

**Comments on
new bounds
on neutrino
magnetic
moment**

**“Results and
Perspectives
in Particle
Physics”**

**La Thuile
28/02/2011**

**Alexander
Studenikin**
MSU R
JINR

Outline

- *Neutrino magnetic moment*
- *Bruno Pontecorvo
Centenary Council*
- *Conclusions and greetings*

Why

magnetic moment

of



?

...Why

*electromagnetic
properties* ✓

*provide a kind of
window / bridge
to*

NEW Physics ?

... in spite of

- *results of terrestrial laboratory experiments on ν EM properties*

as well as

- *data from astrophysics and cosmology are in agreement with " ν ERO" EM*

... However, in course of recent developments in knowledge on ν mixing and oscillations



1 **Carlo Giunti, Alexander Studenikin :**
“Neutrino electromagnetic properties”
Phys.Atom.Nucl. 73, 2089-2125 (2009)
arXiv:0812.3646 v5, Apr 12, 2010

2 **A.Studenikin :**
“Neutrino magnetic moment: a window to new physics”
Nucl.Phys.B (Proc.Supl.) 188, 220 (2009)

3 **C. Giunti, A. Studenikin :**
“Electromagnetic properties of neutrino”
J.Phys.: Conf.Series. 203 (2010) 01210
arXiv:1006.3646 June 8, 2010

4 **C.Giunti, A.Studenikin :** **“Theory and phenomenology of neutrino electromagnetic properties”**
Rev.Mod.Phys. (in preparation)

Carlo Giunti, Alexander Studenikin


“Neutrino electromagnetic prop

Phys.Atom.Nucl. 73, 2089-2125 (2009)


arXiv:0812.3646 v5, Apr 12, 2010

 exhibits unexpected properties (p

W. Pauli, 1930

● neutral "neutron" \Rightarrow  E. Fermi, 1933

● probably $\mu_\nu \neq 0$! ?

...recent claims for
new experimental
bounds on μ_ν
continue chain
of puzzles.. 

● Pauli himself wrote to Baade:

"Today I did something a physicist should
I predicted something which will never be
experimentally...".

...the present status...

to have visible $\mu \neq 0$

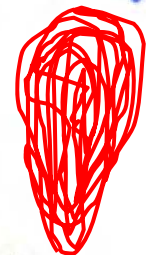
is not an easy task for

theoreticians

and experimentalists

In the Standard Model : $m_\nu = 0$,
there is no $\nu_R \Rightarrow$
 ν magnetic moment $\mu_\nu = 0$.

Thus, $\mu_\nu \neq 0 \leftarrow$ beyond the SM.



$$m_\nu \neq 0$$

Theory (Standard Model with ν_R)

$$\mu_e^\nu = \frac{3eG_F}{8\sqrt{2}\pi^2} m_\nu \sim 3 \cdot 10^{-19} \mu_B \left(\frac{m_\nu}{1\text{eV}} \right), \quad \mu_B = \frac{e}{2m_e}$$

Lee Shrock, 1977; Fujikawa Shrock, 1980

$$a_e = \frac{\alpha_{QED}}{2\pi} \sim 10^{-3}$$



... much greater values are desired

for astrophysical or cosmology

visualization of μ_ν

Astrophysical bounds

$$\mu_\nu \leq 3 \cdot 10^{-12} \mu_B$$

(Red Giant Lumin.)
etc.

G. Raffelt, D. Dearborn,
J. Silk, 1989.

Theory (Standard Model with ν_R)

$$\mu_e = \frac{3eG_F}{8\sqrt{2}\pi^2} m_{\nu_e} \sim 3 \cdot 10^{-19} \mu_B \left(\frac{m_{\nu_e}}{1\text{eV}} \right), \quad \mu_B = \frac{e}{2m_e}$$

Lee Shrock, 1977; Fujikawa Shrock, 1980

... importance μ_ν studies

If diagonal $\mu_\nu \neq 0$

were confirmed

then \checkmark Dirac

... for \checkmark Majorana
non-diagonal = transitional
 $\mu_\nu \neq 0$

... progress
in experimental μ_ν \Rightarrow
studies or

Studies of ν - e scattering - most sensitive method of experimental investigation of μ_ν

Cross-section:

$$\frac{d\sigma}{dT}(\nu + e \rightarrow \nu + e) = \left(\frac{d\sigma}{dT}\right)_{\text{SM}} + \left(\frac{d\sigma}{dT}\right)_{\mu_\nu}$$

where the Standard Model contribution

$$\left(\frac{d\sigma}{dT}\right)_{\text{SM}} = \frac{G_F^2 m_e}{2\pi} \left[(g_V + g_A)^2 + (g_V - g_A)^2 \left(1 - \frac{T}{E_\nu}\right)^2 + (g_A^2 - g_V^2) \frac{m_e T}{E_\nu^2} \right],$$

T is the electron recoil energy

$$\left(\frac{d\sigma}{dT}\right)_{\mu_\nu} = \frac{\pi \alpha_{em}^2}{m_e^2} \left[\frac{1 - T/E_\nu}{T} \right] \mu_\nu^2$$

$$\mu_\nu^2 = \sum_{j = \nu_e, \nu_\mu, \nu_\tau} |\mu_{ij} - \epsilon_{ij}|^2$$

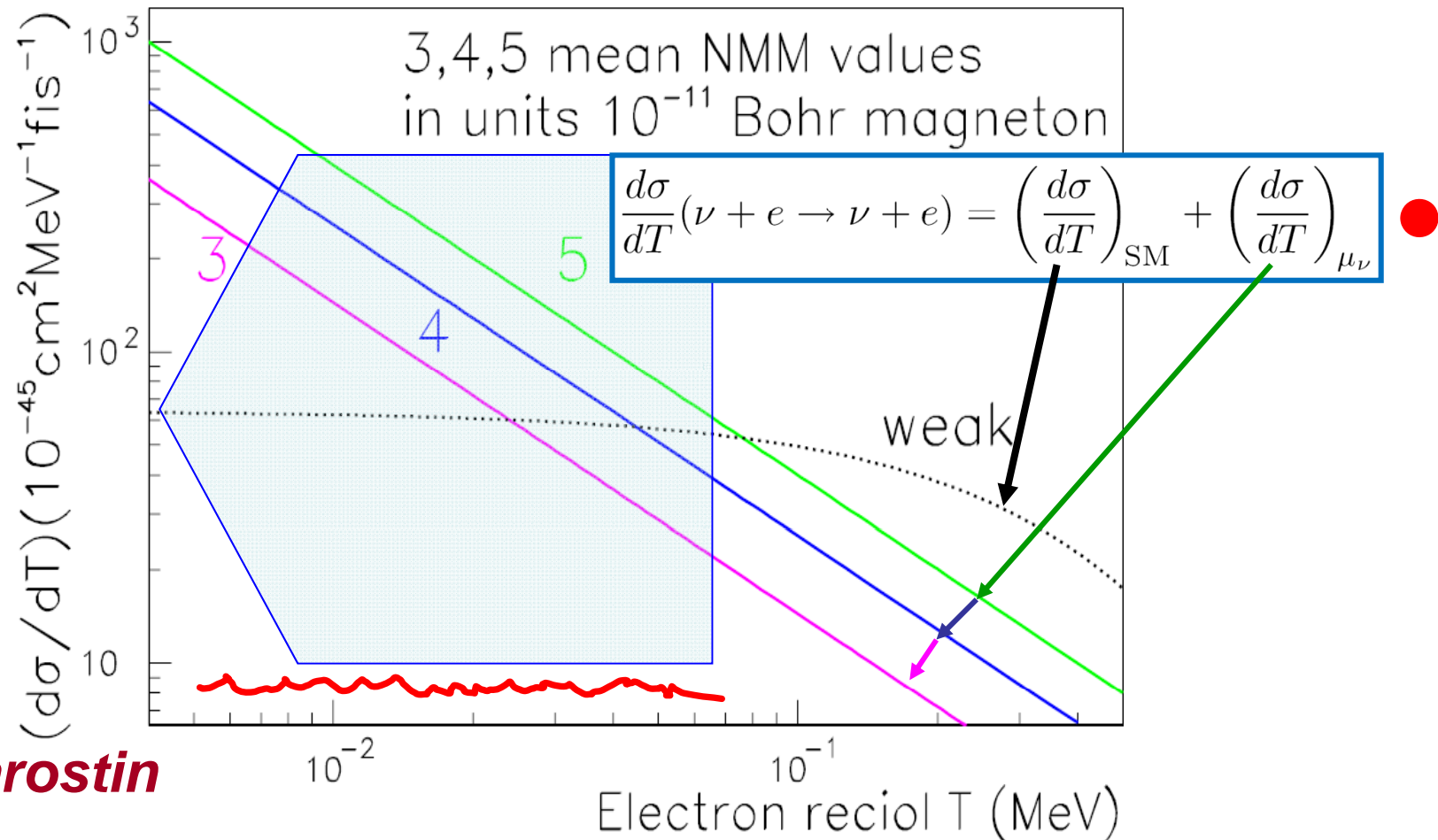
$$g_V = \begin{cases} 2 \sin^2 \theta_W + \frac{1}{2} & \text{for } \nu_e, \\ 2 \sin^2 \theta_W - \frac{1}{2} & \text{for } \nu_\mu, \nu_\tau, \end{cases} \quad g_A = \begin{cases} \frac{1}{2} & \text{for } \nu_e, \\ -\frac{1}{2} & \text{for } \nu_\mu, \nu_\tau \end{cases} \quad \text{for anti-neutrins } g_A \rightarrow -g_A$$

to incorporate charge radii $g_V \rightarrow g_V + \frac{2}{3} M_W^2 \langle r^2 \rangle \sin^2 \theta_W$

Magnetic moment contribution is dominated at low

recoil energies $\nu \left(\frac{d\sigma}{dT} \right)_{\mu\nu} > \left(\frac{d\sigma}{dT} \right)_{SM}$ $\frac{T}{m_e} < \frac{\pi^2 \alpha_{em}}{G_F^2 m_e^4} \mu_\nu^2$ and

... the lower the smallest measurable electron recoil energy, the smaller value μ_ν^2 can be probed in scattering



from A.Starostin



MUNU experiment at Bugey reactor (2005)

$$\mu_{\nu} \leq 9 \times 10^{-11} \mu_B$$

TEXONO collaboration at Kuo-Sheng power plant (2006)

$$\mu_{\nu} \leq 7 \times 10^{-11} \mu_B$$

GEMMA (2007) $\mu_{\nu} \leq 5.8 \times 10^{-11} \mu_B$

GEMMA I 2005 - 2007

BOREXINO (2008) $\mu_{\nu} \leq 5.4 \times 10^{-11} \mu_B$

...was considered as the world best constant

$$\mu_{\nu} \leq 8.5 \times 10^{-11} \mu_B \quad (\nu_{\tau}, \nu_{\mu})$$

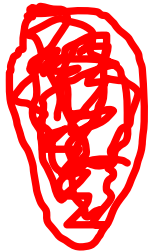
Montanino,
Picariello,
Pulido, PRD 2008

based on first release of
BOREXINO data

GEMMA (2005-2008)

Germanium Experiment on measurement of Magnetic Moment of Antineutrino

R (Dubna) + ITEP (Moscow) at Kalinin Nuclear Power P



$$\mu_\nu < 3.2 \times 10^{-11} \mu_B$$

...till 13 January 2010 and again since 23 August
best limit \checkmark magnetic moment

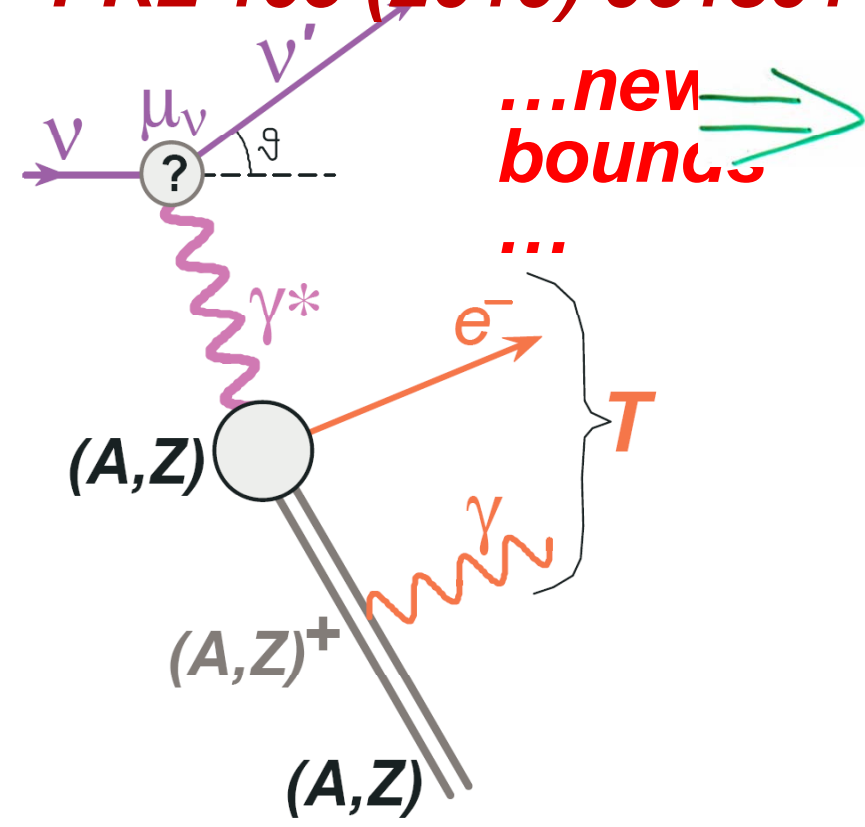
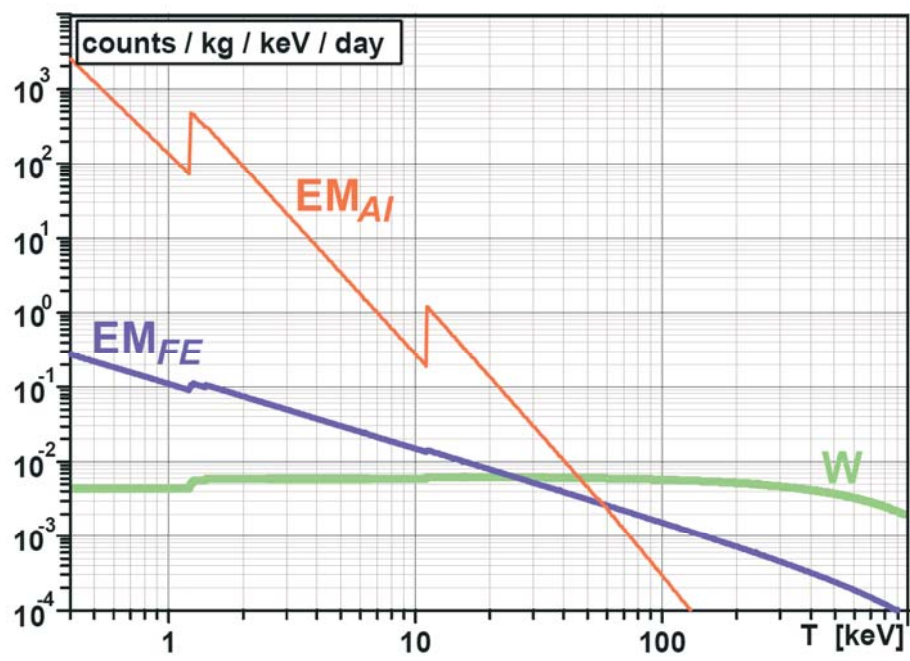
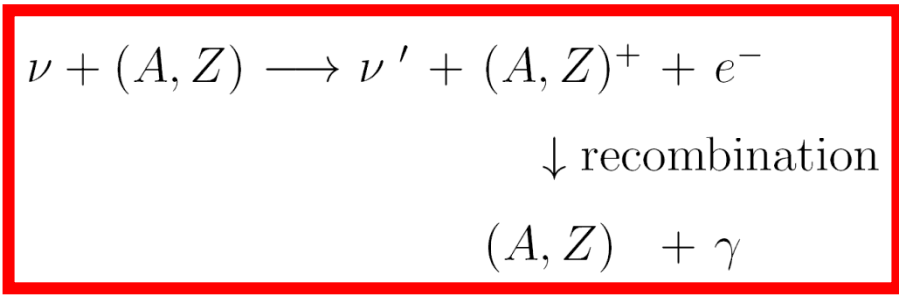
A.Beda, E.Demidova, A.Starostin et al,
arXiv:09.06.1926 , June 10, 2009,

A.Beda, V.Brudanin, E.Demidova et al,
in: "Particle Physics on the Eve of LHC",
ed. A.Studenikin, World Scientific (Singapore),
p.112, 2009 (13th Lomonosov Conference)

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... quite recent **claim**
 that ν - e cross section
 should be increased by
Atomic Ionization effect

H.Wong et al. (TEXONO
 Coll.), arXiv: 1001.2074,
 13 Jan 2010,
 reported at
**Neutrino 2010
 Conference
 (Athens, June 2010),
 PRL 105 (2010) 061801**



...much better limit ν on

effective magnetic

$\mu_\nu < 1.3 \times 10^{-11} \mu_B$



H.Wong et al.,
(TEXONO Coll.)
arXiv: 1001.2074
13 Jan 2010,
PRL 105 (2010)
061801

... atomic ionization effect accounted for ...

Neutrino 2010 Conference,
Athens ... however ...



$\mu_\nu < 5.0 \times 10^{-12} \mu_B$



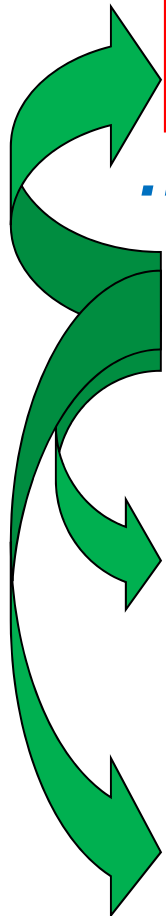
A.Beda et al.
(GEMMA Coll.),
arXiv: 1005.2736,
16 May 2010

... atomic ionization effect accounted for ...

$\mu_\nu < 3.2 \times 10^{-11} \mu_B$



... ν -e scattering on free electrons ...
(without atomic ionization)



K.Kouzakov, A.Studenikin,

- **“Magnetic neutrino scattering on atomic electrons revisited”
Phys.Lett. B 105 (2011) 061801, arXiv: 1011.5847**
- **“Electromagnetic neutrino-atom collisions: The role of electron binding”
to appear in
Nucl.Phys.B (Proc.Supp.) 2011 (Proc. of Neutrino Oscillation Workshop)**

K.Kouzakov, A.Studenikin, M.Voloshin,

- **“Neutrino-impact ionization of atoms in search for neutrino magnetic moment”
arXiv: 1101.4878, 25 Jan 2011, submitted to *Phys.Rev.D***
- **“On neutrino-atom scattering in searches for neutrino magnetic moment”
arXiv: 1102.0643, 3 Feb 2011, to appear in
*Nucl.Phys.B (Proc.Supp.) 2011 (Proc. of Neutrino 2010 Conference)***

M.Voloshin,

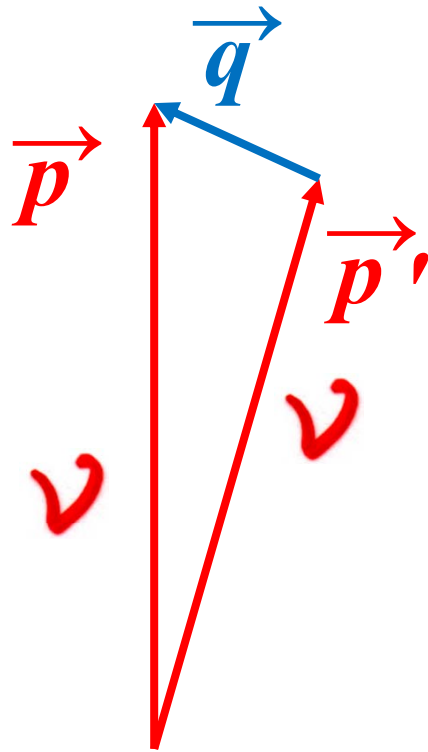
- **“Neutrino scattering on atomic electrons in search for neutrino magnetic moment”
Phys.Rev.Lett. 105 (2010) 201801, arXiv: 1008.2171**



***No important effect of Atomic
ionization on $cr \mu_s$ section in
experiments once all possible
final electronic states accounted
for***

***M.Voloshin, 23 Aug 2010;
K.Kouzakov, A.Studenikin, 26 Nov 2010;
H.Wong et al, arXiv: 1001.2074 V3, 28 Nov
2010***

Neutrino-impact ionization of atoms in search for μ_ν



scattering on atoms (Ge) at low energy T

$T \sim \text{few keV}$ and lower $s \left(\frac{T}{E_\nu} \ll 1 \right)$

Ge atom recoil energy $\ll \frac{2E_\nu^2}{M_{\text{Ge}}} \ll T$ $M_{\text{Ge}} \rightarrow \infty$ ✓

✓ interaction with nucleus is neglected

✓ scattering on atomic **e** is important

Four momentum transfer

$$q = p - p'$$

$$q_\mu = (T, \vec{q}), \quad q^2 = \vec{q}^2$$

energy and spatial momentum transfer from neutrinos to atomic electrons

Kouzakov,
Studenikin, 2010;
Kouzakov,
Studenikin,
Voloshin, 2011

Double differential ν -e cross section

$$\frac{d^2\sigma_{(\mu)}}{dT dq^2} = 4\pi \alpha \frac{\mu_\nu^2}{q^2} \left[\left(1 - \frac{T^2}{q^2}\right) S(T, q^2) + \left(1 - \frac{q^2}{4E_\nu^2}\right) R(T, q^2) \right]$$

$$\frac{d^2\sigma_{(\mu)}}{dT dq^2} = \left(\frac{d^2\sigma_{(\mu)}}{dT dq^2}\right)_{\parallel} + \left(\frac{d^2\sigma_{(\mu)}}{dT dq^2}\right)_{\perp}$$

**Kouzakov,
Studenikin, 201
Kouzakov,
Studenikin,
Voloshin, 2011**

where dynamical structure factor (Van Hove, 1954)

$$S(T, q^2) = \sum_n \delta(T - E_n + E_0) |\langle n | \rho(\vec{q}) | 0 \rangle|^2 \quad \text{and} \quad (\vec{j}_{\perp} \cdot \vec{q}) = 0$$

$$R(T, q^2) = \sum_n \delta(T - E_n + E_0) |\langle n | j_{\perp}(\vec{q}) | 0 \rangle|^2$$

summ is over all states $|n\rangle$ of electron system,

For single-differential inclusive cross section measured

$$\frac{d\sigma_{(\mu)}}{dT} = 4\pi \alpha \mu_\nu^2 \int_{T^2}^{4E_\nu^2} S(T, q^2) \frac{dq^2}{q^2}$$

$$R(T, q^2) = \frac{T^2}{q^2} S(T, q^2)$$

transversal contribution practically for most q^2 is negligible

SM electroweak contribution to cross section

$$\frac{d\sigma_{EW}}{dT} = \frac{G_F^2}{4\pi} (1 + 4 \sin^2 \theta_W + 8 \sin^4 \theta_W) \int_{T^2}^{4E_\nu^2} S(T, q^2) dq^2$$

nonrelativistic limit $\int_{T^2}^{4E_\nu^2} \Rightarrow \int_0^\infty$

For free electron

$$S_{(FE)}(T, q^2) = \delta(T - q^2/2m)$$

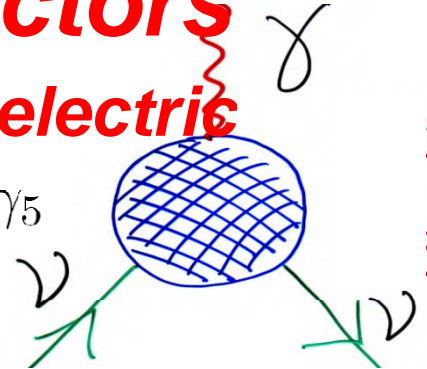
$$\int_0^\infty S_{(FE)}(T, q^2) \frac{dq^2}{q^2} = \frac{1}{T}, \quad \int_0^\infty S_{(FE)}(T, q^2) dq^2 = 2m$$

$$\frac{d\sigma_{(\mu)}}{dT} = 4\pi \alpha \mu_\nu^2 \left(\frac{1}{T} - \frac{1}{E_\nu} \right) = \pi \frac{\alpha^2}{m^2} \left(\frac{\mu_\nu}{\mu_B} \right)^2 \left(\frac{1}{T} - \frac{1}{E_\nu} \right)$$

- **... the same for electron bound in atom... ~~atom~~ free electron approximation is valid**
(ν -e scattering on free electrons)

ν e.m. vertex function \rightarrow **4 form factors**
charge dipole magnetic and electric

● $\Lambda_\mu(q) = f_Q(q^2)\gamma_\mu + f_M(q^2)i\sigma_{\mu\nu}q^\nu + f_E(q^2)\sigma_{\mu\nu}q^\nu\gamma_5 + f_A(q^2)(q^2\gamma_\mu - q_\mu\not{q})\gamma_5$
anapole



● **EM properties** \rightarrow **a way to distinguish Dirac**

● **Standard Model with ν_R ($m_\nu \neq 0$):** $\mu_e = \frac{3eG_F}{8\sqrt{2}\pi^2} m_\nu \sim 3 \cdot 10^{-19} \mu_B \left(\frac{m_\nu}{1\text{eV}}\right)$

● **In extensions of SM** \rightarrow **enhancement of magnetic moment, even electrically millicharged**

● **Limits from reactor ν -e scattering experiments (2009, 2010):** $\mu_\nu < 3.2 \times 10^{-11} \mu_B$
 ● **Limits from astrophysical star cooling (1990):** $\mu_\nu < 3 \times 10^{-12} \mu_B$

$\mu_\nu < 3.2 \times 10^{-11} \mu_B$

A.Beda et al. (GEMMA Coll.)

$\mu_\nu < 3 \times 10^{-12} \mu_B$

G.Raffelt



Бруно Понтекорво

- Since 1950 **Bruno Pontecorvo**, outstanding Italian scientist, lived in Russia and was staff member of Joint Institute for Nuclear Research, Dubna
- **Bruno Pontecorvo** was Head of Department of Particle Physics and member of Scientific Council of Faculty of Physics of Moscow State University
- **Bruno Pontecorvo Laboratory on Neutrino Physics and Astrophysics (PLN)**

has been recently established at Faculty of Physics of Moscow State University

... to provide continuation of long-standing traditions in teaching

and performing scientific researches of neutrinos of relat

3 *proposals*

Faculty of Physics

Moscow State University

Dedicated to the 300th Anniversary of
the birth of Mikhail Lomonosov

FIFTEENTH LOMONOSOV CONFERENCE ON ELEMENTARY PARTICLE PHYSICS



Mikhail Lomonosov
1711-1765

Moscow, August 18 - 24, 2011

Electroweak Theory
Tests of Standard Model & Beyond
Neutrino Physics
Astroparticle Physics
Gravitation and Cosmology
Developments in QCD (Perturbative
and Non-Perturbative Effects)
Heavy Quark Physics

NINTH
INTERNATIONAL
MEETING
ON
August 24, 2011
PROBLEMS
OF INTELLIGENTSIA

Physics at the Future Accelerators
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**15th Lomonosov
Conference on
Elementary Particle
Physics,
Moscow State
University,
August 18-24, 2011**

www.icas.ru

**16th Lomonosov Conference on Elementary
Particle Physics**

MSU Moscow, August 22-28, 2013

August 22, 2013

**is the centenary of
Bruno Pontecorvo birth**



Бруно Понтекорво

1913-1993

... our 3

proposals:

- **16LomCon** will be dedicated to the memory of **Bruno Pontecorvo**
- **Scientific programme of 16LomCon** to be devoted to neutrino physics astroparticle physics and related subjects

- ***To establish***

***International Bruno Pontecorvo
Centenary Council***

***for preparing celebrations
(conferences, workshops etc)***

dedicated to

Bruno Pontecorvo heritage

MEMORANDUM

2nd Round Table Italy-Russia@Dubna:

“SPACE PHYSICS and BIOLOGY”

- In the days from 19th to 23rd December 2010 at the Joint Institute for Nuclear Research (JINR), it took place the second Round Table Italy-Russia at Dubna devoted this year to the problems of Cosmo Physics and Biology. Just as the first Round Table in December 2009, the present event was co-organized by the Italian Embassy in the Russian Federation and by JINR,

- **A) Review and Perspectives of Italo-Russian collaborations in the field of Astrophysics and Cosmology.**

In this field the joint collaboration activities are already very much alive and developed. The participants supported the proposal of devoting to the memory of Bruno Pontecorvo the 16th Lomonosov Conference on Elementary Particle Physics organized for the year 2013 by the Bruno Pontecorvo Neutrino Physics Laboratory of Moscow State University and creating an International Bruno Pontecorvo Centenary Council to prepare celebrations. Similarly the outstanding figure of the Italian physicist Nicola Cabibbo who was full member of the Russian Academy of Sciences should be honored in one of the next possible meetings.



Co-presidents of the Memorandum Committee

Academician A. Grigoriev 

Academician V. Matveev 

Prof. G. Martinelli 

Prof. M. Capaccioli 

Dubna, December 22nd , 2010

RUSSIAN ACADEMY
OF SCIENCES
(RAS)



Institute for Nuclear Research
of the Russian Academy of Sciences


Учреждение Российской академии наук
ИНСТИТУТ ЯДЕРНЫХ ИССЛЕДОВАНИЙ РАН



LXX



DEAR PROFESSOR MARIO GRECO,

 On the occasion of your seventieth jubilee I have a honour to send you warmest greetings on behalf of the Russian Academy of Sciences, the Russian scientific community and myself.

You are a world know scientist and it is a pleasure for me to emphasize your personal great contribution to different branches of high energy and particle physics.

Your brilliant scientific career started in Italy at Laboratori Nazionali di Frascati and in collaborating with Bruno Touschek, then it was continued in the world famous centers of high energy physics, in particular at CERN, SLAC and DESY. Your contribution to electroweak theory, QCD jet and heavy-quark physics are well known and highly appreciated by scientific community.

Your successful scientific research work has been accomplished with great academic teaching activity in universities of Marseille, Pavia, L'Aquila and more than for 20 years in the University of Roma Tre.

We also very much appreciate your determinative role in founding the world's prestigious series of international conference on "Results and Perspectives in Particle Physics" (Les Recontres de Physique de la Vallee d'Aoste) that is successfully running under your supervision for 25 years.

There is a long standing tradition of fruitful scientific and academic relations between Russia and Italy. It is a pleasure for me to emphasize your personal great contribution in promoting these relations and also in support of the Russian scientists during the series of Les Recontres de Physique de la Vallee d'Aoste.

Please be sure that we would be very glad to continue our contacts and collaboration in the future.

We would like to wish you new achievements in your scientific research and academic activity and prosperity to you, Mrs Halina Bilokon and all of your family.

Sincerely yours,



Victor Matveev,
Director
Institute for Nuclear Research of Russian Academy of Sciences,
Academician-Secretary, Division of Physical Sciences
Russian Academy of Sciences