# BiPo: A dedicated radiopurity detector for the SuperNEMO experiment

Héctor Gómez Maluenda on behalf of SuperNEMO Collaboration

Laboratoire de l'Accélérateur Linéaire

gomez@lal.in2p3.fr

- NEMO 3 and SuperNEMO experiments.
  - Motivation: SuperNEMO  $\beta\beta$  source foils radiopurity.
- The BiPo detector:
- BiPo 3 present status.
- Outlook and Prospects.
- Summary and Conclusions.



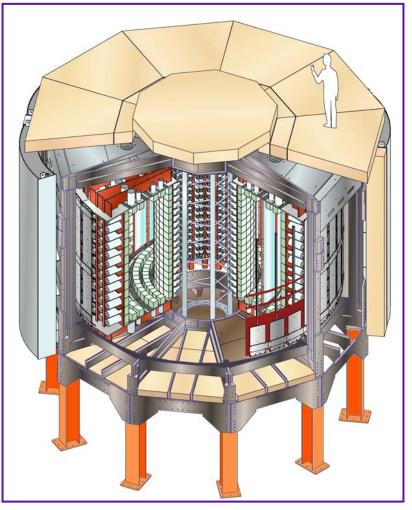


## NEMO 3 and SuperNEMO



#### NEMO 3

- 10 kg of ββ isotopes:
  - 6.9 kg of <sup>100</sup>Mo, 0.9 kg of <sup>82</sup>Se
  - Also <sup>130</sup>Te, <sup>116</sup>Cd, <sup>150</sup>Nd, <sup>96</sup>Zr and <sup>48</sup>Ca
- "Tracko calo" detection:
  - Tracking  $\rightarrow$  Geiger cells
  - Calorimetry → Polystyren + 3" & 5" PMTs
- Located at Modane Underground Laboratory
  - Depth ~4800 mwe
- Running from February 2003 until January 2011.
- Results on  $T_{1/2}^{2\nu\beta\beta}$  and  $T_{1/2}^{0\nu\beta\beta}$  for different isotopes.
- Complete background characterization.

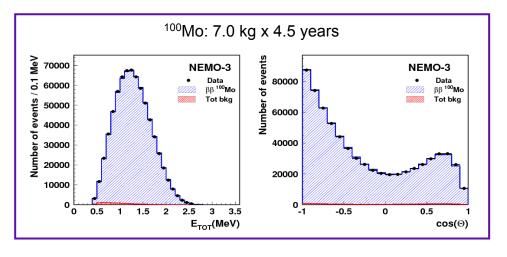




### NEMO 3 and SuperNEMO



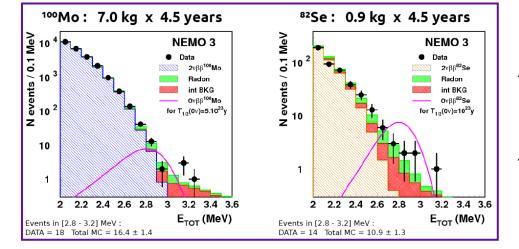
#### **NEMO 3 - Results**



#### 2νββ [Preliminary]

 $T_{1/2}^{2\nu\beta\beta}(^{100}Mo) = (7.16 \pm 0.01 \text{ (stat.)} \pm 0.54 \text{ (sys.)}) \ 10^{18} \text{ y}.$ 

 $T_{1/2}^{2\nu\beta\beta}(^{82}Se) = (9.6 \pm 0.1 \text{ (stat.)} \pm 1.0 \text{ (sys.)}) \ 10^{19} \text{ y}.$ 



#### Ονββ

$$\begin{split} T_{1/2}^{0\nu\beta\beta}(^{100}\text{Mo}, <m_{\nu}>) > 1.0 \ 10^{24} \ \text{y} \ (90\% \ \text{C.L.}). \\ <m_{\nu}> < 310 - 960 \ meV \\ T_{1/2}^{0\nu\beta\beta}(^{82}\text{Se}, <m_{\nu}>) > 3.2 \ 10^{23} \ \text{y} \ (90\% \ \text{C.L.}). \end{split}$$

<*m*<sub>v</sub>> < 940 – 2600 *meV* 



## NEMO 3 and SuperNEMO



-

NEMO 3

 $^{100}\text{Mo},\,^{82}\text{Se}$  and others

**7 kg** 18 %

15 % FWHM @ 1 MeV

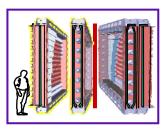
- ~ 100 µBq/kg
- < 300 µBq/kg
- $\sim 5 \text{ mBq/m}^3$

1-2 10<sup>24</sup> y

 $< m_{v} > < 0.3 - 0.9 \text{ eV}$ 

From NEMO 3 to SuperNEMO

Isotope Mass Efficiency Energy Resolution <sup>208</sup>TI source radiopurity <sup>214</sup>Bi source radiopurity Rn level Sensitivity



**SuperNEMO** 

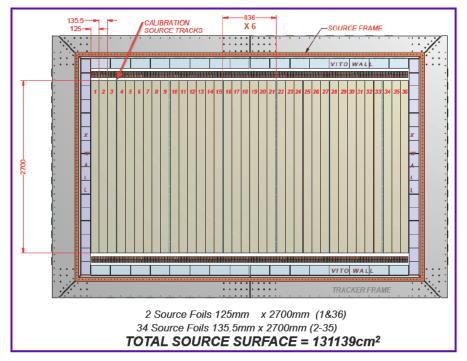
<sup>82</sup>Se (<sup>150</sup>Nd or <sup>48</sup>Ca?) ~100 kg 30 %
7 % FWHM @ 1 MeV
< 2 μBq/kg</p>
< 10 μBq/kg</p>
< 0.1 mBq/m<sup>3</sup>
1 10<sup>26</sup> y
<m,> < 0.04 – 0.1 eV</p>





- SuperNEMO experiment will have ~100 kg of  $\beta\beta$  emitter distributed in the so-called source foils.
- Sensitivity requirements limits the source foil contamination to:
  - < 2 μ**Bq/kg** in <sup>208</sup>TI
  - < 10 μBq/kg in <sup>214</sup>Bi
- Most sensitive HPGe detectors ~50  $\mu$ Bq/kg for <sup>208</sup>Tl.
- In addition:
  - Better to measure the source itself (particular geometry)
  - Non destructive measurement procedure

NECESSITY OF A DEDICATED DETECTOR

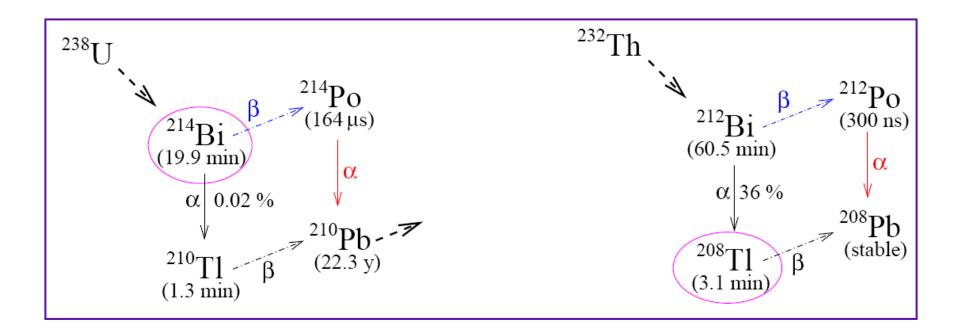




#### The BiPo detector: Detection Principle



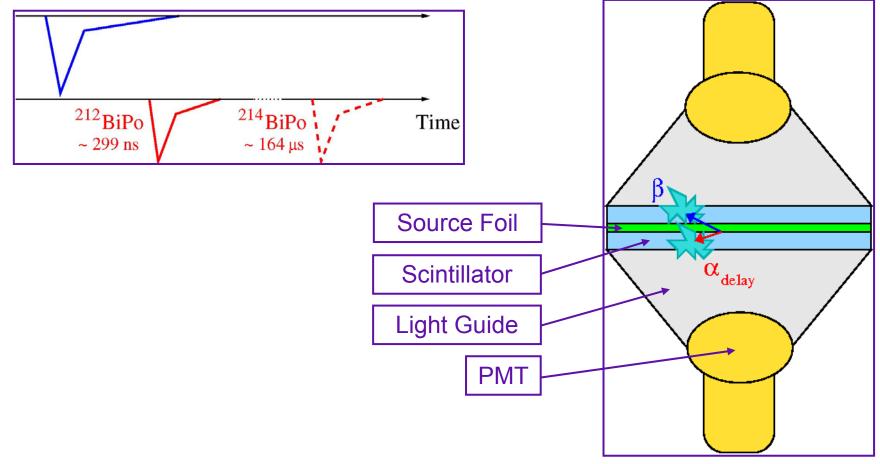
- BiPo is a dedicated detector for the measurements of ultra-low levels of contamination in <sup>208</sup>Tl and <sup>214</sup>Bi present in the SuperNEMO source foils.
  - Detection principle  $\rightarrow$  BiPo  $\beta \alpha$  delayed coincidence detection.







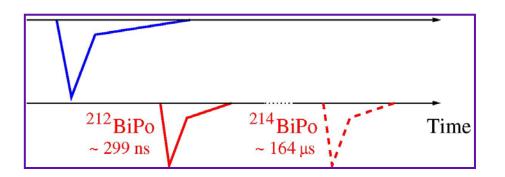
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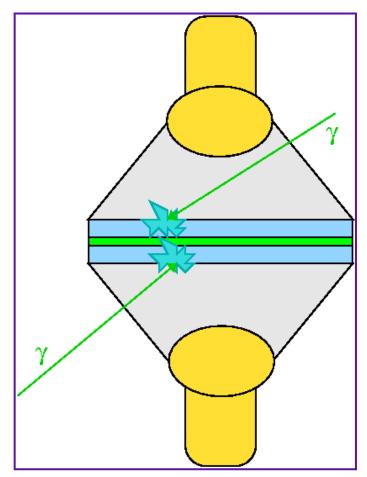




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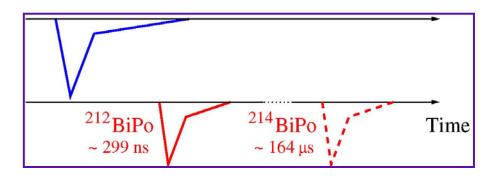
- Background sources:
  - γ-induced random coincidences



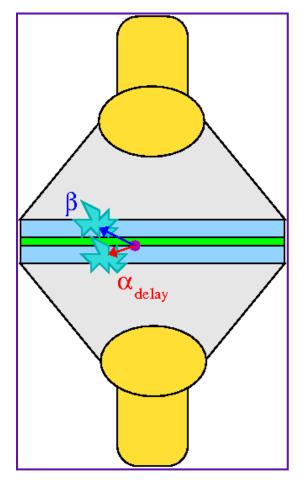




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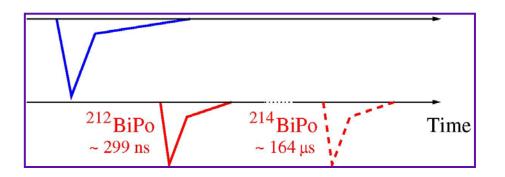
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  - Scintillator surface contamination
  - Radon contamination in the sensitive volume



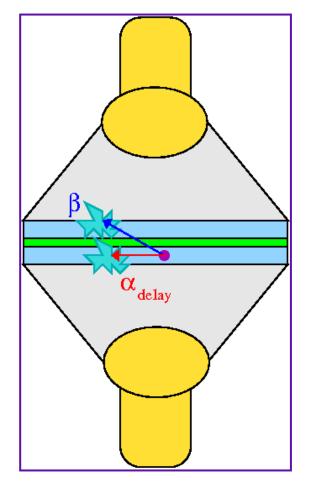




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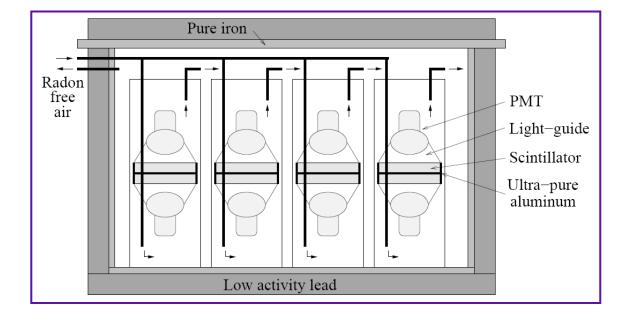
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  - Scintillator bulk contamination







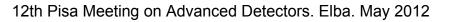
- BiPo prototypes have been really useful to fix different open questions and to validate the technology.
  - **BiPo 1**:
    - 20 modules
    - Face to face scintillators
    - Installed at LSM
  - DAQ tests
    - MatAcq digitizer board
    - Trigger modules
  - Analysis methods tests
  - Radiopurity of the scintillators



Detection principle tested with a calibrated *AI* foil between the scintillators

<sup>208</sup>*TI* contamination level of **1.5**  $\mu$ *Bq/m*<sup>2</sup> of scintillator measured

NIM A 622 (2010) 120-128





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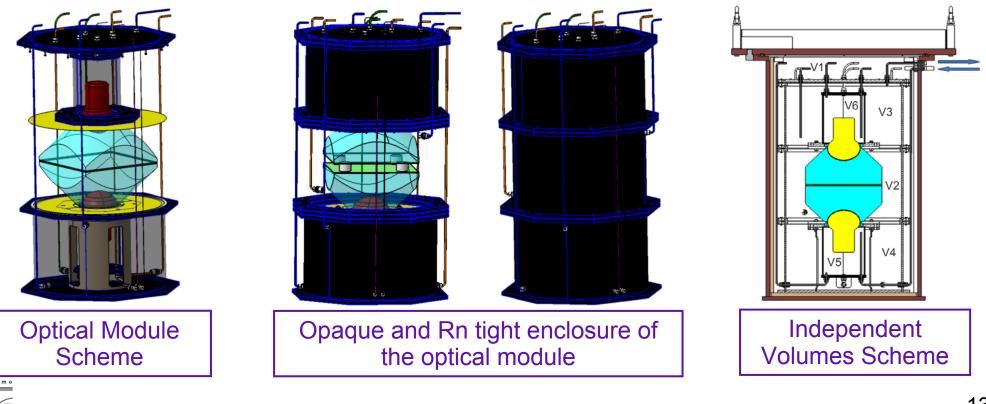


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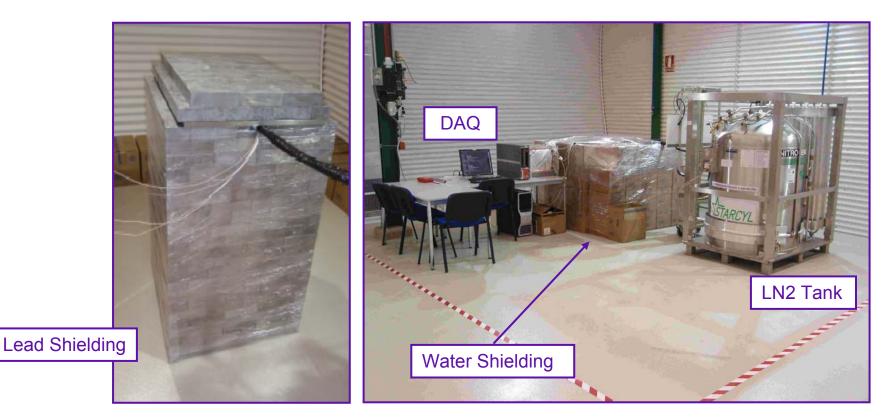


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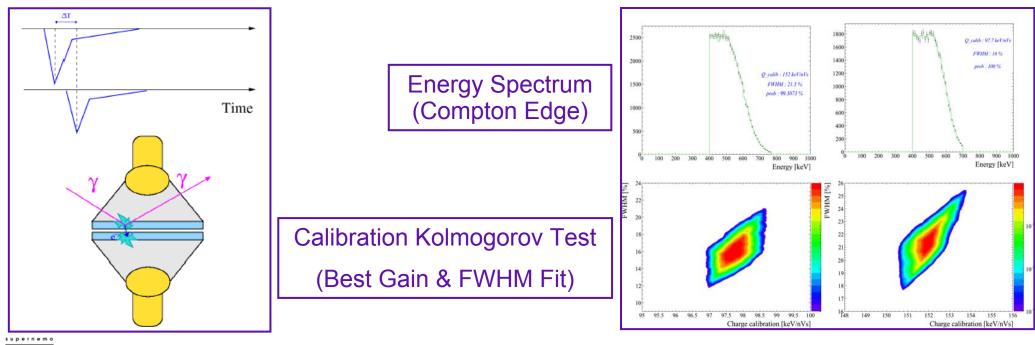




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#### → How low can we go in terms of Rn level and Counting Rate?

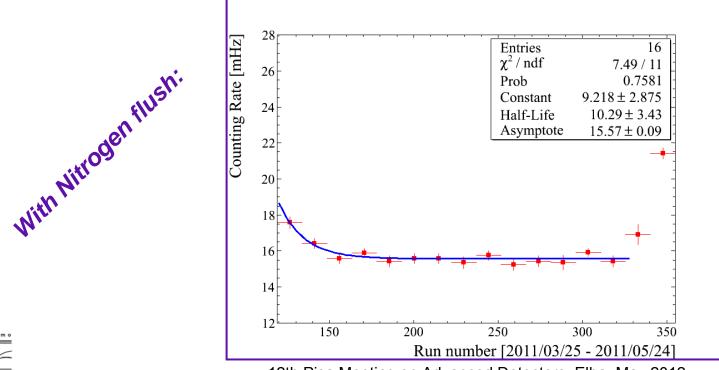
- Calibrations (<sup>54</sup>Mn and <sup>22</sup>Na) for energy, resolution and timing
- Counting Rate monitoring
- BiPo events analysis



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• Quite stable

Dependance with Rn concentration



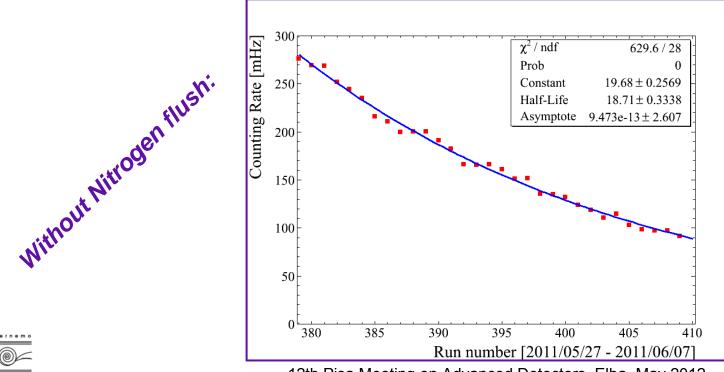




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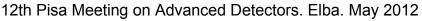
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- Decay rate
- $T_{1/2} \sim 4.7 \pm 0.1$  days
- Rn detection





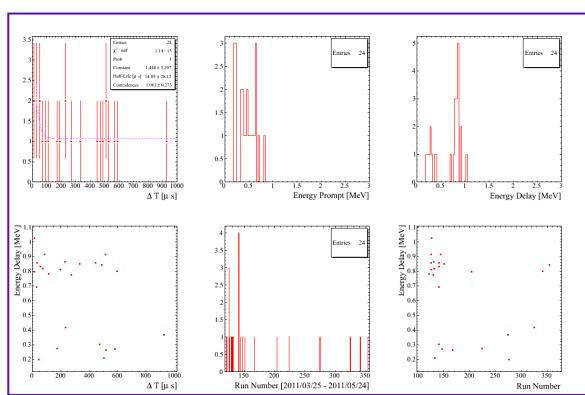




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- Data from one candidate:
  - •Prompt energy (e<sup>-</sup>)
  - •Delay energy ( $\alpha$ )
  - •∆t
- Rate evolution
  - Assure is constant







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• *Results* after ~60 days data taking:

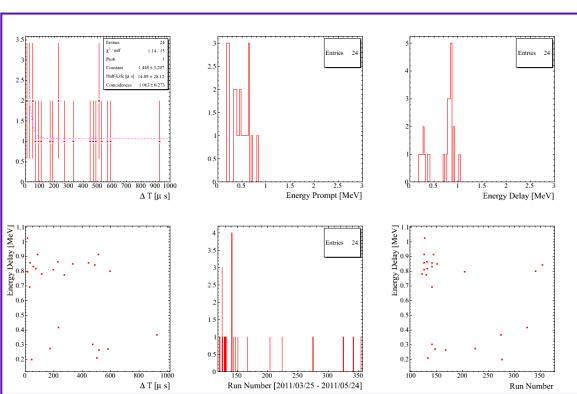
<sup>208</sup>TI:

```
A (^{208}TI) < 5.45 \muBq/m<sup>2</sup> scint @ 90% C.L.
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<sup>214</sup>Bi:

```
0.6 < A (^{214}Bi) = 5.3 < 23 \ \mu Bq/m^2 \ scint
```

**PROMISING RESULTS** 

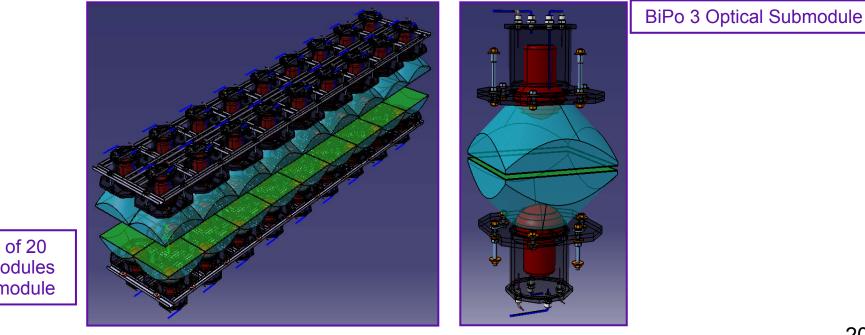




## Final setup: BiPo 3 detector



- Prototypes have shown that required sensitivity levels could be reachable with the face to face detector approximation.
- Main features:
  - 2 modules detector  $\rightarrow$  Possibility to measure 8 SuperNEMO source foils simultaneously
  - 20 30x30 cm<sup>2</sup> optical sub-modules in each module
    - 3.6 m<sup>2</sup> of sensitive surface
  - 2 mm thick scintillator plates
  - Light guide geometry optimized
  - Volume separation and nitrogen flushing for Rn suppression and external shielding



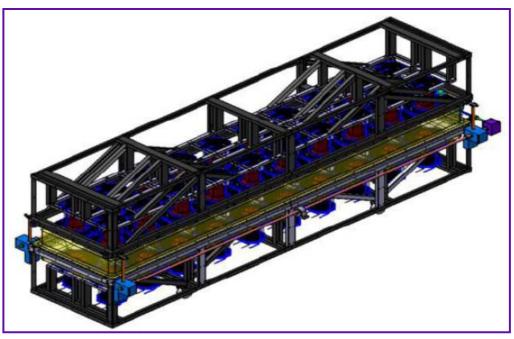
Assembling of 20 optical submodules  $\rightarrow$  1 BiPo 3 module



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BiPo 3 detectors installed in the internal structure

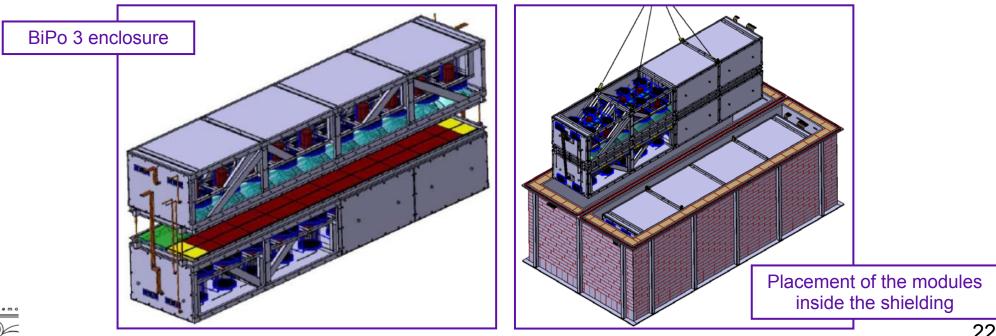


12th Pisa Meeting on Advanced Detectors. Elba. May 2012

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• Almost the 80 optical submodules assembled and characterized in a test bench @ LAL.

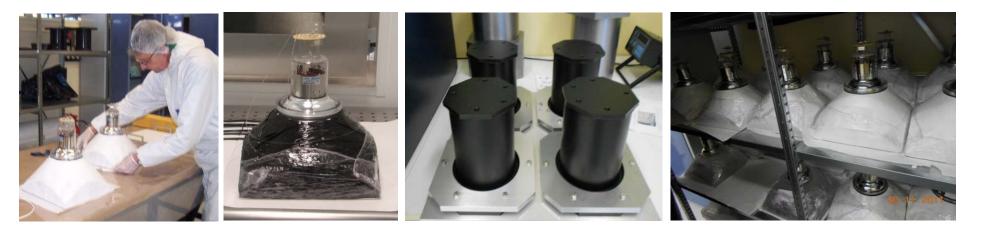


Light guide annealing

Gluing of the scintillator

Aluminization

PMT gluing



Tyvek wrapping

Black film

Light black box

Characterized sub-modules storage

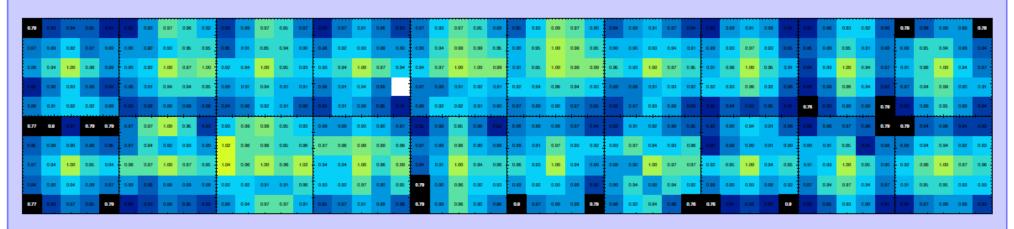




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	1.02	1.00	0.9	17		0.99	0.94	1.00	0.92	0.85	0.		0.94	1.00	0.96	0.97	0.98	0.96	1.00	0.99	1.0	6 0.5	r 0.	97 1.	00 0.	99 D.	93 D.	<b>as</b> 0.	94 1.	00 0.5	e 1.	2 0.	0 0.0	1.00	0.92	0.92	0.99	0.92	1.00	0.94	0.92	0.92	0.95	1.00	0.97	0.95	0.91	0.95	1.00	0.95
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	0.95	0.92	0.9	2	.70	8.0		0.82		0.7	0.		0.94	0.99	0.99	0.93	0.83	0.96	0.90	0.96	0.0	0 0.5	<b>a</b> 0.	96 0.	aa o	95 0	91 D.	<b>as</b> 0.	82 Q.I	<b>IS</b> 0.5	2 0.1	a a	0.0	0.9	0.00	0.05	0.04	0.90	0.99	0.99	0.00	0.06	0.91	0.92	0.99	0.79	0.01	0.86	0.91	0.96
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	0.94	1.00	0.5	6		0.94	0.96	1.00	0.90	0.92	0.1	x0 0	9.94	1.00	0.99	1.02	0.98	0.90	1.00	0.99	0.9	4 03	e 0.	96 1.	00 0.	95 0.	90 O	<b>9</b> 5 0.1	25 1.	80 1.5	o 0.	H 01	86 1.0	0 1.0	0.90	0.92	0.96	0.97	1.00	0.99	0.94	98.0	0.92	1.00	0.96	0.99	0.95	1.00	1.00	0.97
	39.0	0.91			1.74	0.07	0.90	0.94	0.85	0.00		12 0	0.91	0.94	0.95	1.02	0.92	0.22	0.97	0.94	0.0	0 0.5	<b>0</b> 1)	00 1.	0 00	99 O.	96 0.	94 03	90 Q.	a7 0.5	6 0.		н о.9	6 0.92	0.07	0.05	0.92	0.94	0.94	0.92	0.90	10.06	0.95	0.94	0.94	0.96	1.10	1.24	0.94	0.99

#### Bottom Part of BiPo3 Module





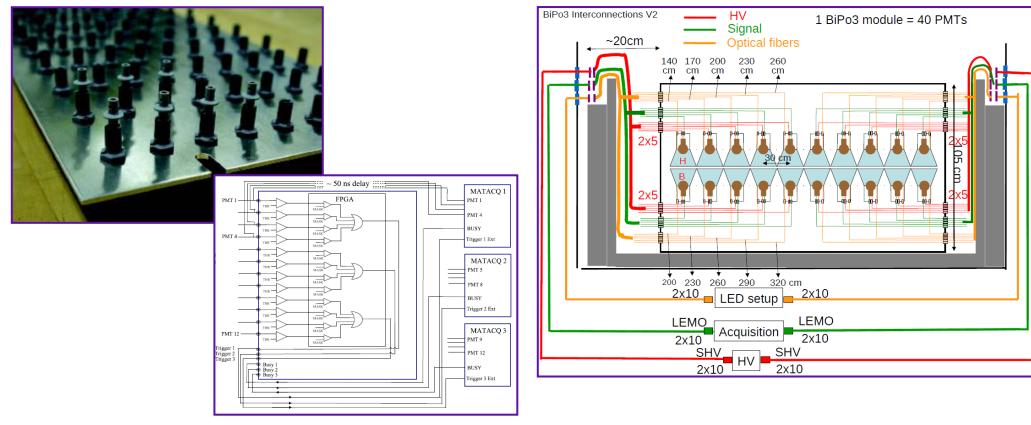


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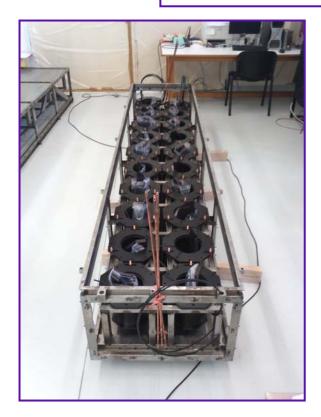
#### ALMOST READY FOR THE COMMISSIONING OF THE FIRST MODULE.





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#### FIRST ASSEMBLING TESTS @ LAL CLEAN ROOM











- Approximate timeline:
  - Module 1:
  - All the required material for the first BiPo 3 module will be tested at LAL before mid June
  - The material will be delivered to Canfranc Underground Laboratory last week of June
  - Installation of the first BiPo 3 module between 2 and 13 of July
  - Commissioning of the detector during summer
  - Calibrations and background measurement: Until end 2012
  - Source measurement: Beginning 2013
  - Module 2:
  - Same schedule that for module 1 with ~3 months of delay

In March 2013 the whole detector will be completely ready, being possible to make source measurements since beginning 2013



#### **Summary and Conclusions**



- **NEMO 3** experiment has obtained really good results in the measurement of  $T_{1/2}^{2\nu\beta\beta}$  and also upper limits of  $T_{1/2}^{0\nu\beta\beta}$  for <sup>82</sup>Se and <sup>100</sup>Mo.
- SuperNEMO experiment will use the same technique with 100 kg of  $\beta\beta$  emitter trying to reach the  $\langle m_{\nu} \rangle \sim 50$  meV level.
- One requirement for SuperNEMO is to have really good *radiopurity* levels for the *source foils*: < 2 μBq/kg in <sup>208</sup>TI and < 10 μBq/kg in <sup>214</sup>Bi.
- **BiPo** is a **dedicated** detector designed to measure these foils with the required sensitivity.
- A set of *prototypes* have showed the *viability* to construct the detector proving the *detection principle*.
- **Results** of these prototypes lead to think that required **sensitivity** will be **reached**.
- *Final setup* is almost *completed* and source foil measurements are expected to *start beginning 2013*.



# BiPo: A dedicated radiopurity detector for the SuperNEMO experiment

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Laboratoire de l'Accélérateur Linéaire

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- The BiPo detector:
- BiPo 3 present status.
- Outlook and Prospects.
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