

Neutron *sd*-shell excitations in exotic nuclei near N=8

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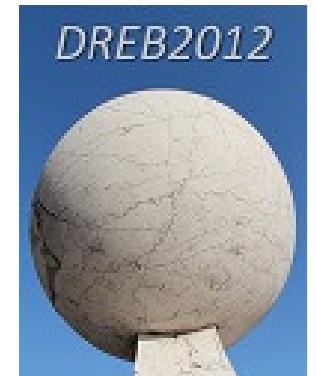
⁴Joint Institute for Nuclear Astrophysics, Michigan State University, East Lansing, Michigan 48824, USA

⁵Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

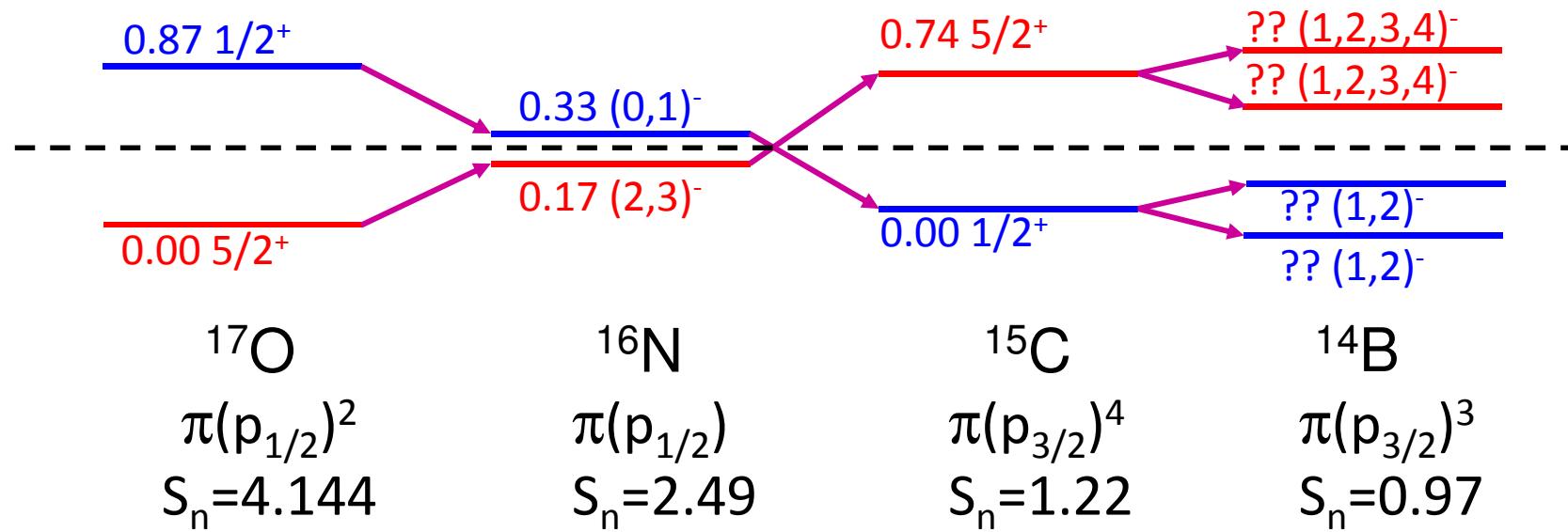
⁶Department of Physics, University of Manchester

⁷LANSCE-NS, Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA

⁸Lawrence Livermore National Laboratory, Livermore, California 94551, USA

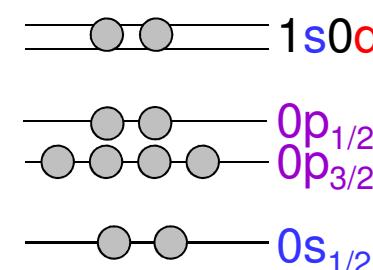
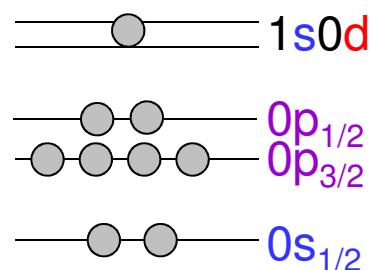
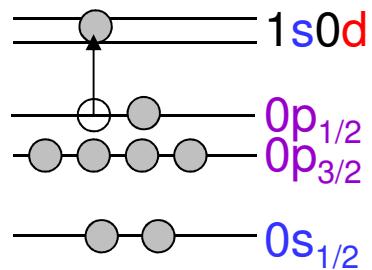


Evolution of $1s_{1/2}$ - $0d_{5/2}$ splitting outside N=8



Gross behavior from p-n tensor attraction/repulsion

Neutron configurations around N=8



N=8 (¹³B)
Positive-parity states
are $v(p_{1/2})^{-1}(sd) \otimes \pi(p_{3/2})^{-1}$

N=9 (¹⁴B)
Negative-parity states
are $v(sd) \otimes \pi(p_{3/2})^{-1}$

N=10 (¹⁶C)
Positive-parity states
are $v(sd)^2 \otimes {}^{14}\text{C}$



¹²B from ¹¹B(d,p)

At 6 MeV/u



¹³B from ¹⁴C(⁹Be, ¹⁰B)

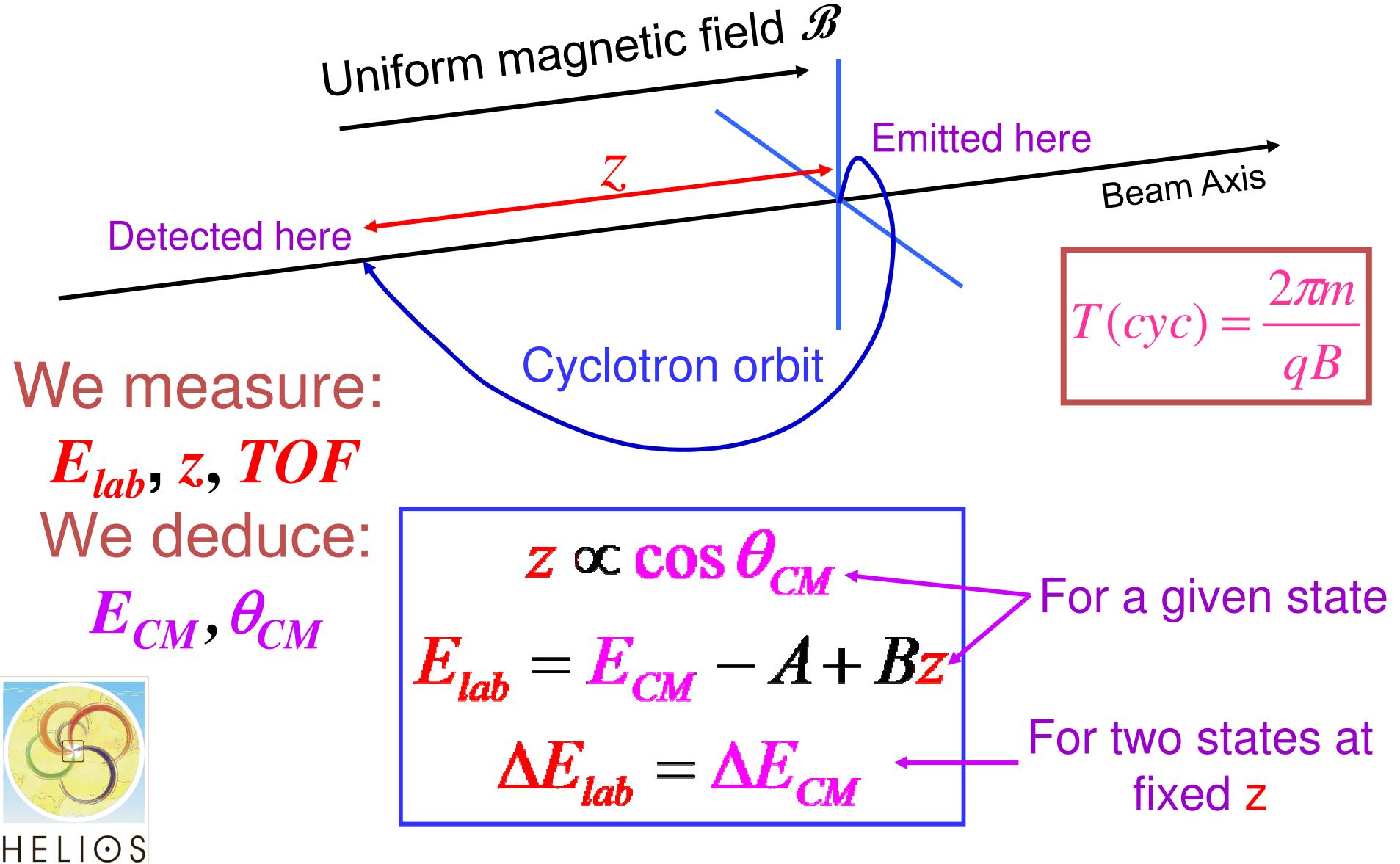
At 15.7 MeV/u



¹⁵C from ¹⁴C(d,p)

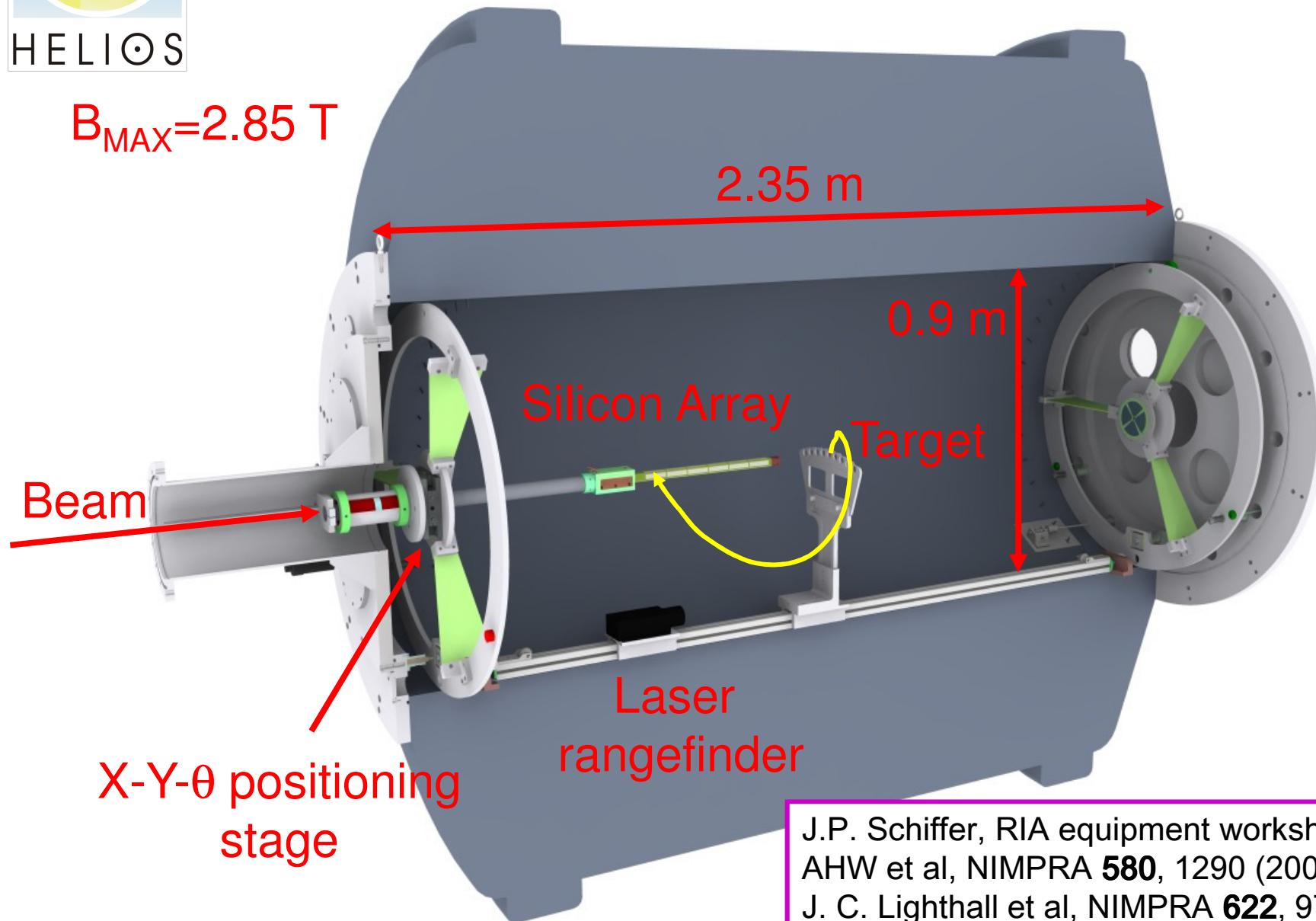
At 8.5 MeV/u

The HELIOS approach to inverse kinematics





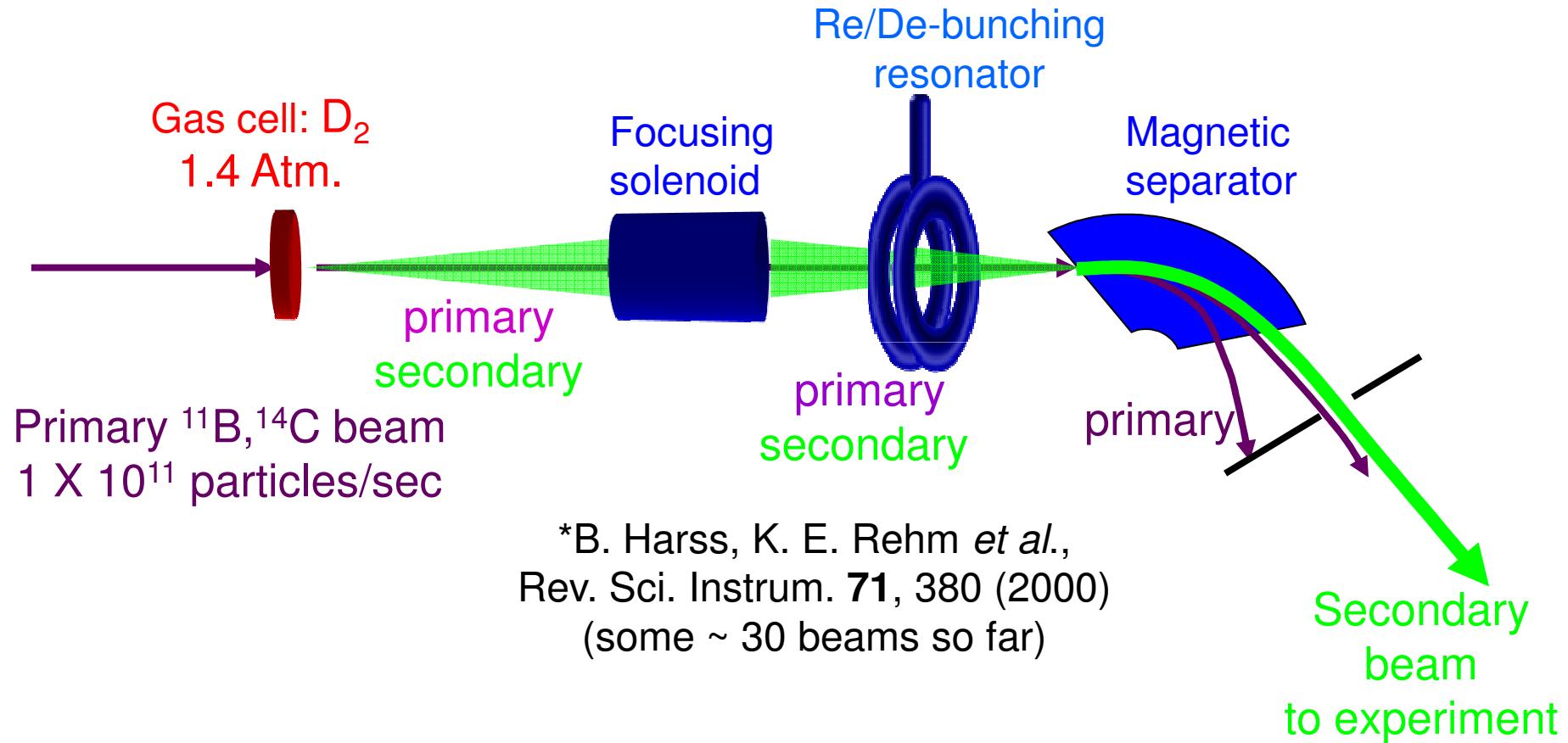
HELIcal Orbit Spectrometer -HELIOS



J.P. Schiffer, RIA equipment workshop 1999,
AHW et al, NIMPRA **580**, 1290 (2007)
J. C. Lighthall et al, NIMPRA **622**, 97 (2010)



Producing secondary beams: “In-flight” production at ANL*



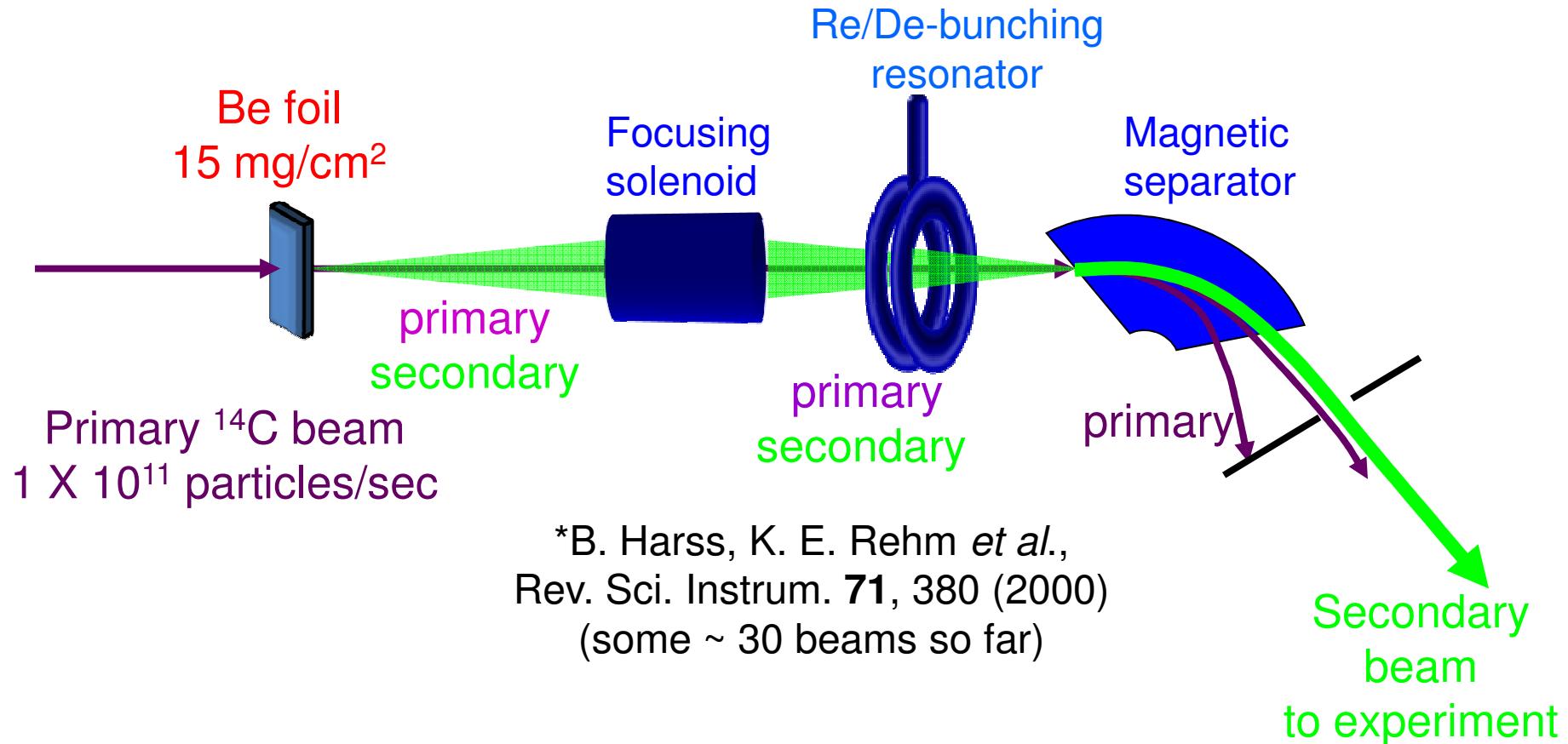
*B. Harss, K. E. Rehm *et al.*,
Rev. Sci. Instrum. **71**, 380 (2000)
(some ~ 30 beams so far)

^{12}B intensity $\sim 6 \times 10^5$ /sec at 6.25 MeV/u

^{15}C intensity $\sim 1.5 \times 10^6$ /sec at 8.2 MeV/u

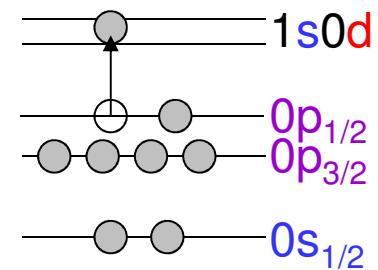
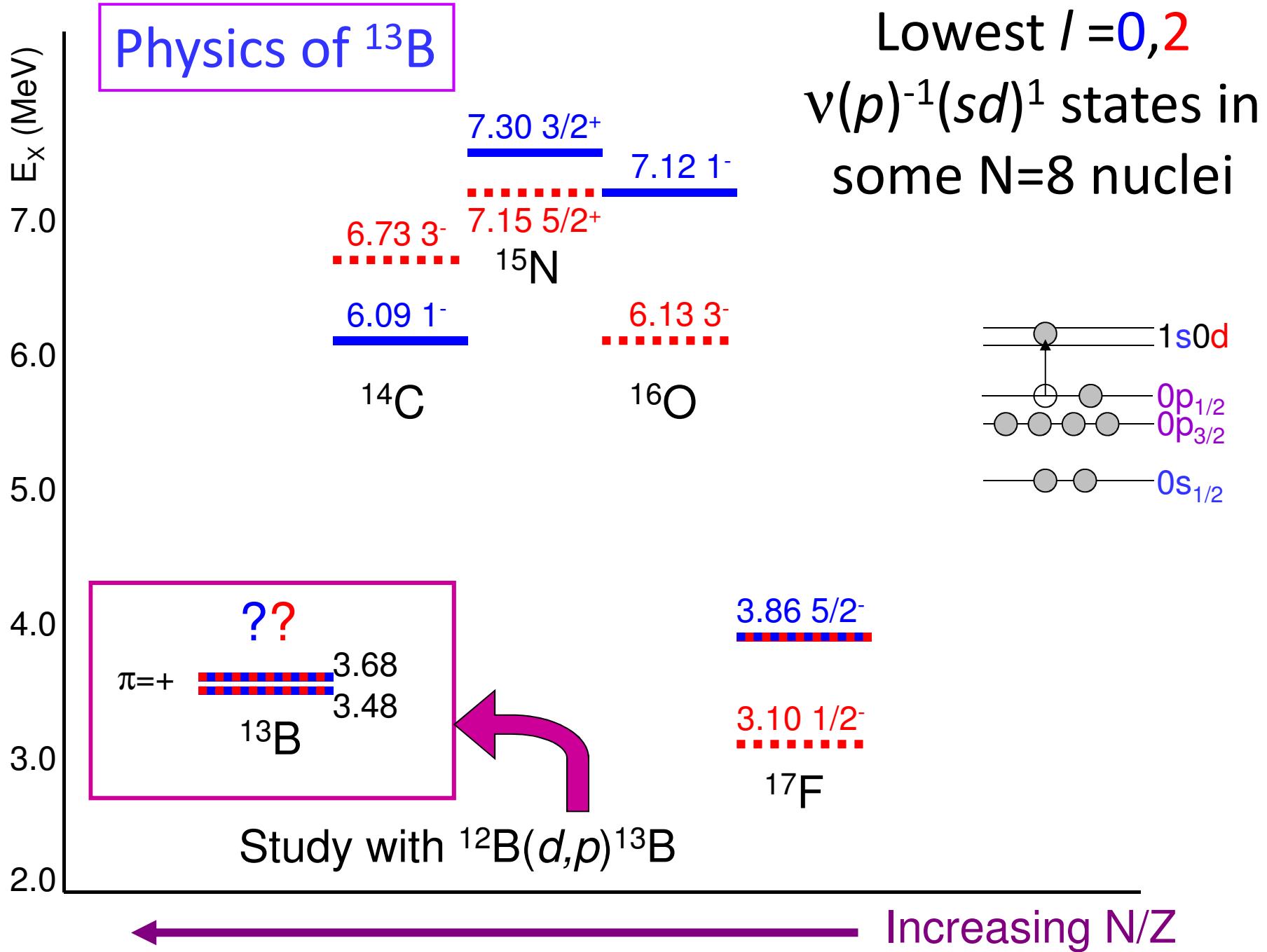


Producing secondary beams: “In-flight” production at ANL*

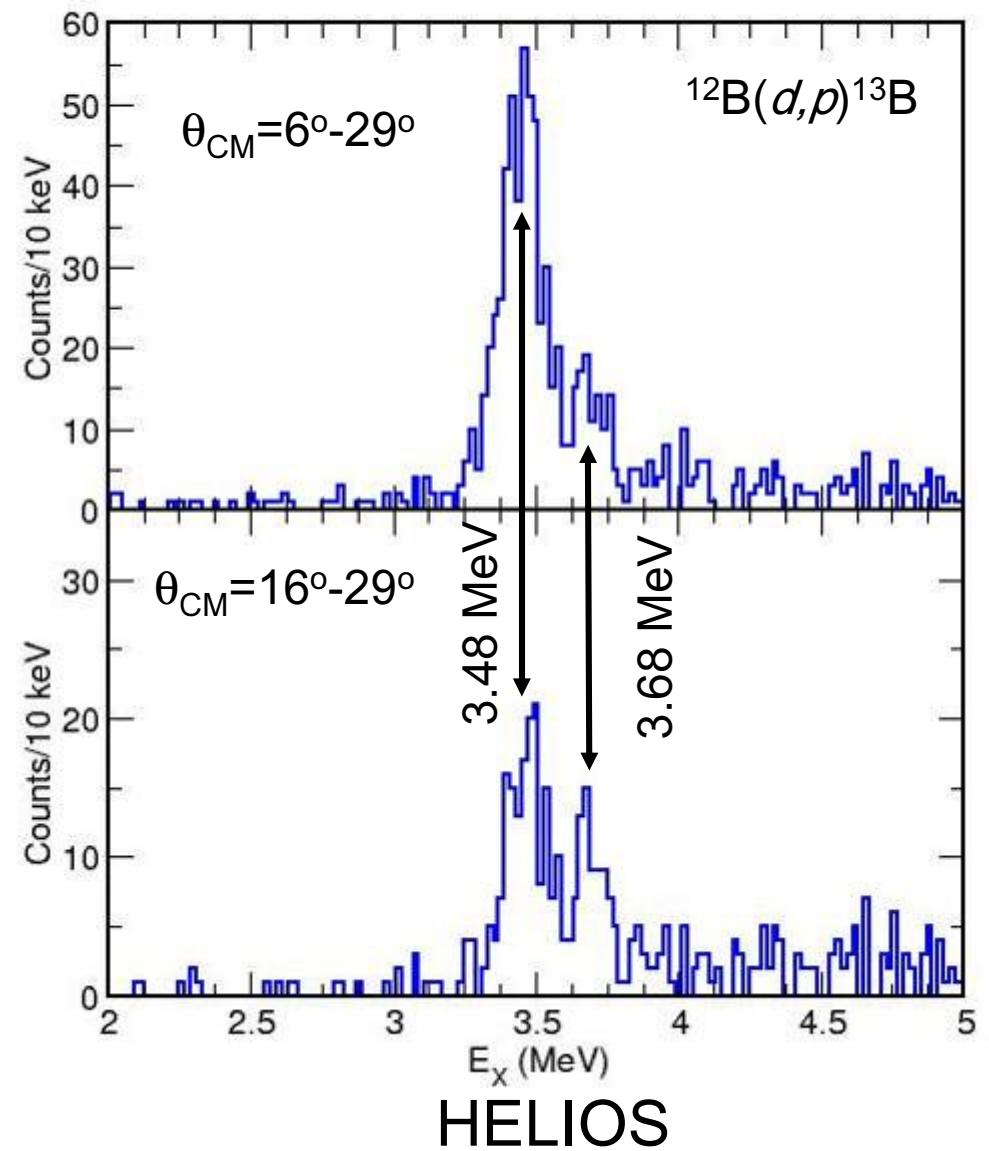
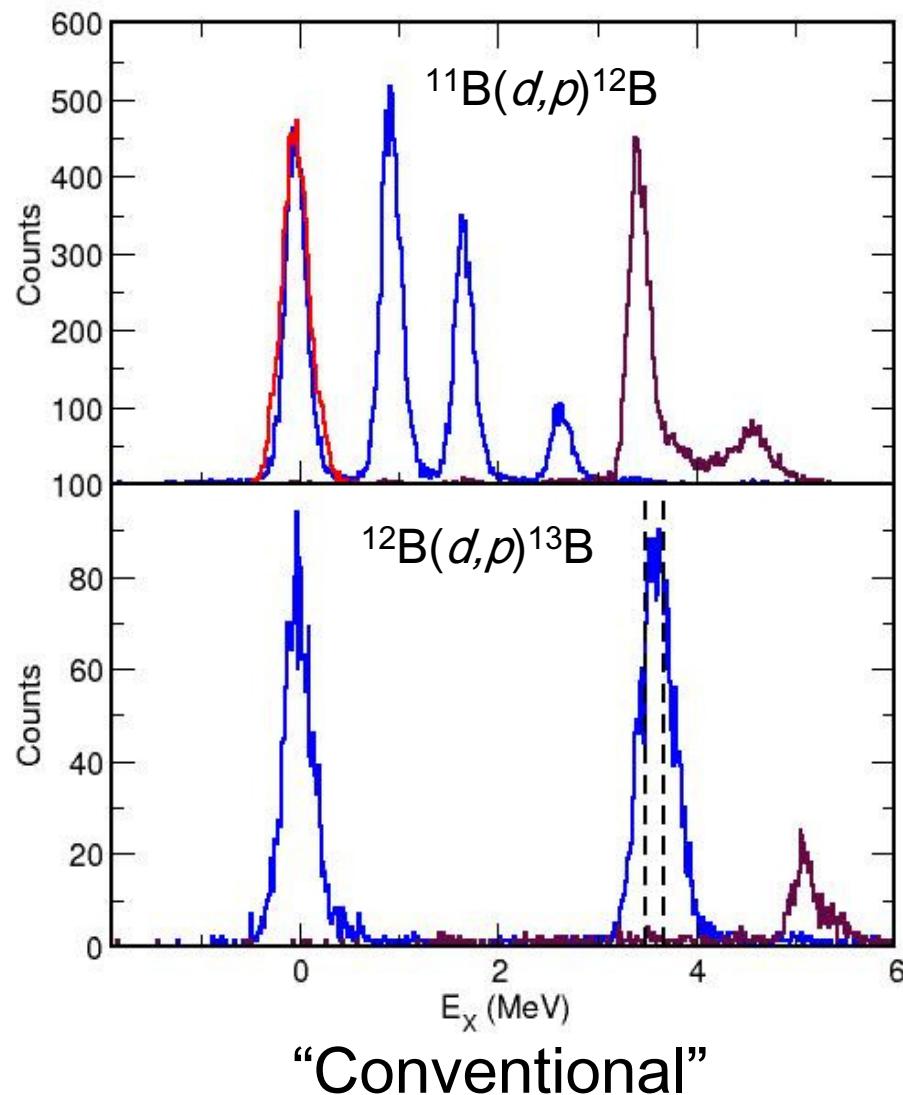


*B. Harss, K. E. Rehm *et al.*,
Rev. Sci. Instrum. **71**, 380 (2000)
(some ~ 30 beams so far)

^{13}B intensity $\sim 4 \times 10^4$ /sec at 15.7 MeV/u



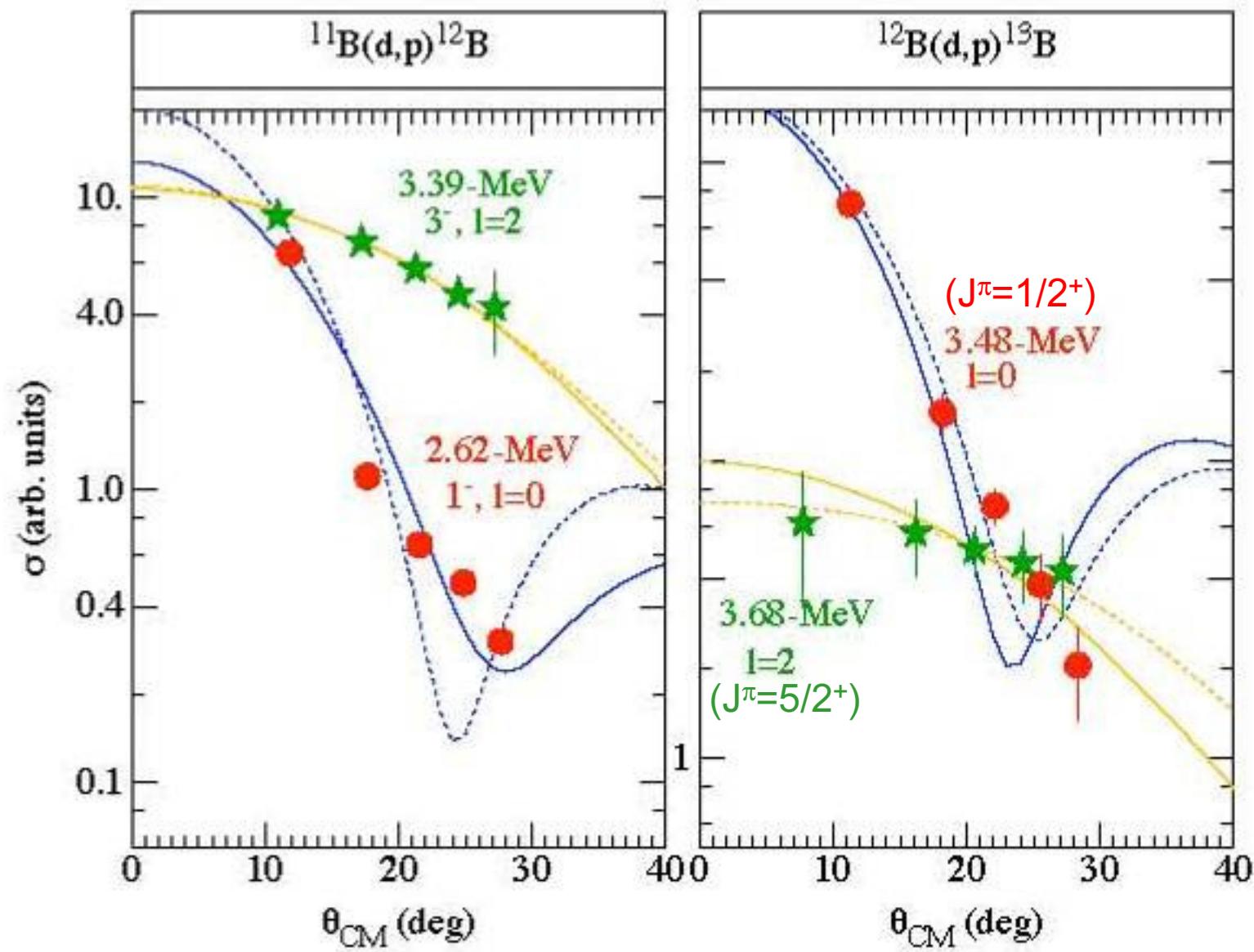
First HELIOS RIB results with $^{12}\text{B}(d,p)^{13}\text{B}$

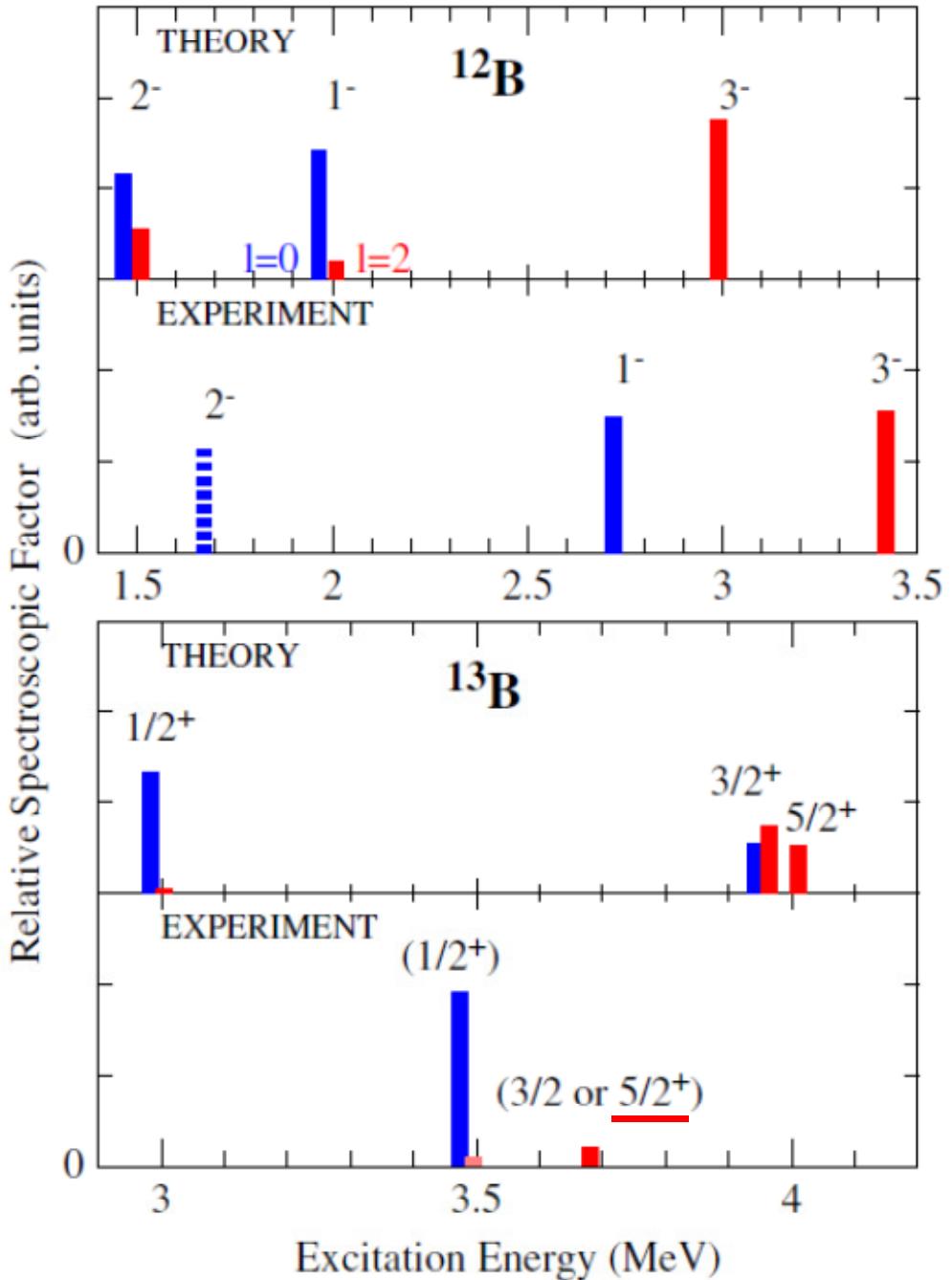


H. Y. Lee et al., PRC 81, 015802 (2010)

B. B. Back et al., PRL 104, 132501 (2010)

$^{11,12}\text{B}(d,p)^{12,13}\text{B}$ angular distributions





Theory versus experiment for ^{13}B

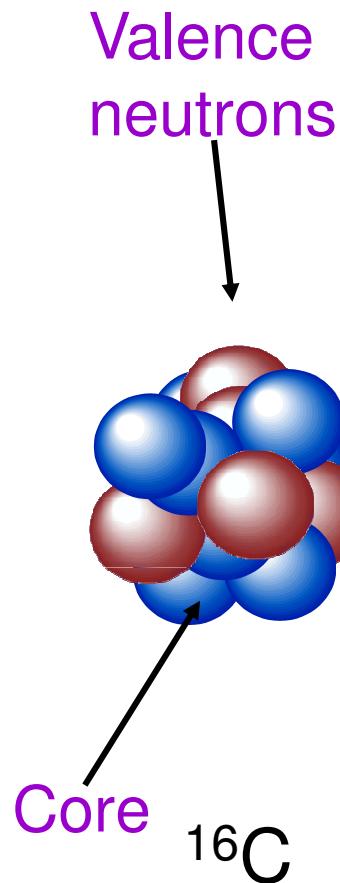
Excitation energies and relative spectroscopic factors from the shell model

Blue: L=0

Red: L=2

5/2⁺ L=2 is reduced,
no nearby 3/2⁺ is observed

Exotic behavior in ^{16}C ?



VOLUME 92, NUMBER 6

PHYSICAL REVIEW LETTERS

week ending
13 FEBRUARY 2004

Anomalously Hindered $E2$ Strength $B(E2; 2_1^+ \rightarrow 0^+)$ in ^{16}C

N. Imai,^{1,*} H. J. Ong,² N. Aoi,¹ H. Sakurai,² K. Demichi,³ H. Kawasaki,³ H. Baba,³ Zs. Dombrádi,⁴ Z. Elekes,^{1,†} N. Fukuda,¹ Zs. Fülöp,⁴ A. Gelberg,⁵ T. Gomi,³ H. Hasegawa,³ K. Ishikawa,⁶ H. Iwasaki,² E. Kaneko,³ S. Kanno,³ T. Kishida,¹ Y. Kondo,⁶ T. Kubo,¹ K. Kurita,³ S. Michimasa,⁷ T. Minemura,¹ M. Miura,⁶ T. Motobayashi,¹ T. Nakamura,⁶ M. Notani,⁷ T. K. Onishi,² A. Saito,³ S. Shimoura,⁷ T. Sugimoto,⁶ M. K. Suzuki,² E. Takeshita,³ S. Takeuchi,¹ M. Tamaki,⁷ K. Yamada,³ K. Yoneda,^{1,‡} H. Watanabe,¹ and M. Ishihara¹

Physics Letters B 586 (2004) 34–40 Decoupling of valence neutrons from the core in ^{16}C

Z. Elekes^{a,1}, Zs. Dombrádi^b, A. Krasznahorkay^b, H. Baba^c, M. Csatlós^b, L. Csige^b, N. Fukuda^a, Zs. Fülöp^b, Z. Gácsi^b, J. Gulyás^b, N. Iwasa^d, H. Kinugawa^c, S. Kubono^e, M. Kurokawa^c, X. Liu^e, S. Michimasa^e, T. Minemura^e, T. Motobayashi^a, A. Ozawa^a, A. Saito^c, S. Shimoura^e, S. Takeuchi^a, I. Tanihata^a, P. Thirolf^f, Y. Yanagisawa^a, K. Yoshida^a

PRL 100, 152501 (2008)

PHYSICAL REVIEW LETTERS

week ending
18 APRIL 2008

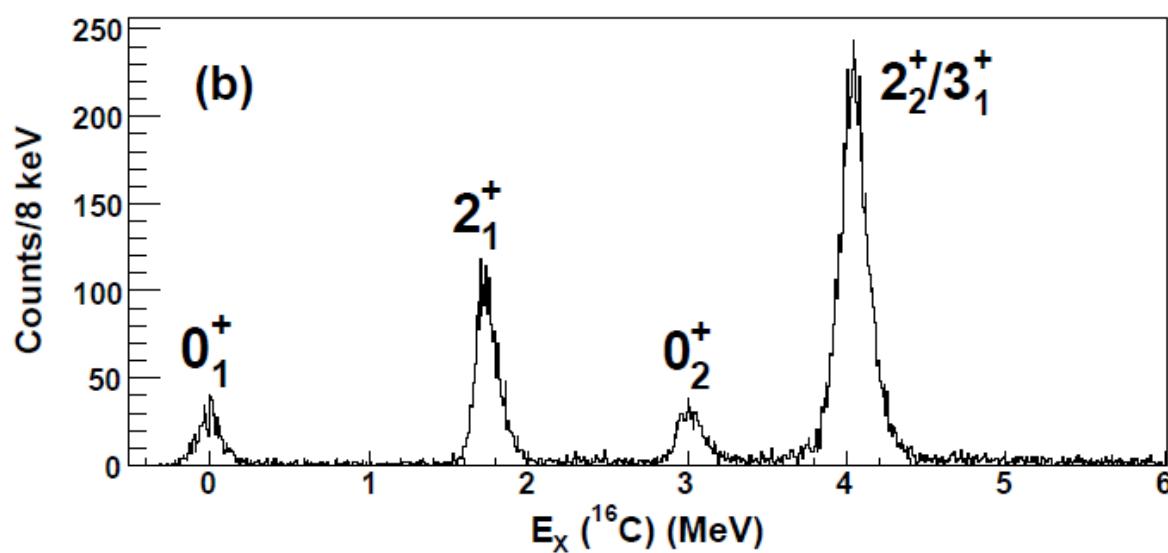
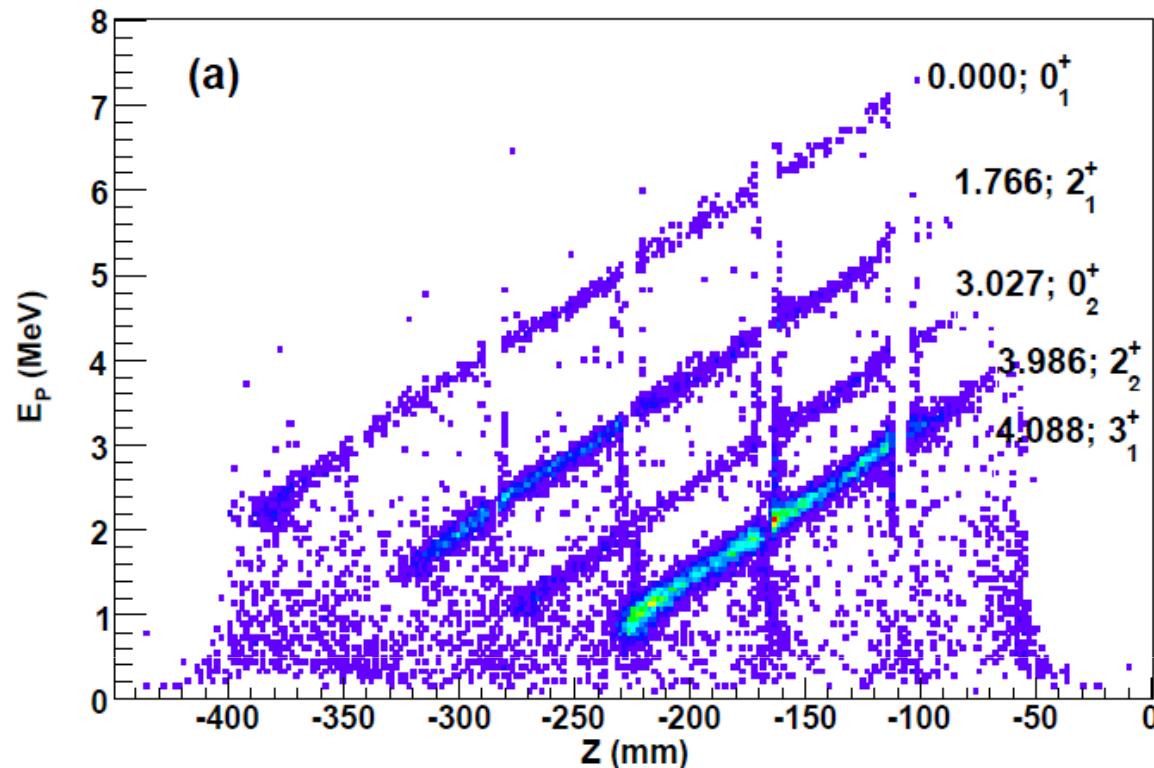
Lifetime Measurement of the First Excited 2^+ State in ^{16}C

M. Wiedeking, P. Fallon, A. O. Macchiavelli, J. Gibelin, M. S. Basunia, R. M. Clark, M. Cromaz, M.-A. Deleplanque, S. Gros, H. B. Jeppesen, P. T. Lake, I.-Y. Lee, L. G. Moretto, J. Pavan, L. Phair, and E. Rodriguez-Vietiez
Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

L. A. Bernstein, D. L. Bleuel, J. T. Burke, S. R. Lesher, B. F. Lyles, and N. D. Scielzo
Lawrence Livermore National Laboratory, Livermore, California 94550, USA
(Received 20 November 2007; published 16 April 2008)

Study with $^{15}\text{C}(\text{d},\text{p})^{16}\text{C}$

No hindrance, and no exotic behavior.

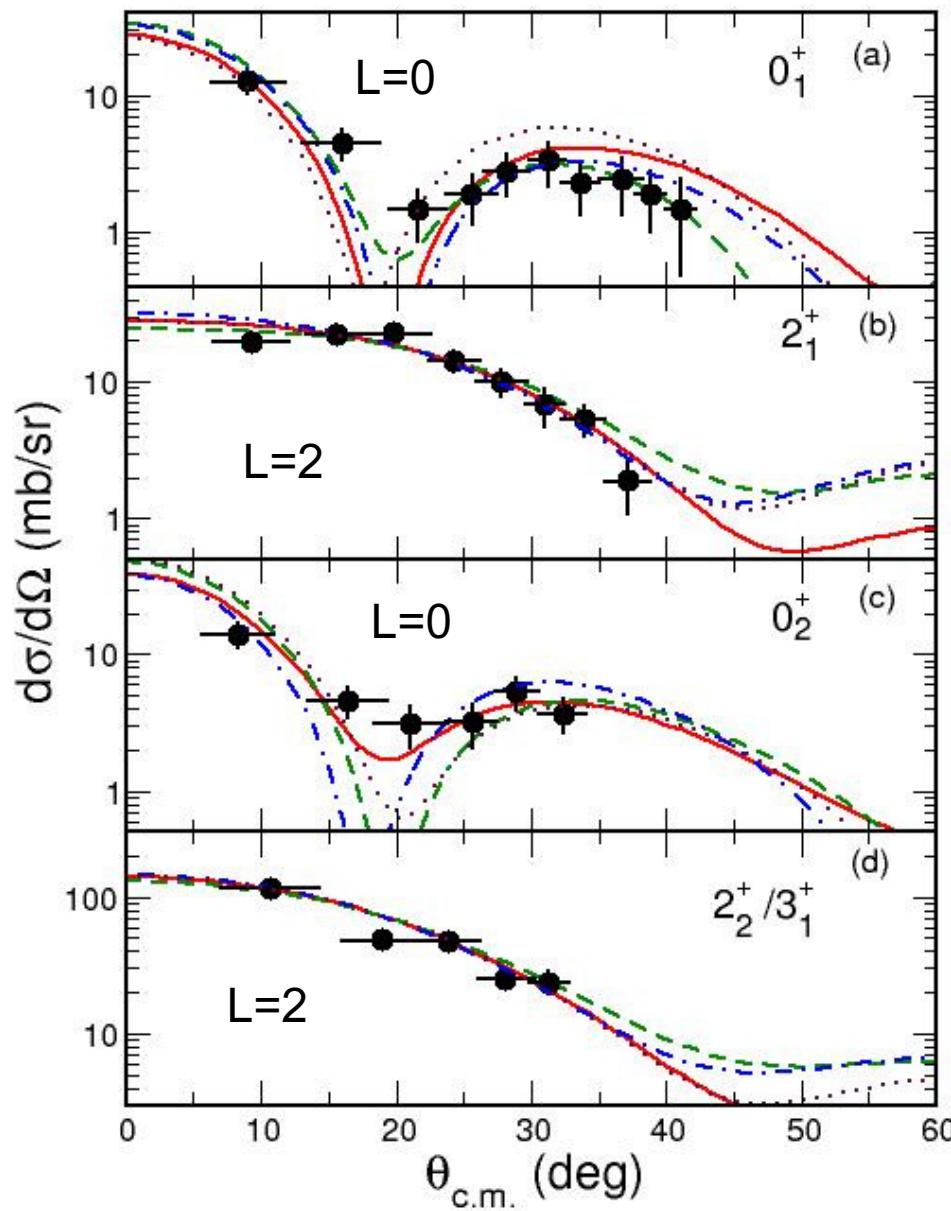


$^{15}\text{C}(\text{d},\text{p})^{16}\text{C}$
with HELIOS

Proton energy-position
correlation

(d,p) samples the
 $\nu(1s_{1/2})$ content of
the wave functions
for positive-parity states

^{16}C Excitation-energy
spectrum



$^{15}\text{C}(d,p)^{16}\text{C}$ angular distributions

Curves are DWBA calculations with various optical-model potentials.

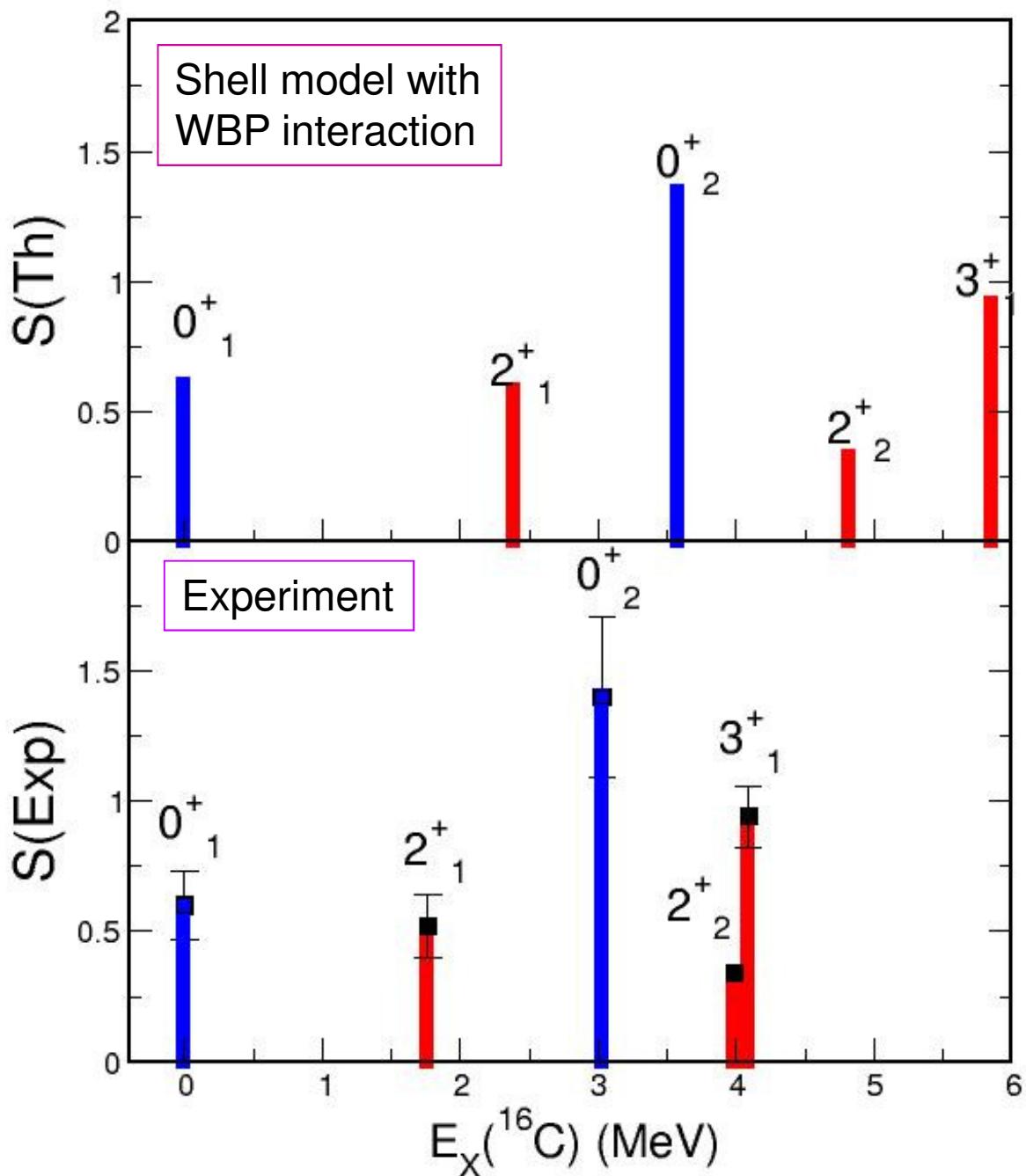
Spectroscopic factors obtained from the average over four sets of OMP.

Relative uncertainties in SF dominated by OMP variations
 Absolute uncertainty ($\sim 30\%$) from beam-integration uncertainty

$^{15}\text{C}(\text{d},\text{p})^{16}\text{C}$

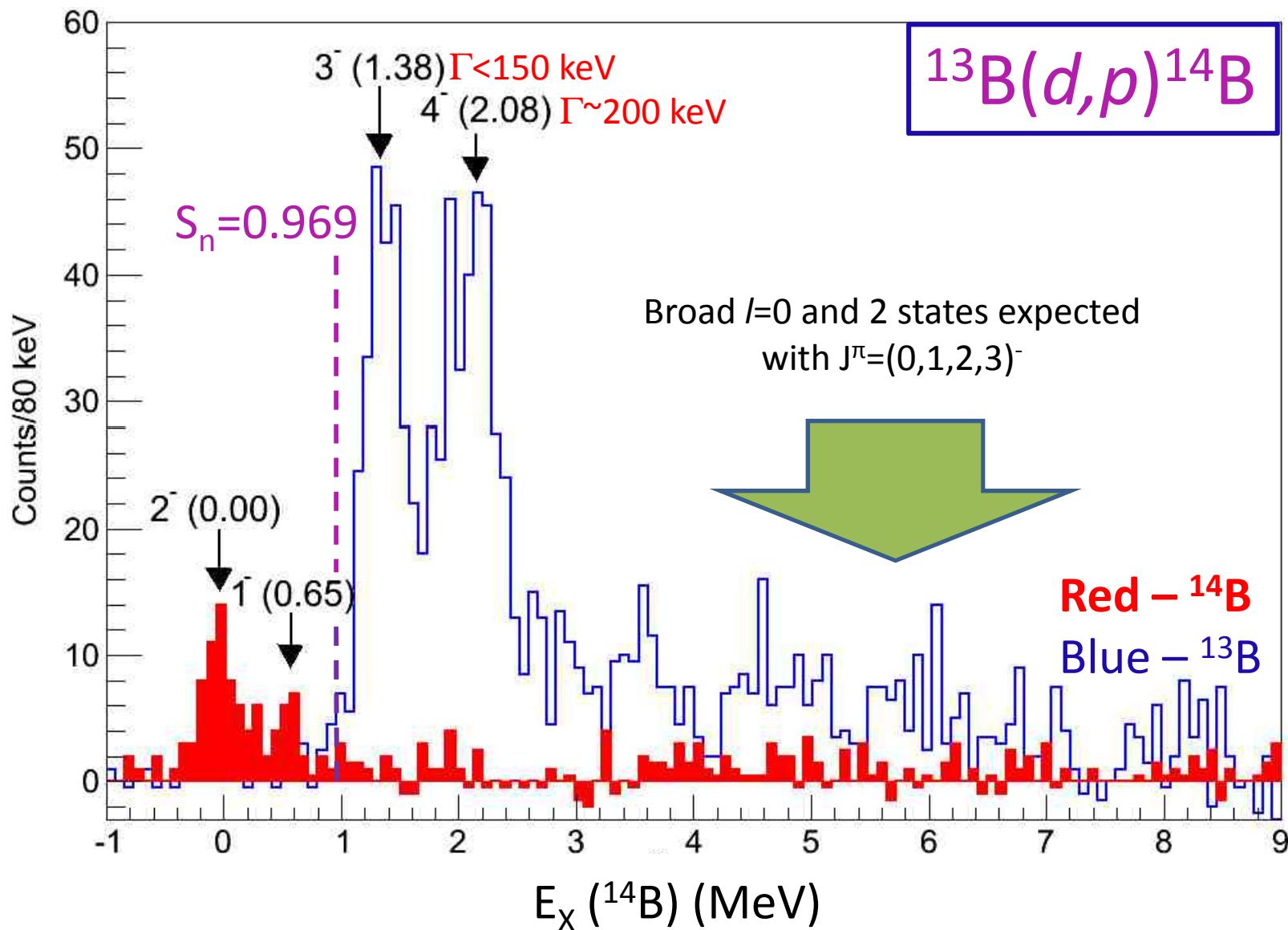
Spectroscopic factors

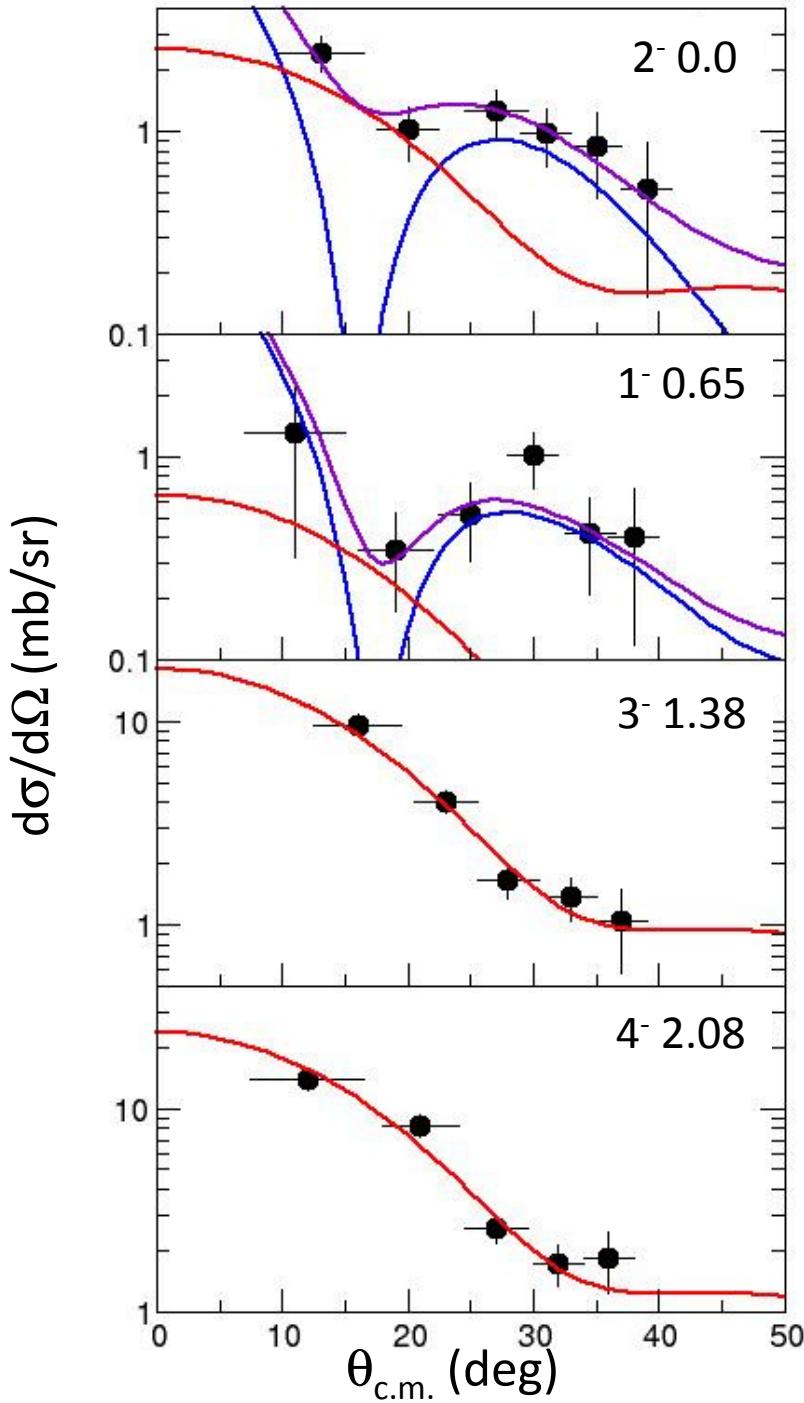
Excitation energies
and relative
spectroscopic
factors from the
shell model



Agreement for SF is excellent!
No need for exotica

Preliminary excitation-energy spectrum





Preliminary

$^{13}\text{B}(d,p)^{14}\text{B}$ angular distributions

Blue: $L=0$

Red: $L=2$

Violet: $L=0 + L=2$

$2^-(0.00)$: $S_0 = .71$ $S_2 = .17$

$1^-(0.65)$: $S_0 = 0.96$ $S_2 = .06$

$3^-(1.38)$: $S_2 = 1.00$ (fixed)

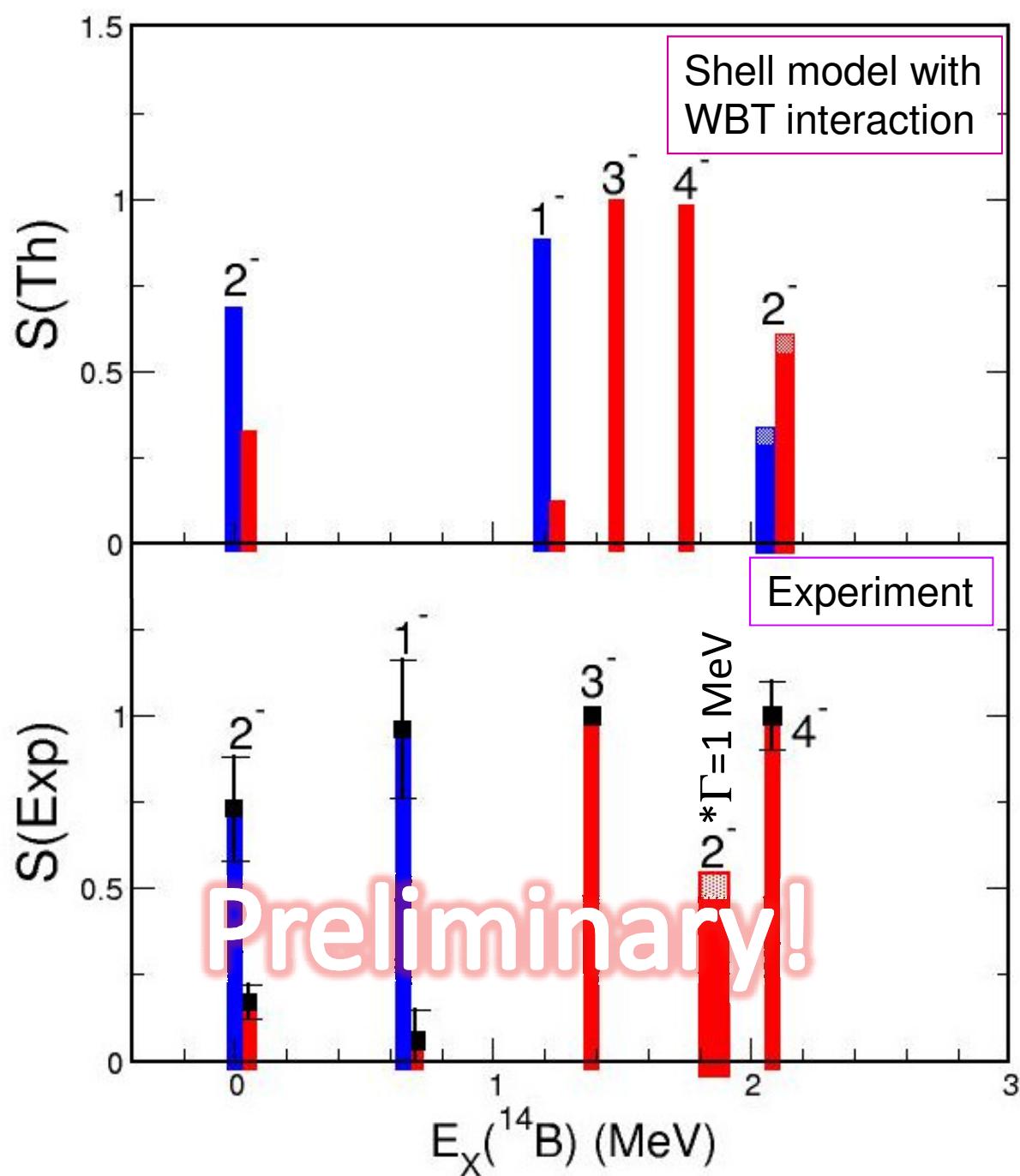
$4^-(2.08)$: $S_2 = 1.00$

OMPs fit 30 MeV $d+^{12}\text{C}$, $p+^{12,13}\text{C}$ elastic scattering

$^{13}\text{B}(d,p)^{14}\text{B}$ Spectroscopic factors

Excitation energies
and relative
spectroscopic
factors from the
shell model

Blue: $L=0, 1s_{1/2}$
Red: $L=2, 0d_{5/2}$



2⁻ mixed L=0+2,
1⁻ pure L=0
Reasonable agreement
But *caveat emptor!*

Summary

- HELIOS provides a new approach to studying reactions in inverse kinematics
- Alleviates problems with light particle identification and gives improved excitation-energy resolution and straightforward determination of CM quantities
- Around N=8, (d,p) nicely probes the evolution of the $1s_{1/2}$ - $0d_{5/2}$ orbitals and the p-n/n-n residual interactions
- $^{14}\text{B}(1^-)$ ($S_n = .319 \text{ MeV}$) is mostly s-wave, so is as good or better a halo state than $^{11}\text{Li}_{\text{g.s.}}$ or $^{11}\text{Be}_{\text{g.s.}}$.
- Structure aspects seem reasonably well in hand, BUT: we still worry about DWBA and weakly (or un-) bound s states.

Acknowledgements



The HELIOS Collaboration



S. Bedoor, J. C. Lighthall, S. T. Marley, D. Shetty, J. R. Winkelbauer
(SULI student), A. H. Wuosmaa

Western Michigan University



B. B. Back, S. Baker, C. M. Deibel, C. R. Hoffman, B. Kay, H. Y. Lee, C. J. Lister, P. Mueller, K.E. Rehm, J. P. Schiffer, K. Teh, A. Vann (SULI student)

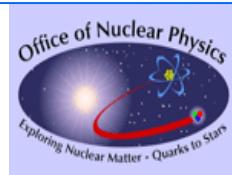
Argonne National Laboratory



S. J. Freeman
University of Manchester



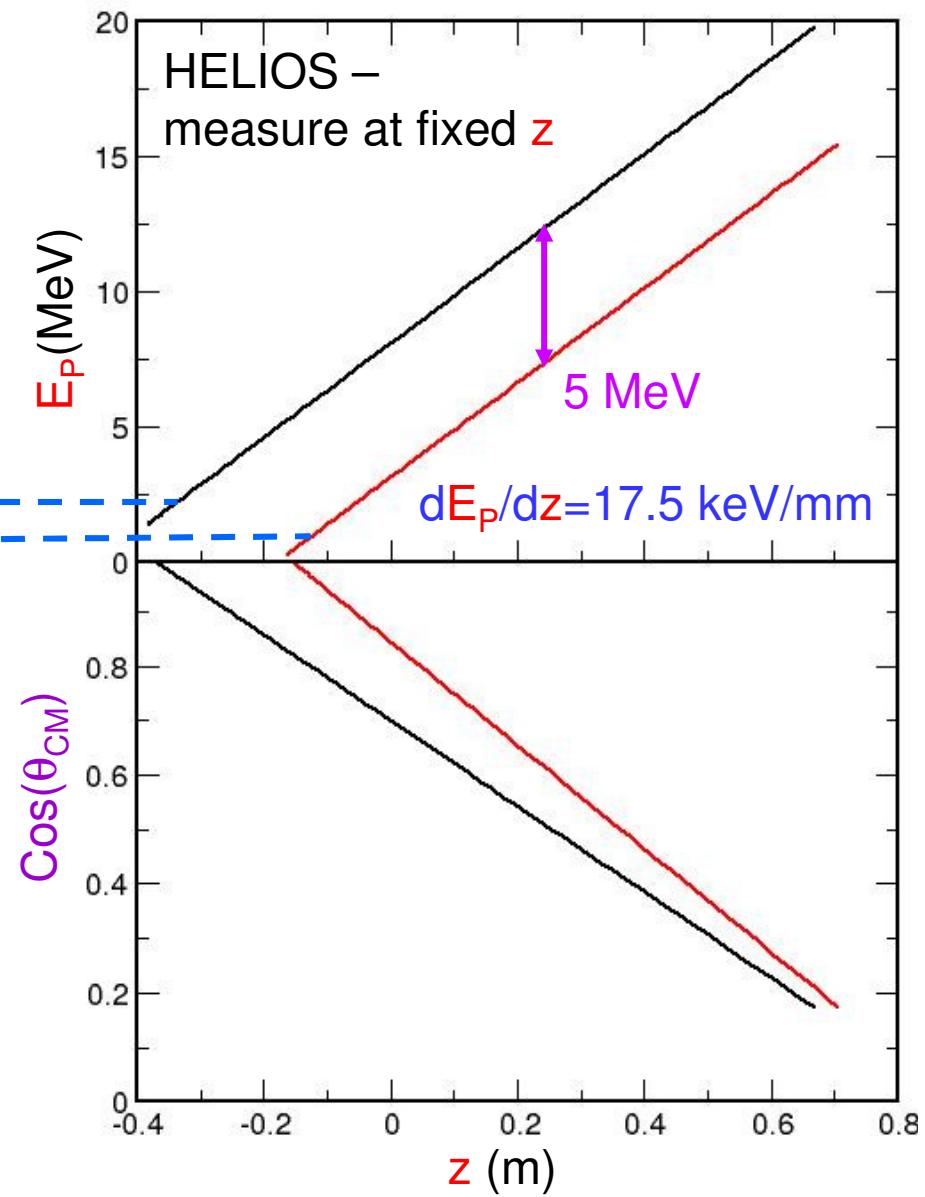
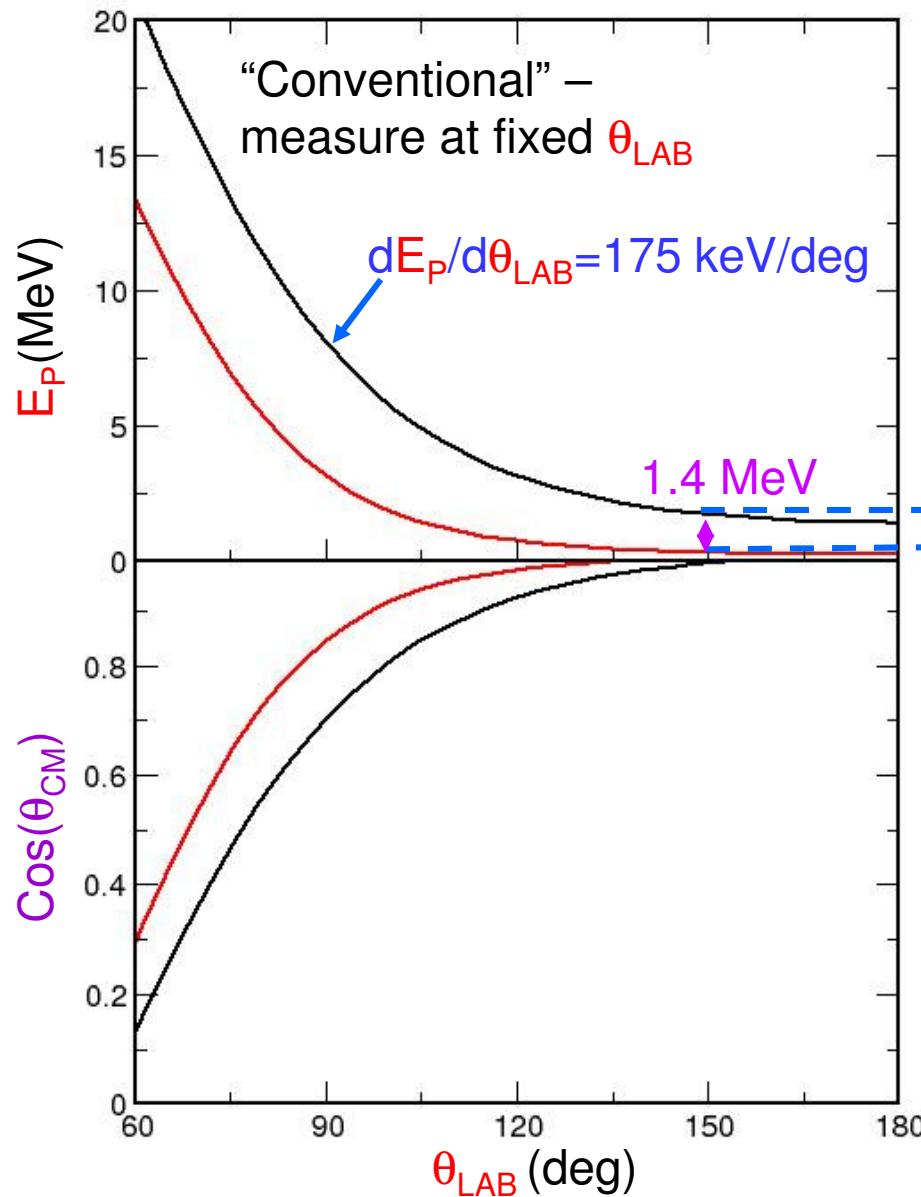
Work supported by the U. S. Department of Energy, Office of Nuclear Physics, under contract numbers DE-FG02-04ER41320 (WMU) and DE-AC02-06CH11357 (ANL)



Also, special thanks to:

N. Antler, Z. Grelewicz, S. Heimsath, J. Rohrer, J. Snyder

Advantages to the HELIOS approach for (d,p)



Empirical $v(sd)^2$ residual interaction for 0^+

$$|0_1^+ \rangle = \alpha |(1s_{1/2})^2\rangle + \beta |(0d_{5/2})^2\rangle$$

$$|0_2^+ \rangle = -\beta |(1s_{1/2})^2\rangle + \alpha |(0d_{5/2})^2\rangle$$

$$\alpha = \sqrt{S(0_1^+) \times [J_f] / [J_i]} = 0.55 \quad [J]=2J+1$$

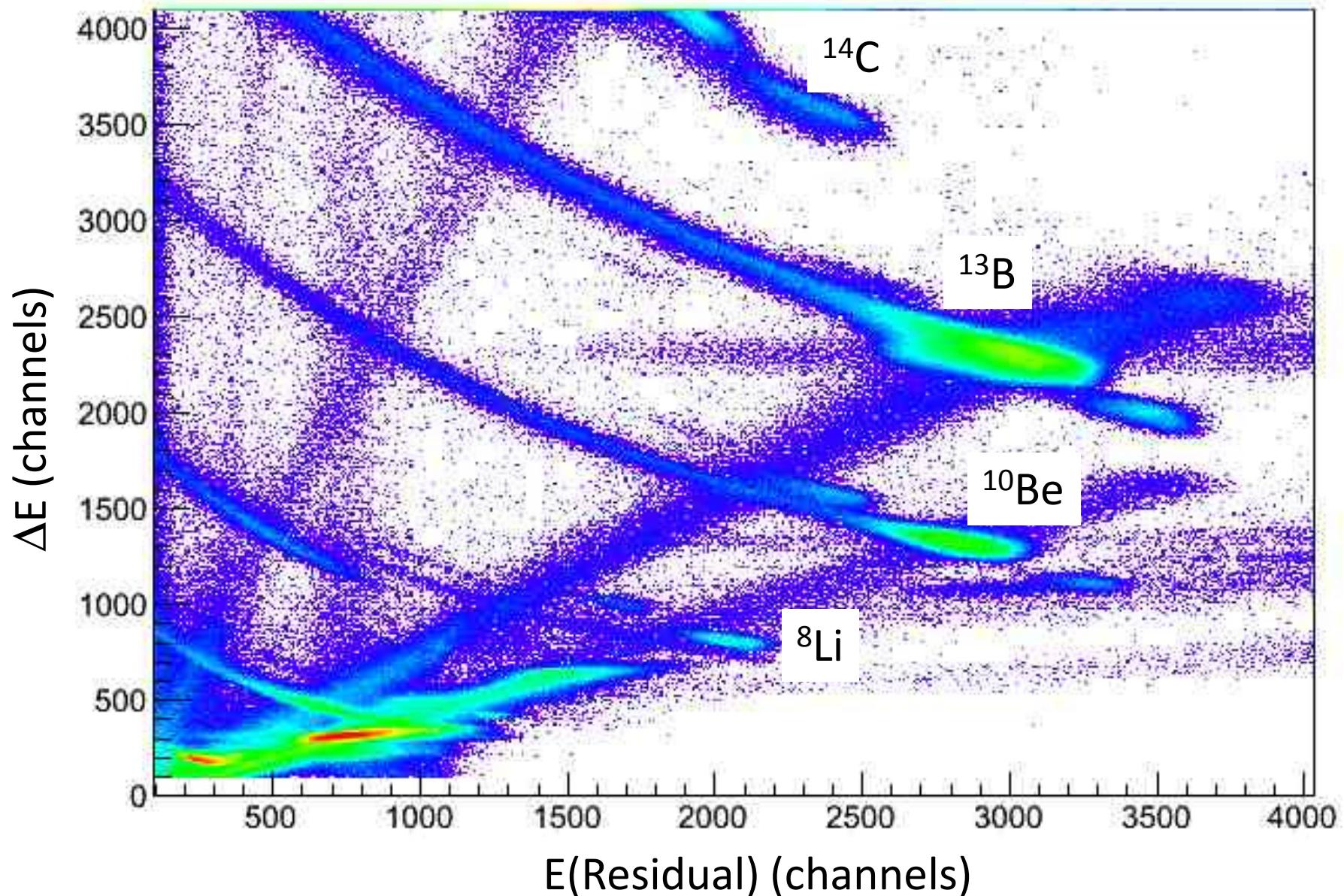
$$\beta = \sqrt{S(0_2^+) \times [J_f] / [J_i]} = 0.84$$

$$\begin{pmatrix} E_{1/2}^0 + \delta_{1/2;1/2} & \delta_{1/2;5/2} \\ \delta_{1/2;5/2} & E_{5/2}^0 + \delta_{5/2;5/2} \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \end{pmatrix} = E_x \begin{pmatrix} \alpha \\ \beta \end{pmatrix}$$

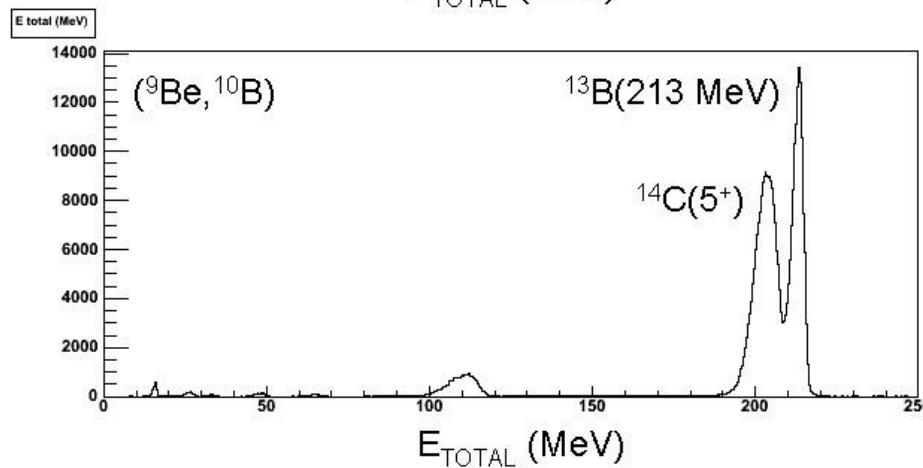
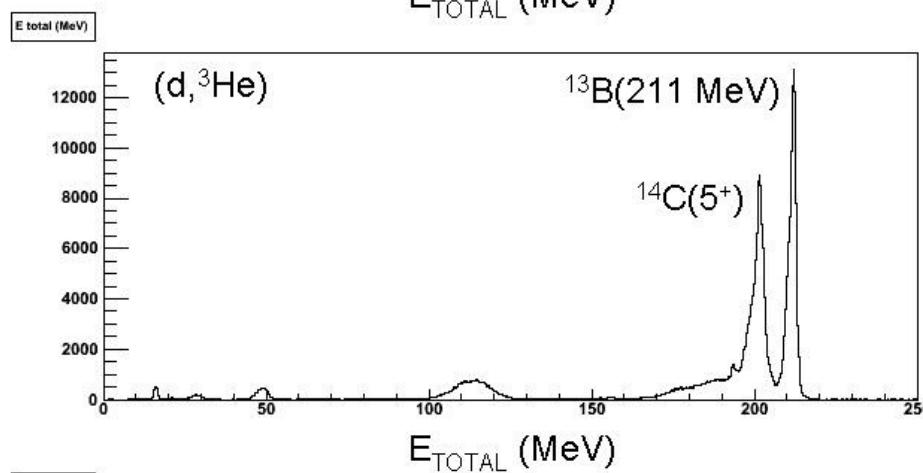
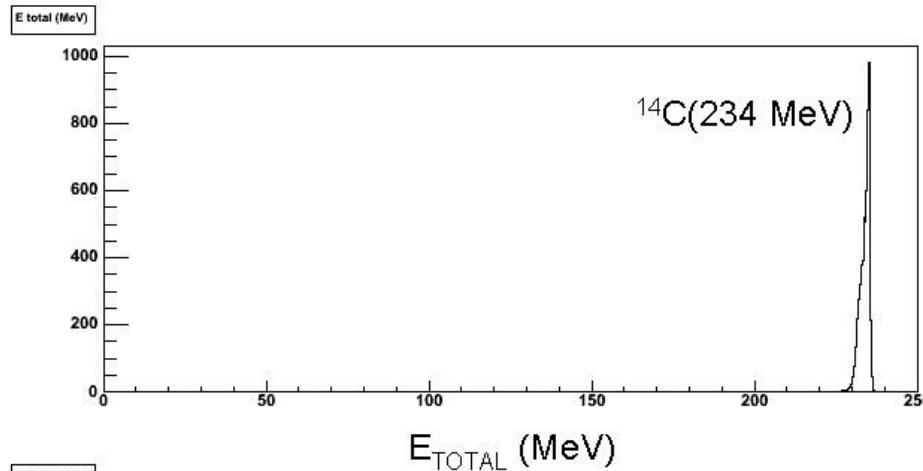
Single-particle energies E^0 from ^{15}C .

		$(j_1 j_2, j'_1 j'_2)$
$\langle j_1 j_2 v j'_1 j'_2 \rangle$		$(1/2 \ 1/2, 1/2 \ 1/2)$ $(5/2 \ 5/2, 5/2 \ 5/2)$ $(1/2 \ 1/2, 5/2 \ 5/2)$
Exp	-0.92(28)	-3.60(28) -1.39(12)
LSF	-1.54	-2.78 -1.72
WBP	-2.12	-2.82 -1.32

^{13}B and friends... from $^{14}\text{C} + ^9\text{Be}$



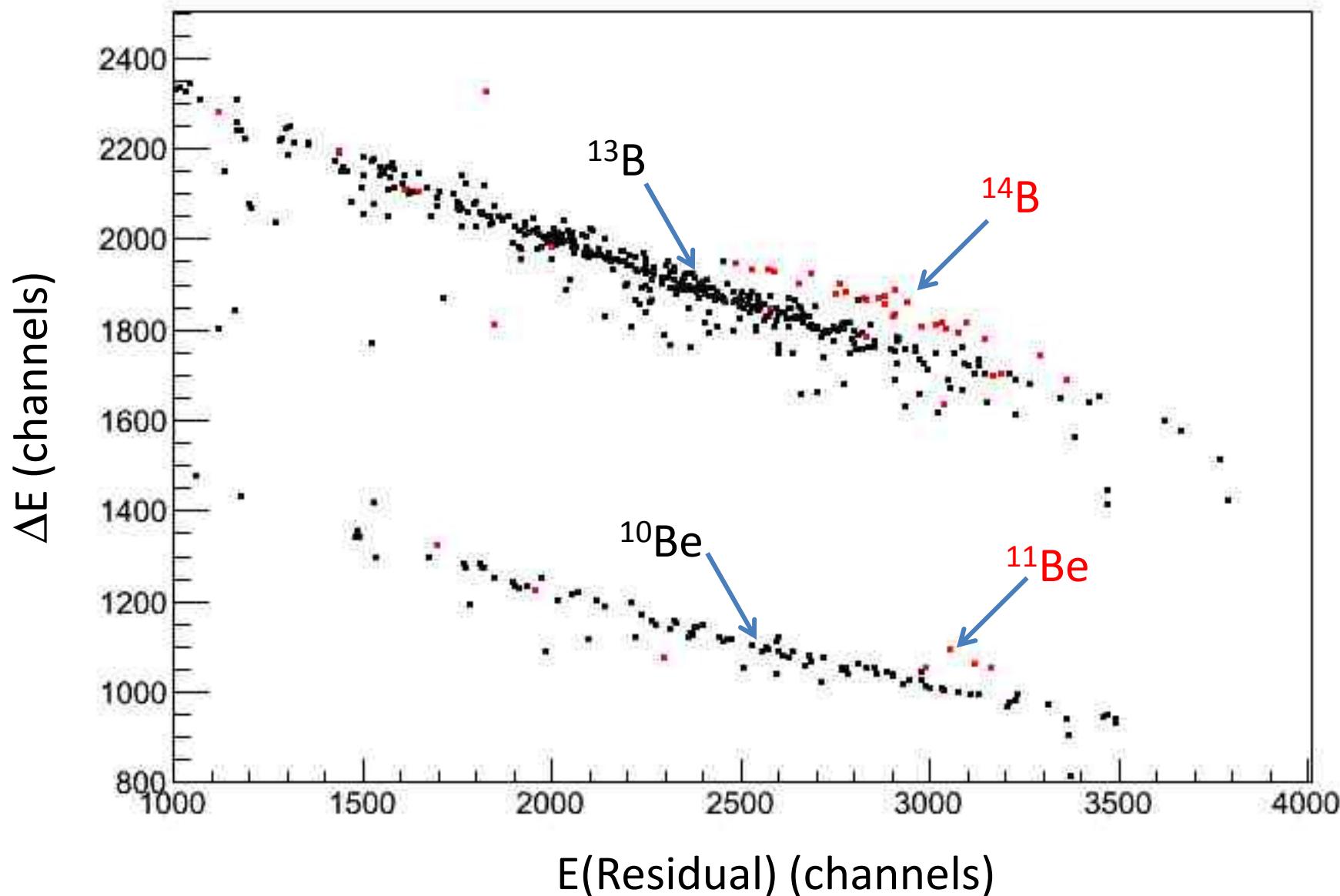
^{13}B beam quality



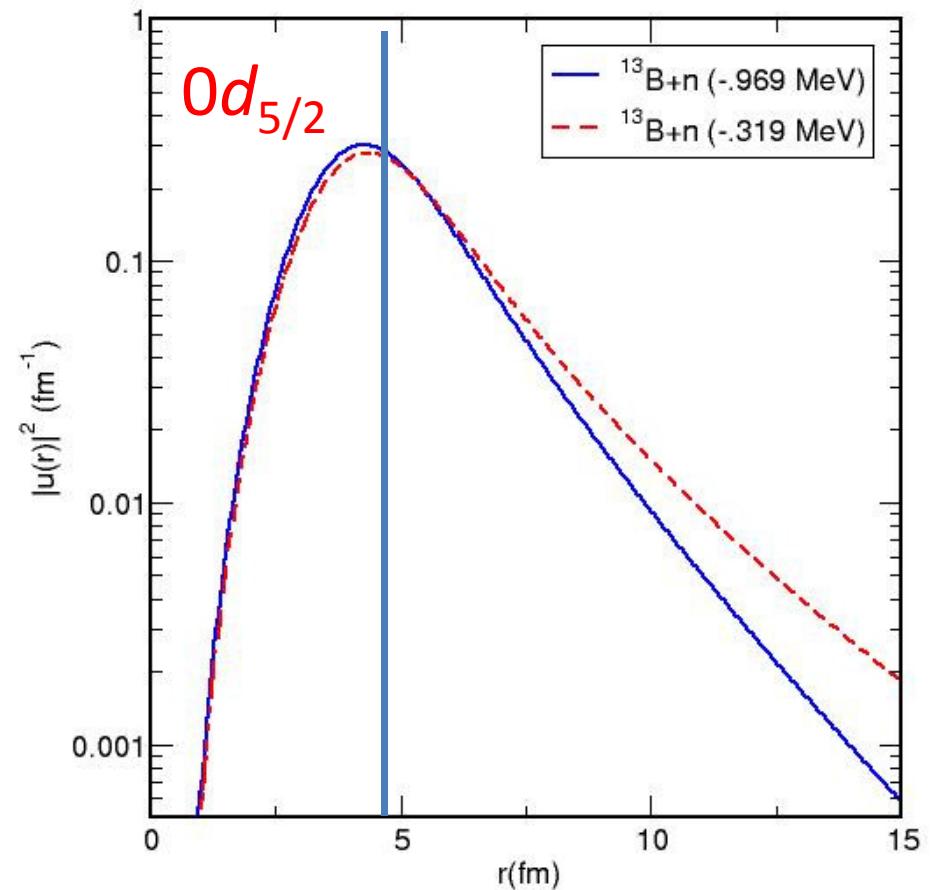
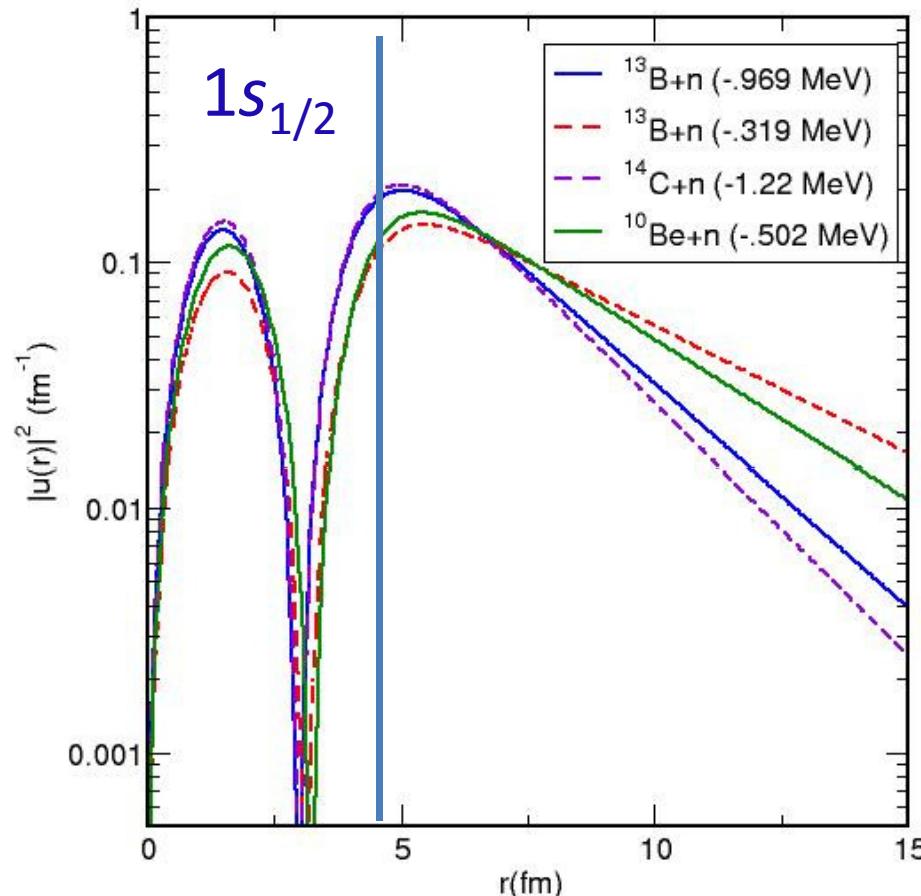
From $d(^{14}\text{C}, ^{13}\text{B})^3\text{He}$
 2×10^4 pps

From $^9\text{Be}(^{14}\text{C}, ^{13}\text{B})^{10}\text{B}$
 4×10^4 pps

Recoil particle identification



$1s_{1/2}$ and $0d_{5/2}$ neutron form factors

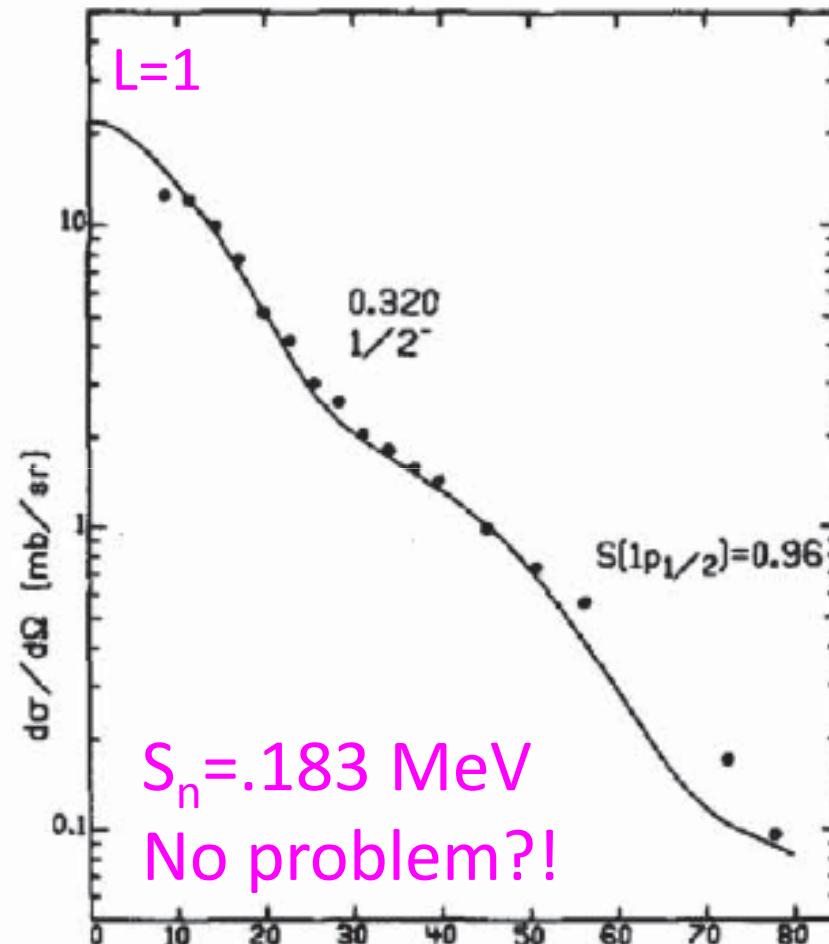
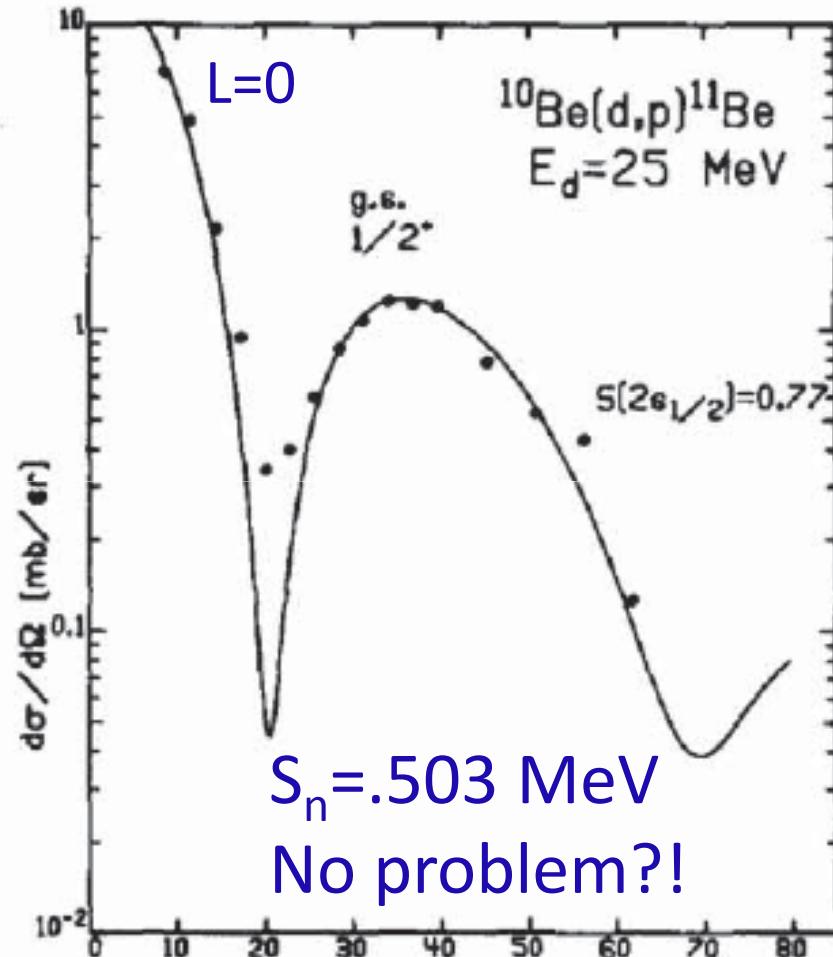


Woods-Saxon potential, $r_0=1.35$, $a=.6$, V_0 adjusted for BE

s-wave tail may cause problems for DWBA!

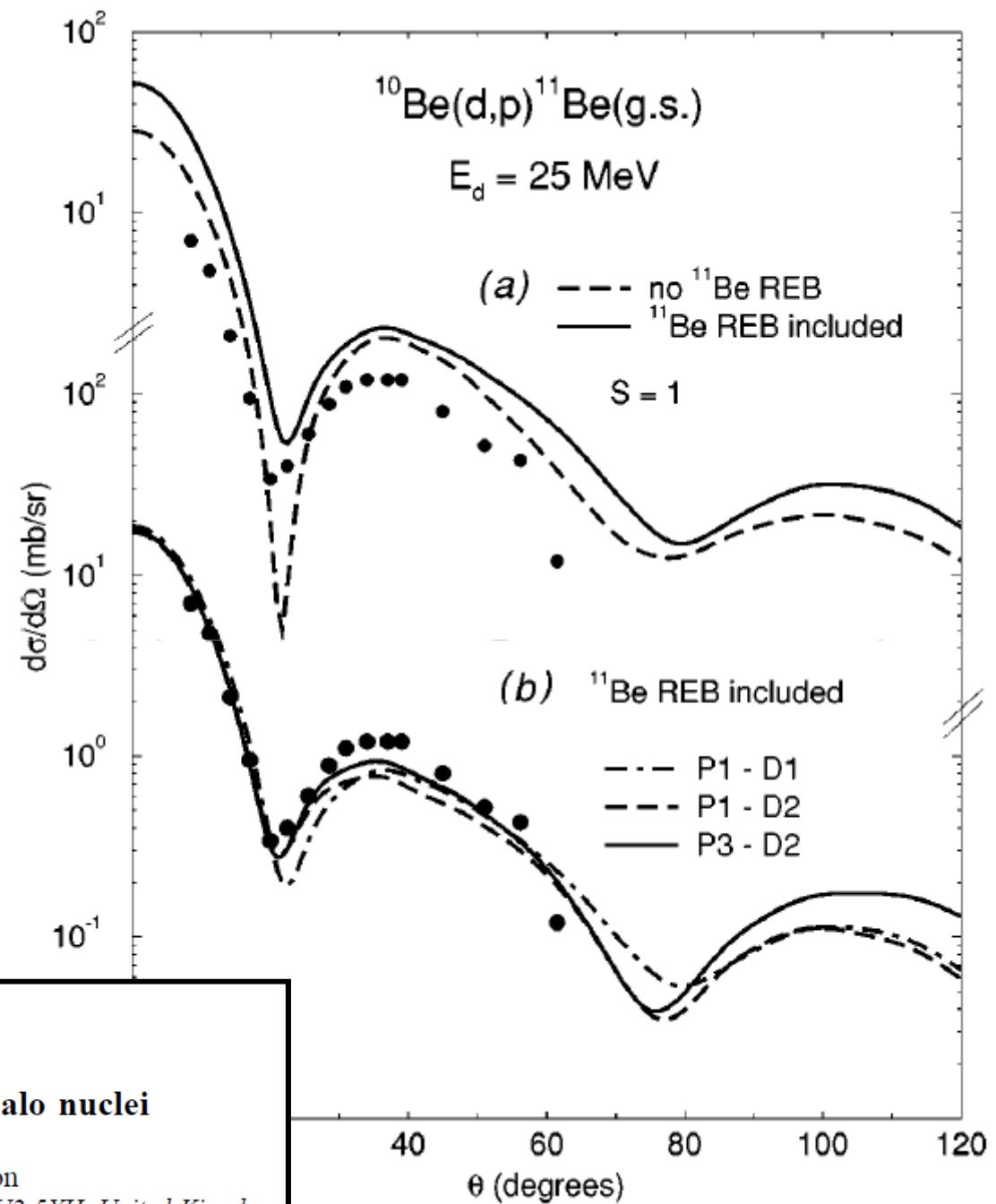
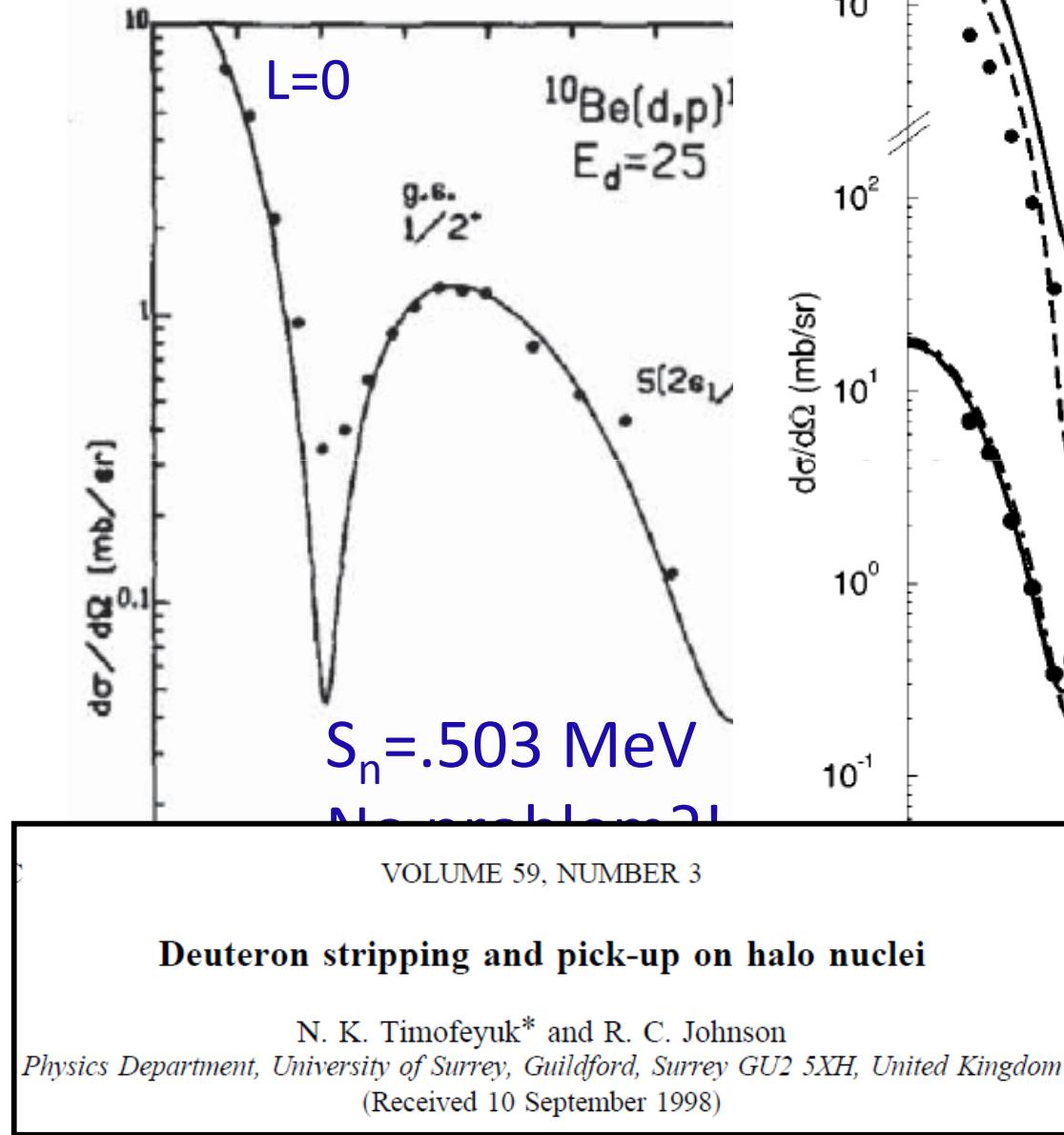
R. Huby, J. Phys. G 11, 931 (1985)

But: Don't forget history- $^{10}\text{Be}(\text{d},\text{p})^{11}\text{Be}$



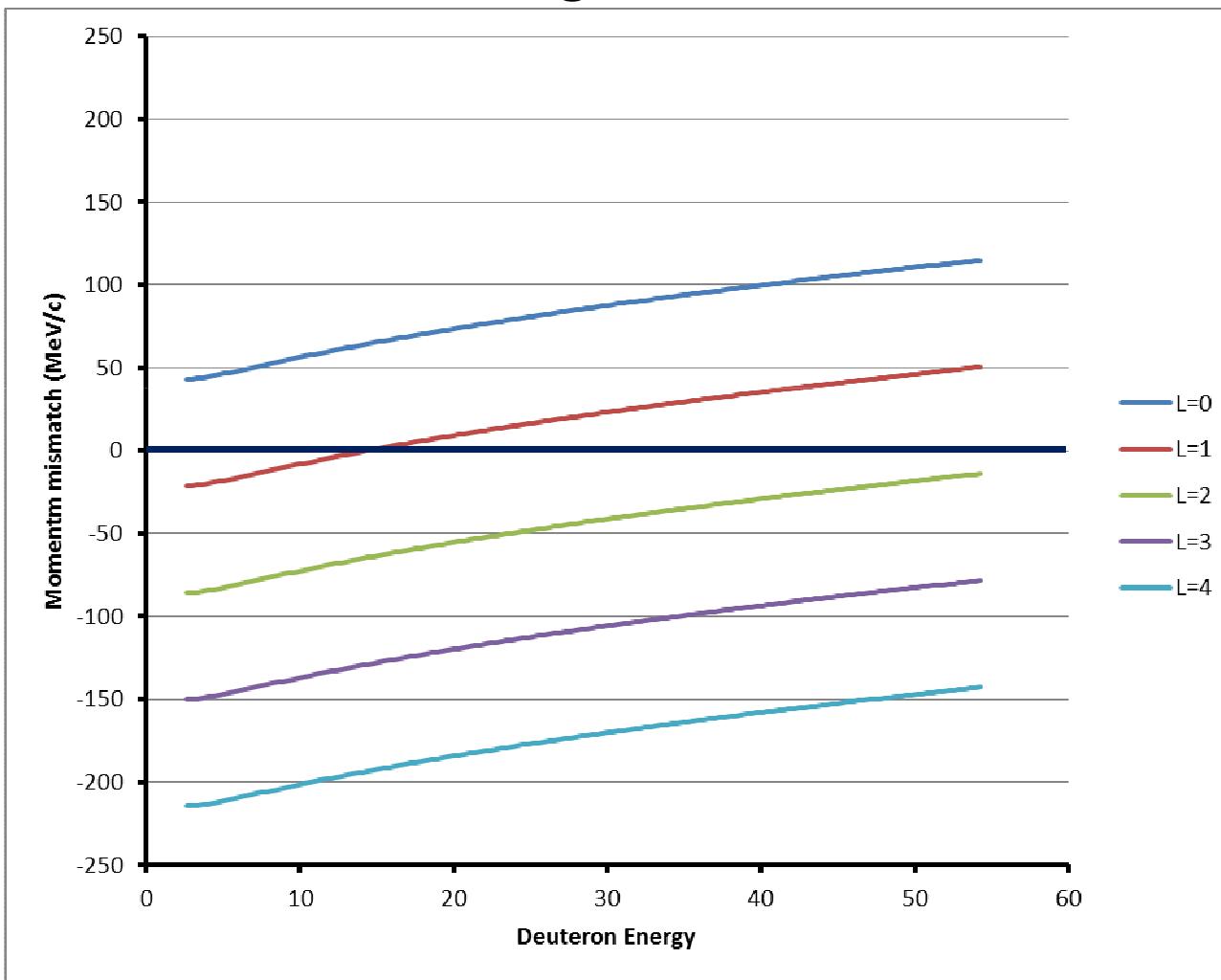
Zwieglinski, Benenson, Robertson, Coker – NP **A315**, 124 (1979)

But: Don't forget



(d,p) momentum mismatch at 0°

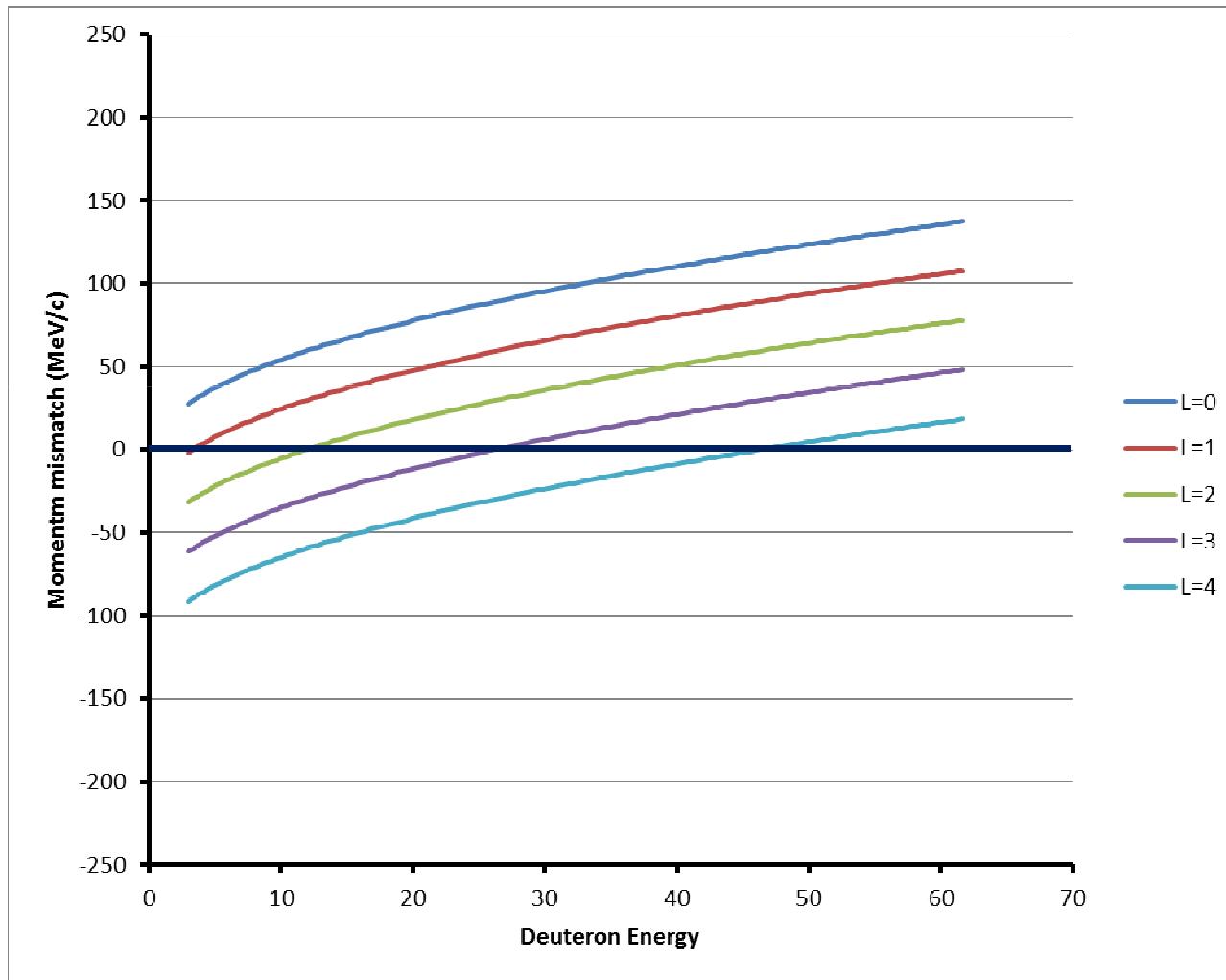
$(A_{tgt}=13)$



$\Delta q(1\hbar) \sim 65 \text{ MeV/c}$

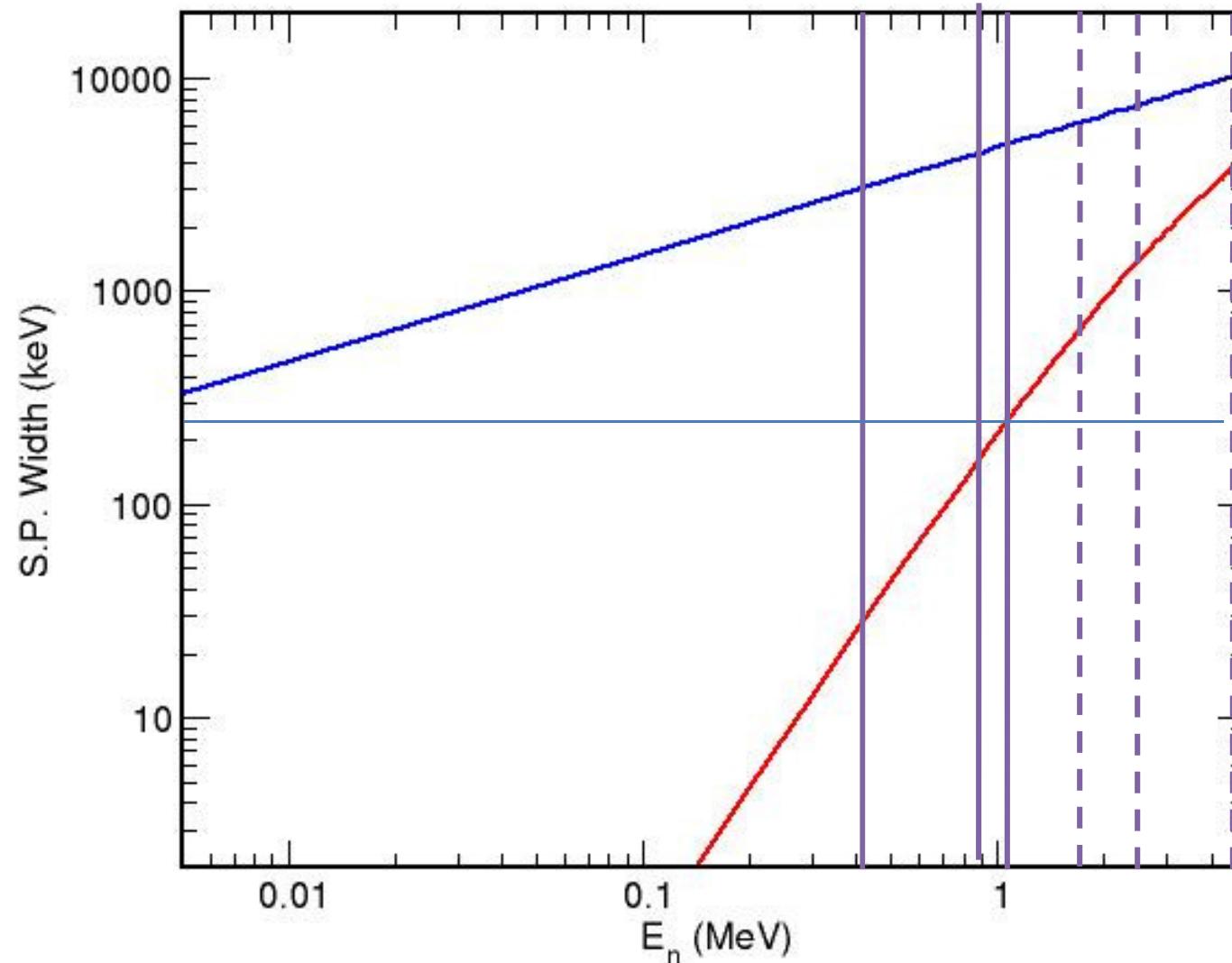
(d,p) momentum mismatch at 0°

$(A_{tgt}=132)$



$\Delta q(1\hbar) \sim 30 \text{ MeV/c}$

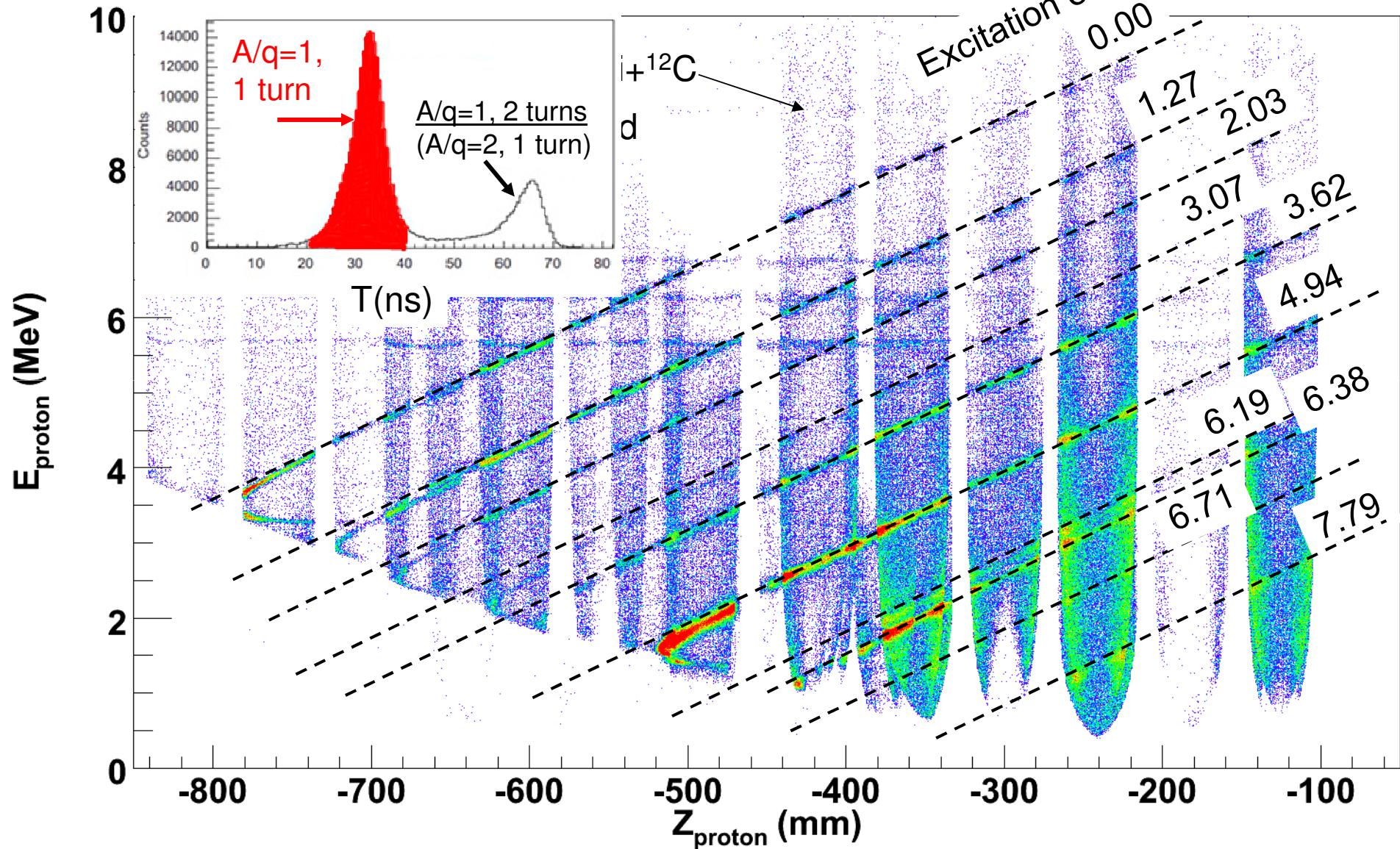
Single-particle widths for $^{13}\text{B}+\text{n}$



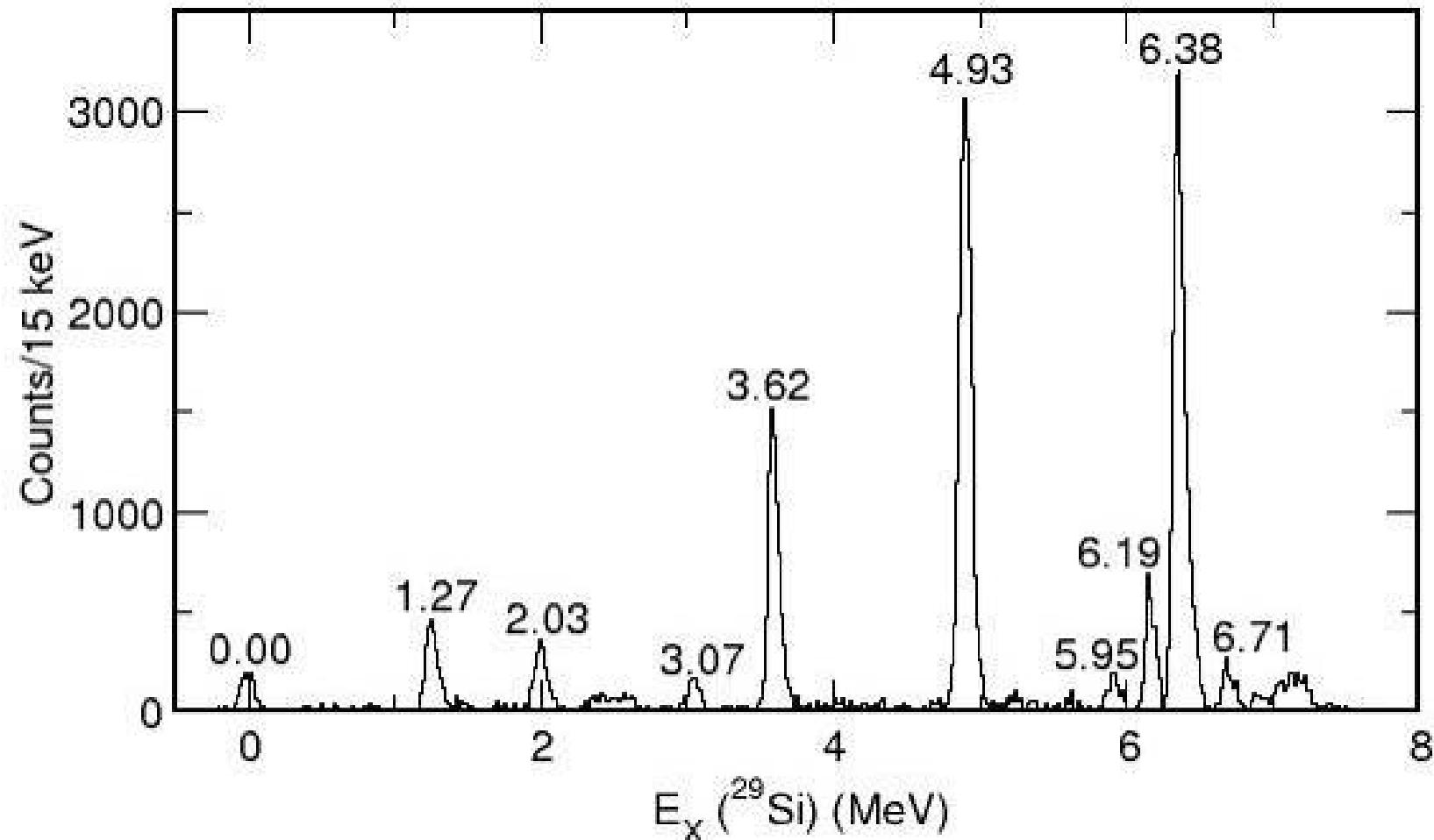
Spectrometer completed in August 2008



$^{28}\text{Si}(\text{d},\text{p})^{29}\text{Si}$: Seems to work!



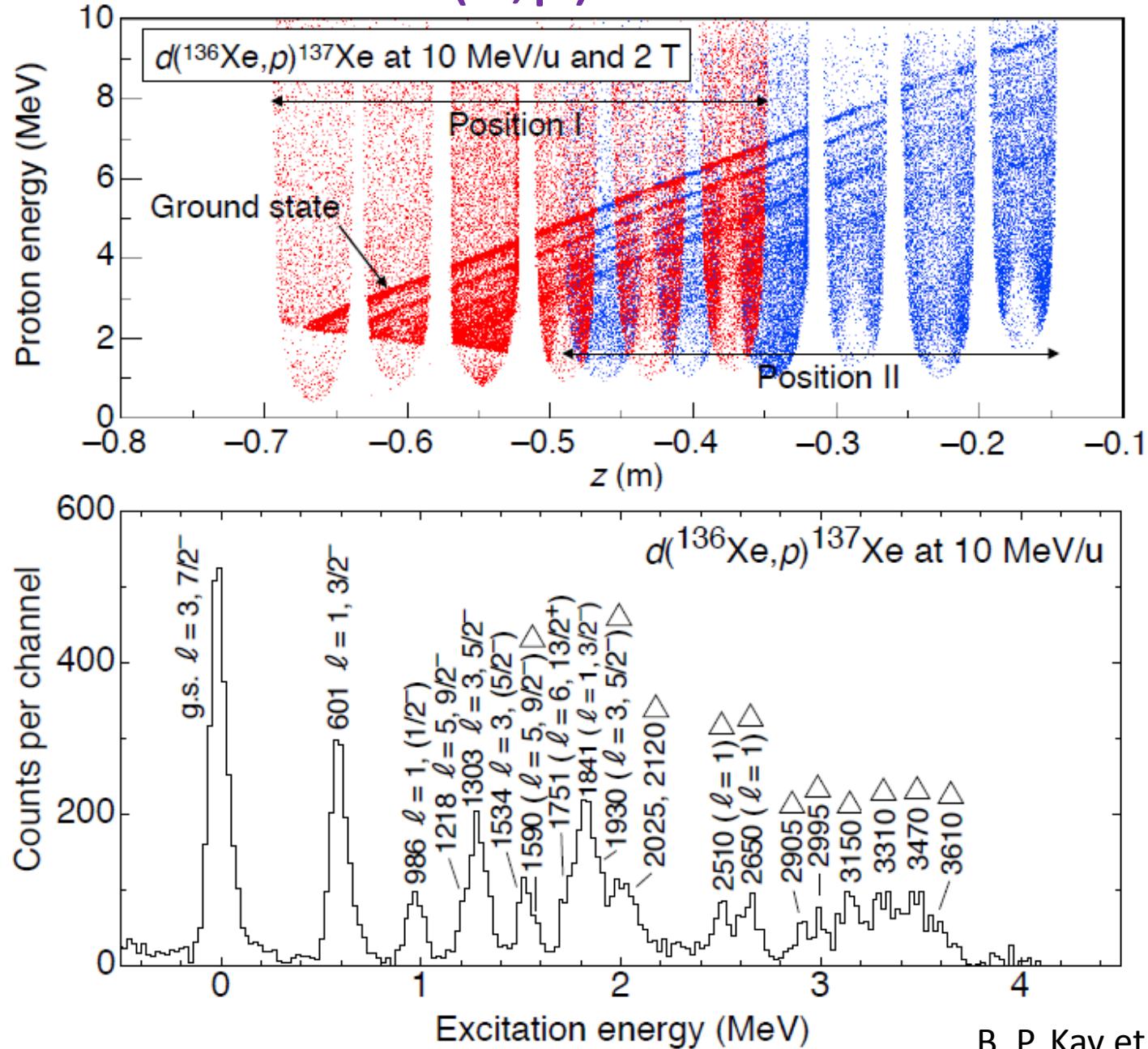
$^{28}\text{Si}(\text{d},\text{p})^{29}\text{Si}$ Excitation-energy spectrum



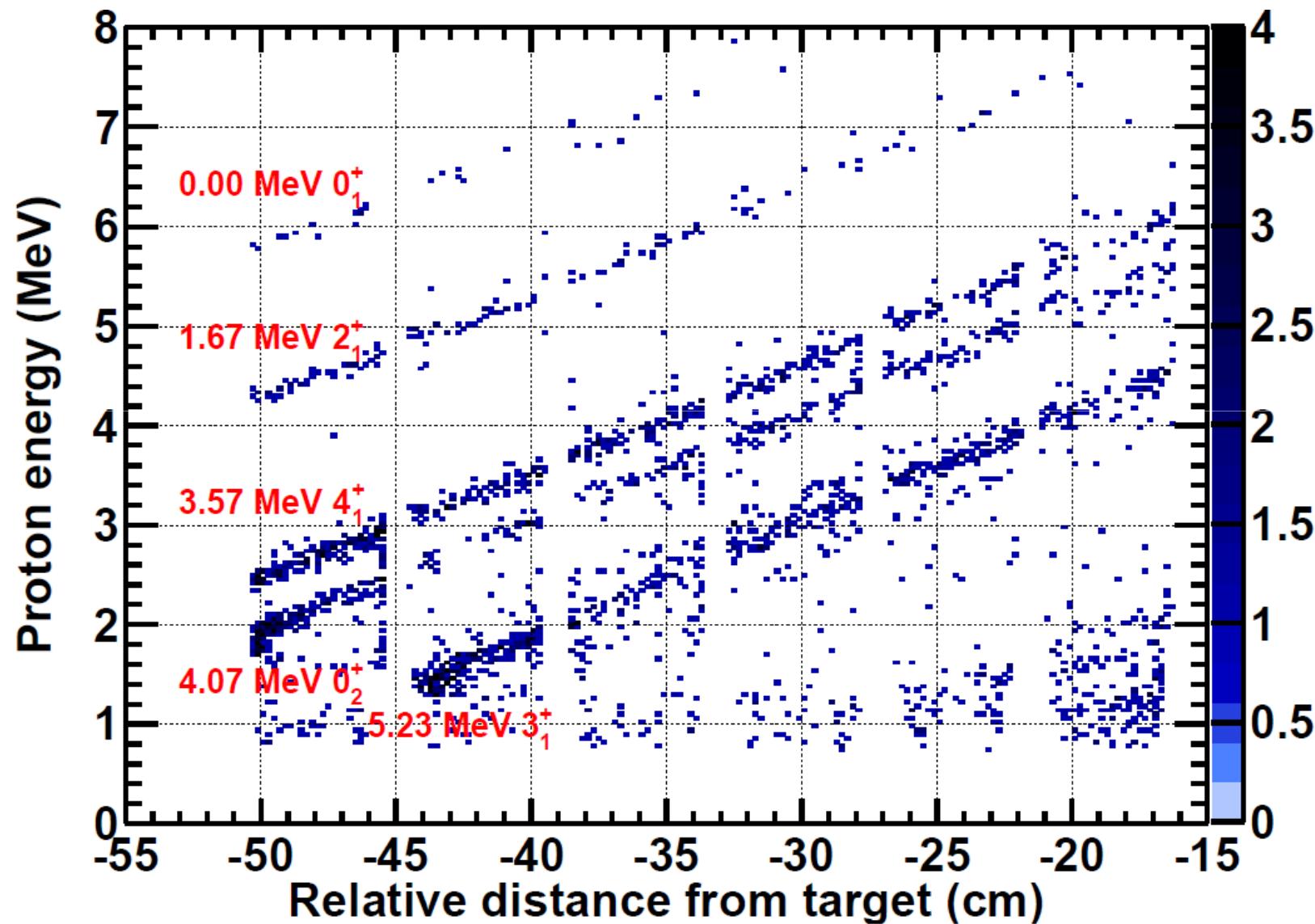
J. C. Lighthall et al,
NIMPRA 622, 97 (2010)

Typical resolution $\sim 120 \text{ keV FWHM}$
Best resolution $\sim 80 \text{ keV FWHM}$

Towards $^{132}\text{Sn}(\text{d},\text{p})^{133}\text{Sn}$ with CARIBU

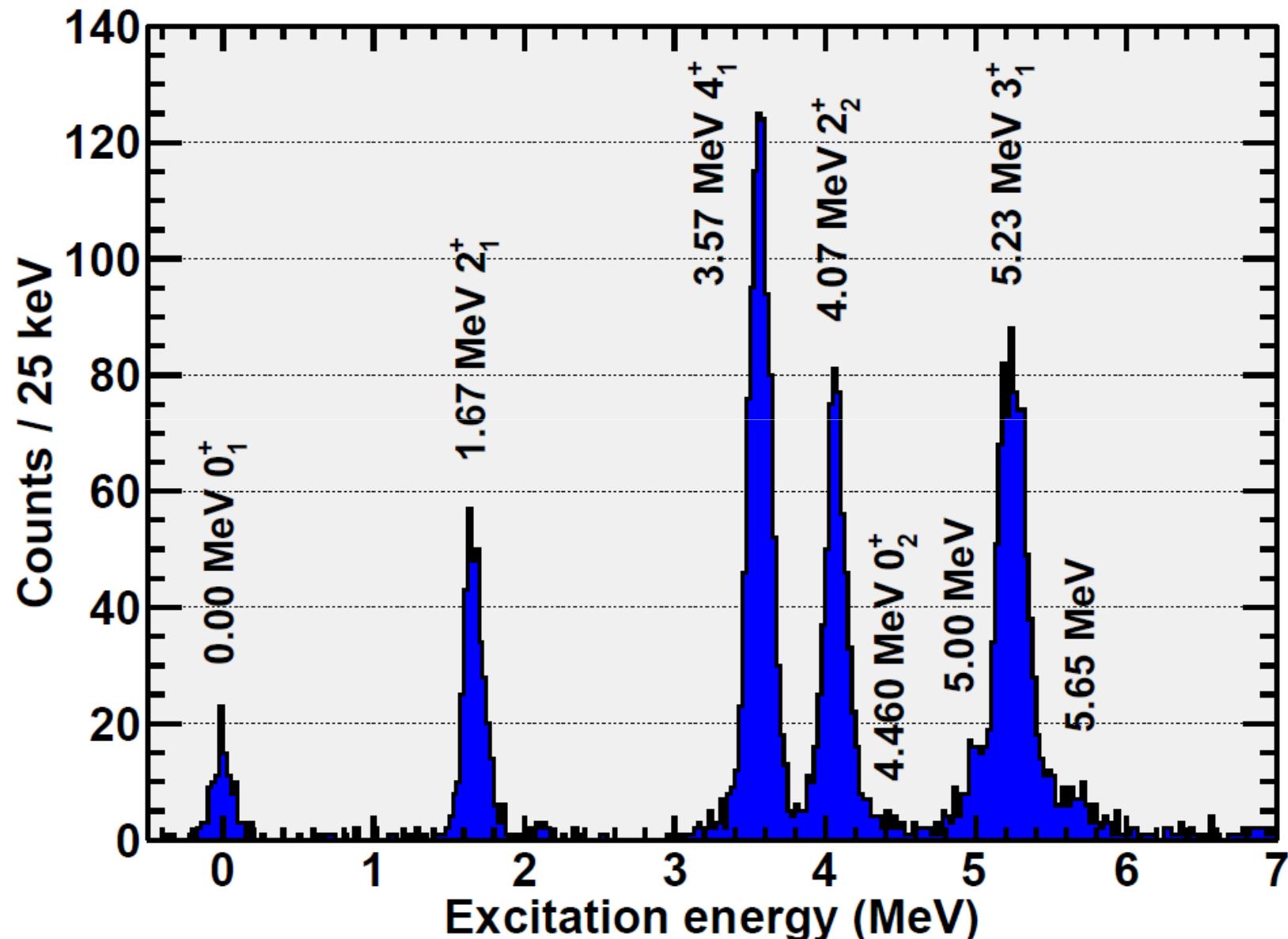


$^{19}\text{O}(\text{d},\text{p})^{20}\text{O}$: Further into the *sd* shell



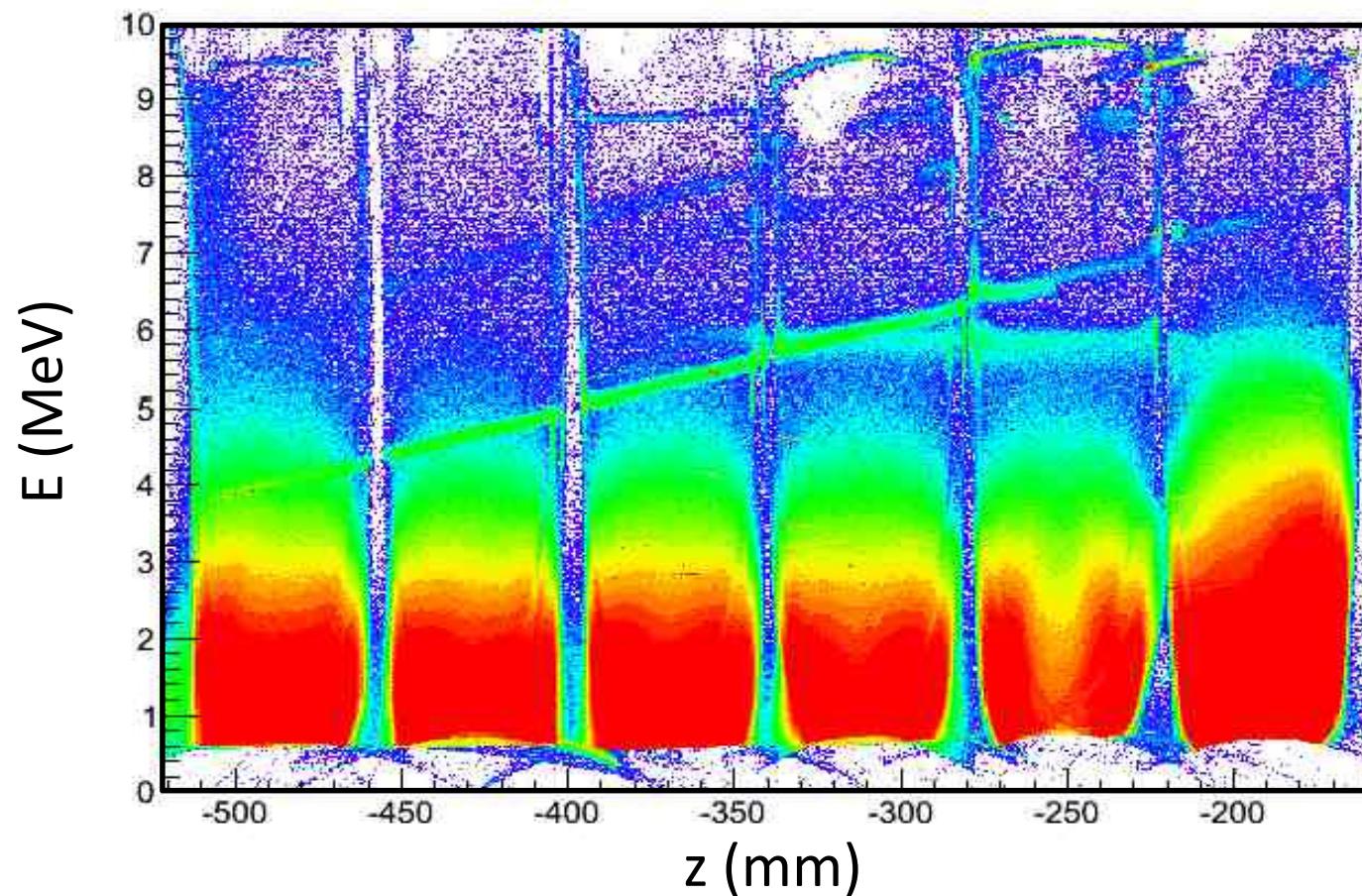
C. R. Hoffman, Submitted to PRC

$\nu(sd) + \nu(d_{5/2})^3$ $5/2_+$ states in ^{20}O

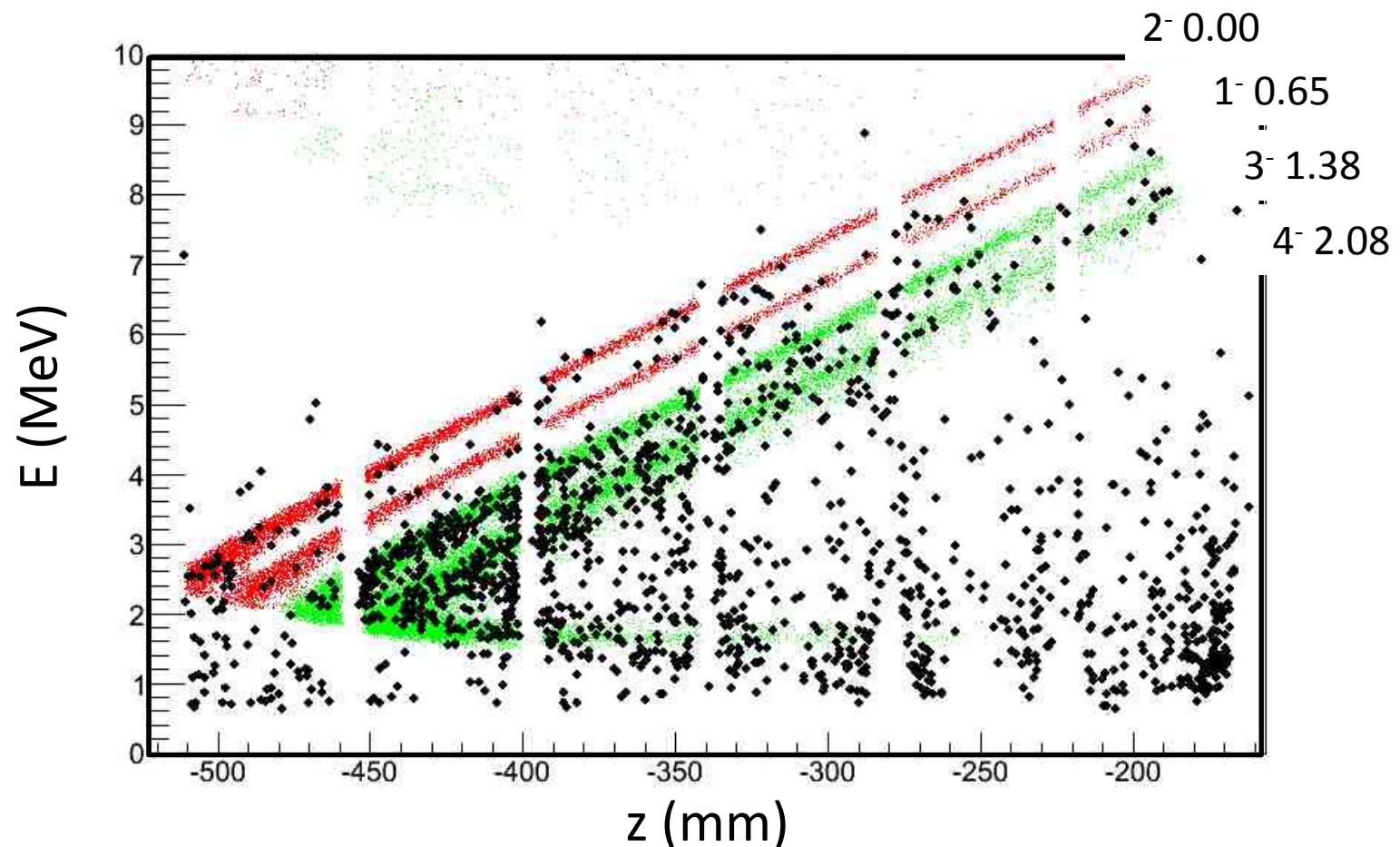


C. R. Hoffman, Submitted to PRC

Proton beam impurity: p-d elastics



E vs Z, data and Monte-Carlo



Red: n bound, p+ ^{14}B

Green: n-unbound p+ ^{13}B

$^{13}\text{B}(\text{d},\text{p})^{14}\text{B}$ in HELIOS

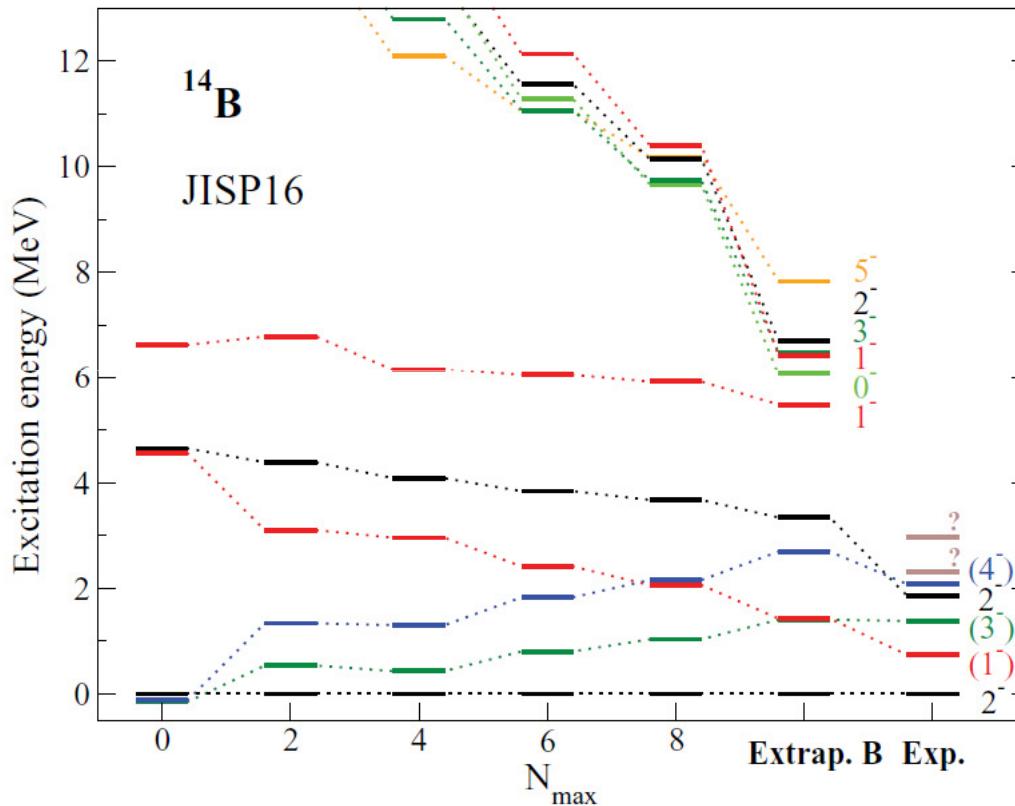
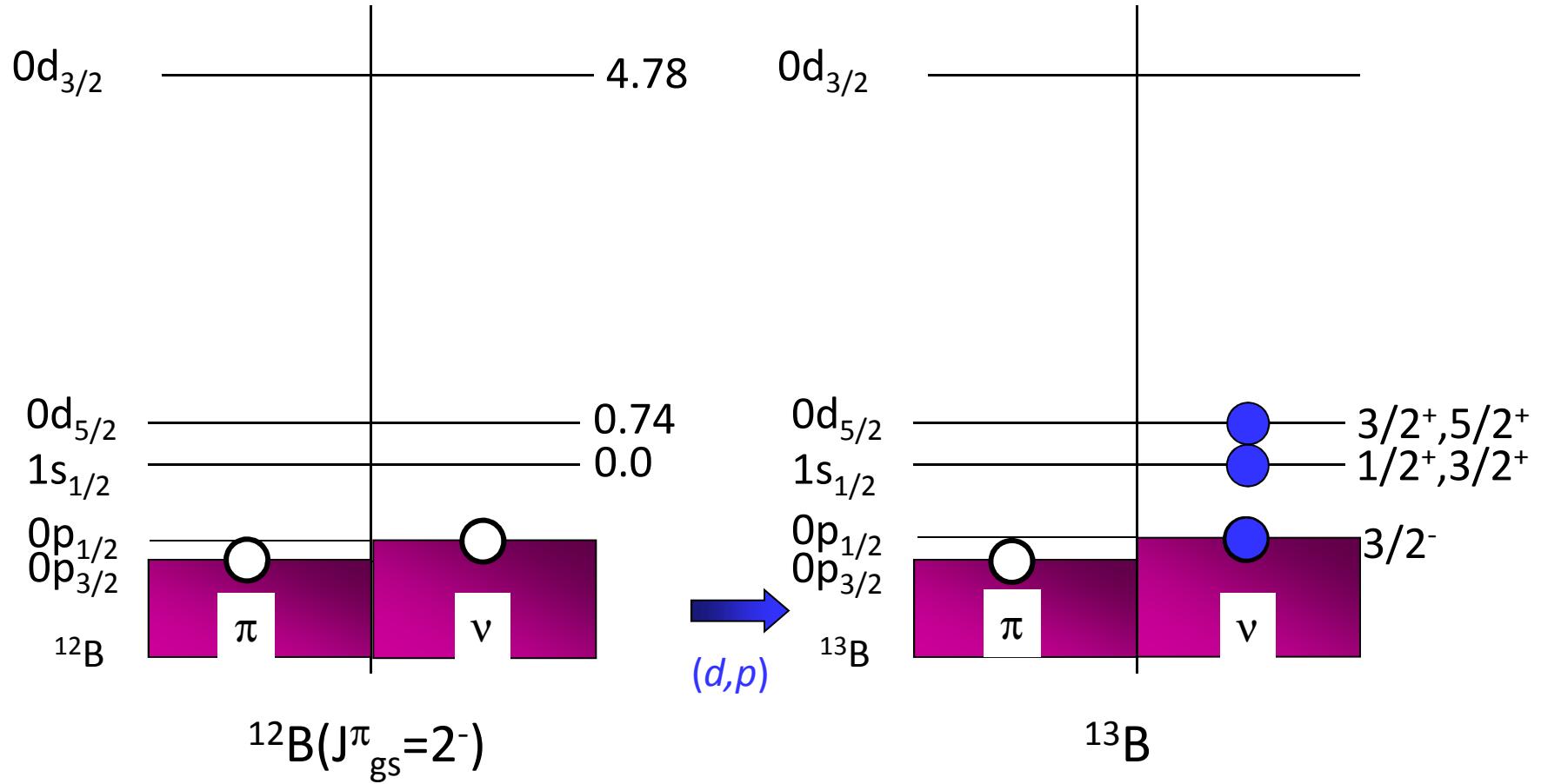
Ab initio nuclear structure simulations: The speculative ^{14}F nucleusP. Maris,¹ A. M. Shirokov,^{1,2,*} and J. P. Vary¹

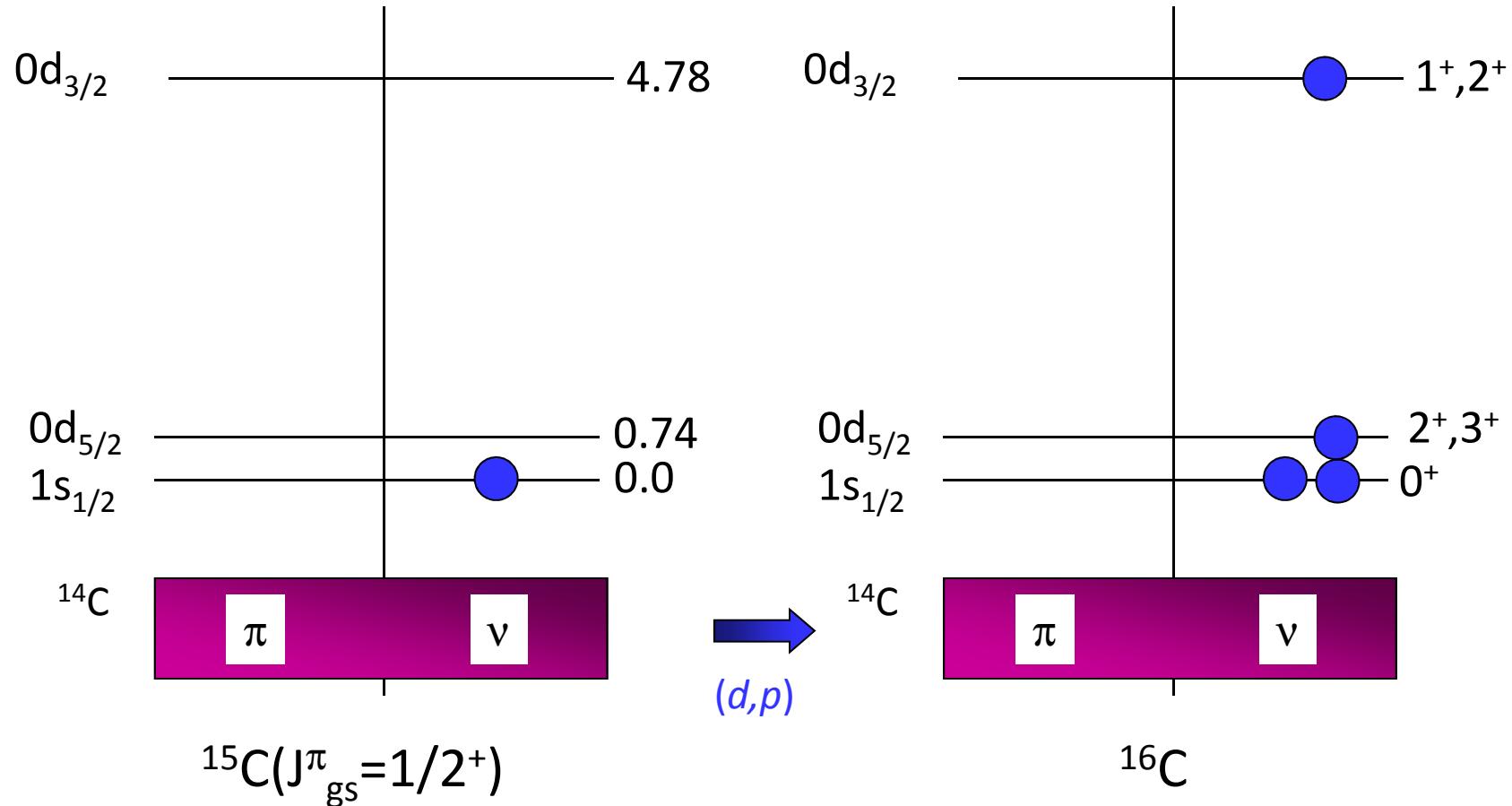
FIG. 3. (Color online) Negative-parity ^{14}B spectrum obtained with JISP16 at fixed $\hbar\Omega = 25$ MeV in successive basis spaces and extrapolated to infinite basis space using extrapolation B. Experimental (exp.) data are taken from Ref. [13].

Simple considerations for $^{12}\text{B}(d,p)^{13}\text{B}$



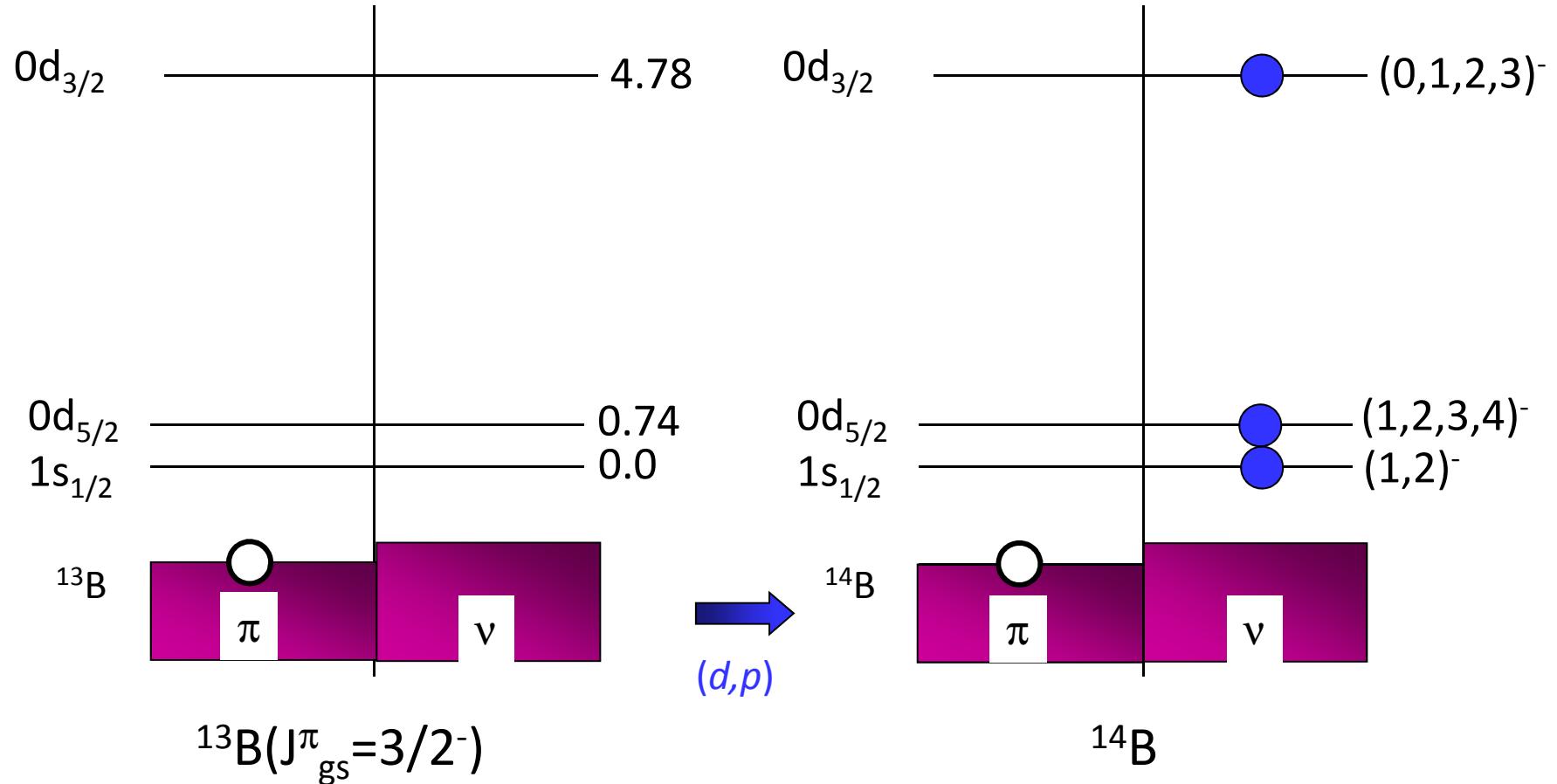
+ parity states are p-h excitations out of the p shell

Simple considerations for $^{15}\text{C}(d,p)^{16}\text{C}$

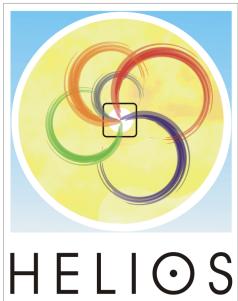


(d,p) samples $\nu(1s_{1/2})$ content of states in ^{16}C

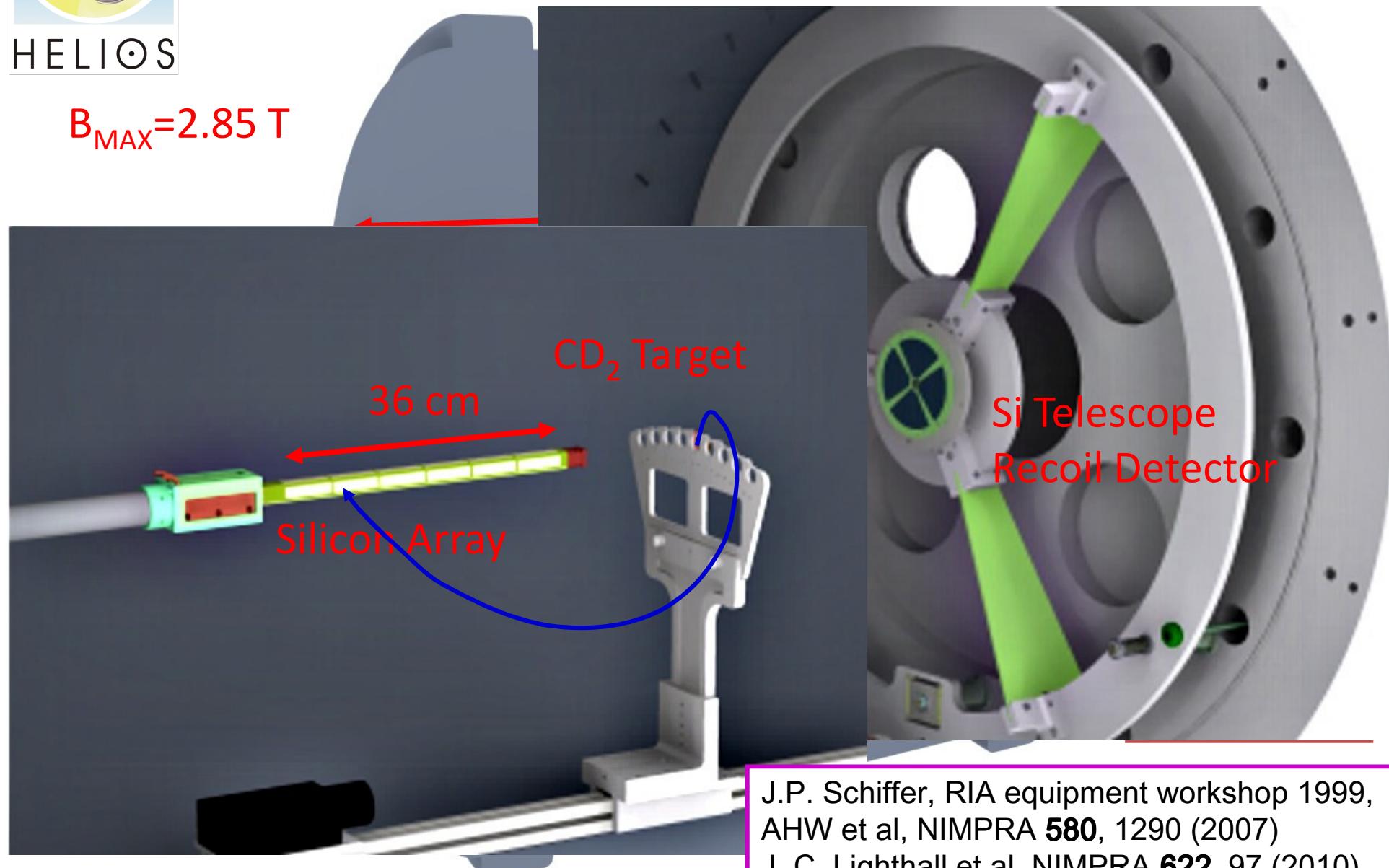
Simple considerations for $^{13}\text{B}(d,p)^{14}\text{B}$



(d,p) populates single-neutron states in ^{14}B



HELIcal Orbit Spectrometer -HELIOS

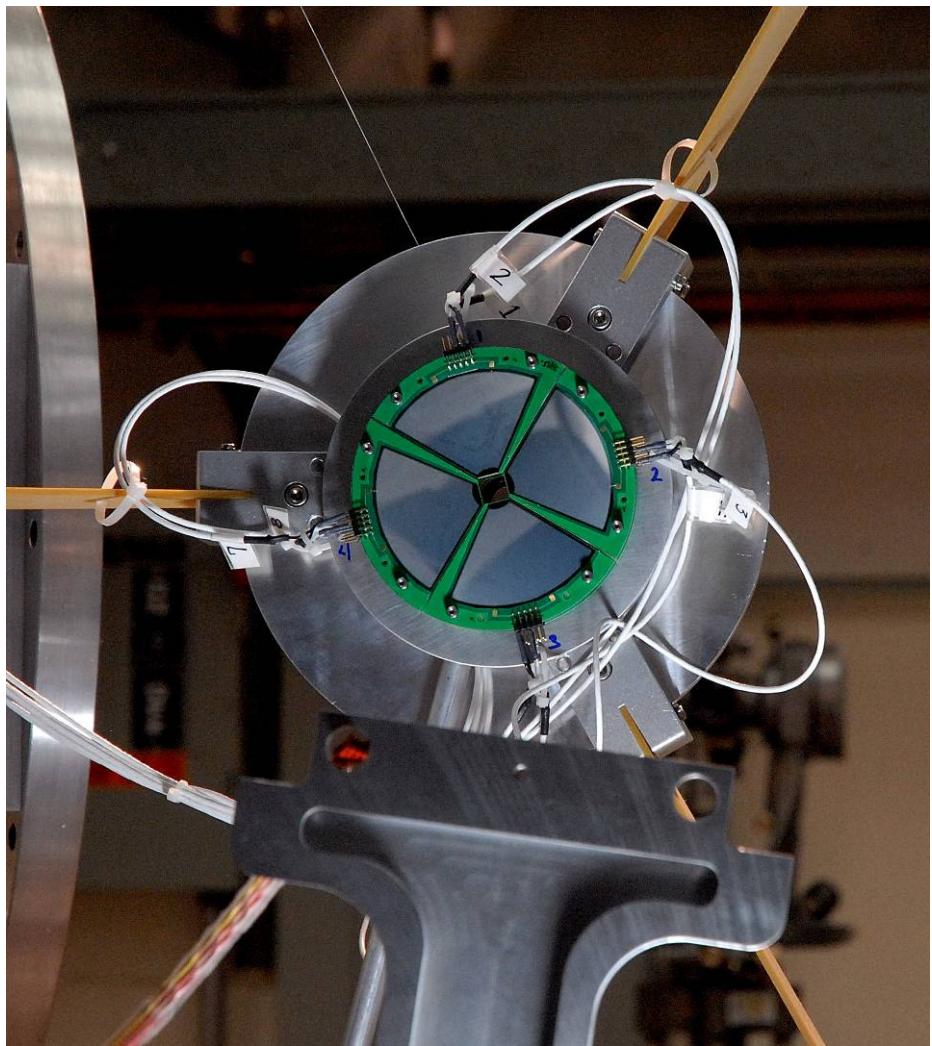
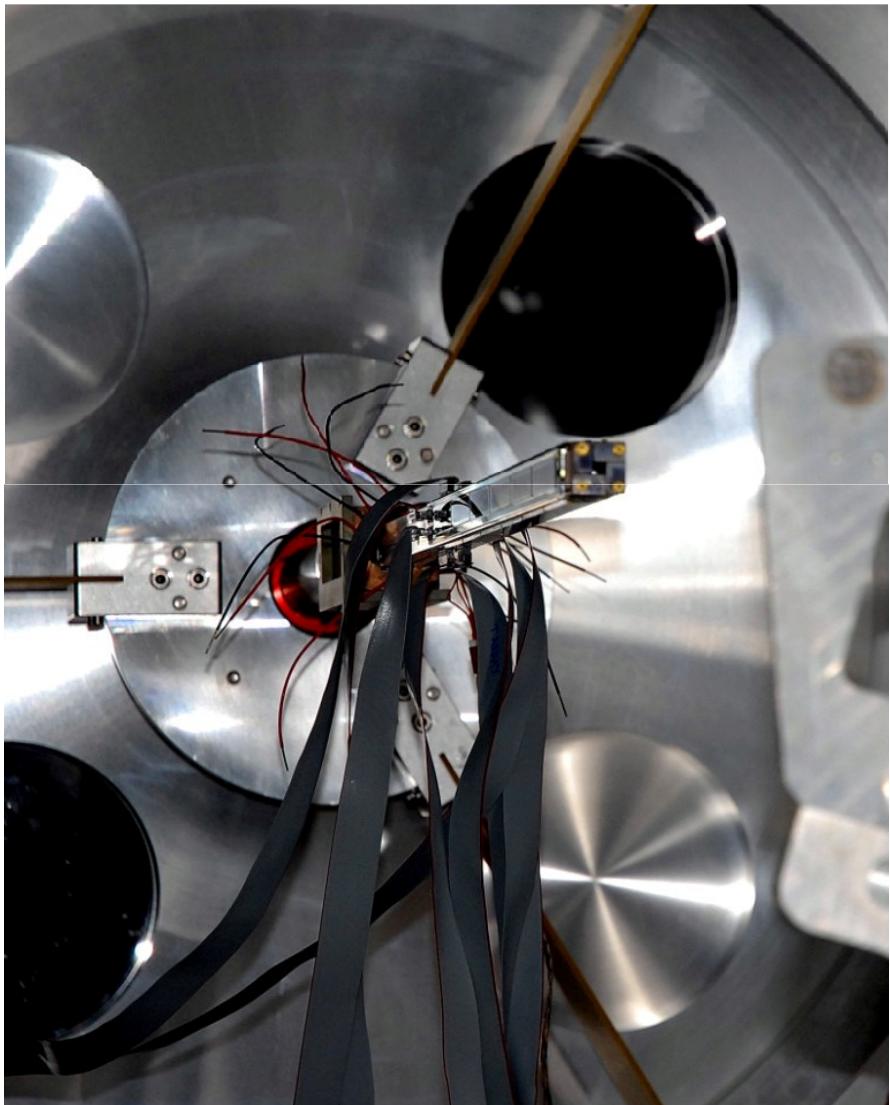


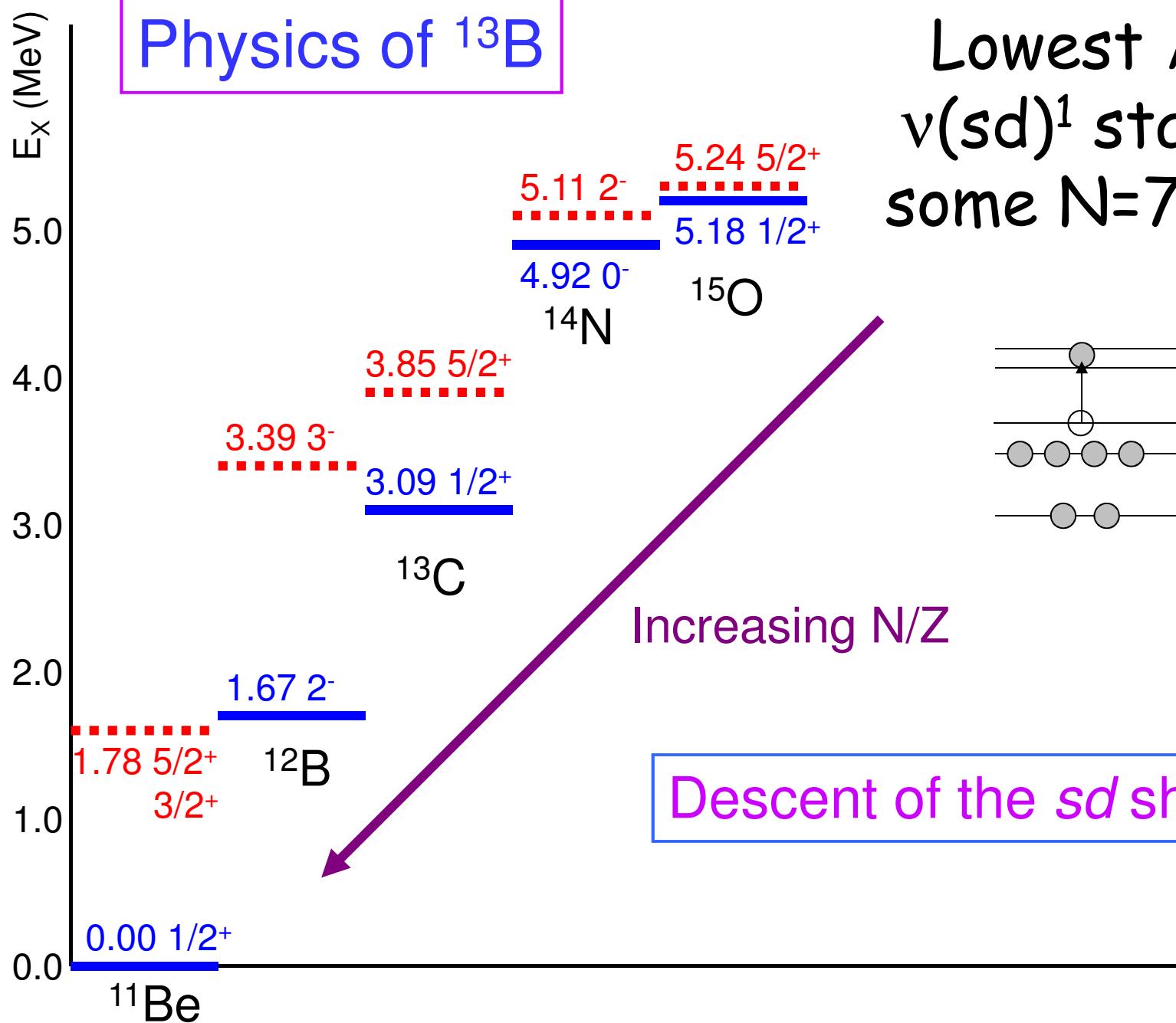
J.P. Schiffer, RIA equipment workshop 1999,
AHW et al, NIMPRA **580**, 1290 (2007)
J. C. Lighthall et al, NIMPRA **622**, 97 (2010)

This you have seen...

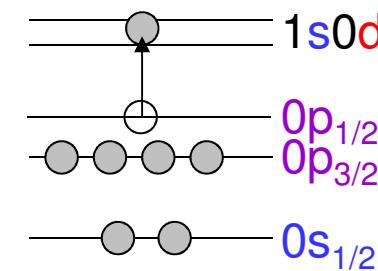


But maybe not this...

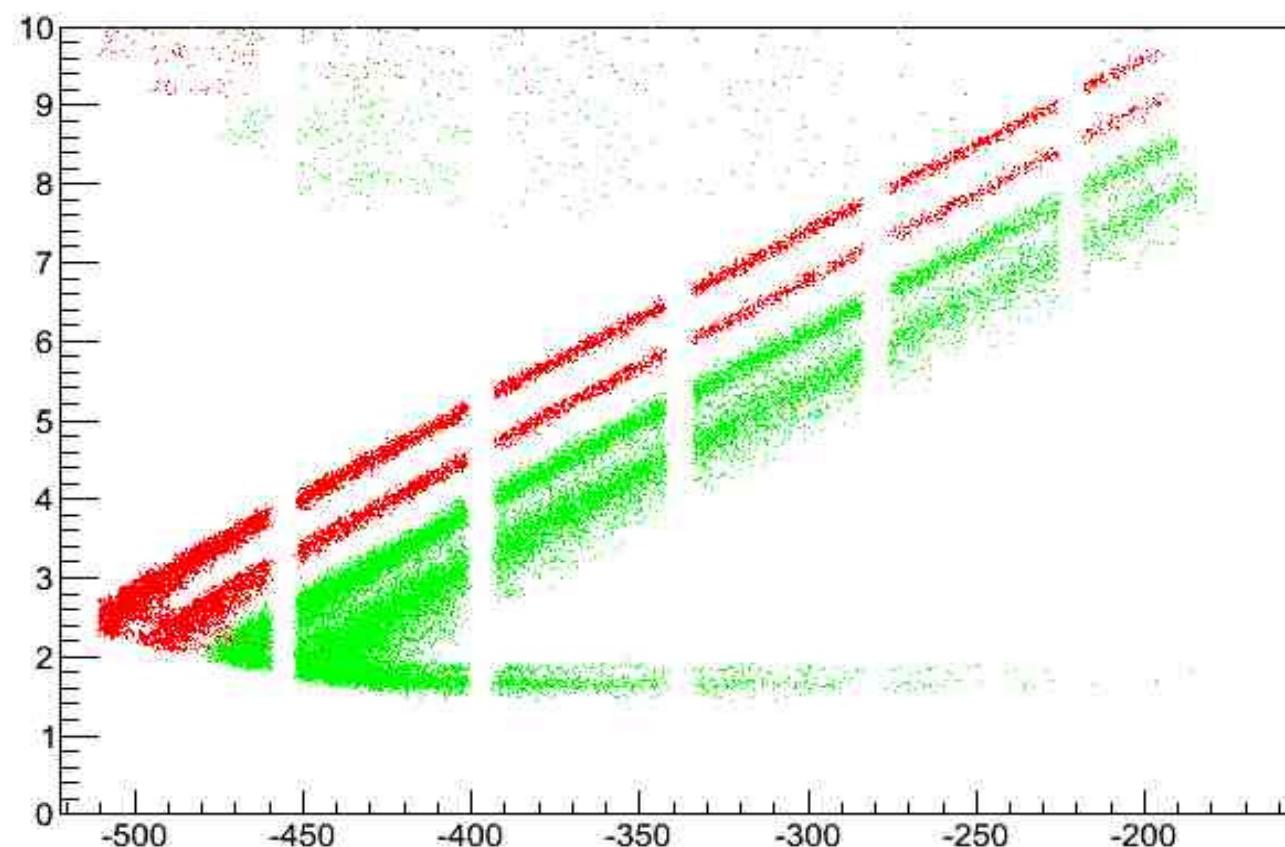




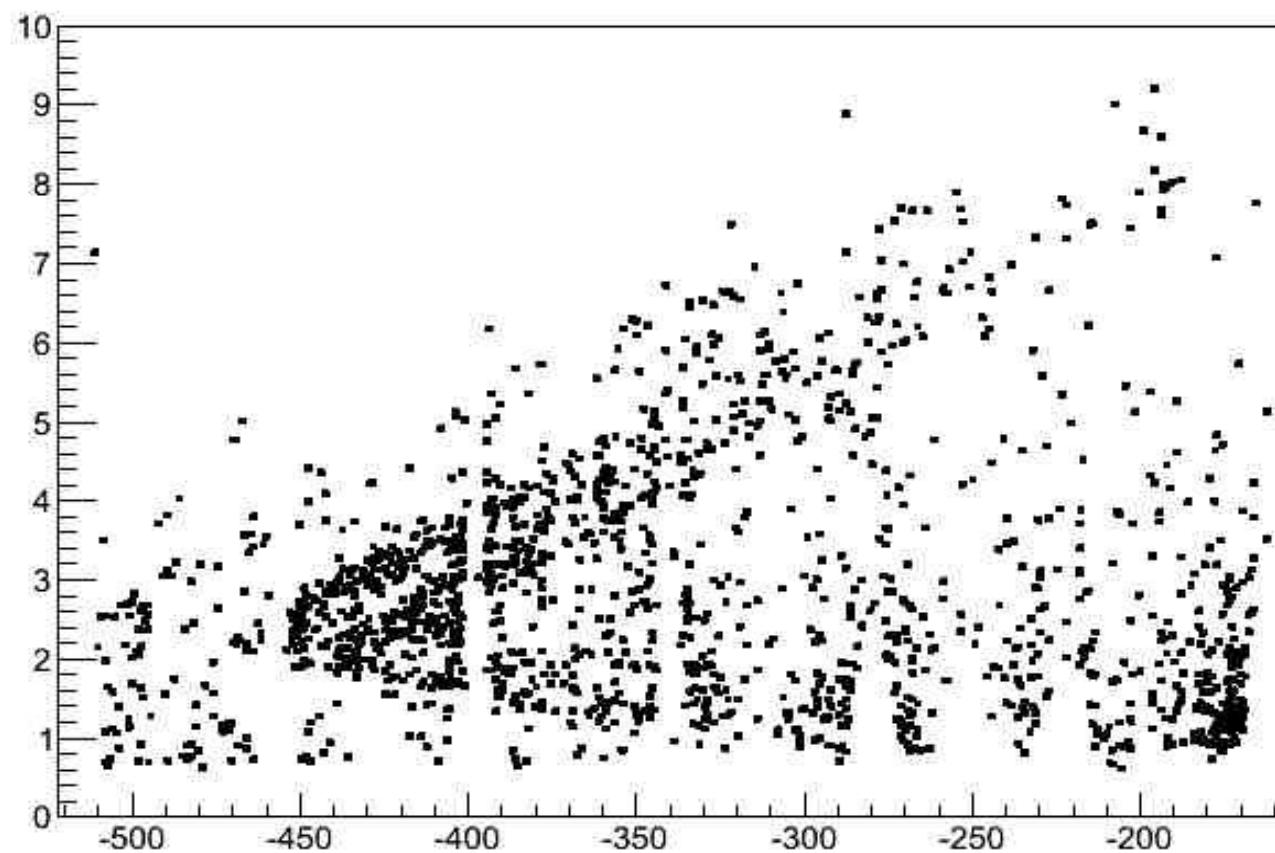
Lowest $J=0,2$
 $\nu(\text{sd})^1$ states in
some $N=7$ nuclei



E vs Z Gated



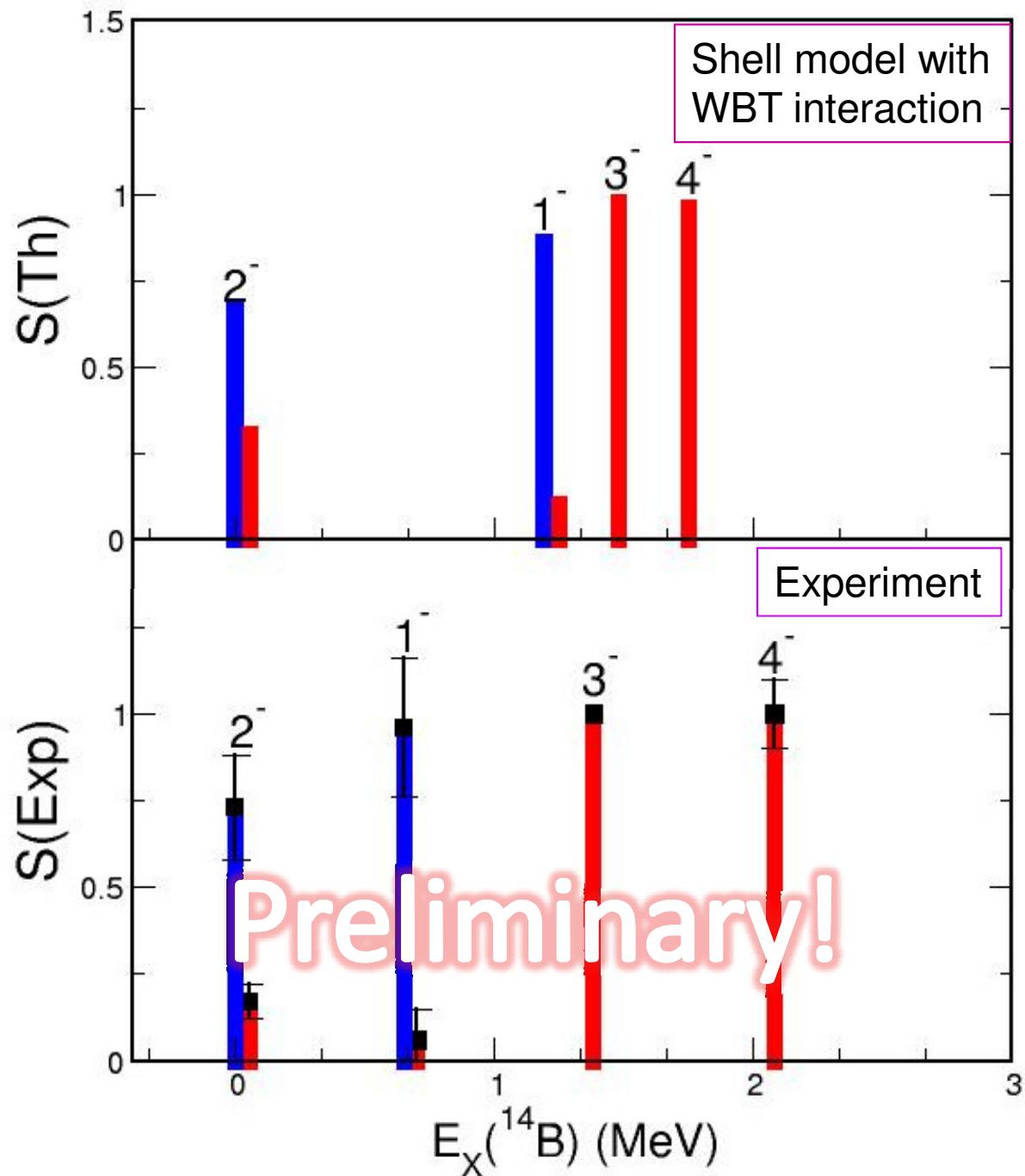
Energy vs Position Boron gated



$^{13}\text{B}(d,p)^{14}\text{B}$ Spectroscopic factors

Excitation energies
and relative
spectroscopic
factors from the
shell model

Blue: $L=0, 1s_{1/2}$
Red: $L=2, 0d_{5/2}$

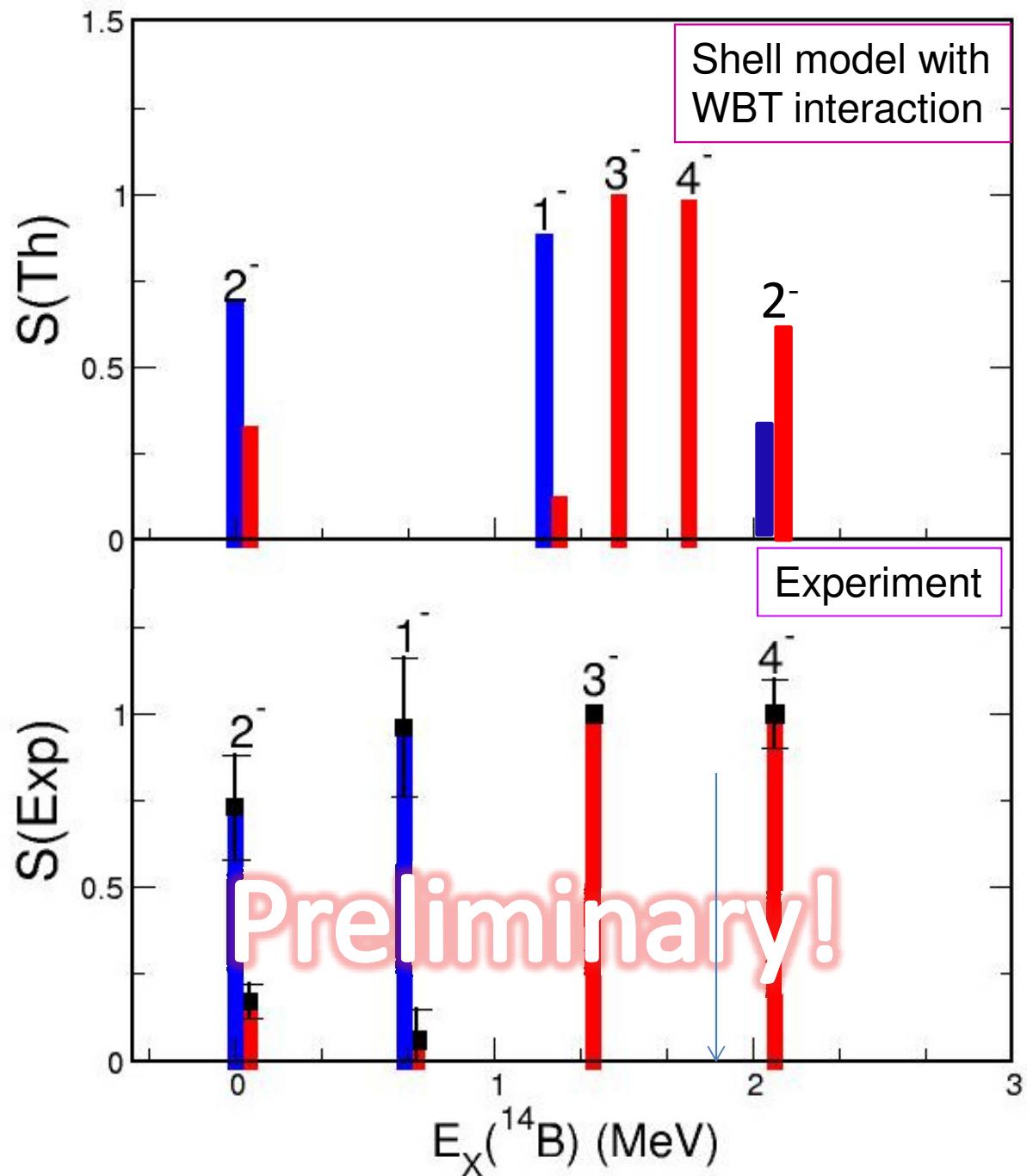


2⁻ mixed $L=0+2$,
1⁻ pure $L=0$
Reasonable agreement
But *caveat emptor!*

$^{13}\text{B}(d,p)^{14}\text{B}$ Spectroscopic factors

Excitation energies
and relative
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2⁻ mixed $L=0+2$,
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Reasonable agreement
But *caveat emptor!*

What's in *your* beam?

