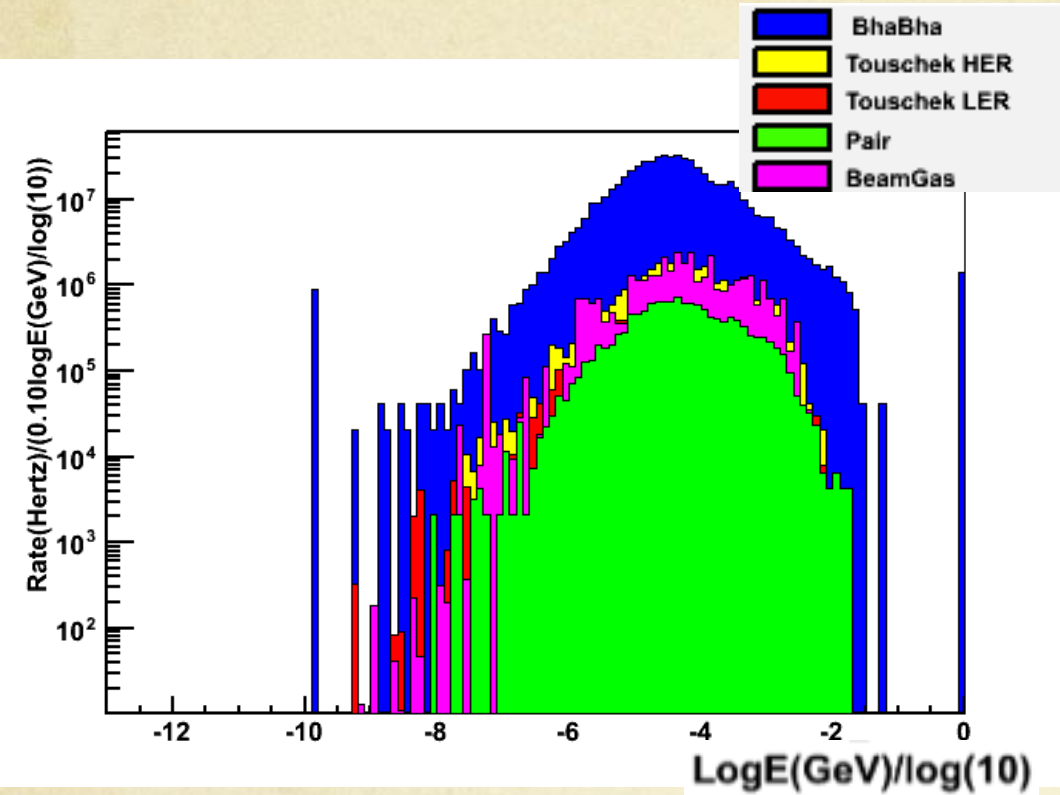
Photons of energy ~ 2.223 MeV are from neutron capture on Hydrogen



Electrons for different background sources

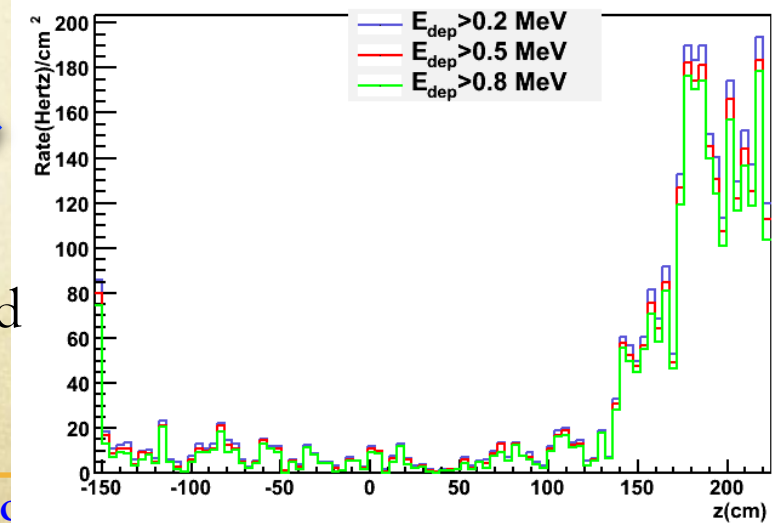
Barrel: Electron Energy Distribution

Rate Layer 0 vs Z-coordinate for Barrel



Rate Layer 0 vs Z-coordinate for Barrel with different cut on deposited Energy

The impact of the electron rates due to the background must be considered on the muon id

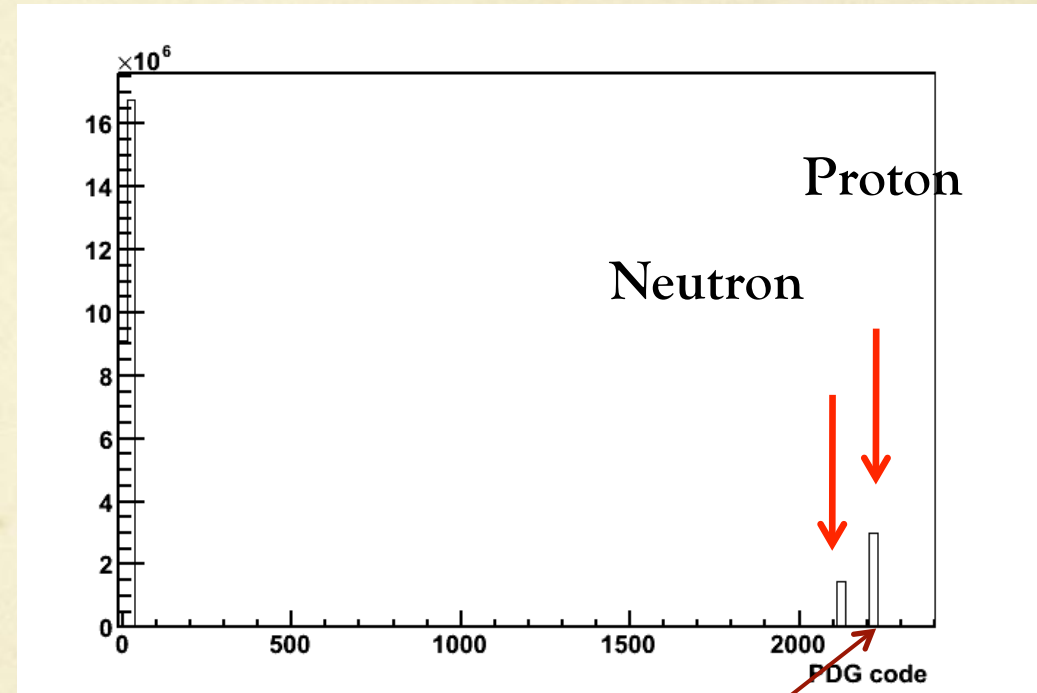
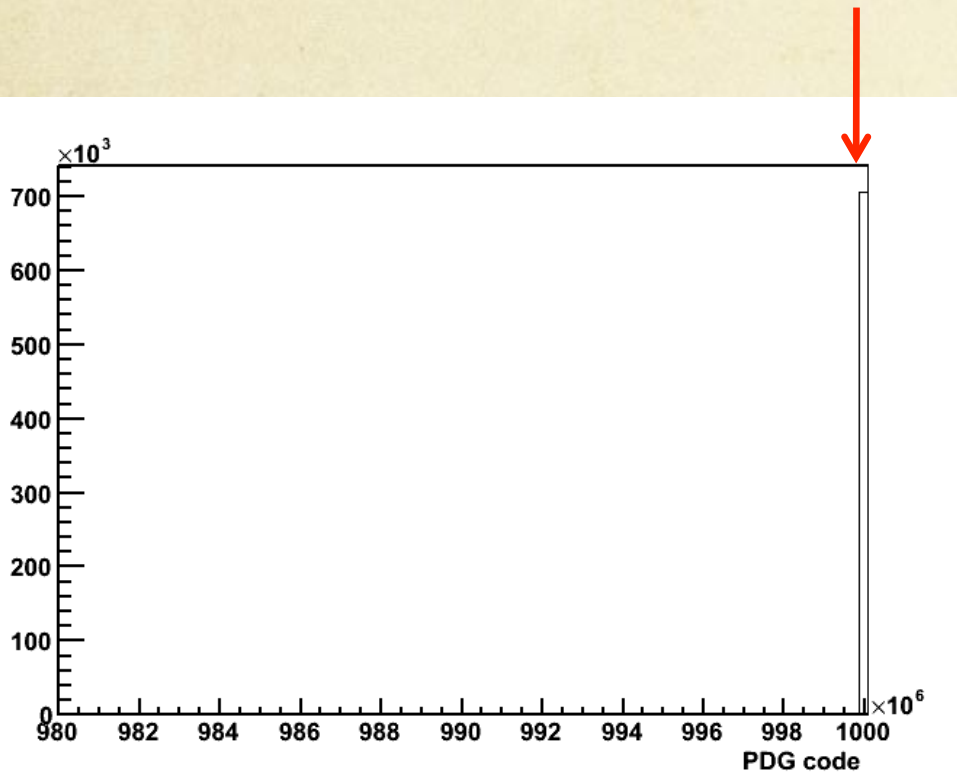


Particle composition of the IFR background



For Bhabha, Touschek Pair, BeamGas events the particle crossing the IFR are photons, electron, protons, neutrons and heavy nuclei

Carbon ion



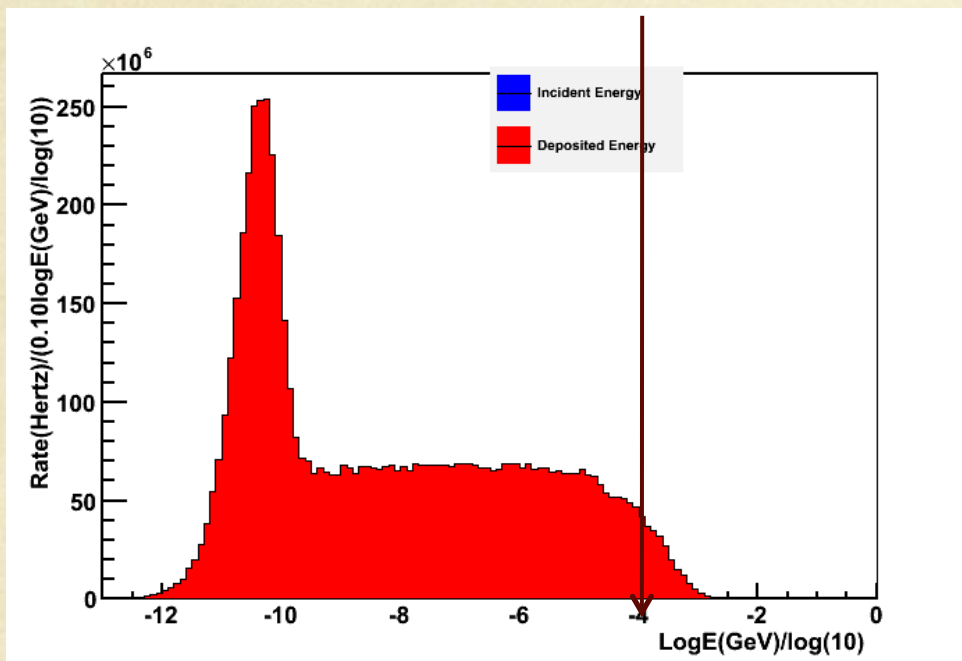
We have a higher number of protons than neutrons

Where are these protons from? (1)

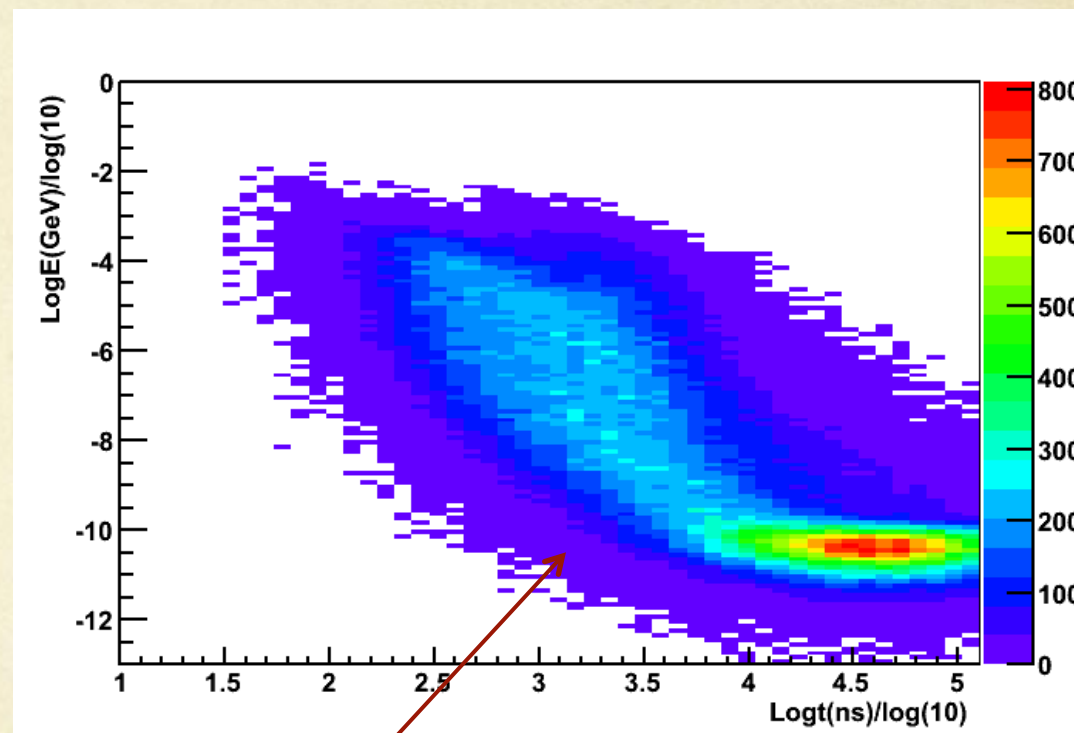


- These protons are not present in the boundaries \rightarrow they are produced inside the IFR

Proton Energy



Proton Energy vs time distributions

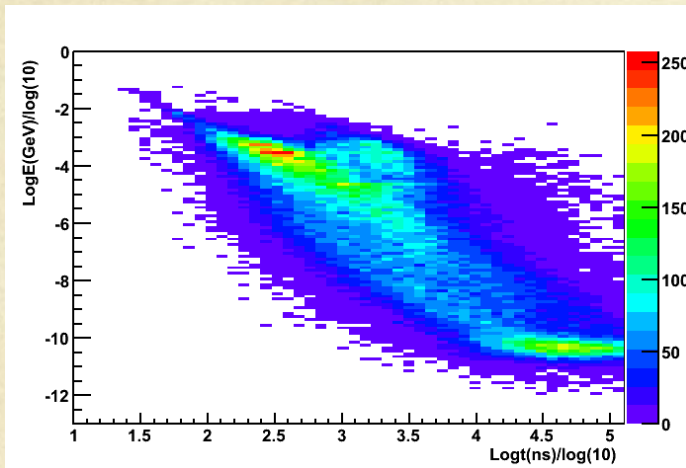


Even if the energy of the proton is very low there is a small fraction of protons that can have energy enough to be considered in the range of charge particle detected in the IFR

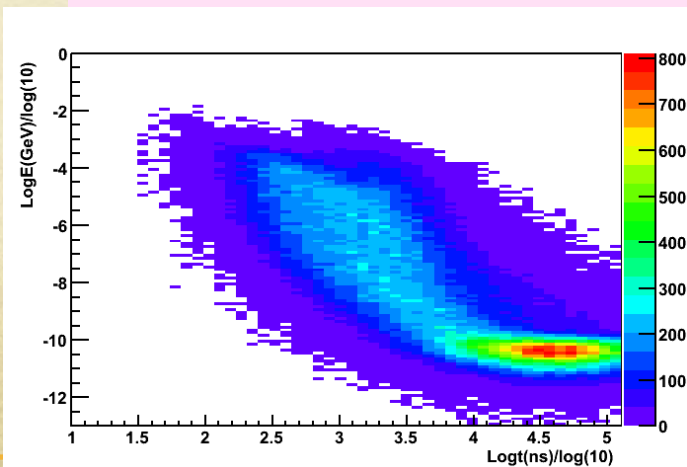
They time evolution of these protons is very peculiar

- The time evolution of the protons remained that one of the neutrons

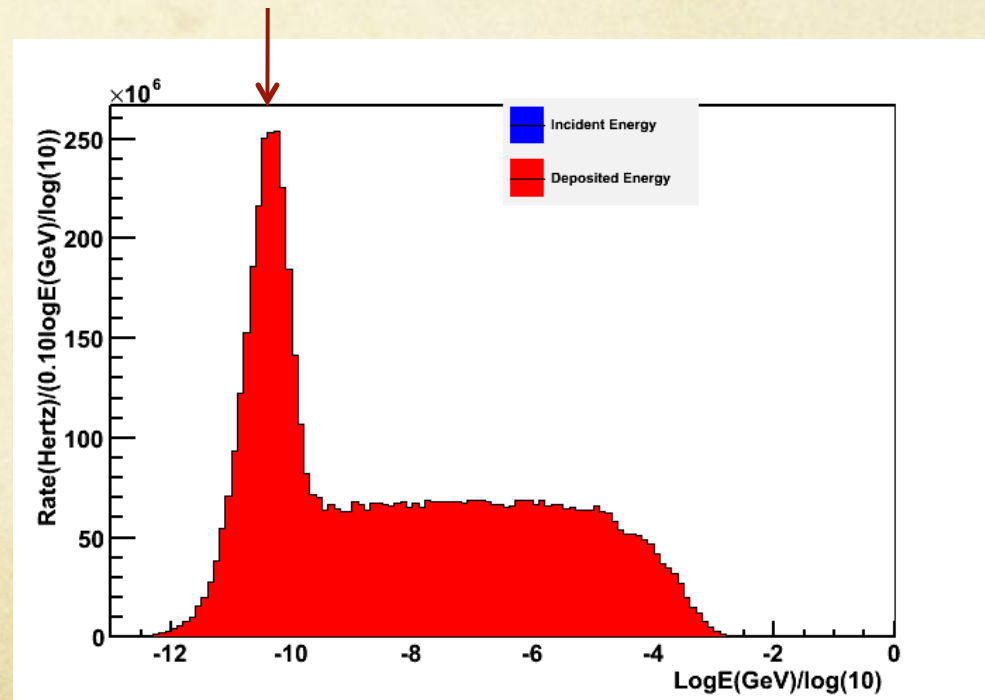
Neutron Energy vs time distributions



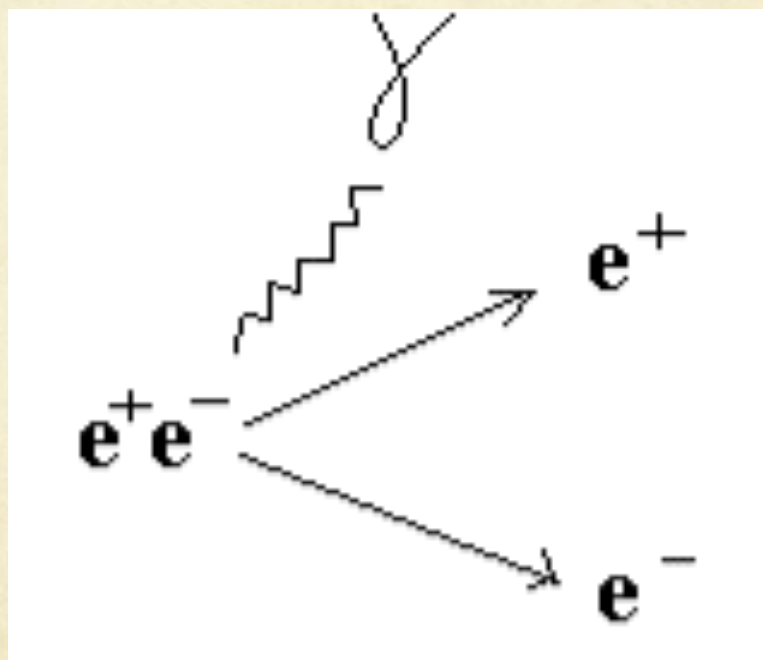
Proton Energy vs time distributions



This means that they are produced by the neutrons through the following process (n,p) in which the Neutron is captured and a proton is emitted. The cross section for this process falls as $1/v$ so it is more likely to happen when the neutron has low energy. This is the reason for the big peak at low proton energy.



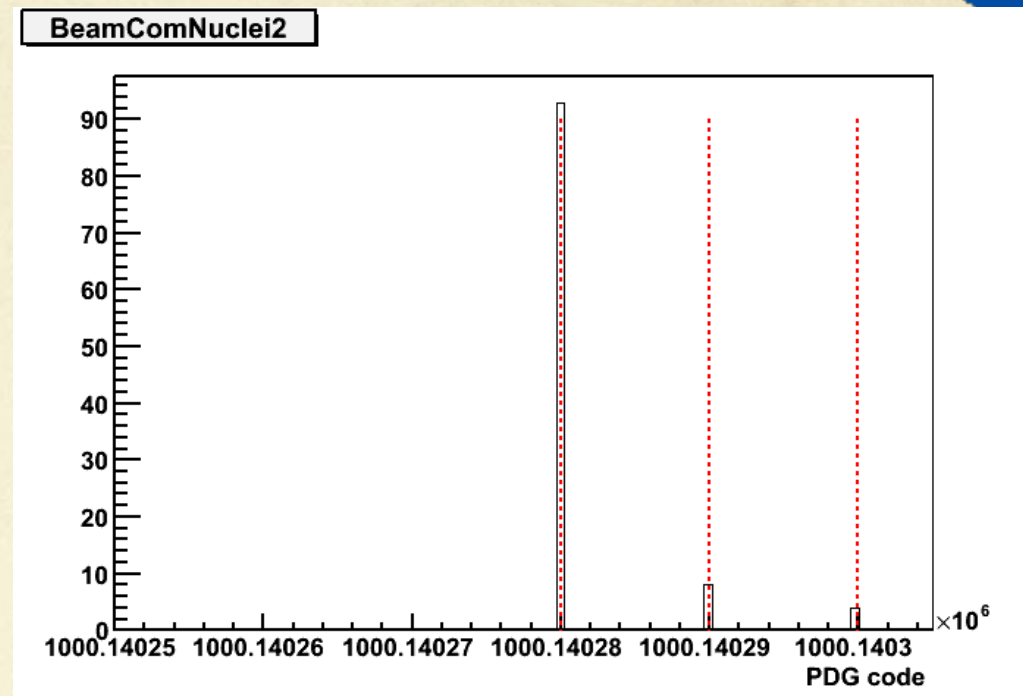
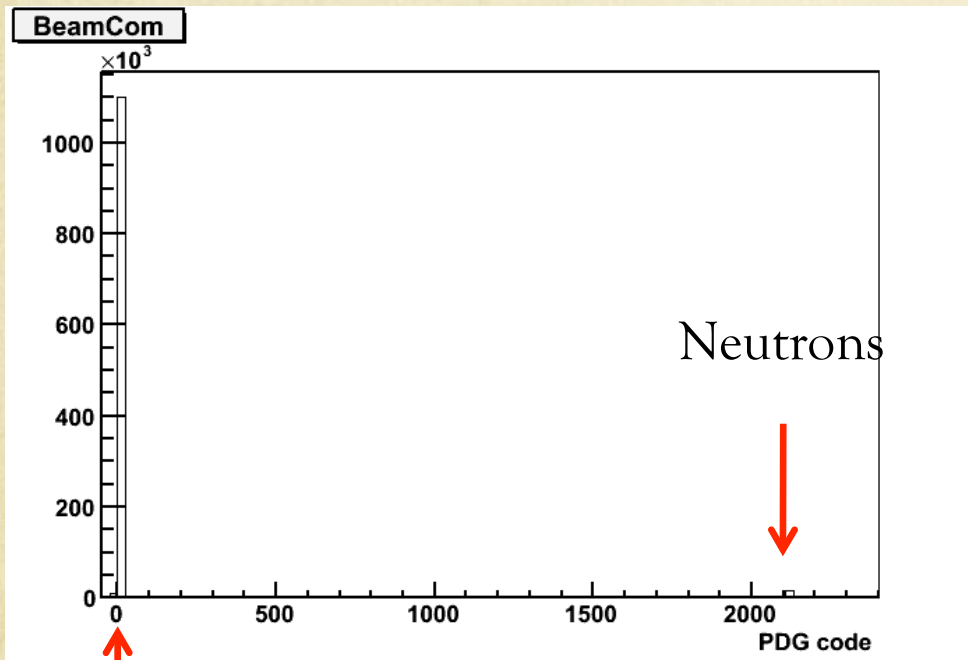
Radiative Bhabha background crossing the IFR FEE boards



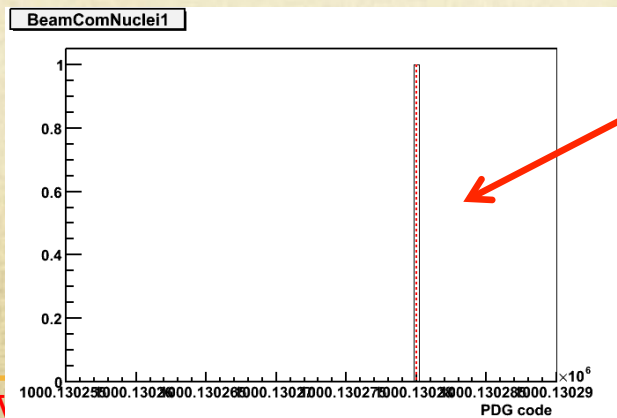
Beam Compositions for FEE electronics



- For Bhabha, Touschek and Pair events the particle crossing the FEE are photons, electron, protons, neutrons and heavy nuclei



Photons, electrons

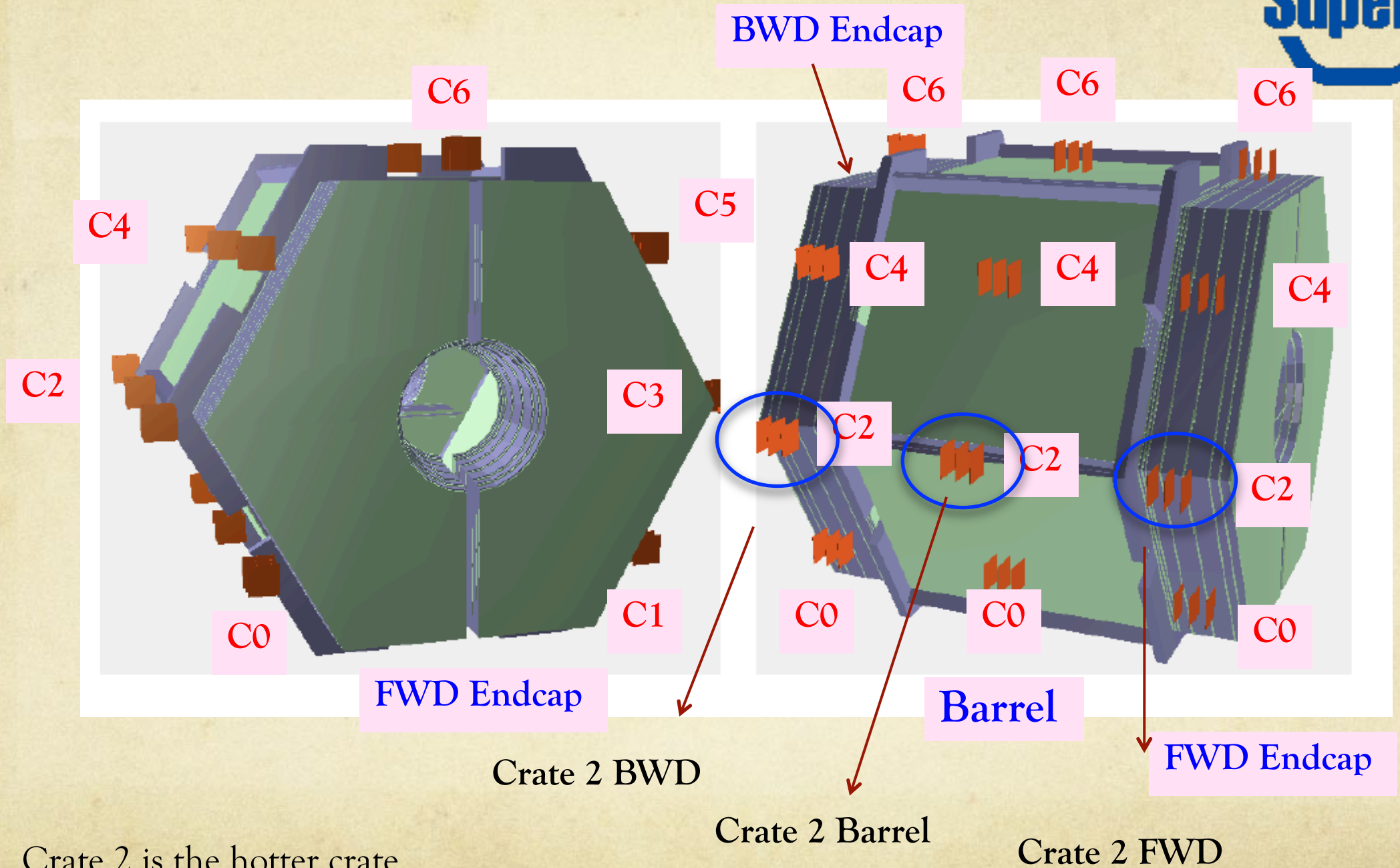


These 3 peaks are Si^{28} , Si^{29} , Si^{30} ions

These are the 3 isotopes that form the natural Silicon that we have in our FEEs

There are also Al^{28} this comes from the reaction $\text{Si}^{28}(n,p) \text{Al}^{28}$, Al^{28} is a radioactive aluminium isotope that decay to Si^{28}

Present layout of the IFR crates



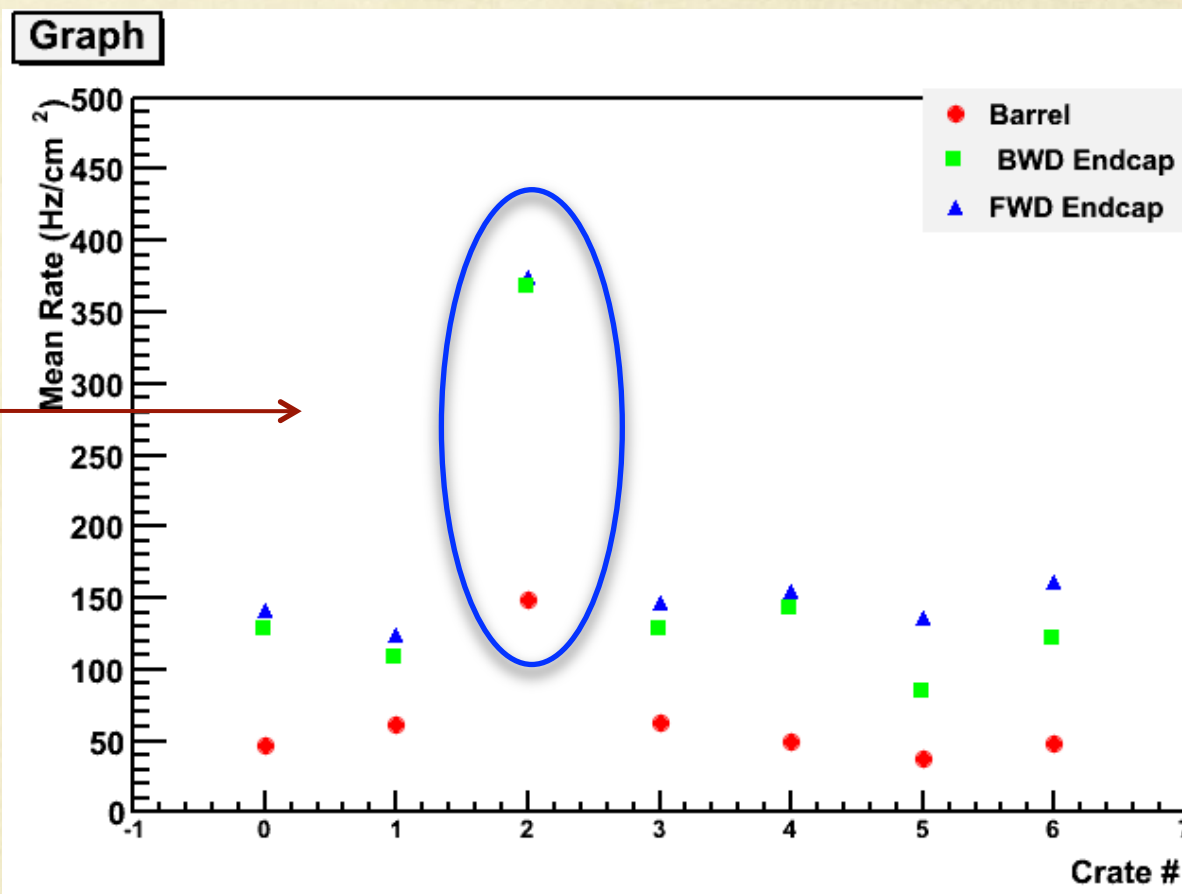
Crate 2 is the hotter crate

- Rate on electronics comparable to that one on the last IFR layer)

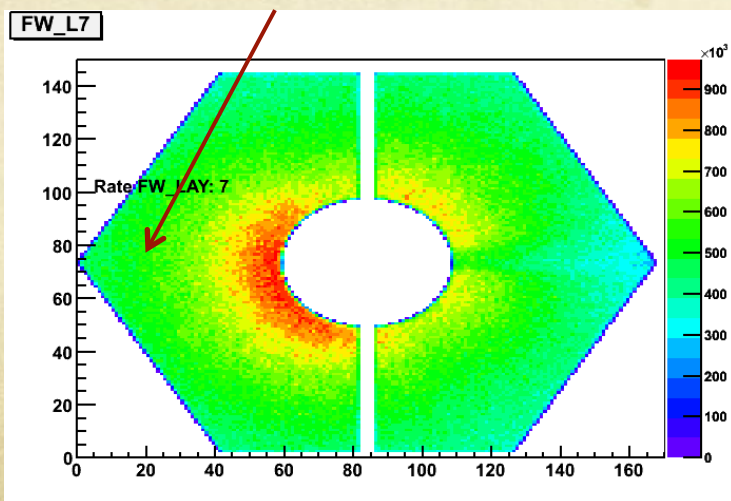
Crates located in the FWD have systematically higher rates compared to that one in the Barrel

Crate 2 very hot compared to other one the crate 2 is located on beam plane negative X

Mean Rate for each FEE in different Crates

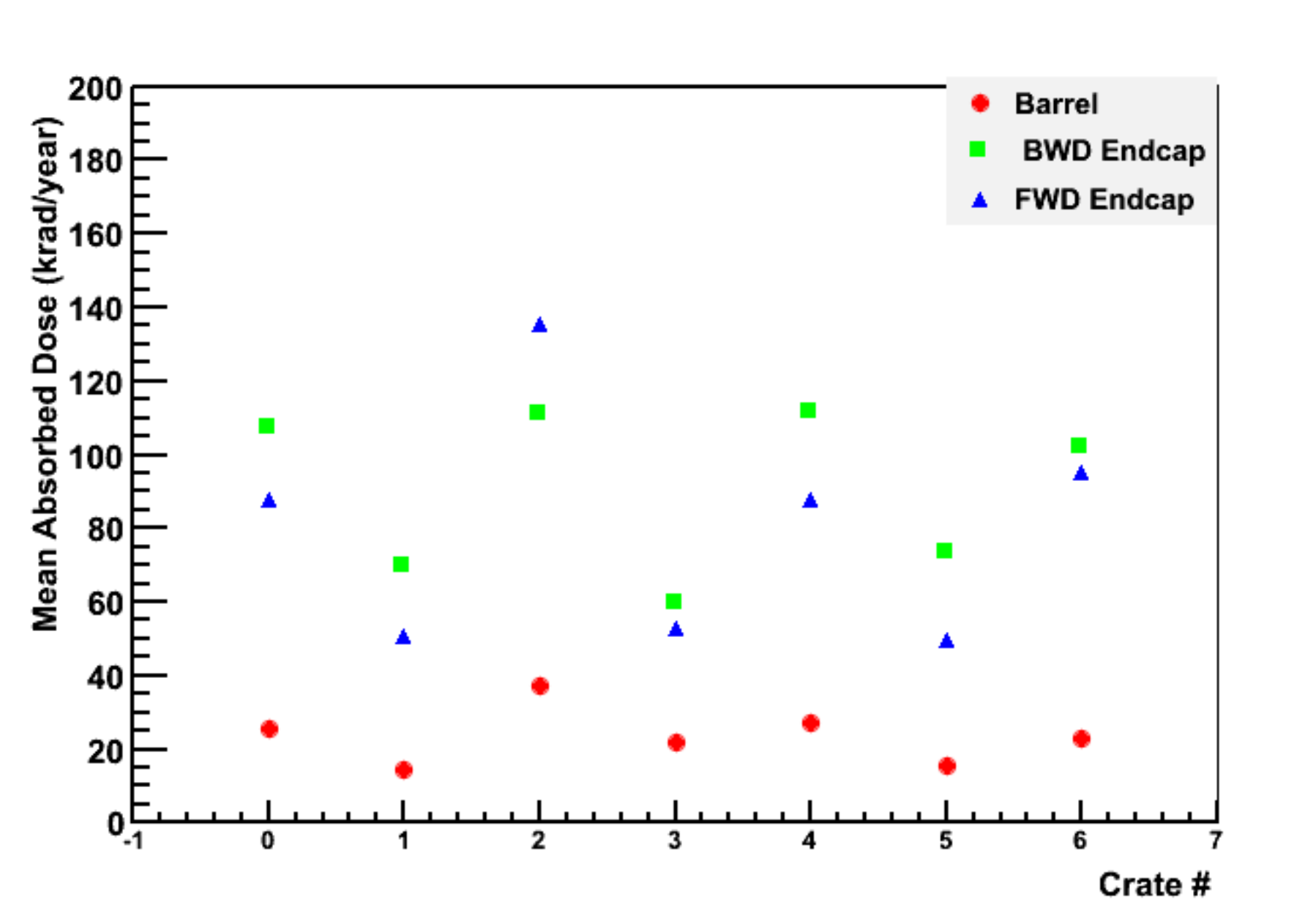


Even the higher rates x5 (safety factor) do not seem to be a problem for our FEEs





Absorbed Dose for each FEE Crates



The absorbed dose on the FEEs x5 (safety factor) does not seem to be a problem

Radiation Shielding



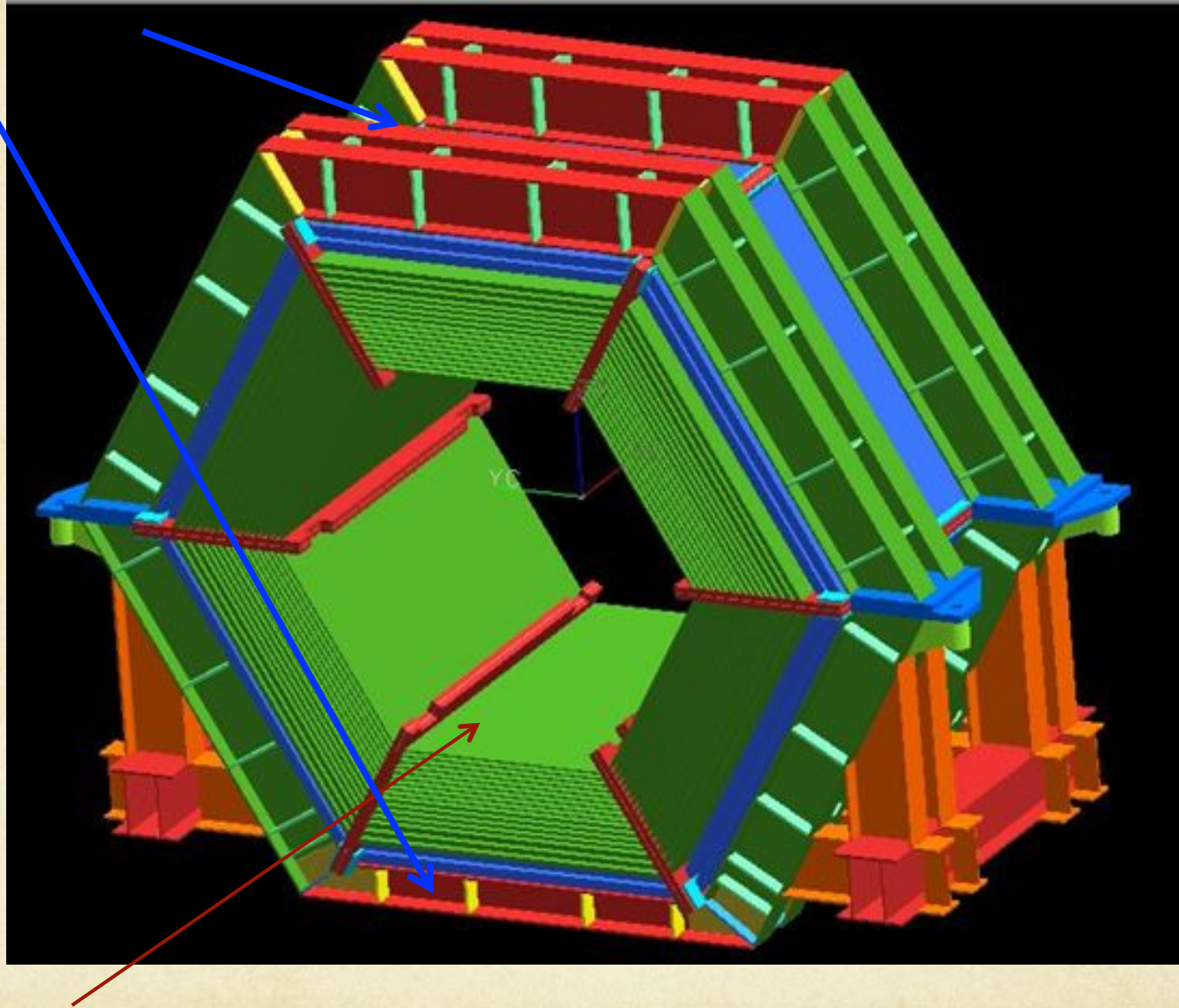


- We would like to implement some Boron-loaded polyethylene, shield for neutrons:
 - A shield between the IFR barrel and the magnet (50mm)
 - Add a shield between the EMC and magnet (21mm available)
 - Add a shield at small radius for the Endcap
- Add the IRON Structures around the IFR envelope (It will shield charged tracks and photons coming from the beam)

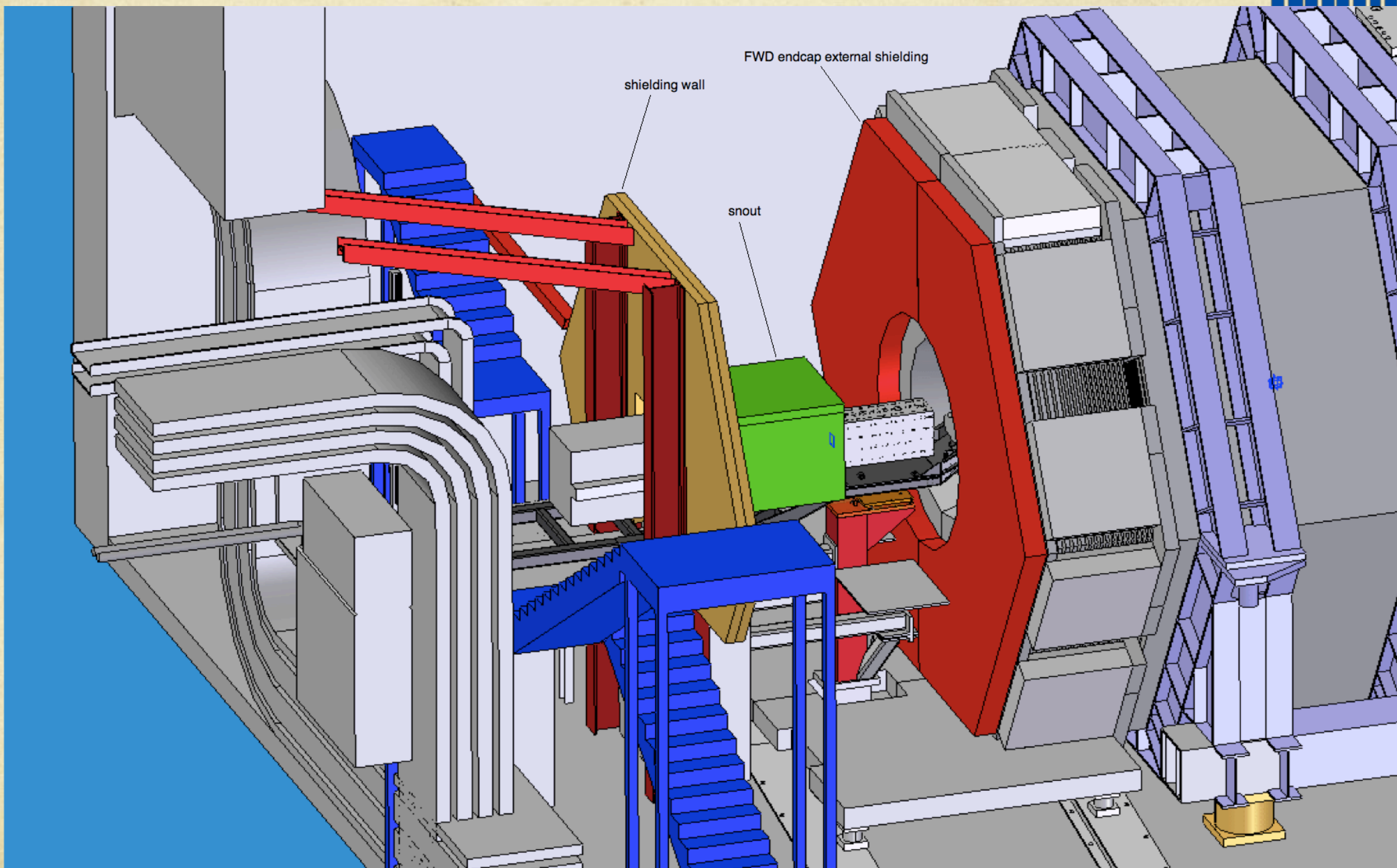
IFR-Shielding Strategy(2)



We have to implement these external structures



Shield between magnet and Barrel will go in that location





- ✓ Radiative Bhabha, Touschek, Pair and Beam-Gas backgrounds have been studied in details.
- ✓ The effect of these backgrounds have been also studied on our FEEs
- ✓ IFR TDR background on writing
- We have to add shielding between EMC and solenoid and between solenoid and IFR layer 0 to moderate neutrons
- We need to add the external iron structure for neutrons and charged tracks
- The effect of background on PID will be studied in details with FullSim
- All the numbers that you have seen in this presentation do not include the safety factor (x5) that must be included to have the final background estimation

BACK-UP SLIDES