



Fermi
Gamma-ray Space Telescope



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The *Fermi* gamma-ray sky: implications for cosmic-ray acceleration and propagation

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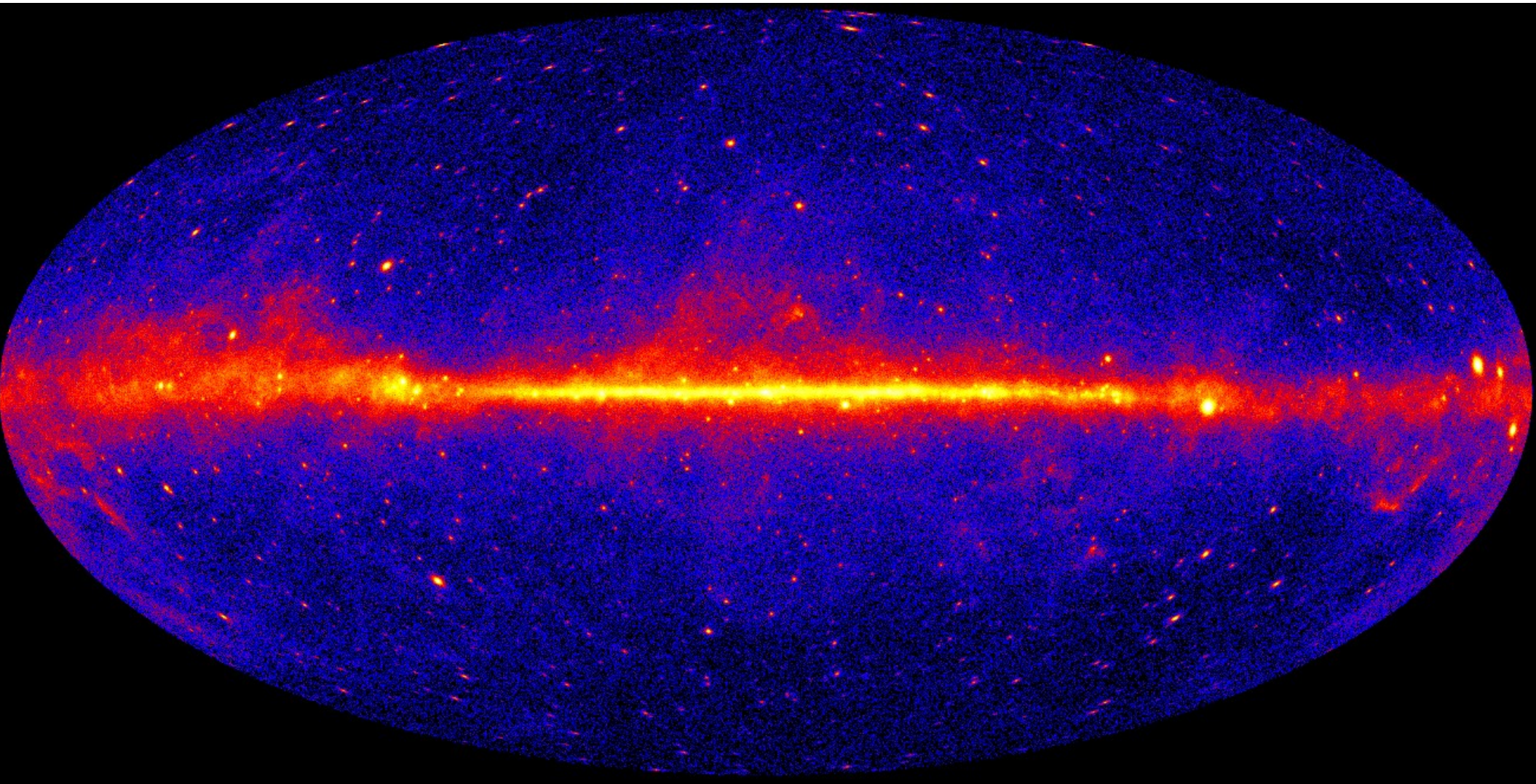
INFN & Dip. di Fisica "G. Galilei", Università di Padova
AIM, Université Paris Diderot/SAP CEA Saclay

on behalf of the
***Fermi* LAT Collaboration**

1. The gamma-ray sky
seen by the *Fermi* Large Area Telescope

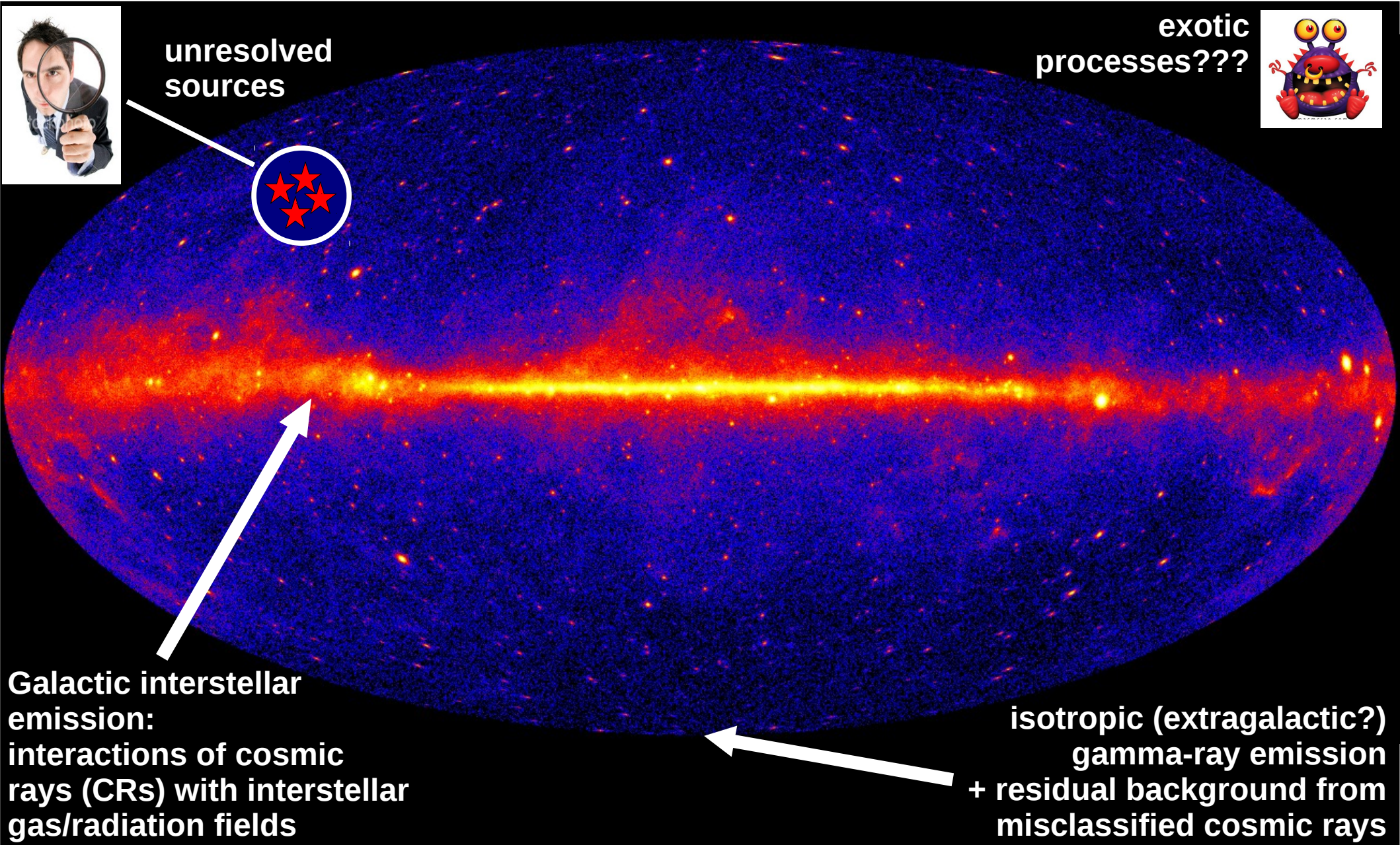
The *Fermi* gamma-ray sky

Two years of LAT data, photons with energies > 1 GeV

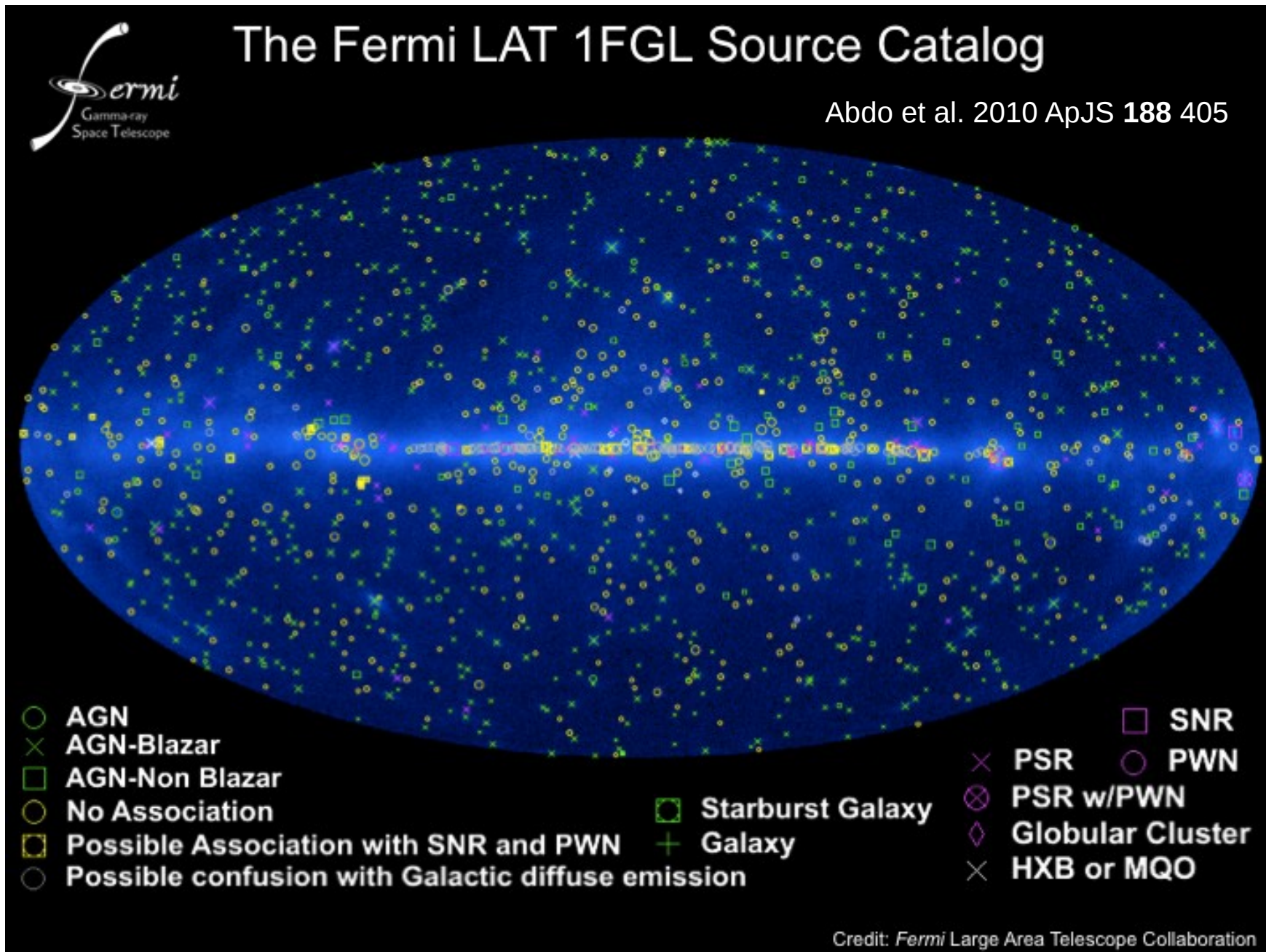


nonthermal \rightarrow high-energy particles

Diffuse emission



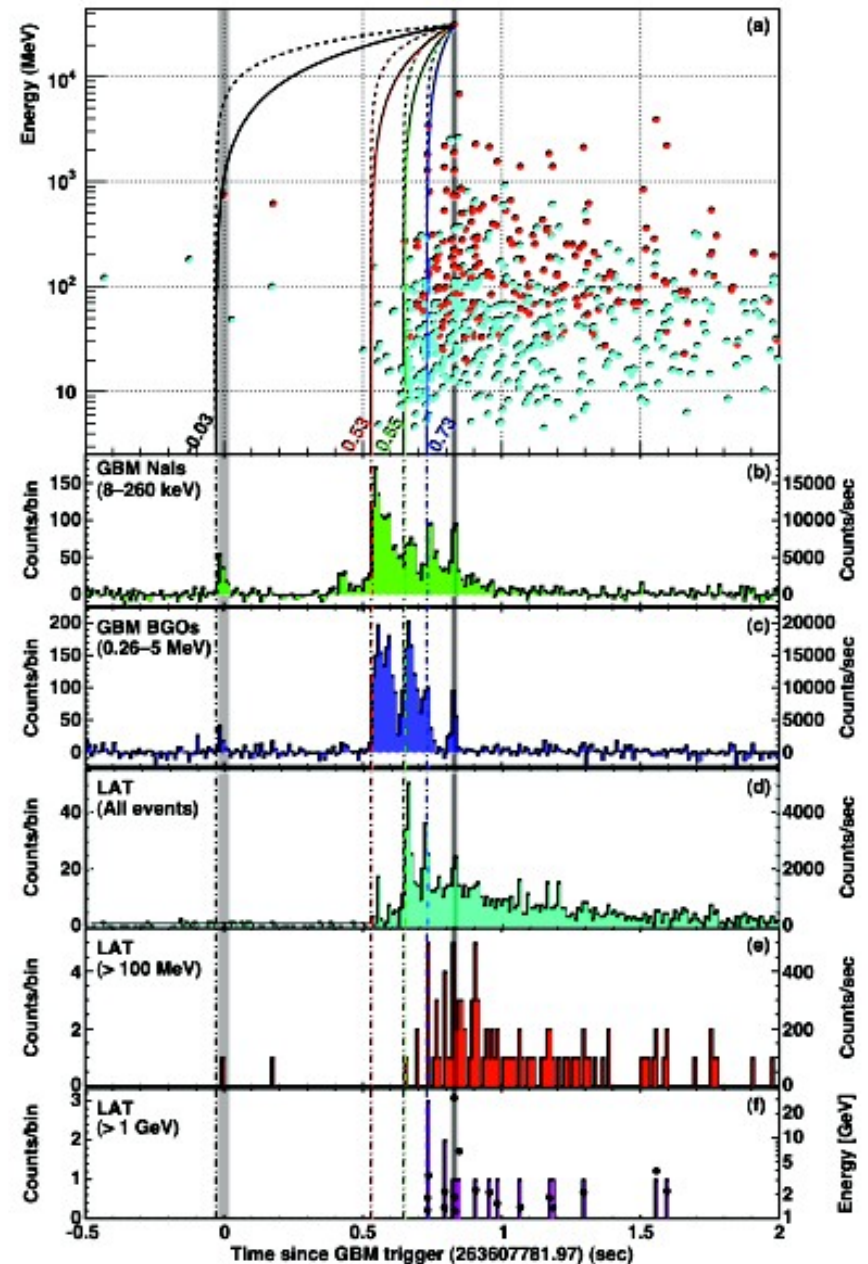
Gamma-ray sources



2. Overview of high-energy gamma-ray sources

Gamma-ray bursts

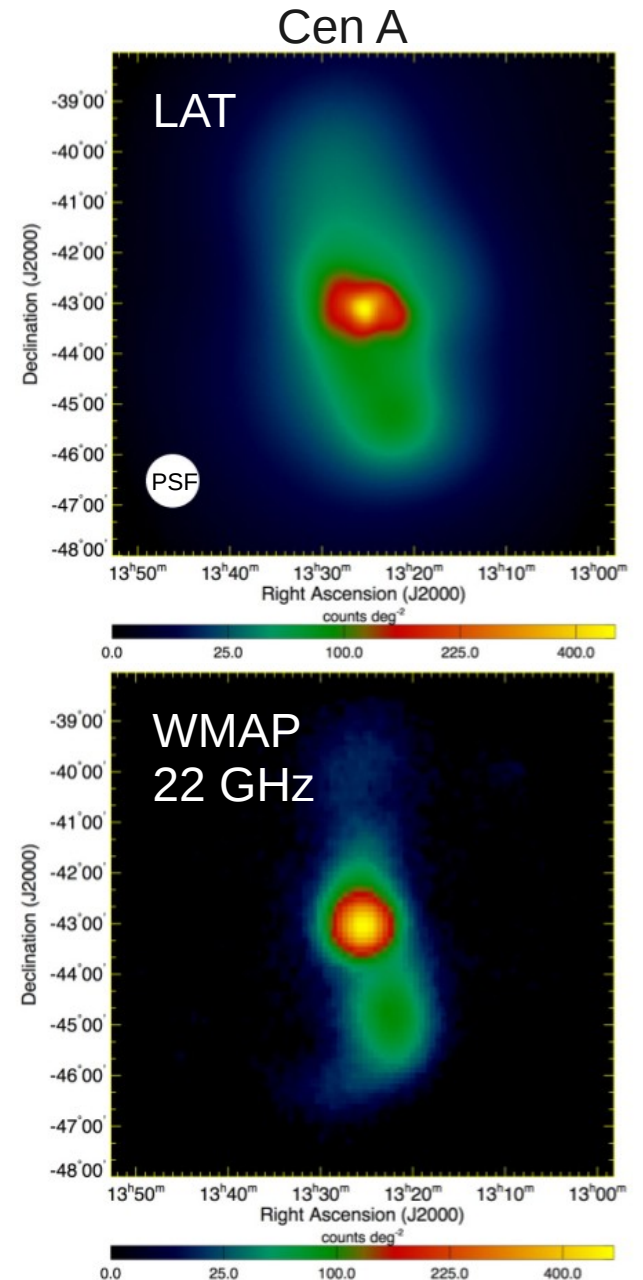
- both long (> 2 s, core-collapse supernovae) and short (< 2 s, compact object merging)
- emission mechanism: leptonic or hadronic?
- 31 GeV photon from GRB 090510 ($z=0.903$)
 - bulk Lorentz factor > 1000
 - limit on breaking of Lorentz invariance: $M > 1.19 M_{\text{Planck}}$



Abdo et al. 2009 Nature **462** 331

Active Galactic Nuclei

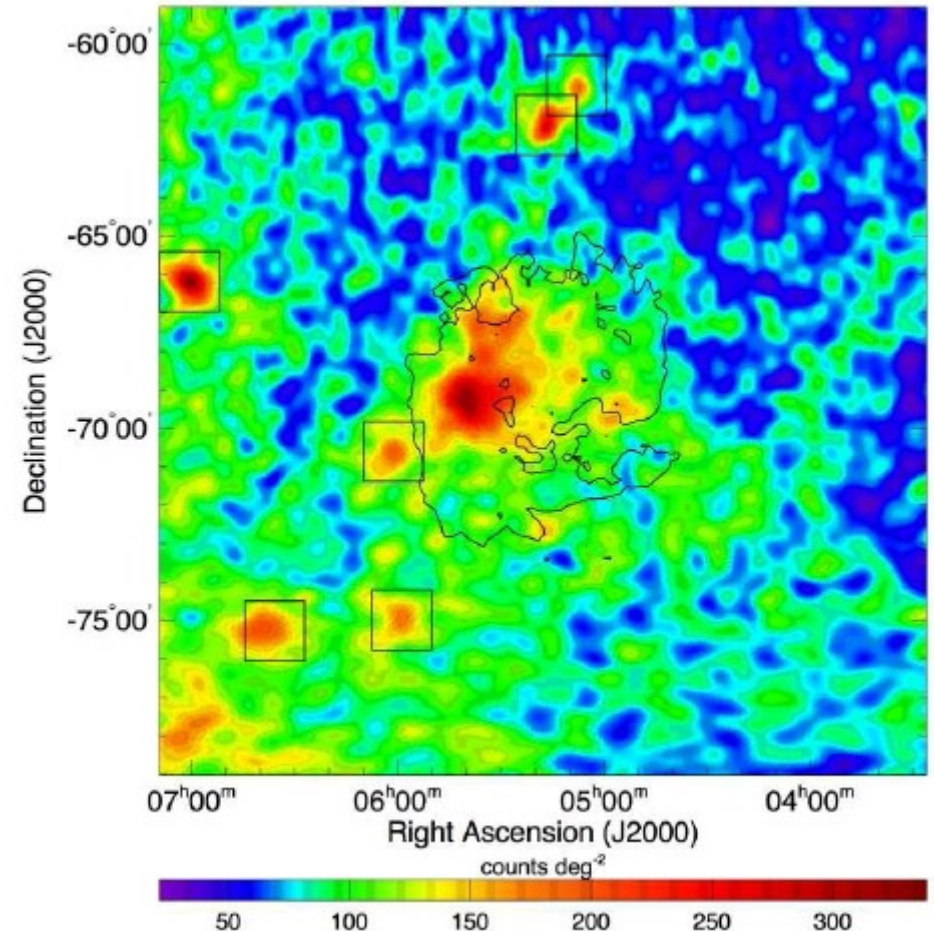
- accretion on super-massive black hole
- first non power-law spectra: internal opacity or cutoff of accelerated particles?
- Centaurus A: first AGN resolved in gammas
- 3C 279: correlated gamma-ray flare and change in optical polarization → emission at $> 10^5$ gravitational radii from BH
- luminosity function → $< 30\%$ of extragalactic background made by unresolved AGNs
- open questions:
 - radiation mechanism: hadronic vs leptonic?
 - emission region and beam formation



Abdo et al. 2010 Science **328** 765

External galaxies (non AGN)

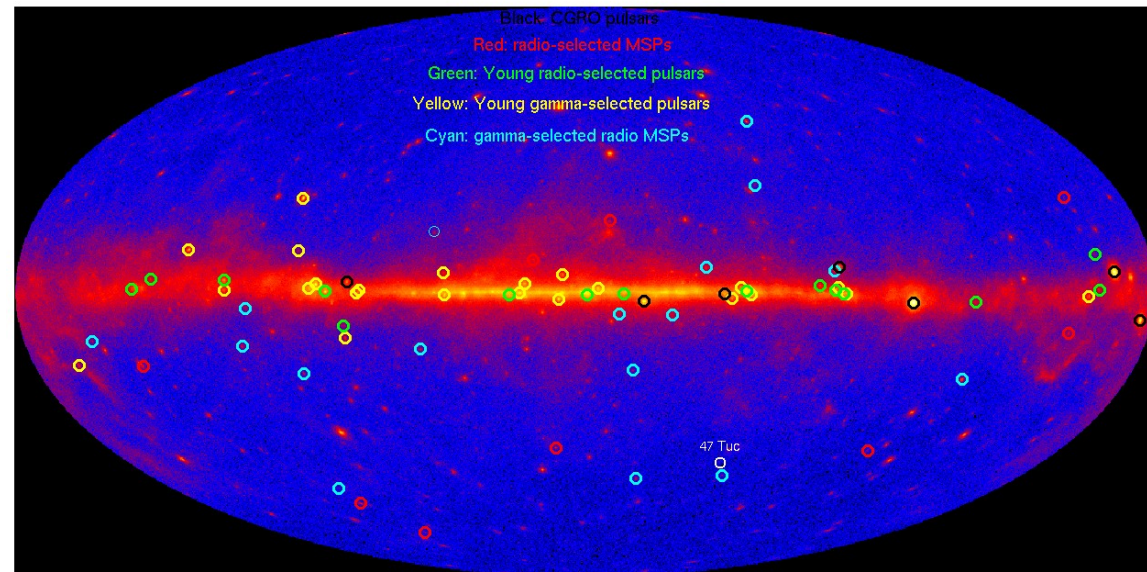
- CR interactions in interstellar space, both nuclei and electrons
- from 1 to 5 detected so far
- local group (LMC,SMC,M31) or nearby starbusts (M82,NGC 253)
- the closest (LMC,SMC) are resolved → CR acceleration/propagation



Abdo et al. 2010 A&A **512** A7

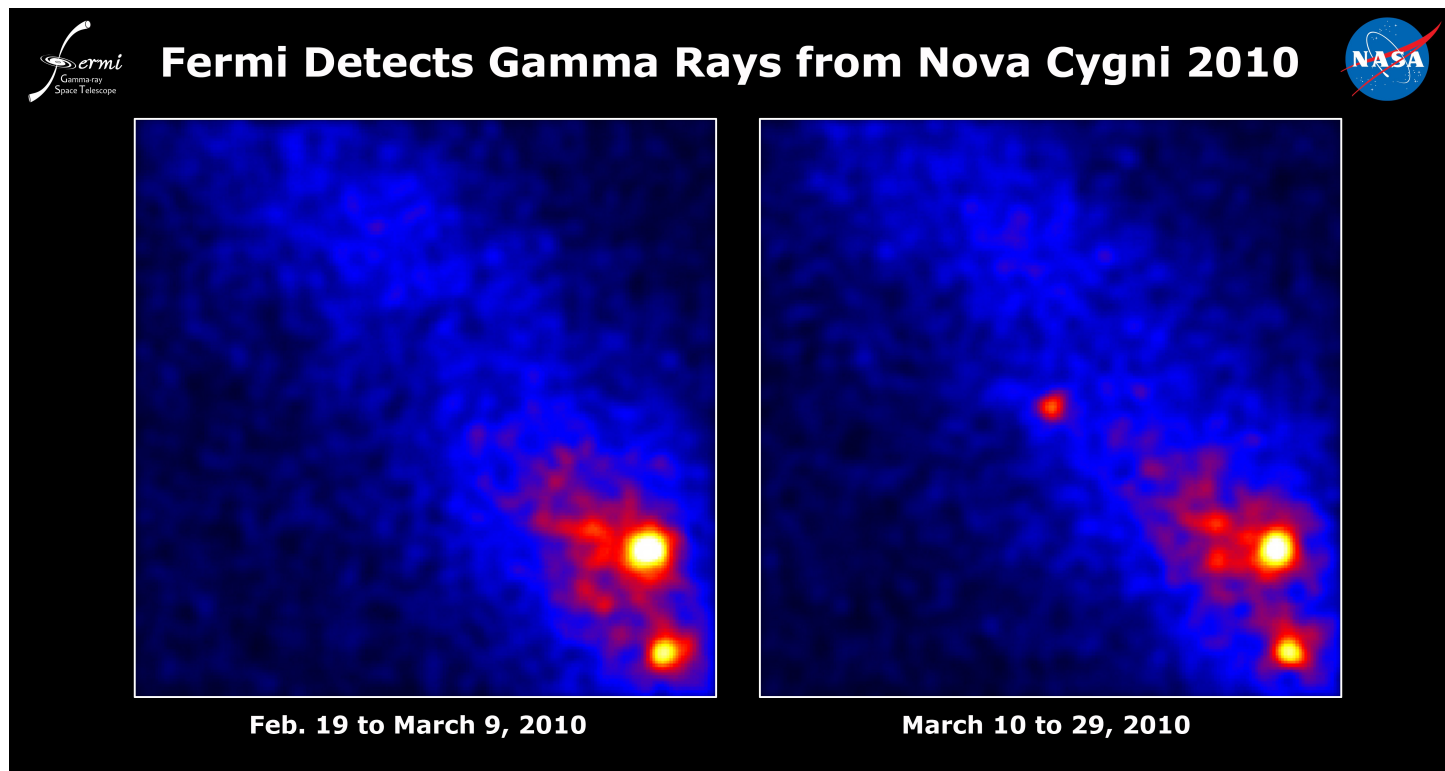
Pulsars and pulsar winds

- PSR: rotating neutron stars with high magnetic fields
 - from 6 to > 70 , including several gamma selected and ms
 - particle acceleration at high B
 - contribution to cosmic rays?
 - pulsar timing arrays \rightarrow gravitational waves
- mechanical energy into PSR wind: particle acceleration at terminal shock
 - a few detected
- gamma-ray production mechanism is **leptonic**
- first variability detection from Crab Nebula:
 - synchrotron from PeV electrons within 17 light days \rightarrow terminal shock



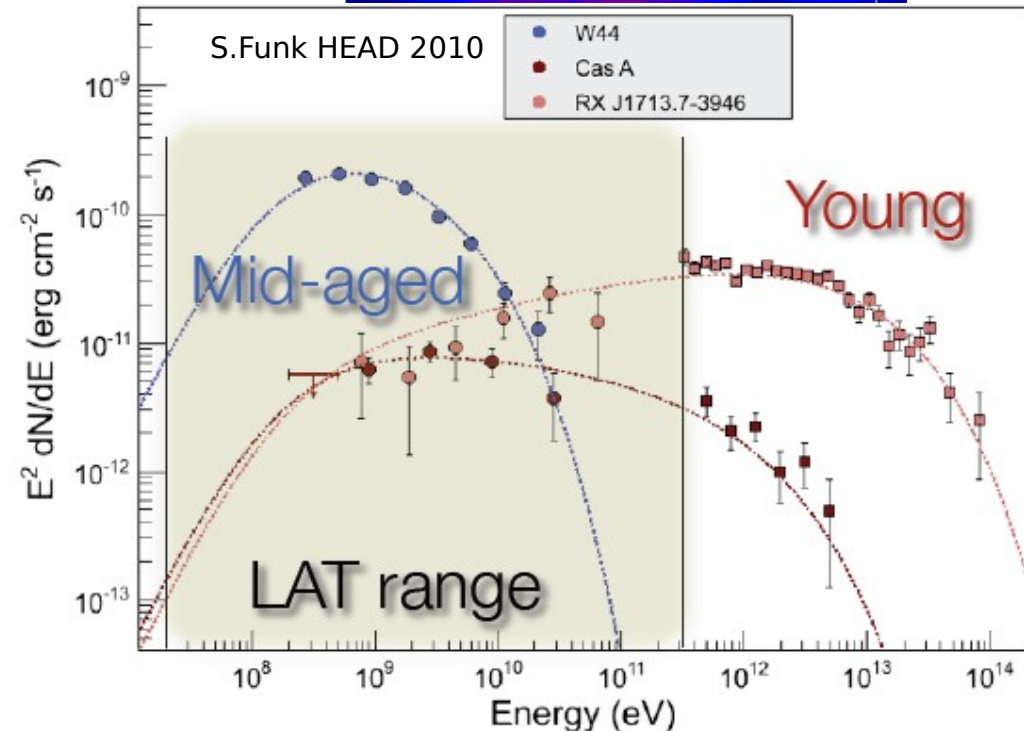
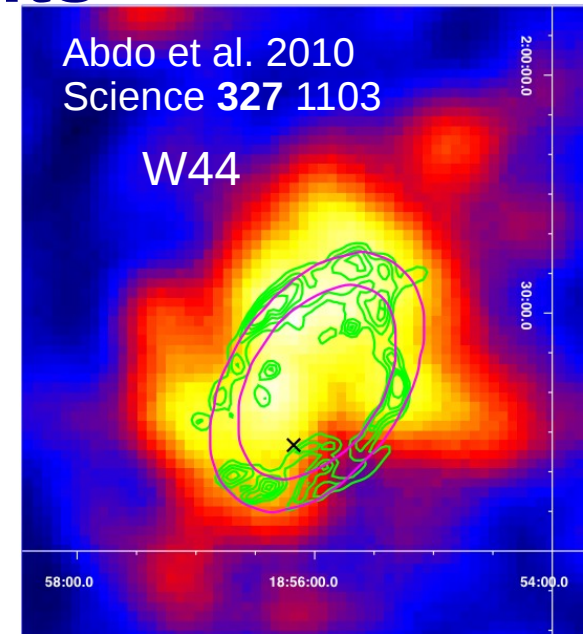
Binaries and Nova(e)

- high-mass binaries (n star or black hole+star) identified as GeV-TeV sources through orbital modulation
 - emission mechanism unclear, most GeV spectra similar to PSR
- unexpected discovery of gamma-rays from a nova (white dwarf+red giant) → supernova little cousin
 - hadronic interactions of shock-accelerated ions?



Supernova remnants

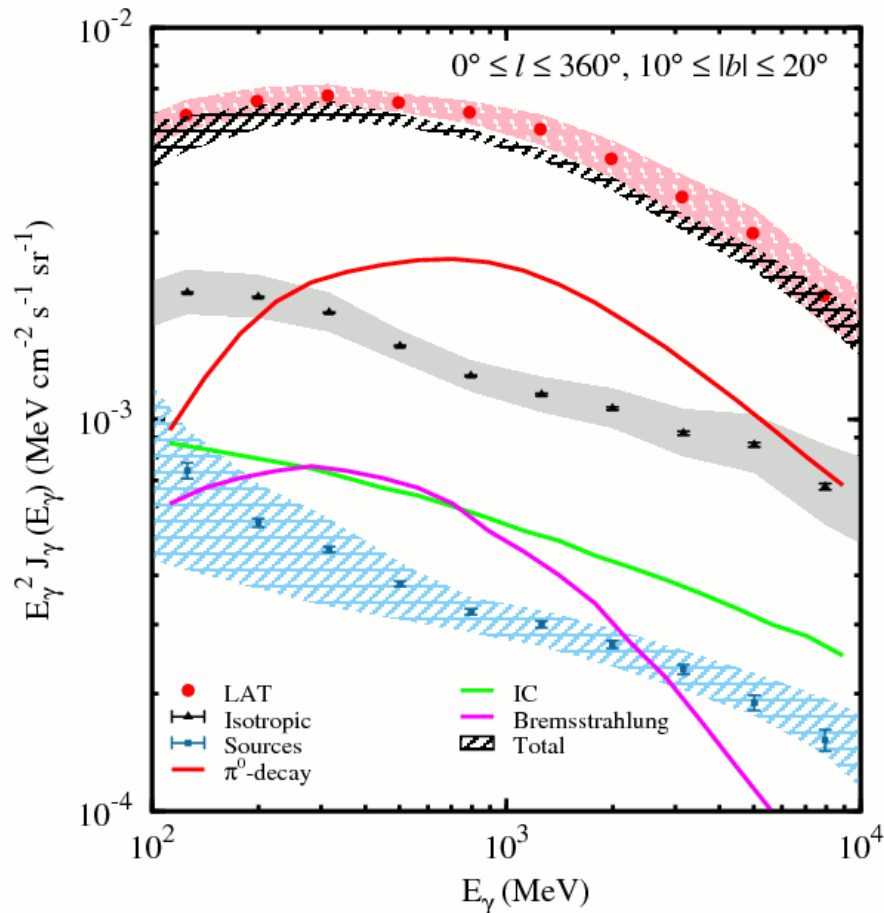
- high-energy particles accelerated by collisionless shock acceleration
 - they do for electrons
 - not so obvious (yet) for ions
- SNRs resolved → association between gamma rays and shock region
- most associated with molecular clouds
- middle-aged SNRs ($> 10^4$ years):
 - show spectral break at few GeV
 - many favor hadronic scenario



3. Interstellar emission and the gamma/cosmic-ray connection

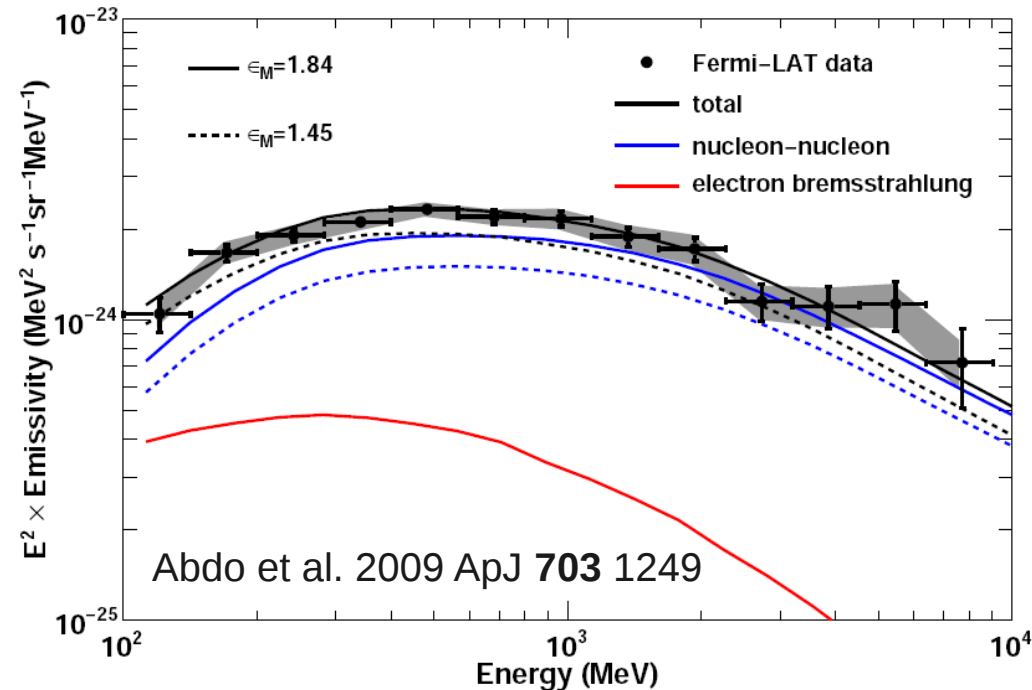
Local interstellar emission

- basic emission mechanisms understood
- no confirmation of GeV excess (related DM scenarios)



Abdo et al. 2009 PRL **103** 251101

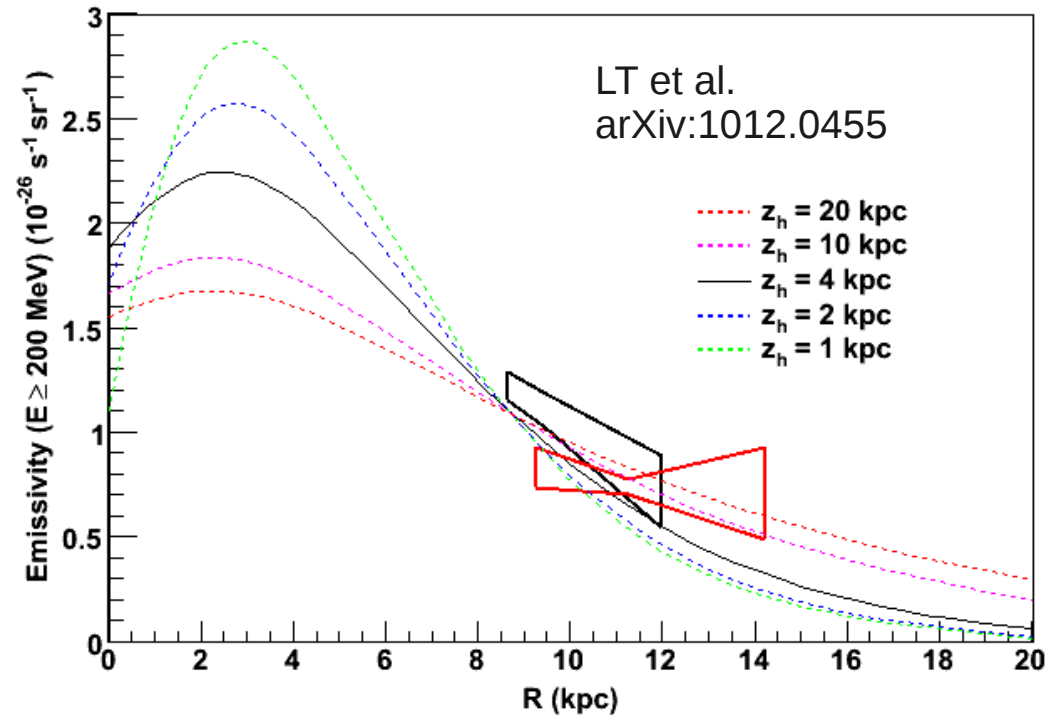
- gamma-ray emission from local interstellar matter → within 1 kpc CR spectra similar to what we measure at the Earth



emissivity = gamma-ray emission rate
per H atom

Large-scale emission from the Milky Way

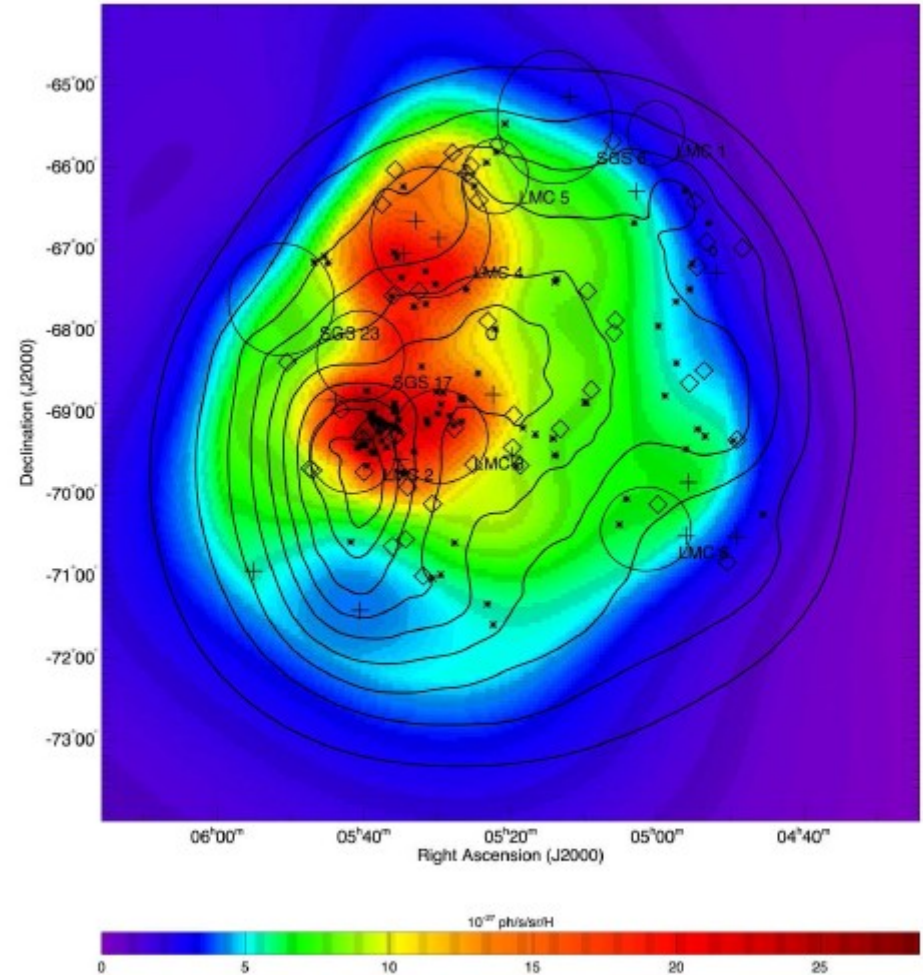
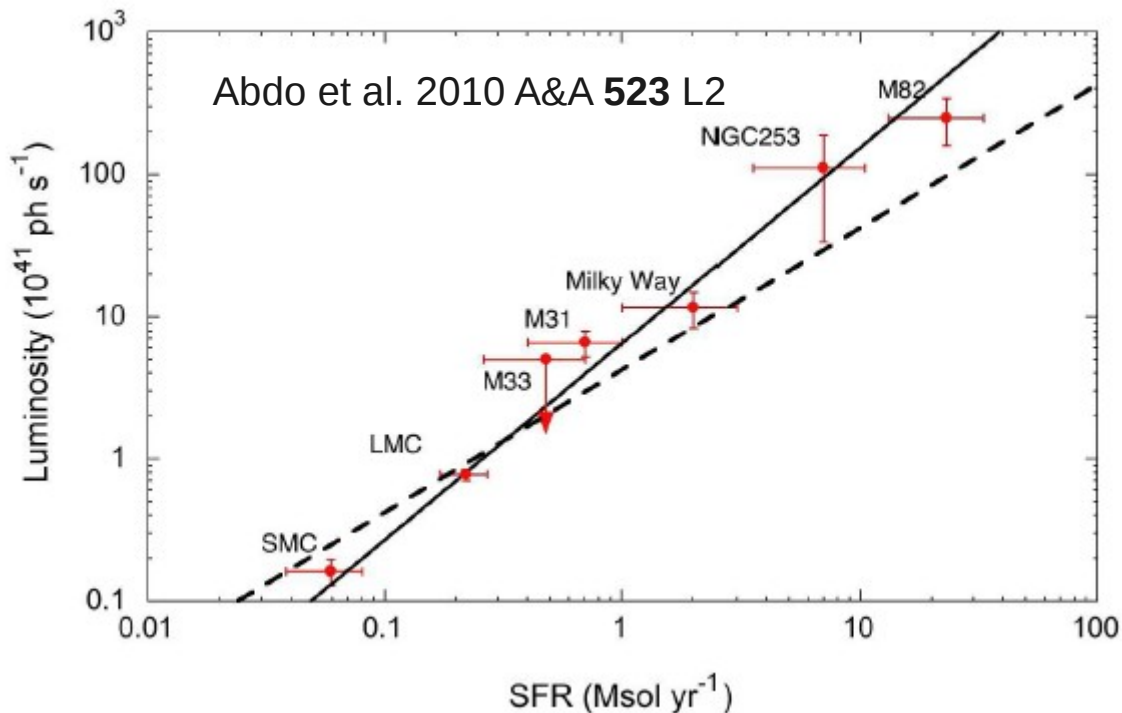
- interstellar emission traces CRs throughout the Galaxy
- first surprise: large CR densities in outer Galaxy
 - large propagation halo (or more complex propagation scenarios)
 - dark accelerators or dark gas
- more to come from comparison with propagation models over the whole Galaxy



External galaxies: the role of massive stars

Emissivity maps of LMC shows correlation with:

- H α line
 - Wolf-Rayet stars and super shells
- massive star formation



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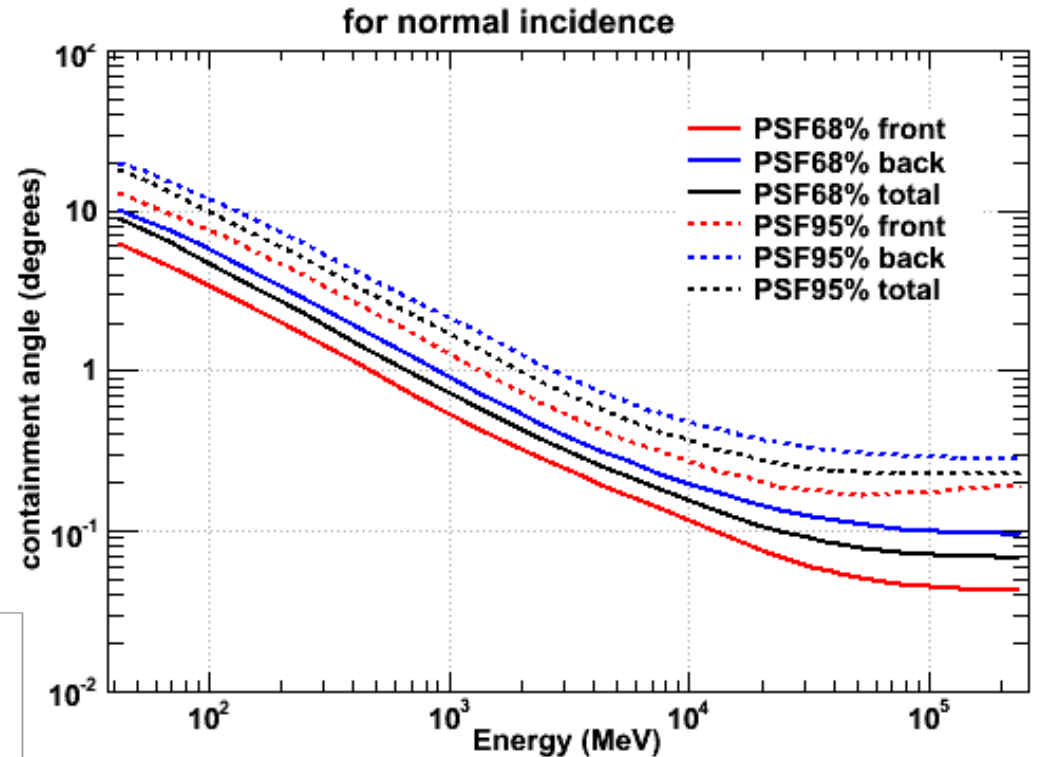
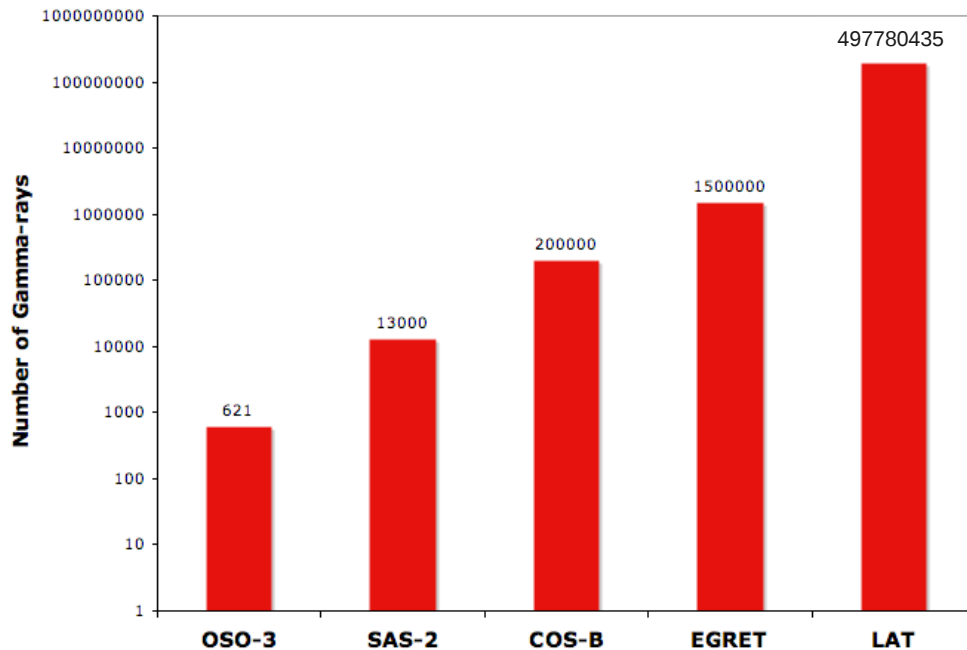
Gamma luminosity-star formation rate correlation.

4. Final remarks: Particle Physics and Astrophysics

The technological link

E.g.: LAT advances due to Si tracking devices

- Improved angular resolution
- Improved detection efficiency (together with geometry ...)



Primarily developed in Particle Phys. experiments

→ large impact on HE Astrophys.

Cosmic-ray Physics

- Interfacing two infinities:
 - birth of subnuclear Physics
 - fundamental component of galaxies
 - signatures of Physics beyond the Standard Model?
- The LAT is contributing to the understanding of HE particles in the Universe
 - studying possible acceleration sites
 - tracing their spreading all around
- Interconnection with:
 - accelerator experiments
 - charged CR detectors
 - neutrino experiments
 - gravitational waves searches

