James Webb Space Telescope: data, problems, and resolution

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Outline

1. Recent problems discovered by HST and LWST.
2. Old cosmological problems
3. PBH solution of new and old problems
4. Antimatter in the MIky Way, including antistars
5. Log-normal mass spectrum of PBHs, confirmed by observations.
6. Black dark matter.
7. Gravitational waves and PBH.

Resolution of the problems by PBH suggested long before they sppeared:
A.Dolgov, J.Silk, PRD 47 (1993) 4244 (DS) "Baryon isocurvature fluctuations at small scale and baryonic **dark matter**".
A.Dolgov, M. Kawasaki, N. Kevlishvili (DKK), NPB807 (2009) 229, ”Inhomogeneous baryogenesis, **cosmic antimatter**, and DM”
Prediction of well evolved early galaxies, quasars (SMBH), rich chemistry (heavy elements) and dust.

JWST and conventional cosmology

Discoveries of several recent months made by JWST, in infrared μm range, created almost panic among traditional cosmologists and astrophysicists. It was observed that the pretty young universe with the age 200-300 million years contains a large lot of bright galaxies which simply **cannot be there** according to the canonical faith or better to say to the standard cosmological model.

As is stated in the JWST publications: "an unexpectedly large density (stellar mass density $\varrho_* \gtrsim 10^6 M_\odot$ Mpc^{-3}) of massive galaxies (stellar masses $M_* \geq 10^{10.5} M_\odot$) are discovered at extremely high redshifts $z \gtrsim 10$."

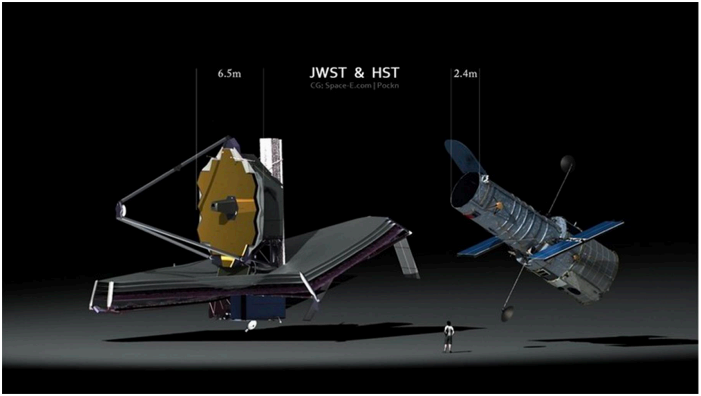
A few examples from CEERS = Cosmic Evolution Early Release Science:

Galaxies at: $z = 14.3 \pm 0.4$, $t_U = 264$ **Myr**; $z = 16.7$, $t_U = 235$ **Myr**

Enforced retreat: 'Bit of panic': Astronomers forced to rethink early JWST findings. Revised instrument calibrations are bedevilling work on the distant Universe. https://www.nature.com/articles/d41586-022-

Confirmation by spectral line identifications is strongly desirable.

JWST infrared telescope and HST



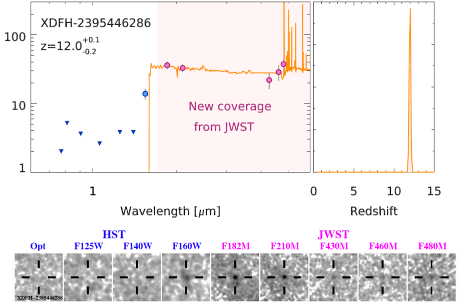
Placing a telescope in space makes it possible to register electromagnetic radiation in the ranges in which the earth's atmosphere is opaque; primarily in the infrared range. Due to the absence of the influence of the atmosphere, the resolution of the telescope is 7-10 times greater than that of a similar telescope located on Earth.

Comparison of JWST and HST

HST: Distance: 570 km
Mirror 2.4 m
Wave length: optical, e.g. blue 450 nm and UV
some IR: 0.8-2.5 microns

JWST: Distance 1.5×10^6 km
Mirror: 6.5 m
Wave length: 0.6 - 28,5 micron.

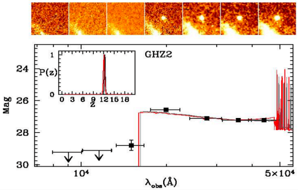
JWST and HST coincidence

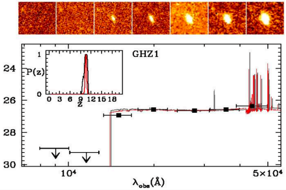


Rychard J.Bouwens et al, Evolution of the UV LF from Z~15 to z~8 Using New JWST NIRCam Medium-Band Observations over the HUDF/XDF. arXiv:2211.02607

Joint observation of object XDFH-2395446286 and measuring its redshift z=12 HST and JWST. This is the most distant galaxy ever discovered by HTS 30 years of observation.

Marco Castellano et al, Early results from GLASS-JWST.III: Galaxy candidates at z~9-15. arXiv:2207.09436

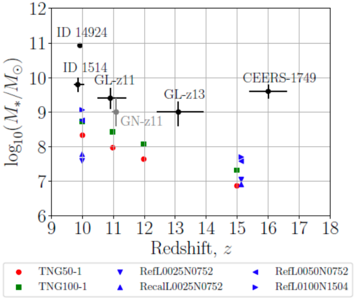




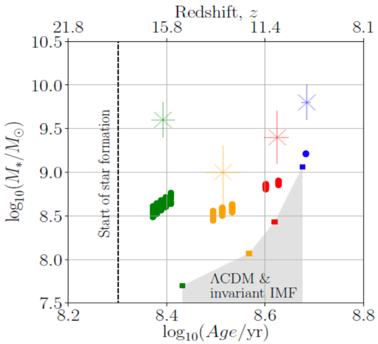
Two more examples of galaxies with z=10.62 and z=12.3 found JWST in a couple of months of observations.

JWST and the conventional ΛCDM cosmology

Moritz Haslbauer et al, Has JWST already falsified dark-matter-driven galaxy formation? arXiv:2210.14915



Comparison of the size of the most massive galaxies, obtained in models of formation and growth of galaxies based on LCDM (colored dots) with JWST observations (black dots with errors) depending on the redshift of the observed galaxies.



Early galaxies, spectroscopic confirmation

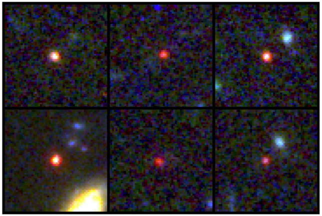
Only continuum in micron range was measured till February. New data now:
S. Tacchella, et al arXiv:2302.07234 14 Feb 2023
JADES Imaging of GN-z11: Revealing the Morphology and Environment of a Luminous Galaxy 430 Myr After the Big Bang. The JWST NIRCam 9-band near-infrared imaging of the luminous $z = 10.6$ galaxy GN-z11 from the JWST Advanced Deep Extragalactic Survey (JADES) of the GOODS-N field is made. **A spectral energy distribution (SED) is determined entirely consistent with the expected form of the high-redshift galaxy.**
A.J. Bunker, *et al* arXiv:2302.07256, 14 Feb 2023 JADES NIRSpec Spectroscopy of GN-z11: Lyman-α emission and possible enhanced **nitrogen** abundance in a $z = 10.60$ luminous galaxy, The spectroscopy of GN-z11, the most luminous candidate $z > 10$ Lyman break galaxy in the GOODS-North field with $M_{UV} = -21.5$ is presented. Redshift of $z = 10.603$ is derived (lower than previous determinations) based on multiple emission lines in low and medium resolution spectra over $0.8 - 5.3\,\mu$m. The spatially-extended Lyman-α in emission is observed. The NIRSpec spectroscopy confirms that GN-z11 is a remarkable galaxy with extreme properties seen 430 Myr after the Big Bang.

Early galaxies, spectroscopic confirmation

Age of Most Distant Galaxy is confirmed with Oxygen observation.
The radio telescope array ALMA has pin-pointed the exact cosmic age of a distant JWST-identified galaxy, GHZ2/GLASS-z12, at 367 million years after the Big Bang. ALMA's deep spectroscopic observations revealed a spectral emission line associated with ionized Oxygen near the galaxy, which has been shifted in its observed frequency due to the expansion of the Universe since the line was emitted. This observation confirms that the JWST is able to look out to record distances, and heralds a leap in our ability to understand the formation of the earliest galaxies in the Universe.
NASA / ESA / CSA / T. Treu, UCLA / NAOJ / T. Bakx, Nagoya U.
MNRAS, 22.02,2023.

Impossible galaxies

I. Labbé et al, A population of red candidate massive galaxies 600 Myr after the Big Bang, Nature, published online 22.02.2023, Six candidate massive galaxies (stellar mass $> 10^{10}$ solar masses) at $7.4 \lesssim z \lesssim 9.1$ 500–700 Myr after the Big Bang, including one galaxy with a possible stellar mass of $\sim 10^{11} M_\odot$, too massive to be created in so early universe. According to the 'science' it is impossible to create so well developed galaxies. "May be they are **supermassive black holes** of the kind never seen before. That might mean a revision of our understanding of black holes." **Well agrees with our predictions of PBHs.**



The six candidate galaxies identified in the JWST data. (NASA, ESA, CSA, I. Labbe/Swinburne University of Technology)

Ultra-massive early QSO observed by ALMA

ALMA confirmation of an obscured hyper-luminous radio-loud AGN at z = 6.853 associated with a dusty starburst in the 1.5 deg2 COSMOS field,
R. Endsley et al, Monthly Notices of the Royal Astronomical Society,
Volume 520, Issue 3, April 2023, Pages 4609–4620, Published: 24.02.2023
VIRCam and IRAC photometry perhaps suggests that COS-87259 is an extremely massive reionization-era galaxy with $M_* = 1.7 \times 10^{11} M_\odot$
Such a very high AGN luminosity suggests that this object is powered by $\sim 1.6 \times 10^9 M_\odot$ black hole if accreting near the Eddington limit.
Nearly impossible, but PBH could seed such monster.

Rich chemistry

B. Peng, et al, The Astrophysical Journal Letters, Volume 944, Issue 2, id.L36, 8 pp. 'Discovery of a Dusty, Chemically Mature Companion to $z \sim 4$ Starburst Galaxy in JWST Early Release Science Data,'
Most surprising about the companion galaxy, considering its age and mass, was its mature metallicity— amounts of elements heavier than helium and hydrogen, such as carbon, oxygen and nitrogen.
The team estimated that as comparable to our sun, which is more than 4 billion years old and inherited most of its metals from previous generations of stars that had 8 billion years to build them up.
High abundances of heavy elements may be a result of BBN with large baryon-to-gamma ratio, as predicted in DJ and DKK

Rich chemistry

Nitrogen enhancements 440 Myr after the Big Bang: super-solar N/O, a tidal disruption event or a dense stellar cluster in GN-z11? A.J. Cameron, et al, arXiv:2302.10142, 20.02.2023.
Observations of GN-z11 with JWST/NIRSpec revealed numerous oxygen, carbon, nitrogen, and helium emission lines at $z = 10.6$. Abundance ratios of individual elements within the interstellar medium (ISM) of this galaxy are found. The data prefers log(N/O)>-0.25, greater than 4 times solar. The derived log(C/O)>-0.78, ($\approx$30 solar) is also elevated with respect to galaxies of similar metallicity (12+log(O/H)$\approx$ 7.82). Nitrogen enhancement in GN-z11 cannot be explained by enrichment from metal-free Population III stars. Yields from runaway stellar collisions in a dense stellar cluster or a tidal disruption event provide promising solutions to give rise to these unusual emission lines at $z = 10.6$, and explain the resemblance between GN-z11 and a nitrogen-loud quasar. These recent observations showcase the new frontier opened by JWST to constrain galactic enrichment and stellar evolution within 440 Myr of the Big Bang.

Problems prior to JWSP data

Similar serious problems are known already for many years. The Hubble space telescope (HST) discovered that the early universe, at $z = 6 - 7$ is too densely populated with quasars, alias SMBH, supernovae, gamma-bursters and it is very dusty. **No understanding how all these creature were given birth in such a short time is found in conventional cosmology.** Moreover great lots of phenomena in the **present day universe** are also in strong tension with canonical cosmological expectations.

A.D. "Massive and supermassive black holes in the contemporary and early Universe and problems in cosmology and astrophysics Phys. Usp. 61 (2018) 2, 115.

"Hubble"sees the universe up to $z = 6 - 7$, but accidentally a galaxy at $z \approx 12$ has been discovered for which both Hubble and Webb are in good agreement. Still only after publications of JWST data astronomy establishment became seriously worried.

All the problems are neatly solved if the universe is populated by primordial black holes (PBH) and the astrophysically large bubbles with very high baryonic density

BH types by formation mechanisms

1. Astrophysical black holes,

created by the collapse of a star which exhausted its nuclear fuel. The expected masses should start immediately above the neutron star mass, i.e. about $3\boldsymbol{M}_\odot$, but noticeably below $100M_\odot$. Instead we observe that the BH mass spectrum in the galaxy has maximum at $\boldsymbol{M} \approx 8\boldsymbol{M}_\odot$ with the width $\sim (1-2)\boldsymbol{M}_\odot$. The result is somewhat unexpected but an explanations in the conventional astrophysical frameworks is possible.

Recently LIGO/Virgo discovered BHs with masses close to $100M_\odot$. Their astrophysical orrgin was considered **impossible.** Now some, quite exotic, formation mechanisms are suggested.

2. Formation by accretion on the mass excess in the galactic center.

In any large galaxy there exists a supermassive BH (SMBH) at the center, with masses varying from a few millions $M_\odot$ (e,g, Milky Way) up to almost hundred billions $M_\odot$. **However, the conventional accretion mechanisms are not efficient enough to create such monsters during the universe life-time, $t_U \approx 14.6$ Gyr. At least 10-fold longer time is necessary,** to say nothing about SMBH in 10 times younger universe.

BH types by formation mechanisms

3. Primordial black holes (PBH) created during pre-stellar epoch

The idea of the primordial black hole (PBH) i.e. of black holes which could be formed the early universe prior to star formation was first put forward by Zeldovich and Novikov: "The Hypothesis of Cores Retarded During Expansion and the Hot Cosmological Model Astronomicheskij Zhurnal, 43 (1966) 758, Soviet Astronomy, AJ.10(4):602–603;(1967).

According to their idea, the density contrast in the early universe inside the bubble with radius equal to the cosmological horizon might accidentally happen to be large, $\delta\varrho/\varrho \approx 1$, then that piece of volume would be inside its gravitational radius i.e. it became a PBH, which decoupled from the cosmological expansion.

Elaborated later in S. Hawking, "Gravitationally collapsed objects of **very low mass** Mon. Not. Roy. Astron. Soc. **152**, 75 (1971).

B. J. Carr and S. W. Hawking, "Black holes in the early Universe," Mon. Not. Roy. Astron. Soc. **168**, 399 (1974).

BH types by masses

There is the following conventional division of black holes by their masses:
1. Supermassive black holes (SMBH): $M = (10^6 - 10^{10})M_\odot$.
2. Intermediate mass black holes (IMBH): $M = (10^2 - 10^5)M_\odot$.
3. Solar mass black holes: masses from a fraction of $M_\odot$ up to $100M_\odot$.

The origin of most of these BHs is unclear, except maybe of the BHs with masses of a few solar masses, which may be astrophysical.
Highly unexpected was abundance of IMBH which are appearing during last few years in huge numbers.
The assumption that (almost) all black holes in the universe are primordial strongly reduce or even eliminate the tension.

Problems of the contemporary universe. Summary.

1. SMBH in all large galaxies. Too short time for their formation through the usual accretion mechanism.

2. SMBH in small galaxies and even in (almost) empty space. No material for their creation. Pushed out of large galaxies? Wandering BHs?

A striking example: the Hobby-Eberly Telescope at Texas's McDonald Observatory suggested the presence of a black hole with a mass of about 17 billion $M_\odot$ equivalent to 14% of the total stellar mass of the galaxy. Usually the mass of the central BH is about 0.1 % of the galaxy mass.

3. Too old stars, older than the Galaxy and maybe older that the universe?

4. MACHOs, non-luminous objects with masses $\sim 0.5 M_\odot$ obsserved through microlensing; origin unknown.

5. Problems with the BH mass spectrum in the Galaxy, masses are concentrated in the narrow interval $(7.8 \pm 1.2) M_\odot$.

6. Origin and properties of the sources of the observed gravitational waves.

7. IMBH, with $M \sim (10^3 - 10^5) M_\odot$, in dwarfs and globular clusters, disoovered but unexpected..

8. Strange stars in the Galaxy, too fast and with unusual chemistry. Observed during the last decade..

Solution of all the problems by PBH

To summarise, a large amount of observational data are at odds with the conventional model but nicely fits the model of creation of primordial black holes and primordial stars proposed in DS and DKK The proposed mechanism is the first where inflation and Affleck-Dine baryogenesis are applied to PBH formation, repeated now in many works.
The striking feature of it is the **log-normal** mass spectrum which is the only known spectrum **tested by "experiment" in a good agreement.**

$$\frac{dN}{dM} = \mu^2 \exp\left[-\gamma \ln^2 (M/M_0)\right],$$

$M_0 \sim 10 M_\odot$, is predicted, A.Dolgov, K.Postnov, "Why the mean mass of primordial black hole distribution is close to $10 M_\odot$". JCAP 07 (2020) 063.
The horizon mass at QCD p.t. is $10 M_\odot$, for $\mu = 0$. At larger chemical potential the T_{pt} is smaller and M_{hor} is larger.

Seeding of galaxy formation by PBH

The hypothesis by DS (1993) and DKK (2006), that SMBH seeded galaxy formation allows to explain presence of SMBH in all large and several small galaxies accessible to observation

This mechanism explains how might be created the galaxies observed by JWST in the very young universe.

This statement was recently rediscovered by B. Liu, V. Bromm, "Accelerating early galaxy formation with primordial black holes'', arXiv:2208.13178, 28 Aug 2022: Recent observations with JWST have identified several bright galaxy candidates at $z \gtrsim 10$, some of which appear unusually massive (up to $\sim 10^{11}\ \mathrm{M}_{\odot}$). Such early formation of massive galaxies is difficult to reconcile with standard $\Lambda\mathrm{CDM}$ predictions, ...we show that the observed massive galaxy candidates can be explained with lower SFE than required in $\Lambda\mathrm{CDM}$, if structure formation is accelerated by massive ($\gtrsim 10^{9}\ \mathrm{M}_{\odot}$) primordial black holes that enhance primordial density fluctuations.

Problems solved by PBH

The origin of IMBH is unknown in all mass range, though plenty of them are discovered everywhere.
Moreover, BH with $\mathrm{M} \approx 100\mathrm{M}_\odot$ is strictly forbidden but nevertheless observed by LIGO/Virgo.

The described above model of PBH formation excellently solves all the inconsistencies. **The inverted picture of galaxy formation is assumed: first SMPBH are created and later they seed galaxy formation.**

Primordial IMBHs with masses of a few thousand solar mass explain formation of globular clusters (GCs). In the last several years such IMBH inside GSs are observed. Similar features are predicted for dwarf galaxies.
A. Dolgov, K. Postnov, "Globular Cluster Seeding by Primordial Black Hole Population JCAP 04 (2017) 036, e-Print: 1702.07621 [astro-ph.CO]. This prediction is confirmed by the recent data, e.g. in the dwarf galaxy SDSS J1521+1404 the BH is discovered with the mass $M \sim 10^5 M_\odot$.

BHs in dwarf galaxies

This prediction is confirmed by the recent data, e.g. in the dwarf galaxy SDSS J1521+1404 the BH is discovered with the mass $M \sim 10^5 M_\odot$.

Two Candidates for Dual AGN in Dwarf-Dwarf Galaxy Mergers, M. Mićić, et al, arXiv:2211.04609 [astro-ph.GA]. For the first time, astronomers have spotted evidence of a pair of dwarf galaxies featuring GIANT black holes on a collision course with each other. In fact, they haven't just found just one pair – they've found two.

Intermediate-mass black holes: finding of episodic, large-scale and powerful jet activity in a dwarf galaxy SDSS J090613.77+561015.2. Jun Yang et al, e-Print: 2302.06214 [astro-ph.GA,astro-ph.HE]. Discovery of an intermediate-mass black hole (IMBH) with a mass of $M_{BH} = 3.6^{+5.9}_{-2.3} \times 10^5 M_\odot$ which surely cannot be created by accretion but might seed the dwarf formation.

Strange stars

The model also predicts strange stars, too fast moving, too old (older than the universe), with unusual chemical content which are also observed.

Gravitational waves from BH binaries

- GW discovery by LIGO strongly indicate that the sources of GW are PBHs. see e.g. S.Blinnkov, A.D., N.Porayko, K.Postnov, JCAP 1611 (2016), 036 "Solving puzzles of GW150914 by primordial black holes,"

1. Origin of heavy BHs ($\sim 30 M_{\odot}$); there appeared much more striking problem of BH with $M \sim 100 M_{\odot}$. See however, J. Ziegler, K. Freese, arXiv:2010.00254: DM annihilation inside stars
2. Formation of BH binaries from the original stellar binaries.
3. Low spins of the coalescing BHs .

To form so heavy BHs, the progenitors should have $M > 100 M_{\odot}$. and a low metal abundance to avoid too much mass loss during the evolution. Such heavy stars might be present in young star-forming galaxies but they are not observed in the necessary amount. PBHs with the observed by LIGO masses may be easily created with sufficient density.

Chirp mass

Two rotating gravitationally bound massive bodies are known to emit gravitational waves. In quasi-stationary inspiral regime, the radius of the orbit and the rotation frequency are approximately constant and the GW frequency is twice the rotation frequency. The luminosity of the GW radiation is:

$$\boldsymbol{L = \frac{32}{5}\, m_{Pl}^2 \left(\frac{M_c\, \omega_{orb}}{m_{Pl}^2} \right)^{10/3}},$$

where M_1, M_2 are the masses of two bodies in the binary system and $\boldsymbol{M_c}$ is the so called chirp mass:

$$\boldsymbol{M_c = \frac{(M_1\, M_2)^{3/5}}{(M_1 + M_2)^{1/5}}},$$

and

$$\boldsymbol{\omega_{orb}^2 = \frac{M_1 + M_2}{m_{Pl}^2 R^3}}.$$

Chirp mass distribution - overlap with 2022 talk

A.D. Dolgov, A.G. Kuranov, N.A. Mitichkin, S. Porey, K.A. Postnov, O.S. Sazhina, I.V. Simkine On mass distribution of coalescing black holes, JCAP 12 (2020) 017, e-Print: 2005.00892.

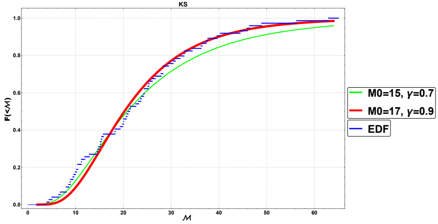
The available data on the chirp mass distribution of the black holes in the coalescing binaries in O1-O3 LIGO/Virgo runs are analyzed and compared with theoretical expectations based on the hypothesis that these black holes are primordial with log-normal mass spectrum.

The inferred best-fit mass spectrum parameters, $M_0 = 17 M_\odot$ and $\gamma = 0.9$, fall within the theoretically expected range and shows excellent agreement with observations.

On the opposite, binary black hole formation based **on massive binary star evolution** require additional adjustments to reproduce the observed chirp mass distribution.

Chirp mass distribution - overlap with 2022 talk

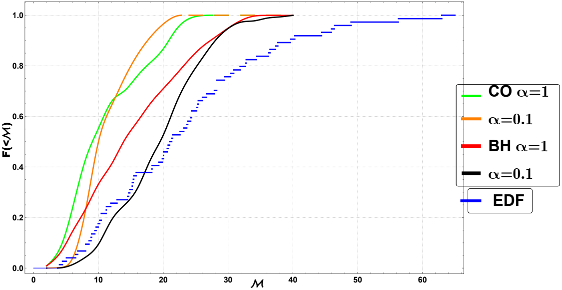
Model distribution $F_{PBH}(< M)$ with parameters $M_0 \approx 17 M_\odot$ and $\gamma \sim 1$ for two best Kolmogorov-Smirnov tests. EDF= empirical distribution function.



Similar value of the parameters are obtained in M. Raidal et al, JCAP.,2019. Feb. V. 2019, no. 2. P. 018. arXiv:1812.01930 and L. Liu, et al arXiv:2210.16094.
See also K. Postnov and N. Mitichkin, e-Print: 2302.06981.

Chirp mass distribution - overlap with 2022 talk

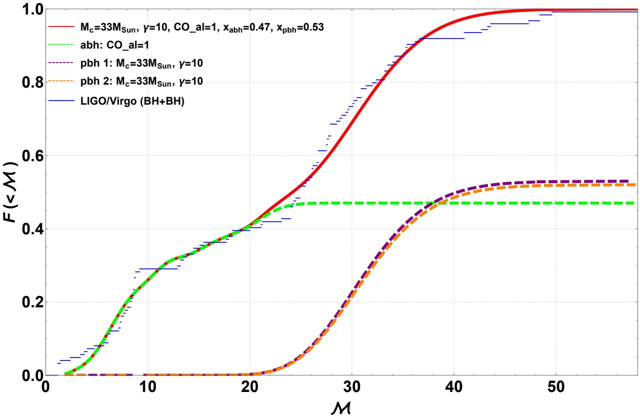
Cumulative distributions $F(< M)$ for several **astrophysical** models of binary BH coalescences.



Conclusion: PBHs with log-normal mass spectrum perfectly fit the data. Astrophysical BHs seem to be disfavoured.

Analysis of recent Ligo-Virgo-Kagra (LVK) data

A new analysis of the Ligo-Virgo-Kagra data was performed recently by K. Postnov and N. Mitichkin, "'On the primordial binary black hole mergings in LVK data', e-Print: 2302.06981 [astro-ph.CO]. They concluded that the chirp-mass distribution of LVK GWTC-3 BH+BH binaries with distinct two bumps can be explained by two different populations of BH+BH binaries:
1) the low-mass bump at $M_0 \sim 10 M_\odot$ due to the astrophysical BH+BH formed in the local Universe from the evolution of massive binaries
2) the PBH binaries with log-normal mass spectrum with $M_0 \simeq 10 M_\odot$ and $\gamma \simeq 10$. The central mass of the PBH distribution is larger than the expected PBH mass at the QCD phase transition ($\sim 8 M_\odot$) but still can be accommodated with the mass of the cosmological horizon provided that the temperature $T_{QCD} \sim 70$ MeV, possible for non-zero chemical potential at QCD p.t.



The observed (blue step-like curve) and model (red solid curve) distribution function of the chirp-masses of coalescing binary BHs from the LVK GWTC-3 catalogue. The model includes almost equal contributions from coalescences of astrophysical binary BHs (green dashed curve) and primordial BHs with the initial log-normal mass spectrum with parameters $M_0 = 33M_\odot$, $\gamma = 10$ - with such γ heavier PBH practically are not created.

PBH and inflation

Inflation allows for formation of PBH with very large masses. It was first applied to PBH production in DS paper, PRD 47 (1993) 4244,
a year later in: B.J. Carr, J.H. Hilbert, J.E. Lidsey, "Black hole relics and inflation: Limits on blue perturbation spectra'', Phys.Rev.D 50 (1994) 4853, astro-ph/9405027;
and soon after in P. Ivanov, P. Naselsky, I. Novikov (May 10, 1994), Inflation and primordial black holes as dark matter, PRD 50 (1994) 7173.
Presently inflationary mechanism of PBH production is commonly used. It allows to create PBH with very high masses, but the predicted spectrum is multi-parameter one and quite complicated
The only exception is the log-normal spectrum of DS and DKK tested by observatons.

Black Dark Matter

The first suggestion PBH might be dark matter "particles"was made by S. Hawking in 1971 "Gravitationally collapsed objects of very low mass Mon. Not. R. astr. Soc. (1971) 152, 75 and later by G. Chapline in 1975 who noticed that low mass PBHs might be abundant in the present-day universe with the density comparable to the density of dark matter. G.F. Chapline, Nature, 253, 251 (1975) "Cosmological effects of primordial black holes". Assumed flat mass spectrum in log interval:

$$dN = N_0(dM/M)$$

with maximum mass $M_{max} \lesssim 10^{22}$ g, which hits the allowed mass range. The next one: A. Dolgov, J. Silk (Mar 13, 1992), Baryon isocurvature fluctuations at small scales and baryonic dark matter, PRD 47 (1993) 4244 with more realistic masses. **first paper with inflation applied to PBH formation, so PBH masses as high as $10^6 M_\odot$, and even higher can be created, log-normal mass spectrum was predicted.**

Black Dark Matter

Constraints on PBHs - B.Carr, F. Kuhnel "Primordial Black Holes as Dark Matter: Recent Developments arXiv:2006.02838, June 2020
Primordial black holes as dark matter candidates B. Carr, F. Kuhnel SciPost Phys.Lect.Notes 48 (2022), e-Print: 2110.02821 [astro-ph.CO]
For monochromatic mass spectrum of PBHs (caution, model-dependent).

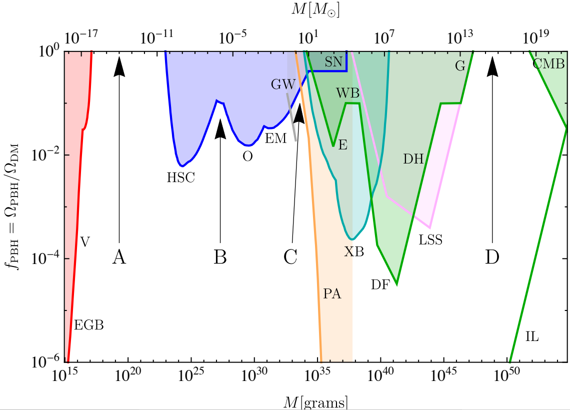


Figure caption

Constraints on $f(M)$ for a **monochromatic** mass function, from evaporations (red), lensing (blue), gravitational waves (GW) (gray), dynamical effects (green), accretion (light blue), CMB distortions (orange) and large-scale structure (purple). Evaporation limits from the extragalactic gamma-ray background (EGB), the Voyager positron flux (V) and annihilation-line radiation from the Galactic centre (GC). Lensing limits from microlensing of supernovae (SN) and of stars in M31 by Subaru (HSC), the Magellanic Clouds by EROS and MACHO (EM) and the Galactic bulge by OGLE (O). Dynamical limits from wide binaries (WB), star clusters in Eridanus II (E), halo dynamical friction (DF), galaxy tidal distortions (G), heating of stars in the Galactic disk (DH) and the CMB dipole (CMB). Large scale structure constraints(LSS). Accretion limits from X-ray binaries (XB) and Planck measurements of CMB distortions (PA). The incredulity limits (IL) correspond to one PBH per relevant environment (galaxy, cluster, Universe). There are four mass windows (A, B, C, D) in which PBHs could have an appreciable density.

Black Dark Matter

Carr, 2019: all limits are model dependent and have caveats.

Eliminating the LIGO bounds on primordial black hole dark matter,
C. Boehm, et al arXiv:2008.10743 .... reopens the possibility for dark matter in the form of LIGO-mass PBHs.
C. Corianò, P.H. Frampton, arXiv:2012.13821 [astro-ph.GA]
Does CMB Distortion Disfavour Intermediate Mass Dark Matter?
The most questionable step in this chain of arguments is the use of overly simplified accretion models. We compare how the same accretion models apply to X-ray observations from supermassive black holes SMBHs, M87 and Sgr A^*. The comparison of these two SMBHs with **intermediate mass MACHOs suggests that the latter could, after all, provide a significant constituent of all the dark matter.**

BH clustering and DM

As is argued by S.G. Rubin, A.S. Sakharov, M.Y. Khlopov, in "The Formation of Primary Galactic Nuclei during Phase Transitions in the Early Universe", Soviet Journal of Experimental and Theoretical Physics. 2001, V. 92, no. 6. 921. arXiv:hep-ph/0106187 PBHs can be formed in clusters. Dynamical interactions in PBH clusters offers additional channel for the orbital energy dissipation thus increasing the merging rate of PBH binaries, and the constraints on f_{PBH} obtained by assuming a homogeneous PBH space distribution can be weaker. A recent analysis by Y. Eroshenko, V. Stasenko, "Gravitational waves from the merger of two primordial black hole clusters" arXiv:2302.05167 based on the PBH formation model M. Sasaki et al "Primordial Black Hole Scenario for the Gravitational-Wave Event GW150914", PRL. 2016. V. 117, no. 6. P. 061101, arXiv:1603.08338 and T. Nakamura, et al "Gravitational Waves from Coalescing Black Hole MACHO Binaries", ApJL 1997, V. 487, no. 2, P. L139, arXiv:astro-ph/9708060. shows that even $f_{PBH} = 0.1 - 1$ is not excluded. Thanks to K. Postnov for these important references.

Intermediate summary and antimatter in the Galaxy

The mechanism of AD and DKK solves the problem of the observed population of the universe at high redshifts by SMBH (QSO), galaxies, SN, and of a large amount of dust.

The predicted log-normal spectrum of PBH is tested and confirmed by the observations (the only one existing in the literature).

The existence of IMBH in GCs is confirmed.

Antimatter in the Galaxy, especially antistars, is predictes and observed.

PBH Creation Mechanism

SUSY motivated baryogenesis, Affleck and Dine (AD).
SUSY predicts existence of scalars with B $\neq 0$. Such bosons may condense along flat directions of the quartic potential:

$$\boldsymbol{U_\lambda}(\chi) = \boldsymbol{\lambda}|\chi|^4 \left(1 - \mathbf{cos}\, 4\boldsymbol{\theta}\right)$$

and of the mass term, $\boldsymbol{U_m} = \boldsymbol{m}^2\chi^2 + \boldsymbol{m}^{*\,2}\chi^{*\,2}$:

$$\boldsymbol{U_m}(\chi) = \boldsymbol{m}^2|\chi|^2[1 - \mathbf{cos}\,\boldsymbol{(}2\boldsymbol{\theta} + 2\alpha)]\,,$$

where $\chi = |\chi|\,\mathbf{exp}\,\boldsymbol{(i\theta)}$ and $\boldsymbol{m} = |\boldsymbol{m}|\boldsymbol{e}^{\alpha}$. If $\alpha \neq 0$, C and CP are broken.
In GUT SUSY baryonic number is naturally non-conserved - non-invariance of $\boldsymbol{U}(\chi)$ w.r.t. phase rotation.

Creation Mechanism

Initially (after inflation) χ is away from origin and, when inflation is over, starts to evolve down to equilibrium point, $\chi = 0$, according to Newtonian mechanics:

$$\ddot{\chi} + 3H\dot{\chi} + U'(\chi) = 0.$$

Baryonic charge of χ:

$$B_\chi = \dot{\theta}|\chi|^2$$

is analogous to mechanical angular momentum. χ decays transferred baryonic charge to that of quarks in B-conserving process.

AD baryogenesis could lead to baryon asymmetry of order of unity, much larger than the observed 10^{-9}.

Creation Mechanism

If $m \neq 0$, the angular momentum, B, is generated by a different direction of the quartic and quadratic valleys at low χ. If CP-odd phase α is small but non-vanishing, both baryonic and antibaryonic domains might be formed with possible dominance of one of them.
Matter and antimatter objects may exist but globally $B \neq 0$.

Affleck-Dine field χ with CW potential coupled to inflaton Φ (AD and Silk; AD, Kawasaki, Kevlishvili):

$$U = g|\chi|^2(\Phi - \Phi_1)^2 + \lambda|\chi|^4 \ln\left(\frac{|\chi|^2}{\sigma^2}\right)$$
$$+\lambda_1(\chi^4 + h.c.) + (m^2\chi^2 + h.c.).$$

Coupling to inflaton is the general renormalizable one.
When the window to the flat direction is open, near $\Phi = \Phi_1$, the field χ slowly diffuses to large value, according to quantum diffusion equation derived by Starobinsky, generalized to a complex field χ.

Creation Mechanism

If the window to flat direction, when $\Phi \approx \Phi_1$ is open only during a short period, cosmologically small but possibly astronomically large bubbles with high β could be created, occupying a small fraction of the universe, while the rest of the universe has normal $\beta \approx 6 \cdot 10^{-10}$, created by small χ. The mechanism of massive PBH formation quite different from all others. The fundament of PBH creation is build at inflation by making large isocurvature fluctuations at relatively small scales, with practically vanishing density perturbations.

Initial isocurvature perturbations are in chemical content of massless quarks. Density perturbations are generated rather late after the QCD phase transition.

The emerging universe looks like a piece of Swiss cheese, where holes are high baryonic density objects occupying a minor fraction of the universe volume.

Results

- PBHs with log-normal mass spectrum - confirmed by the data!
- Compact stellar-like objects, as e.g. cores of red giants.
- Disperse hydrogen and helium clouds with (much) higher than average n_B density. Strange stars with unusual chemistry and velocity.
- **β may be negative leading to creation of (compact?) antistars which could survive annihilation with the homogeneous baryonic background.**
- Extremely old stars would exist even, "older than universe star"is found; the older age is mimicked by the unusual initial chemistry. Several such stars are observed.

The mechanism of PBH creation pretty well agrees with the data on the mass spectrum and on existence of antimatter in the Galaxy, especially of antistars. So we may expect that it indeed solves the problems created by HST and JWST.

More data are expected and coming