

BiPo: A dedicated radiopurity detector for the SuperNEMO experiment

Héctor Gómez Maluenda on behalf of SuperNEMO Collaboration

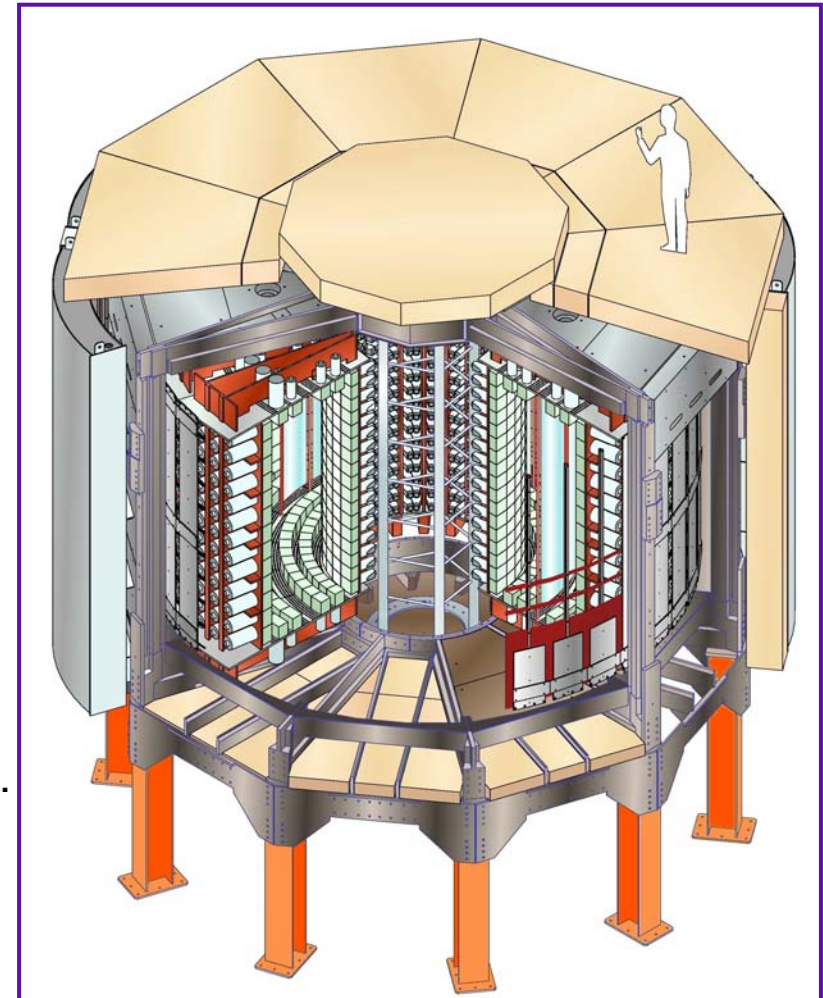
Laboratoire de l'Accélérateur Linéaire

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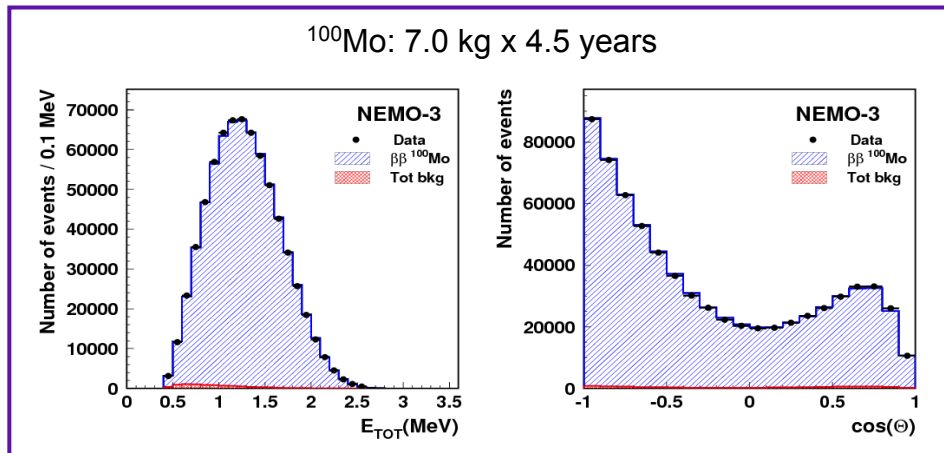
- NEMO 3 and SuperNEMO experiments.
 - Motivation: SuperNEMO $\beta\beta$ source foils radiopurity.
- The BiPo detector:
- BiPo 3 present status.
- Outlook and Prospects.
- Summary and Conclusions.

NEMO 3

- **10 kg** of $\beta\beta$ isotopes:
 - 6.9 kg of ^{100}Mo , 0.9 kg of ^{82}Se
 - Also ^{130}Te , ^{116}Cd , ^{150}Nd , ^{96}Zr and ^{48}Ca
- “Tracko – calo” detection:
 - Tracking \rightarrow Geiger cells
 - Calorimetry \rightarrow Polystyren + 3” & 5” PMTs
- Located at Modane Underground Laboratory
 - Depth ~ 4800 mwe
- Running from February 2003 until January 2011.
- Results on $T_{1/2}^{2\nu\beta\beta}$ and $T_{1/2}^{0\nu\beta\beta}$ for different isotopes.
- Complete background characterization.



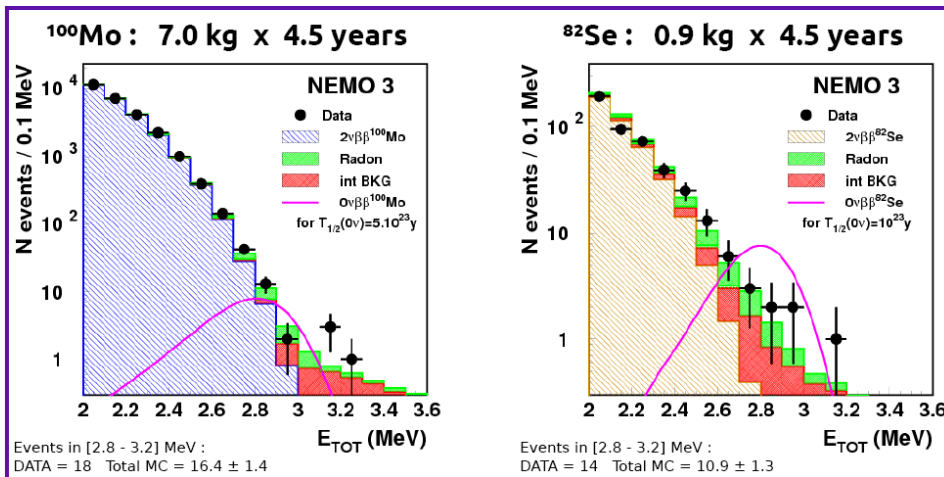
NEMO 3 - Results



$2\nu\beta\beta$ [Preliminary]

$$T_{1/2}^{2\nu\beta\beta}(^{100}\text{Mo}) = (7.16 \pm 0.01 \text{ (stat.)} \pm 0.54 \text{ (sys.)}) 10^{18} \text{ y.}$$

$$T_{1/2}^{2\nu\beta\beta}(^{82}\text{Se}) = (9.6 \pm 0.1 \text{ (stat.)} \pm 1.0 \text{ (sys.)}) 10^{19} \text{ y.}$$



$0\nu\beta\beta$

$$T_{1/2}^{0\nu\beta\beta}(^{100}\text{Mo}, \langle m_\nu \rangle) > 1.0 10^{24} \text{ y (90\% C.L.).}$$

$$\langle m_\nu \rangle < 310 - 960 \text{ meV}$$

$$T_{1/2}^{0\nu\beta\beta}(^{82}\text{Se}, \langle m_\nu \rangle) > 3.2 10^{23} \text{ y (90\% C.L.).}$$

$$\langle m_\nu \rangle < 940 - 2600 \text{ meV}$$

NEMO 3 and SuperNEMO



NEMO 3

^{100}Mo , ^{82}Se and others

7 kg

18 %

15 % FWHM @ 1 MeV

$\sim 100 \mu\text{Bq/kg}$

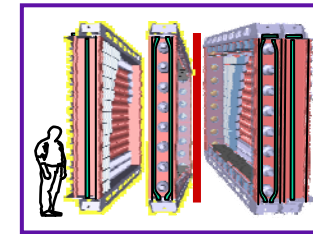
$< 300 \mu\text{Bq/kg}$

$\sim 5 \text{ mBq/m}^3$

$1\text{-}2 \cdot 10^{24} \text{ y}$

$\langle m_\nu \rangle < 0.3 - 0.9 \text{ eV}$

From NEMO 3 to SuperNEMO



SuperNEMO

^{82}Se (^{150}Nd or ^{48}Ca ?)

$\sim 100 \text{ kg}$

30 %

7 % FWHM @ 1 MeV

$< 2 \mu\text{Bq/kg}$

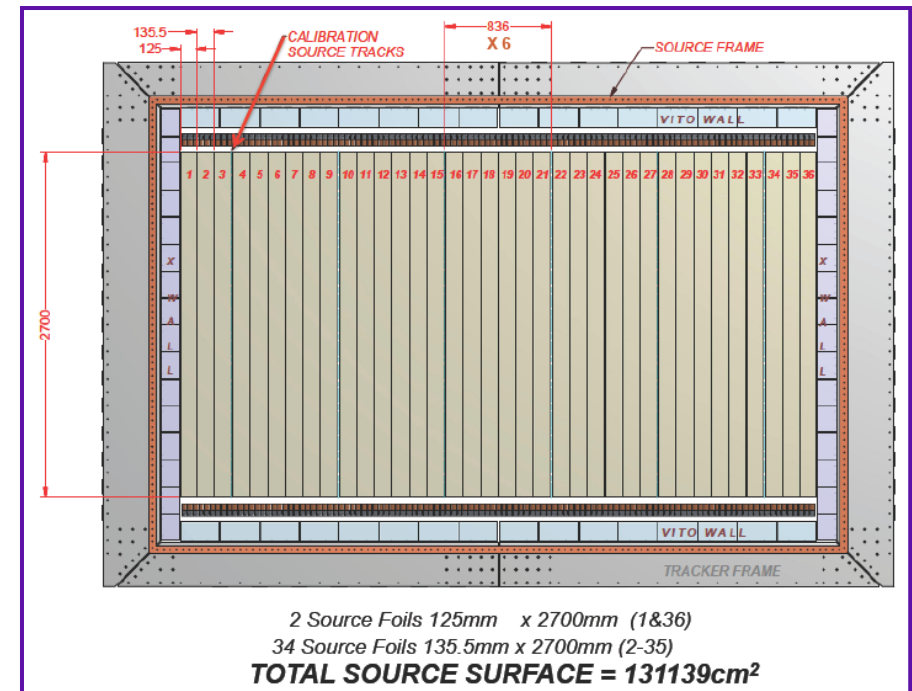
$< 10 \mu\text{Bq/kg}$

$< 0.1 \text{ mBq/m}^3$

$1 \cdot 10^{26} \text{ y}$

$\langle m_\nu \rangle < 0.04 - 0.1 \text{ eV}$

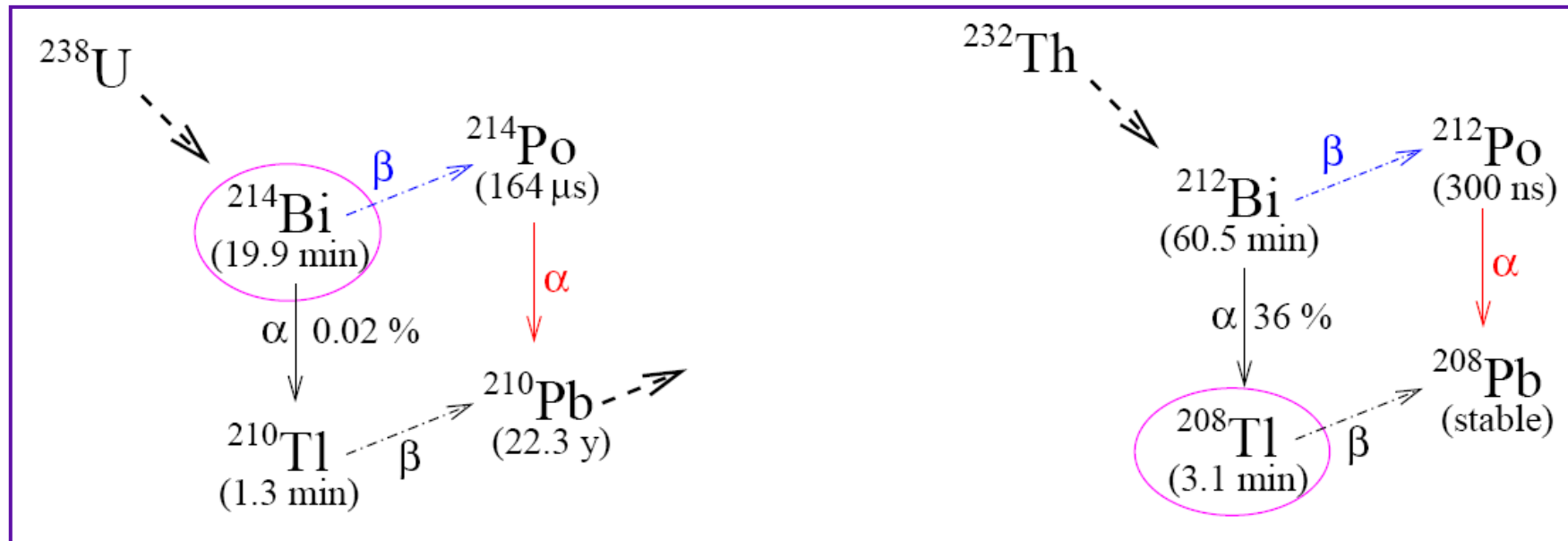
- SuperNEMO experiment will have ~ 100 kg of $\beta\beta$ emitter distributed in the so-called source foils.
- Sensitivity requirements limits the source foil contamination to:
 - $< 2 \mu\text{Bq/kg}$ in ^{208}Tl
 - $< 10 \mu\text{Bq/kg}$ in ^{214}Bi
- Most sensitive HPGe detectors $\sim 50 \mu\text{Bq/kg}$ for ^{208}Tl .
- In addition:
 - Better to measure the source itself (particular geometry)
 - Non destructive measurement procedure



**NECESSITY OF A DEDICATED
DETECTOR**

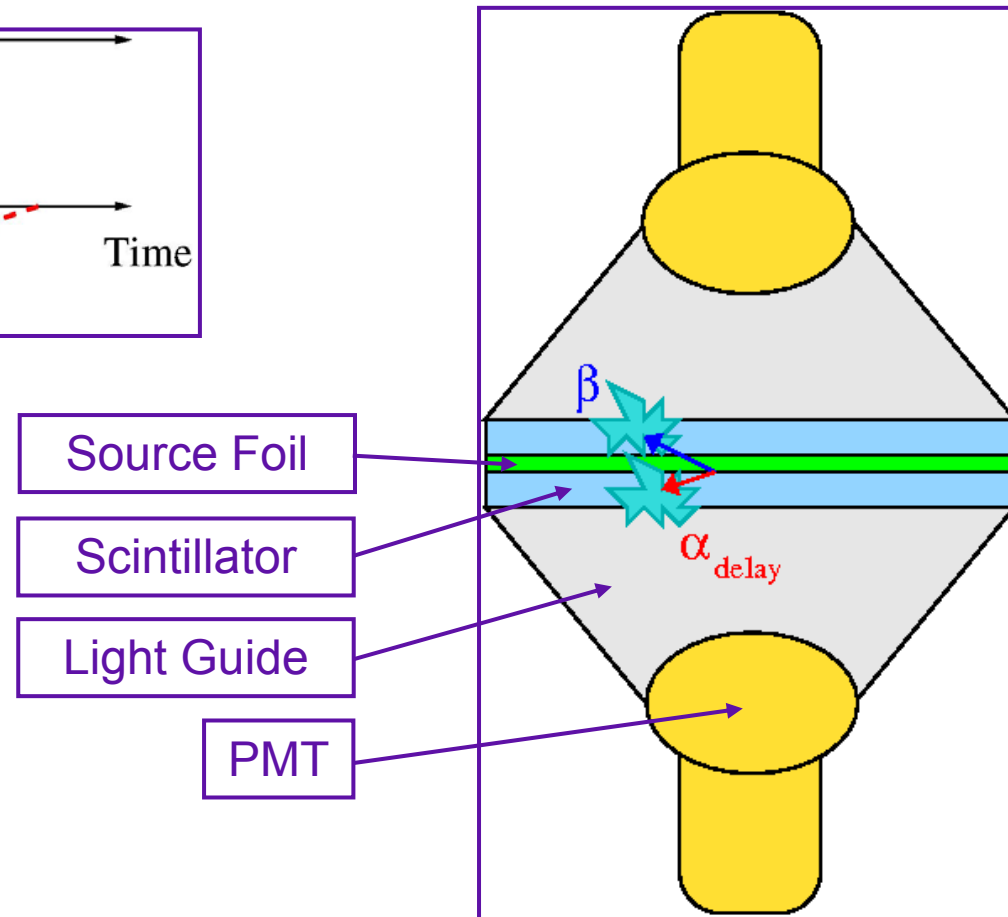
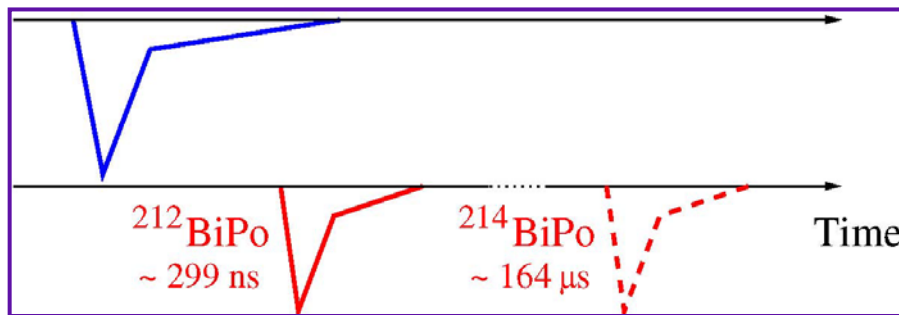
The BiPo detector: Detection Principle

- BiPo is a dedicated detector for the measurements of ultra-low levels of contamination in ^{208}Tl and ^{214}Bi present in the SuperNEMO source foils.
- Detection principle \rightarrow BiPo $\beta - \alpha$ delayed coincidence detection.



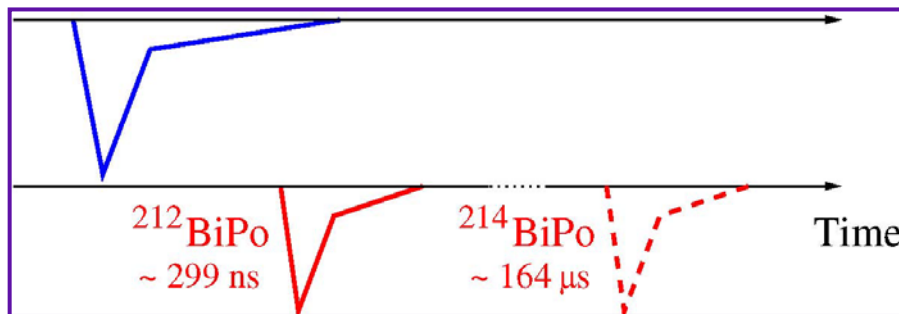
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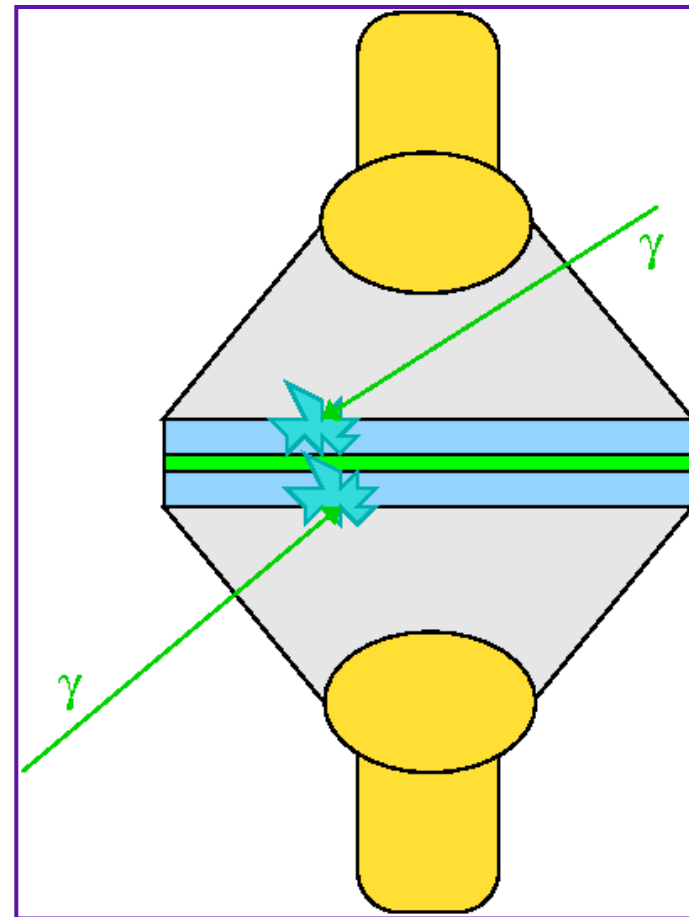


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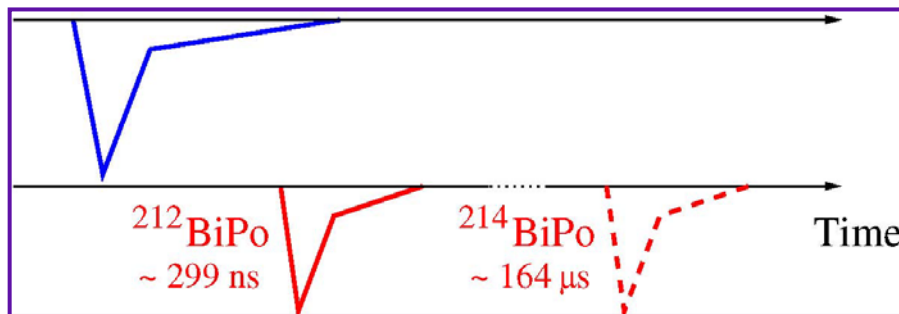


- **Background sources:**
 - γ -induced random coincidences



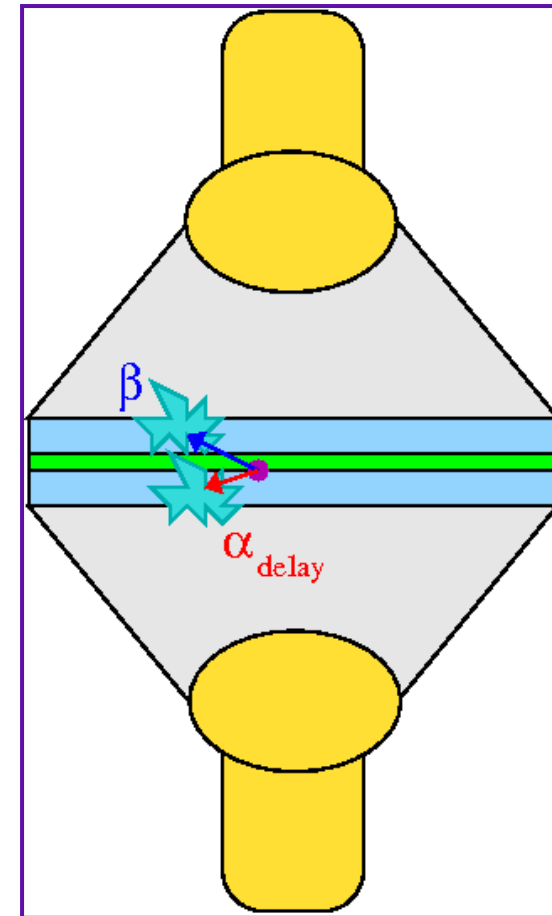
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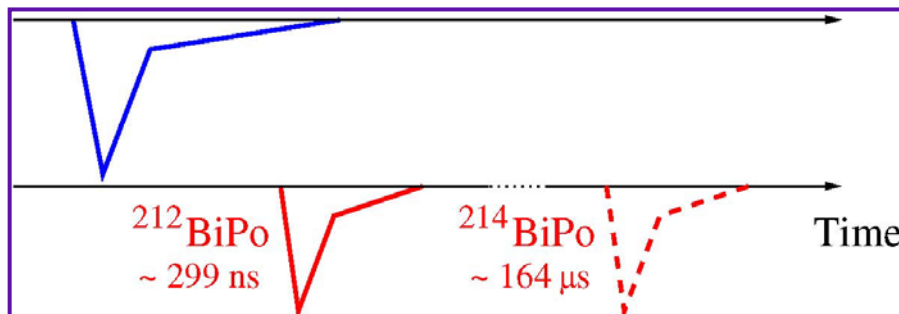
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- Radon contamination in the sensitive volume



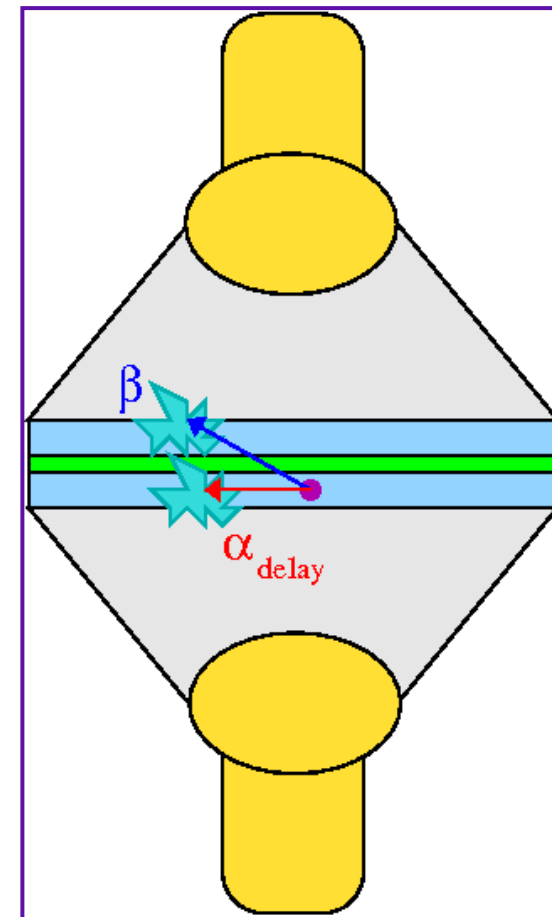
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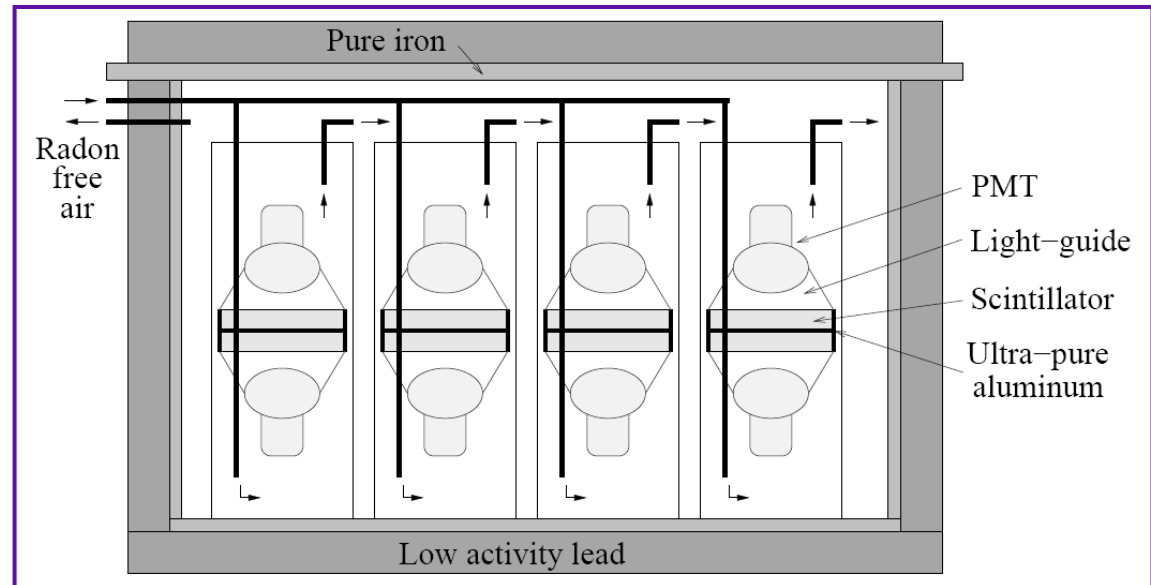
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- Scintillator surface contamination
- Radon contamination in the sensitive volume
- Scintillator bulk contamination



The BiPo detector: Previous prototypes

- BiPo prototypes have been really useful to fix different open questions and to validate the technology.
 - **BiPo 1:**
 - 20 modules
 - Face to face scintillators
 - Installed at LSM
 - DAQ tests
 - MatAcq digitizer board
 - Trigger modules
 - Analysis methods tests
 - Radiopurity of the scintillators

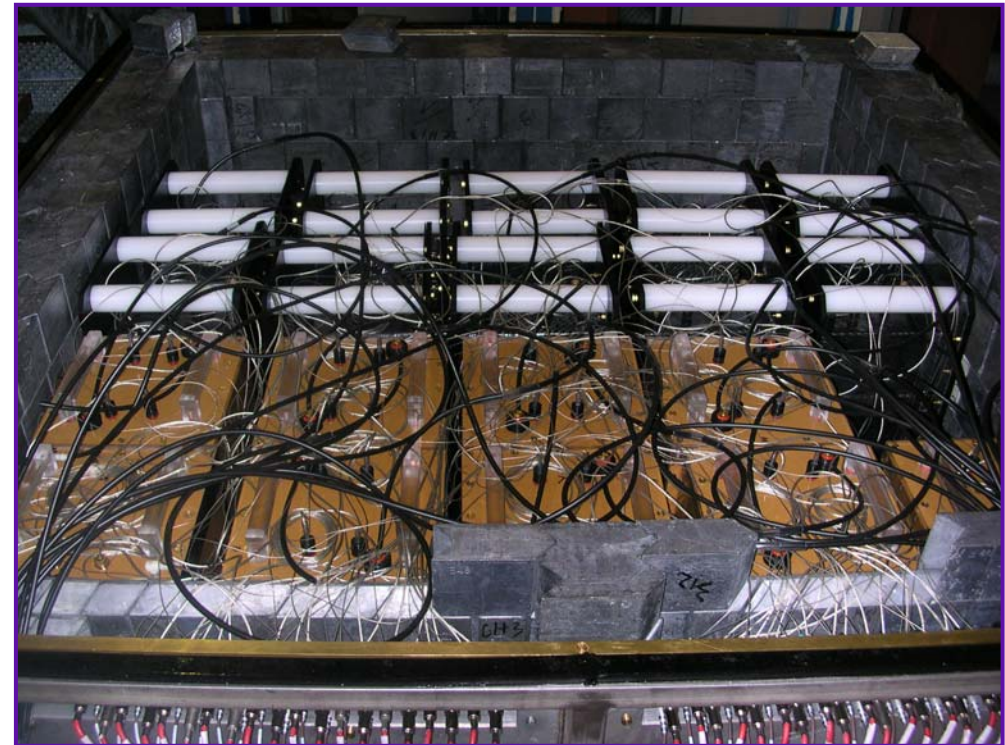


Detection principle tested with a calibrated **Al** foil between the scintillators

^{208}Tl contamination level of **$1.5 \mu\text{Bq}/\text{m}^2$ of scintillator** measured

NIM A 622 (2010) 120-128

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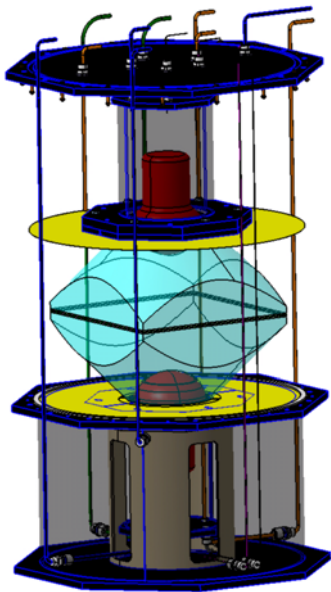
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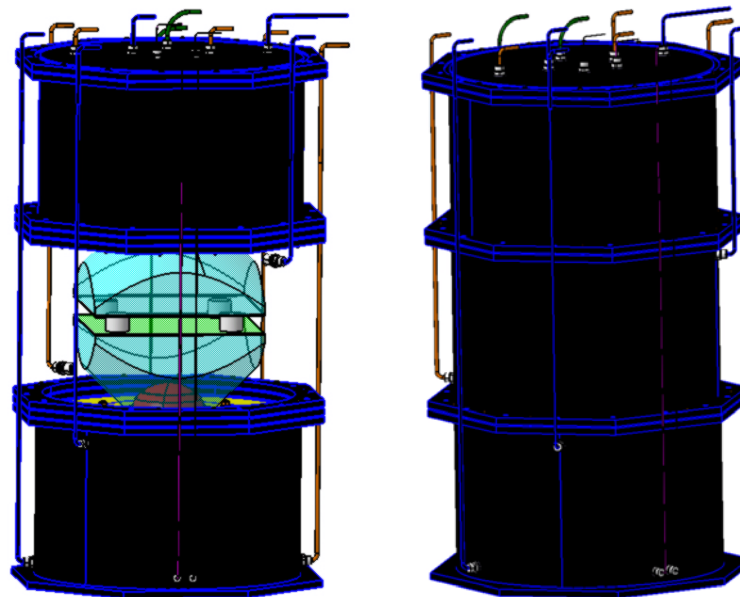
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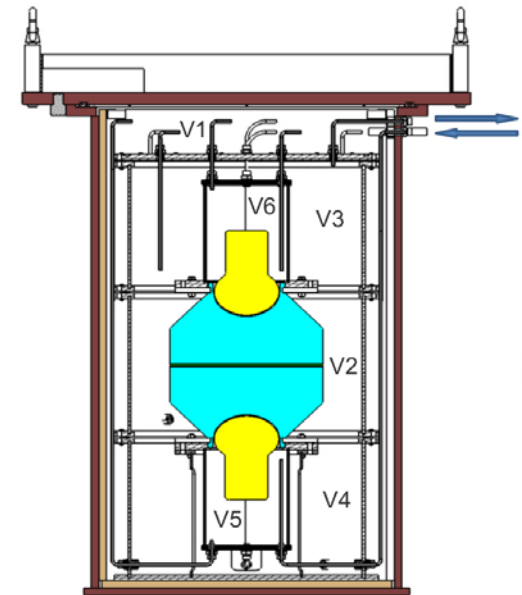
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 - **BiPo 3 prototype:**
 - 1 module previous to the final setup.
 - Installed @ Canfranc Underground Laboratory (Depth ~ 2500 mwe).



Optical Module
Scheme



Opaque and Rn tight enclosure of
the optical module



Independent
Volumes Scheme

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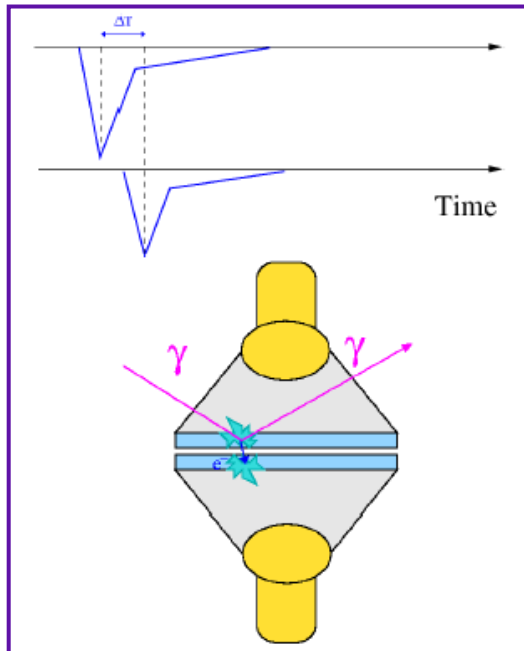


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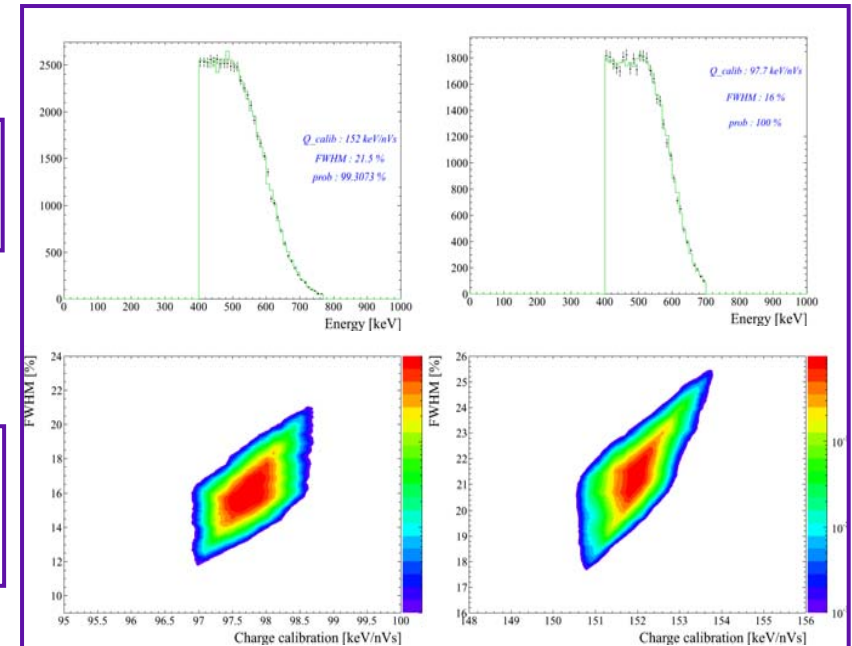
→ **How low can we go in terms of Rn level and Counting Rate?**

- **Calibrations (^{54}Mn and ^{22}Na) for energy, resolution and timing**
- Counting Rate monitoring
- BiPo events analysis



Energy Spectrum
(Compton Edge)

Calibration Kolmogorov Test
(Best Gain & FWHM Fit)



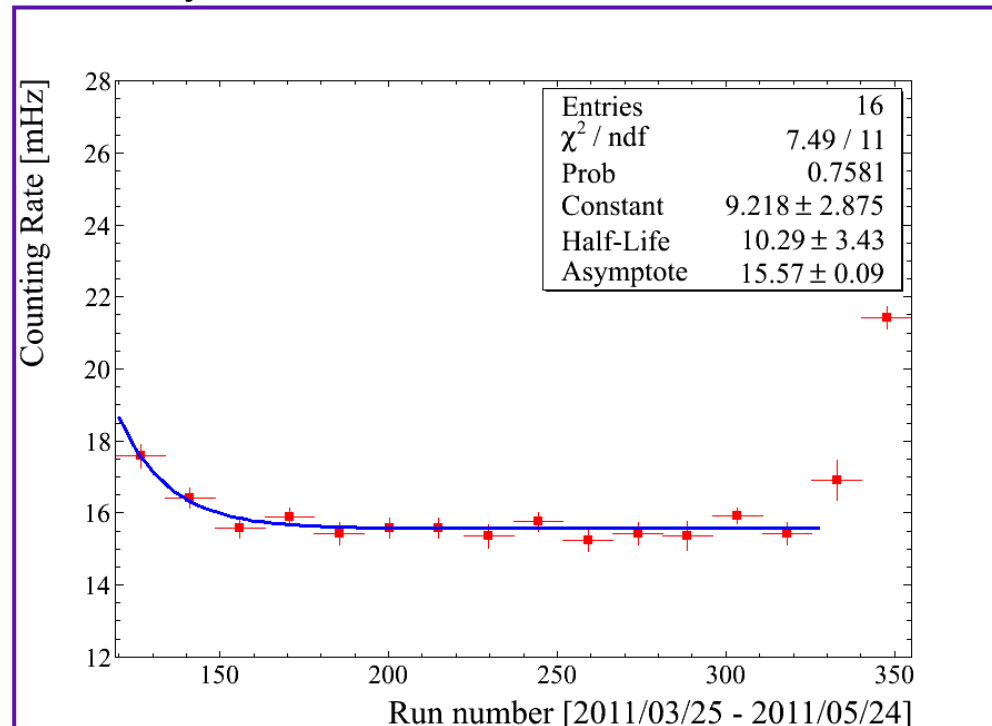
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With Nitrogen flush:



- **15-20 mHz** integral rate
- Quite stable
- Dependance with Rn concentration

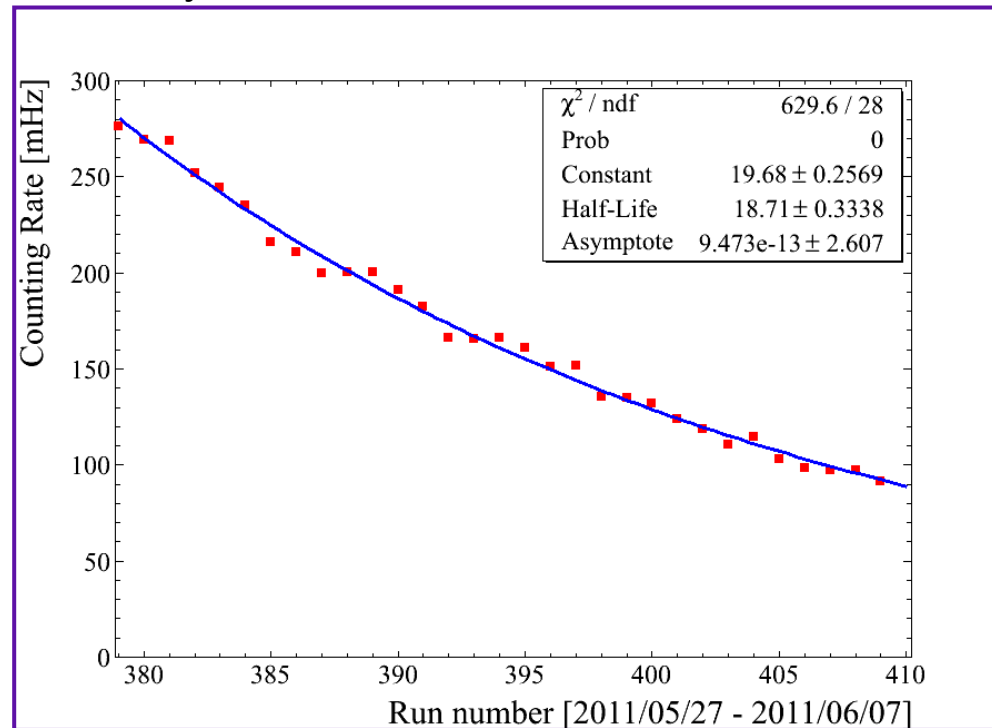
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Without Nitrogen flush:



- Decay rate
- $T_{1/2} \sim 4.7 \pm 0.1$ days
- Rn detection

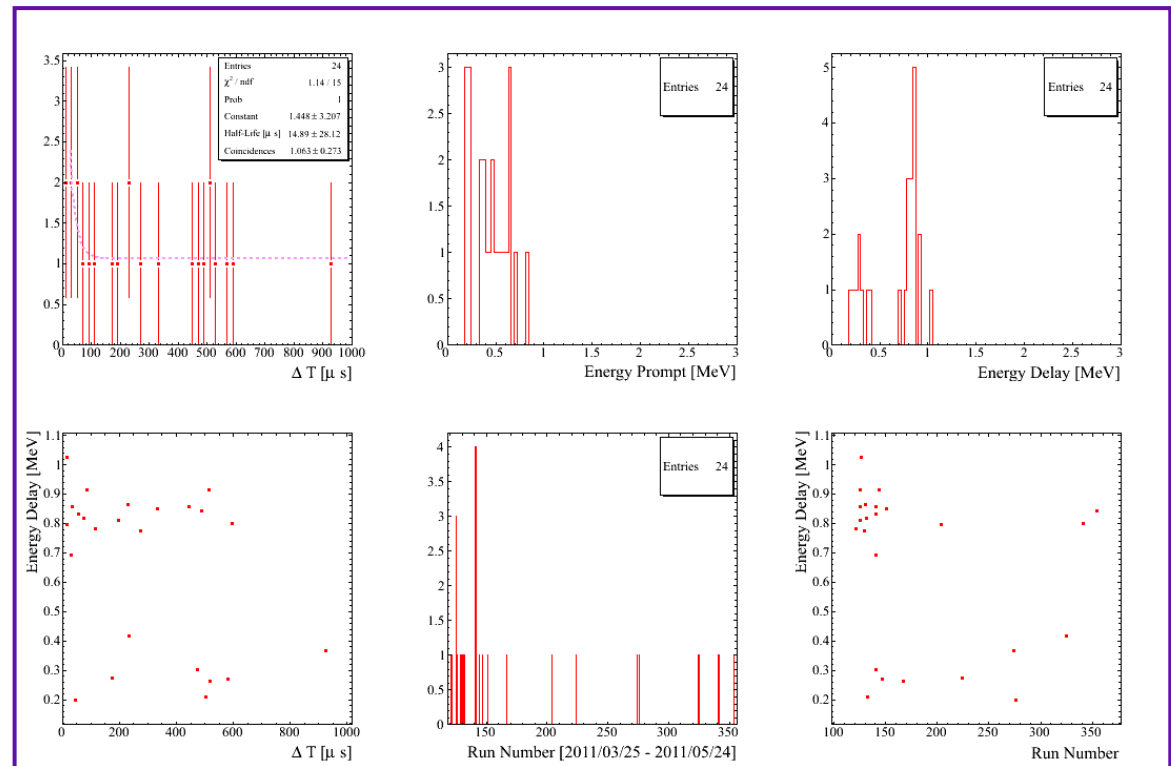
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- Data from one candidate:
 - Prompt energy (e^-)
 - Delay energy (α)
 - Δt
 - Rate evolution
 - Assure is constant



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- **Results** after ~60 days data taking:

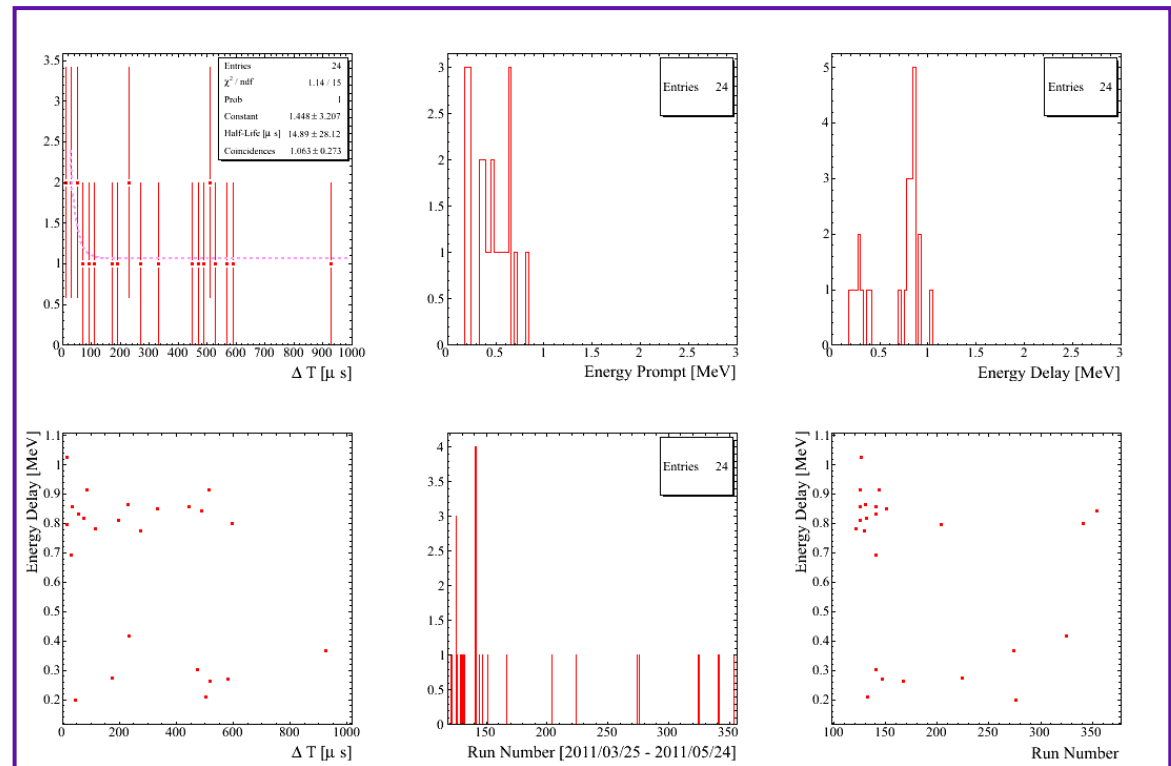
^{208}Tl :

$A(^{208}\text{Tl}) < 5.45 \mu\text{Bq}/\text{m}^2 \text{ scint @ } 90\% \text{ C.L.}$

^{214}Bi :

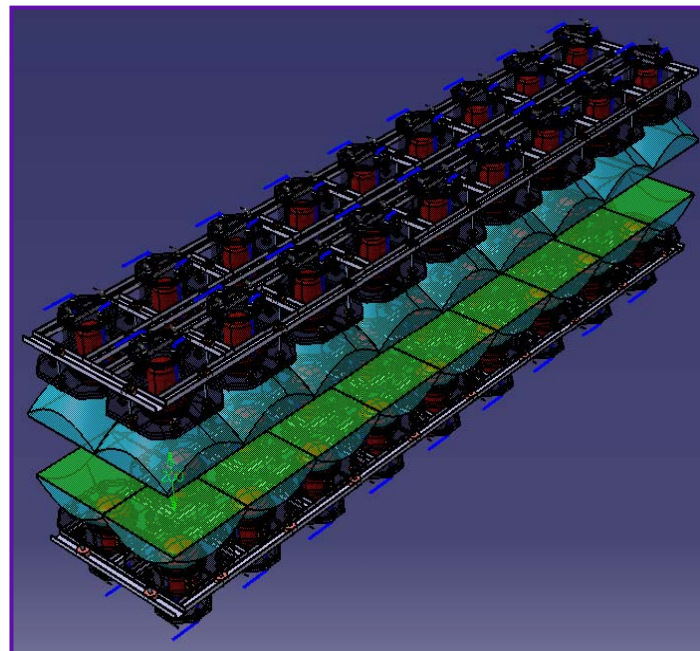
$0.6 < A(^{214}\text{Bi}) = 5.3 < 23 \mu\text{Bq}/\text{m}^2 \text{ scint}$

PROMISING RESULTS

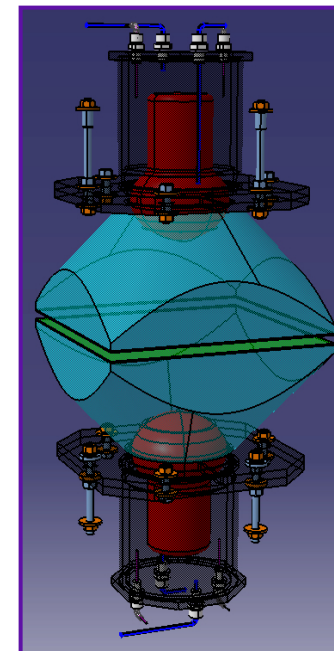


Final setup: BiPo 3 detector

- Prototypes have shown that required sensitivity levels could be reachable with the face to face detector approximation.
- Main features:
 - 2 modules detector → Possibility to measure 8 SuperNEMO source foils simultaneously
 - 20 30x30 cm² optical sub-modules in each module
 - 3.6 m² of sensitive surface
 - 2 mm thick scintillator plates
 - Light guide geometry optimized
 - Volume separation and nitrogen flushing for Rn suppression and external shielding



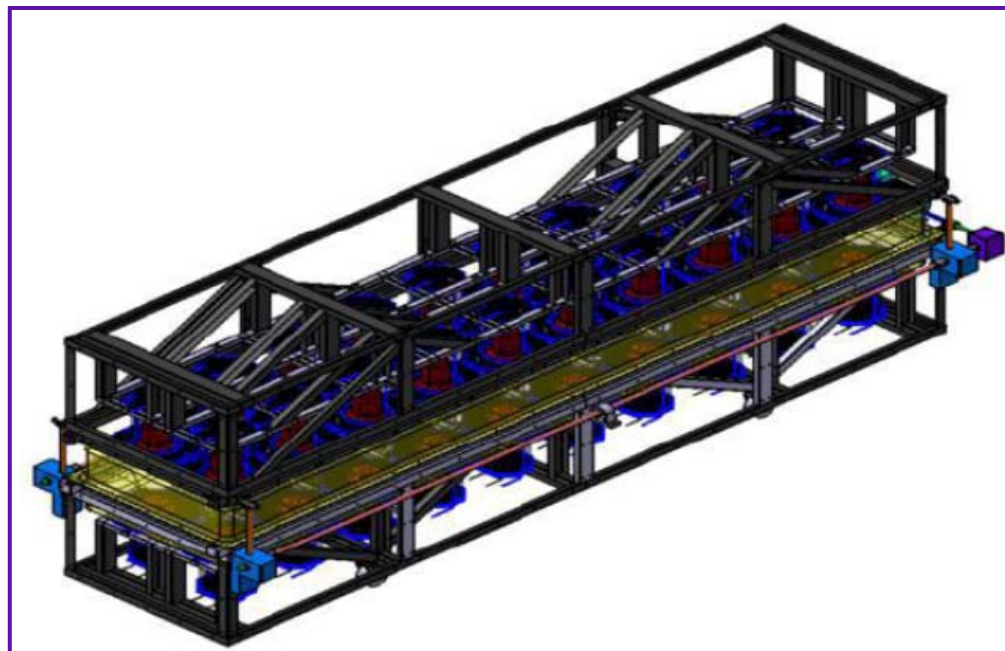
Assembling of 20
optical submodules
→ 1 BiPo 3 module



BiPo 3 Optical Submodule

Final setup: BiPo 3 detector

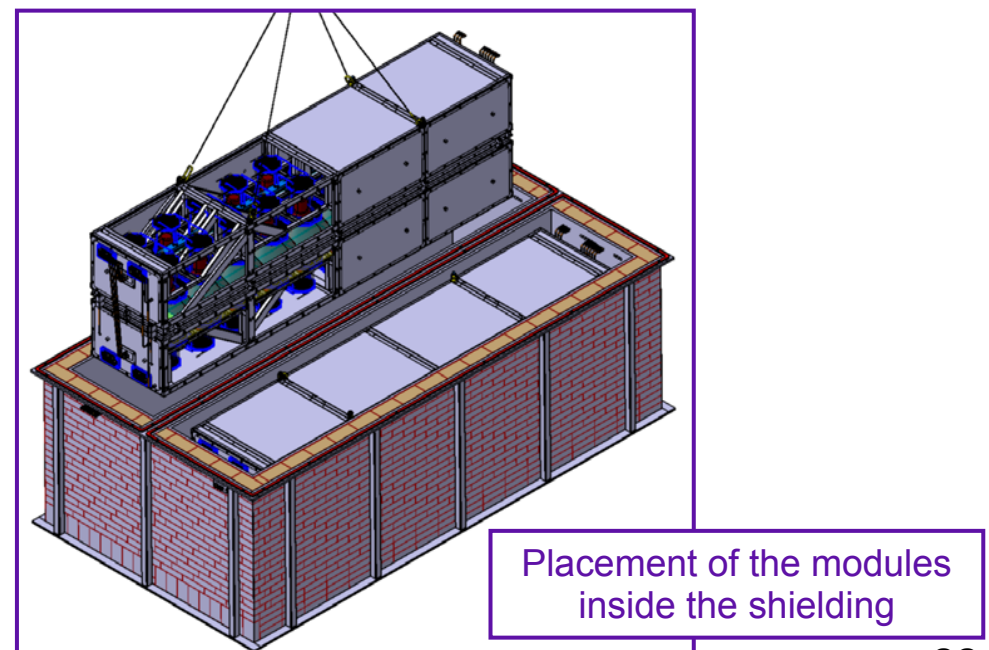
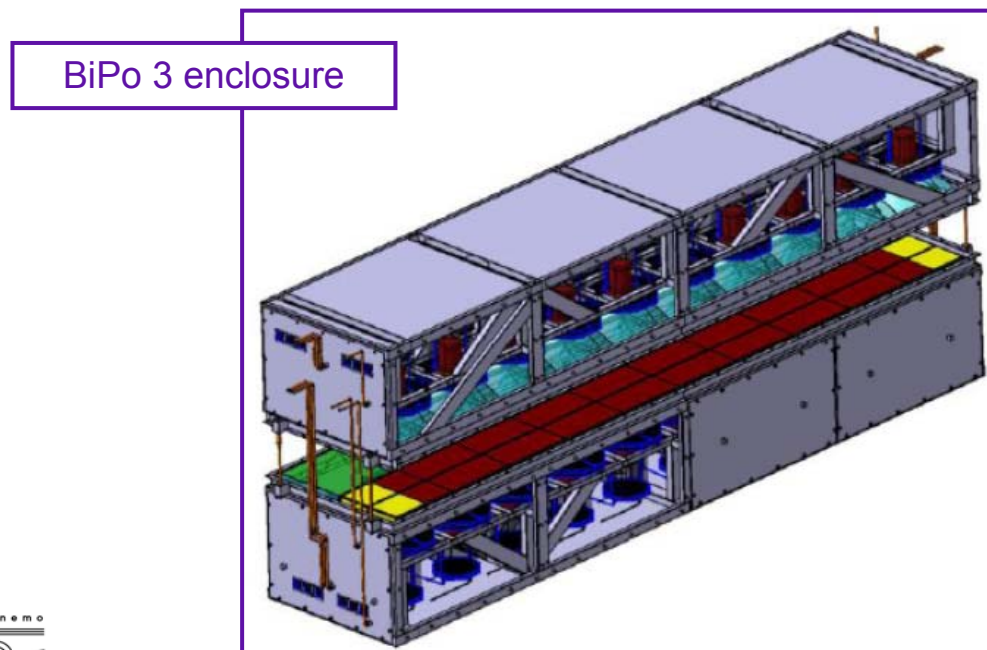
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BiPo 3 detectors installed
in the internal structure

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BiPo 3 present status

- Almost the 80 optical submodules assembled and characterized in a test bench @ LAL.



Light guide annealing



Gluing of the scintillator



Aluminization



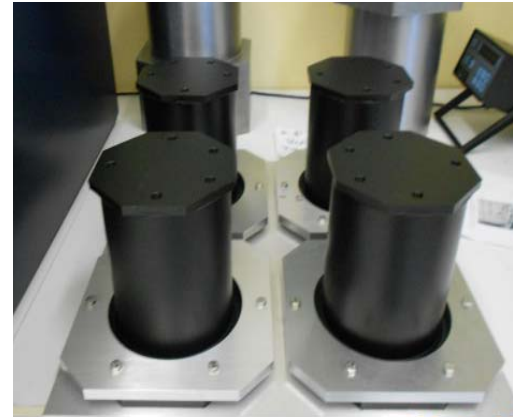
PMT gluing



Tyvek wrapping



Black film



Light black box



Characterized sub-modules
storage

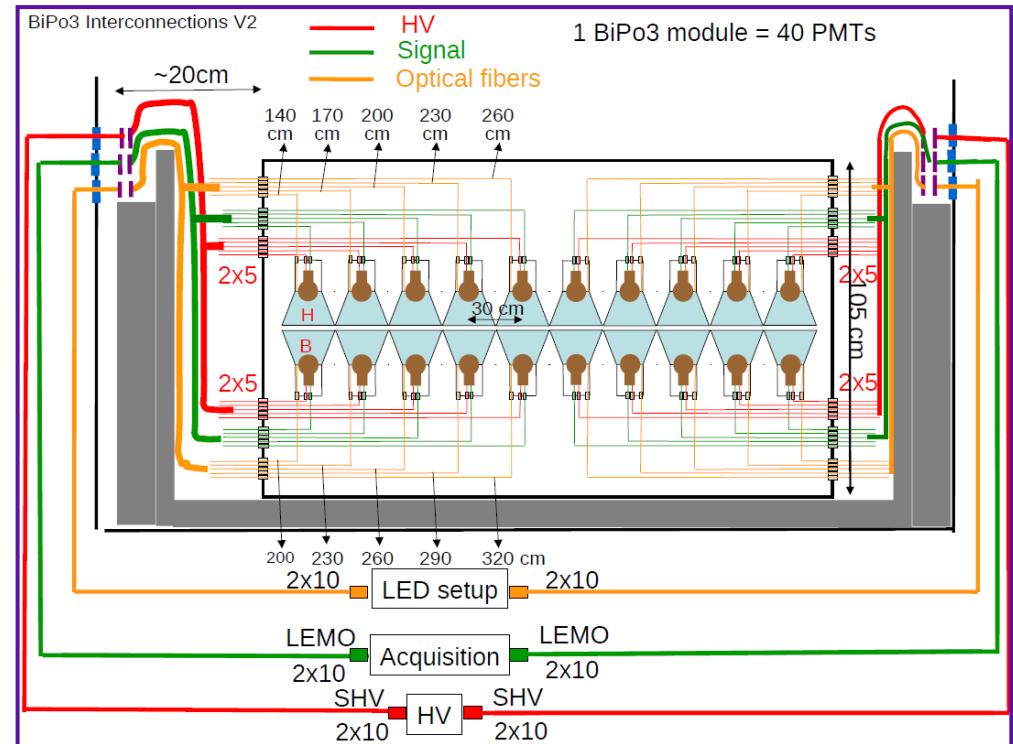
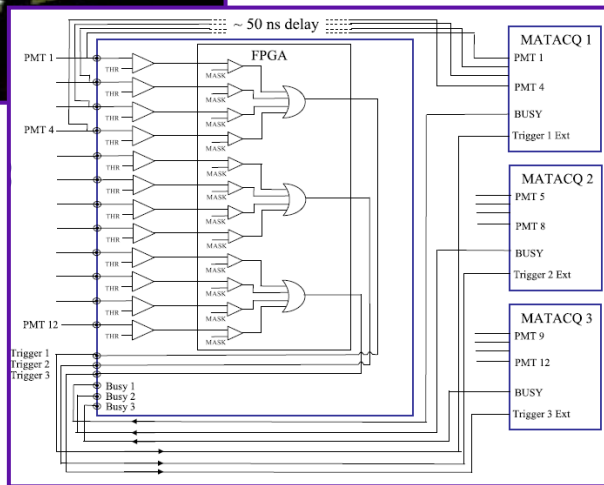
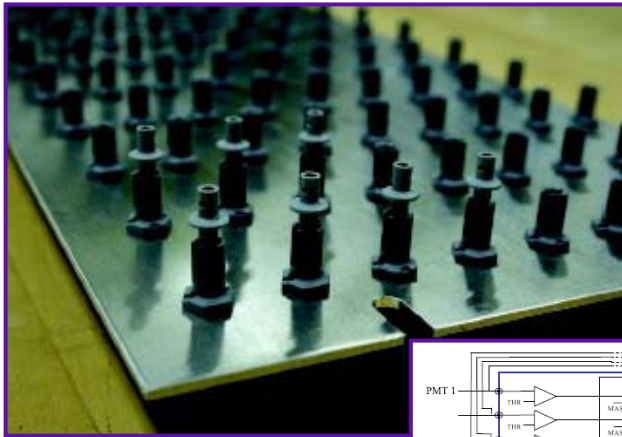
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- Preliminary test and commissioning for DAQ system, Nitrogen flushing feedthroughs, calibrations...

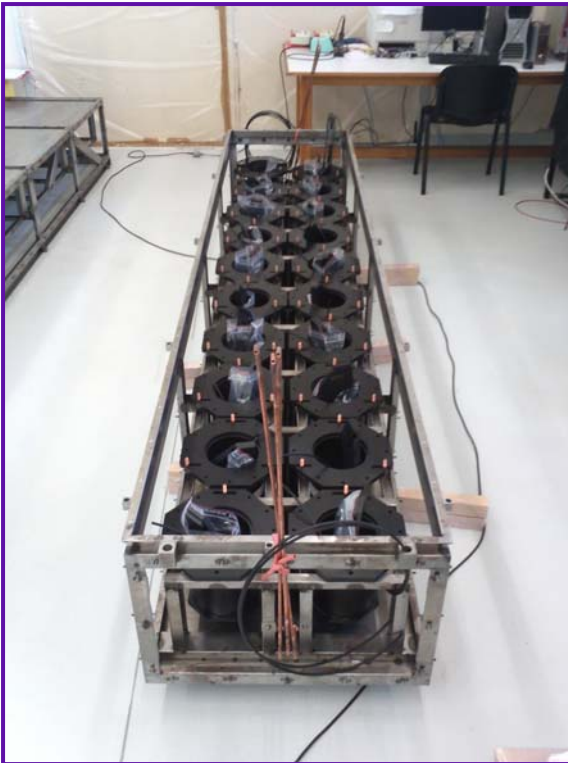


ALMOST READY FOR THE COMMISSIONING OF THE FIRST MODULE.

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FIRST ASSEMBLING TESTS @ LAL CLEAN ROOM



- Approximate timeline:
 - **Module 1:**
 - All the required material for the first BiPo 3 module will be tested at LAL before mid – June
 - The material will be delivered to Canfranc Underground Laboratory last week of June
 - **Installation** of the first BiPo 3 module between **2 and 13 of July**
 - **Commissioning** of the detector during **summer**
 - **Calibrations** and background measurement: Until **end 2012**
 - **Source** measurement: **Beginning 2013**
 - **Module 2:**
 - **Same** schedule that for **module 1** with **~3 months of delay**

***In March 2013 the whole detector will be completely ready,
being possible to make source measurements since beginning
2013***

- **NEMO 3** experiment has obtained really good results in the measurement of $T_{1/2}^{2\nu\beta\beta}$ and also upper limits of $T_{1/2}^{0\nu\beta\beta}$ for ^{82}Se and ^{100}Mo .
- **SuperNEMO** experiment will use the same technique with **100 kg** of $\beta\beta$ emitter trying to reach the $\langle m_\nu \rangle \sim 50 \text{ meV}$ level.
- One requirement for SuperNEMO is to have really good **radiopurity** levels for the **source foils**: $< 2 \mu\text{Bq/kg}$ in ^{208}Tl and $< 10 \mu\text{Bq/kg}$ in ^{214}Bi .
- **BiPo** is a **dedicated** detector designed to measure these foils with the required sensitivity.
- A set of **prototypes** have showed the **viability** to construct the detector proving the **detection principle**.
- **Results** of these prototypes lead to think that required **sensitivity** will be **reached**.
- **Final setup** is almost **completed** and source foil measurements are expected to **start beginning 2013**.

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