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On behalf of the Double Chooz collaboration

Content

- Neutrino physics:
 θ_{13} - neutrino oscillations and reactor ν
- Status quo:
The CHOOZ experiment
- Improving CHOOZ:
“Next generation” concepts
- Focus on Double Chooz:
Design, status, time schedule and sensitivity
- The θ_{13} game is on:
Reno and Daya Bay
- Summary

Θ_{13}

- Flavor vs mass eigenstates of (Dirac) neutrinos
 → **Mixing matrix**

$$c_{ij} = \cos \Theta_{ij} \quad \text{and} \quad s_{ij} = \sin \Theta_{ij}$$

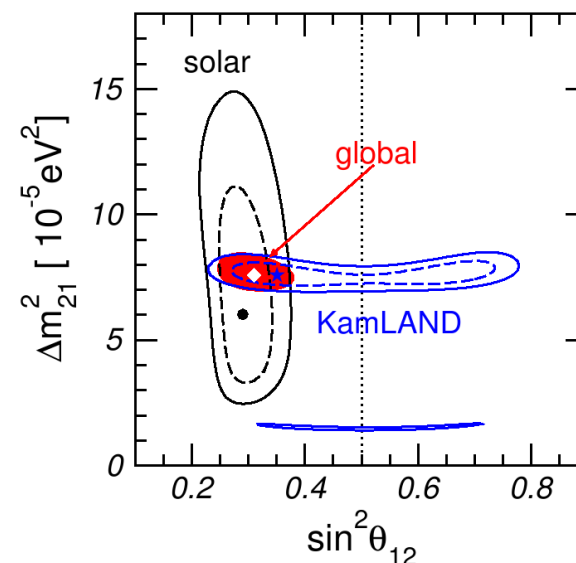
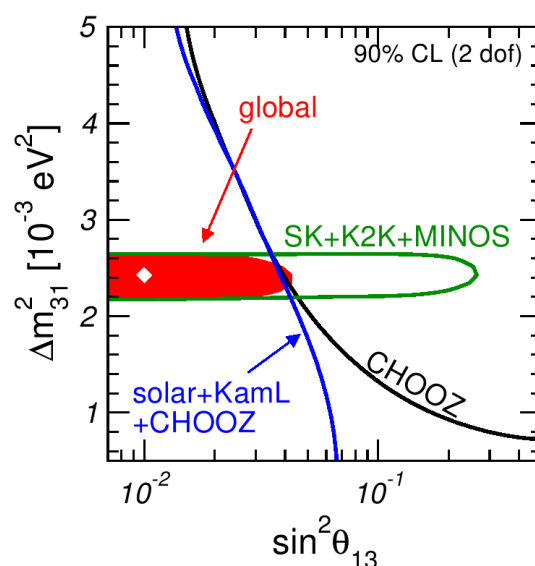
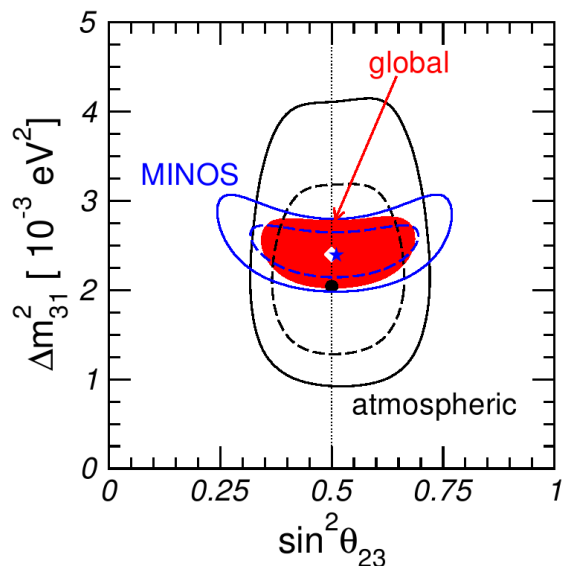
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} \mathbf{1} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & c_{23} & s_{23} \\ \mathbf{0} & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmosph. + Accel.}} \underbrace{\begin{pmatrix} c_{13} & \mathbf{0} & s_{13} e^{-i\delta} \\ \mathbf{0} & \mathbf{1} & \mathbf{0} \\ -s_{13} e^{i\delta} & \mathbf{0} & c_{13} \end{pmatrix}}_{\text{reactor}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & \mathbf{0} \\ -s_{12} & c_{12} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{1} \end{pmatrix}}_{\text{Solar + KamLAND}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Atmosph. + Accel.

reactor

Solar + KamLAND

Maltoni and Schwetz,
 PoS idm2008:072.2008



Flavor vs mass eigenstates of (Dirac) neutrinos

→ **Mixing matrix**

$$c_{ij} = \cos \Theta_{ij} \quad \text{and} \quad s_{ij} = \sin \Theta_{ij}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} \mathbf{1} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & c_{23} & s_{23} \\ \mathbf{0} & -s_{23} & c_{23} \end{pmatrix}}_{\text{Atmosph. + Accel.}} \underbrace{\begin{pmatrix} c_{13} & \mathbf{0} & s_{13} e^{-i\delta} \\ \mathbf{0} & \mathbf{1} & \mathbf{0} \\ -s_{13} e^{i\delta} & \mathbf{0} & c_{13} \end{pmatrix}}_{\text{reactor}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & \mathbf{0} \\ -s_{12} & c_{12} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{1} \end{pmatrix}}_{\text{Solar + KamLAND}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Atmosph. + Accel.

reactor

Solar + KamLAND

$$\Theta_{23} = 42.3^{+5.3}_{-2.8}$$

$$\Theta_{13} = 6.8^{+2.6}_{-3.6}$$

$$\Theta_{12} = 34.4 \pm 1.0$$

$$\Delta m_{12}^2 = 7.59 \pm 0.20 \cdot 10^{-5} \text{ eV}^2$$

$$\Delta m_{31}^2 = -2.40 \pm 0.11 \cdot 10^{-3} \text{ eV}^2 \text{ (inverted)}$$

$$\Delta m_{31}^2 = 2.51 \pm 0.12 \cdot 10^{-3} \text{ eV}^2 \text{ (normal)}$$

1 σ bounds

[Gonzalez-Garcia et al. arXiv:1001.4524v1 \[hep-ph\]](https://arxiv.org/abs/1001.4524v1)

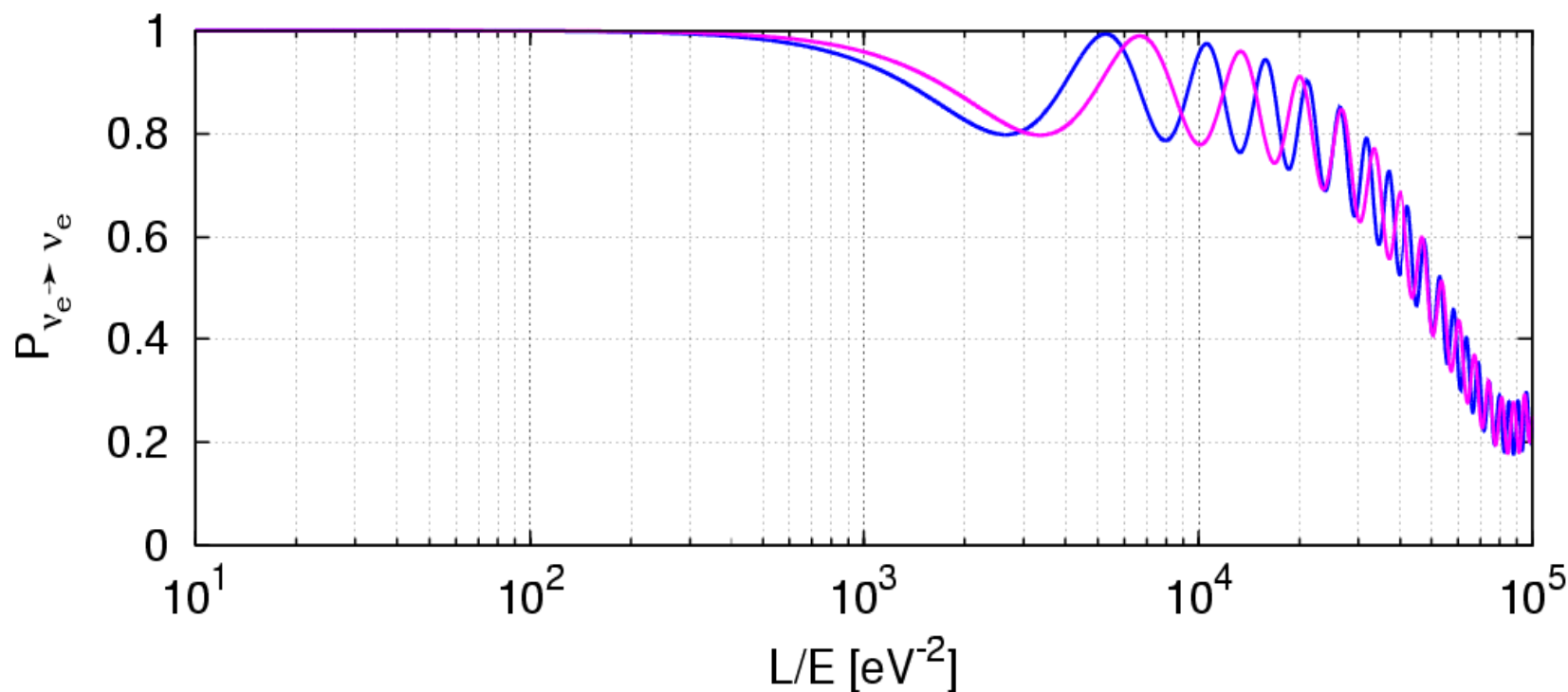
$\bar{\nu}_e$

- Reactor $\bar{\nu}$: pure $\bar{\nu}_e$ from β decay, $E_{\nu} < 10$ MeV
- neutrino oscillation – flavor **detection probability $P(t)$**

$$P(\nu_{\alpha} \rightarrow \nu_{\beta}) = |\langle \nu_{\beta} | \nu_{\alpha} \rangle|^2 = \sum_{k,j} U_{\alpha k}^* U_{\beta k} U_{\alpha j} U_{\beta j}^* \exp\left(-\frac{i}{\hbar}(E_k - E_j)t\right)$$

- With ultrarelativistic approximation and $c=\hbar=1$

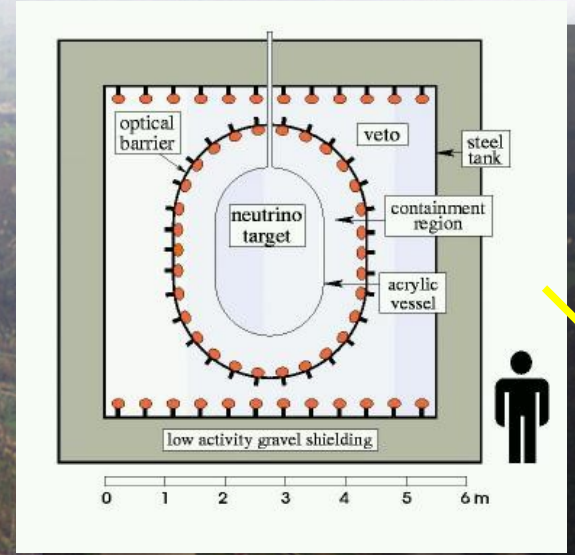
$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \sin^2 2\Theta_{13} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E}\right) - \cos^4 \Theta_{13} \sin^2 2\Theta_{12} \sin^2\left(\frac{\Delta m_{21}^2 L}{4E}\right)$$



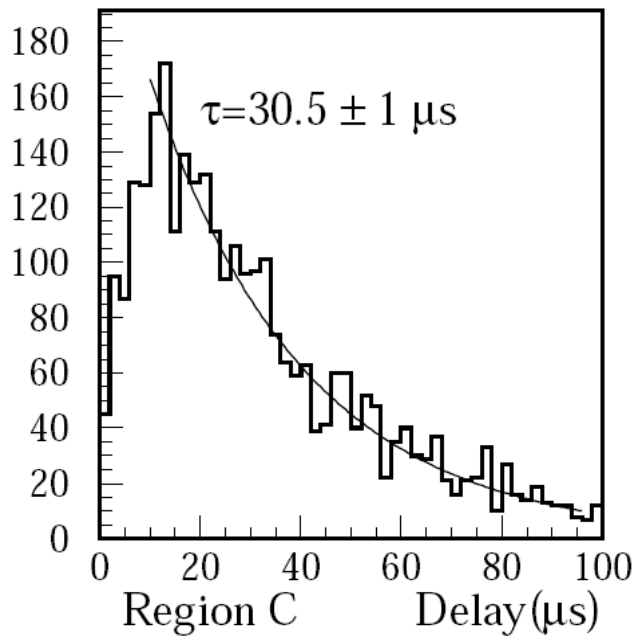
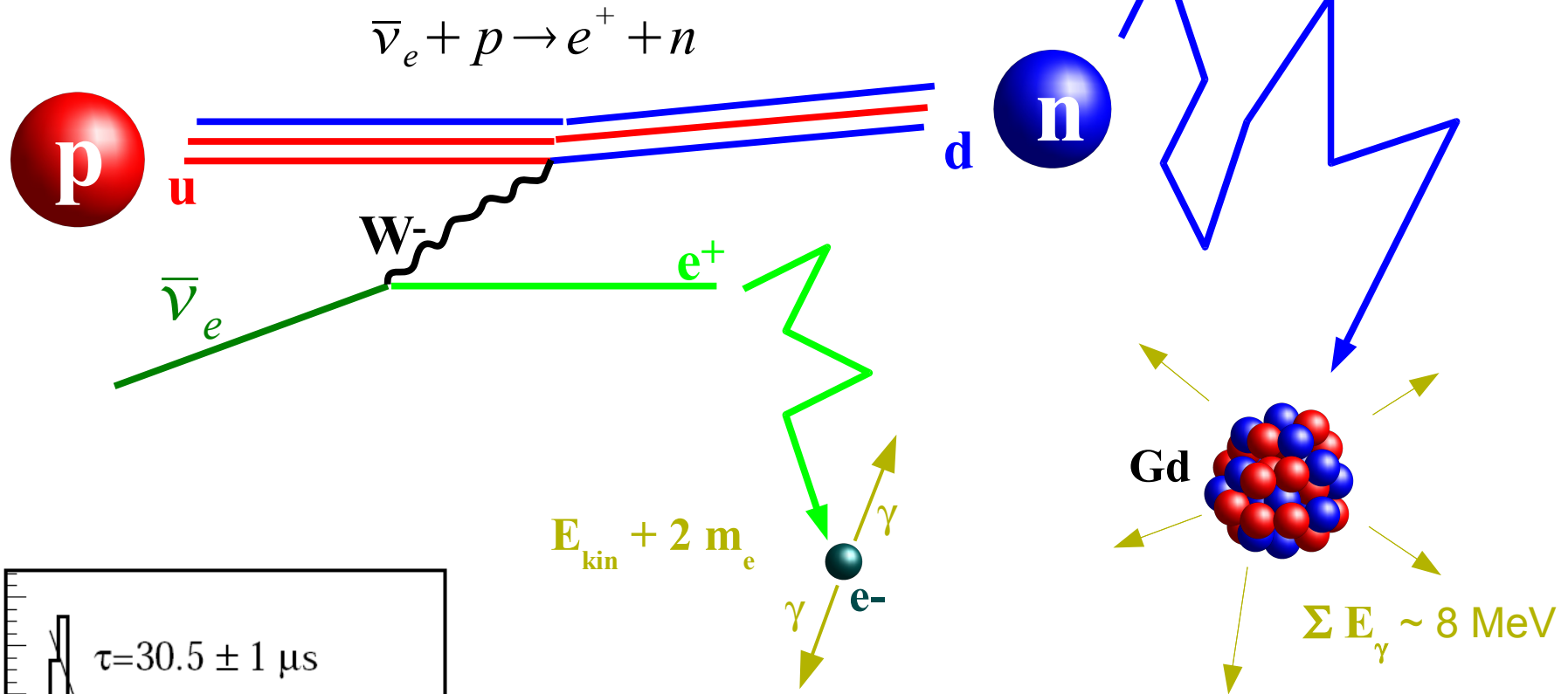
Double Chooz collaboration,
[arXiv:0606025 \[hep_ex\]](https://arxiv.org/abs/0606025)

The CHOOZ experiment

- **Liquid scintillator detector**
Target mass 5t, **Gd-doped**
- ~ 1000 m from Chooz B1 and B2 reactor cores (4 GW_{th} each)
- Looking for $\bar{\nu}_e \rightarrow \bar{\nu}_x$
disappearance
- 300 mwe overburden
- Total run time ~ 15 months
half of that background only
- Detection by neutrino cc interaction (**inverse beta decay**, delayed coincidence)



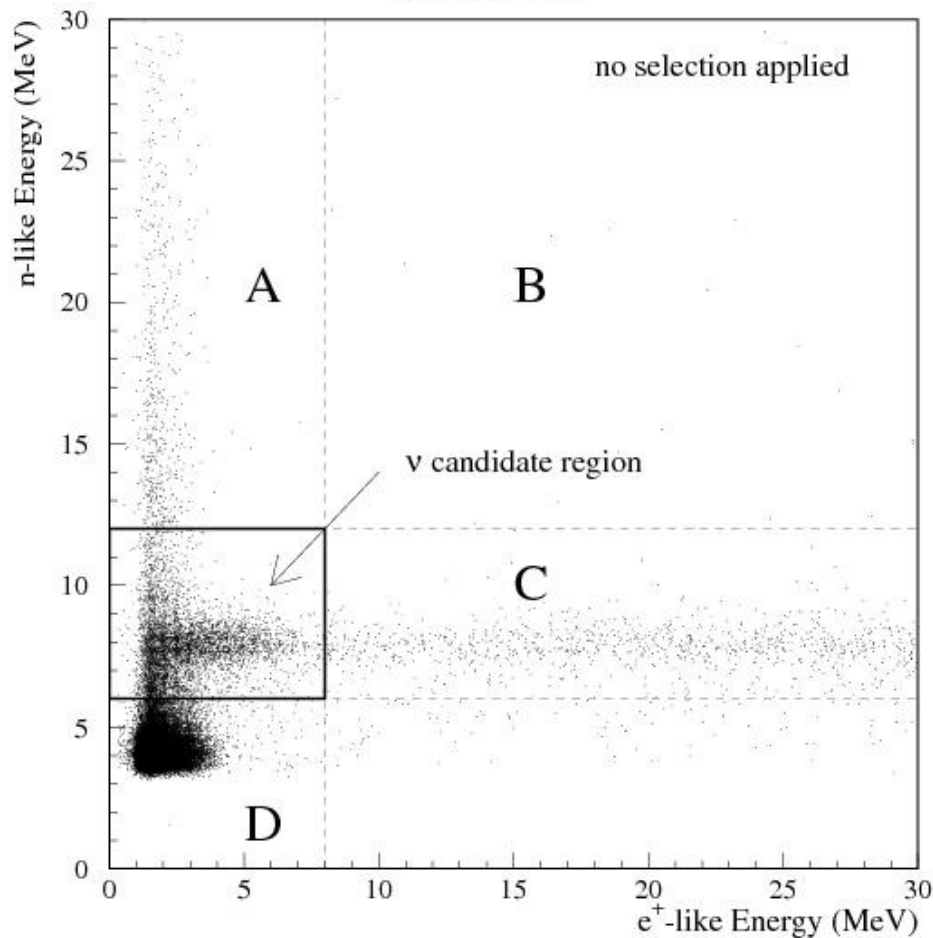
- Neutrino detection by inverse β decay



- **Delayed coincidence:** radiative capture on Gd with $\tau \sim 30 \mu\text{s}$

[Apollonio et al, arXiv:0301017 \[hep-ex\]](https://arxiv.org/abs/0301017)

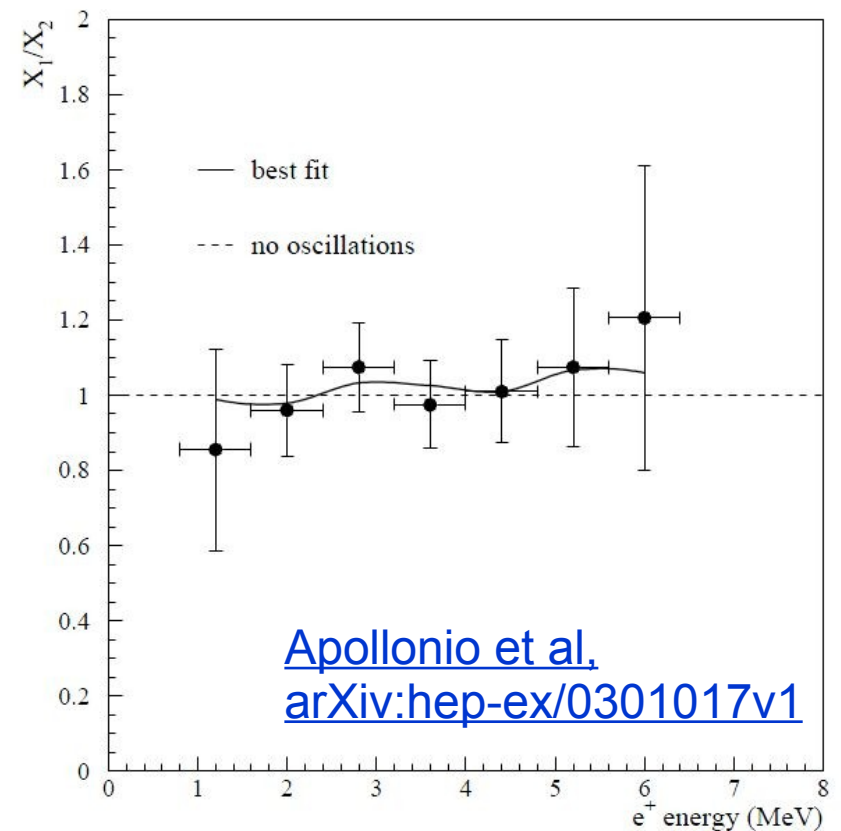
Reactor ON



- ~ 3000 ν candidates after cuts
- Result compatible with **no oscillation** hypothesis @ 90 % CL

| parameter | relative error (%) |
|-----------------------------|--------------------|
| reaction cross section | 1.9% |
| number of protons | 0.8% |
| detection efficiency | 1.5% |
| reactor power | 0.7% |
| energy absorbed per fission | 0.6% |
| combined | 2.7% |

[Apollonio et al.,
PLB 466 \(1999\) 415](#)



[Apollonio et al.,
arXiv:hep-ex/0301017v1](#)

Improving CHOOZ

- **Statistics limitations**

- larger target mass
- longer run time (CHOOZ limited by scintillator degradation)

- **Systematics limitations**

- Dominating uncertainty from neutrino flux
 → Do **relative measurement** with (at least) two detectors at different distances from the source(s)



Eliminates ν production cross section, reactor power and energy per fission errors (if detectors are identical)

- **Reduce number of cuts** (CHOOZ used 7, resulting in 1.5 % relative error)
- **Background reduction** – strict material selection, remove PMTs from scintillating fluid, better shielding and deeper site

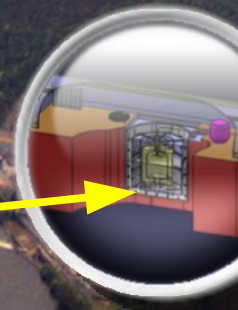
[Anderson et al.](#)
[arXiv:0402041v1\[hep-ex\]](#)

Double Chooz: Design

The EDF Chooz power plant is located in the French département des Ardennes, near Givet, close to the Belgian border.



400 m



1000 m

Near site:

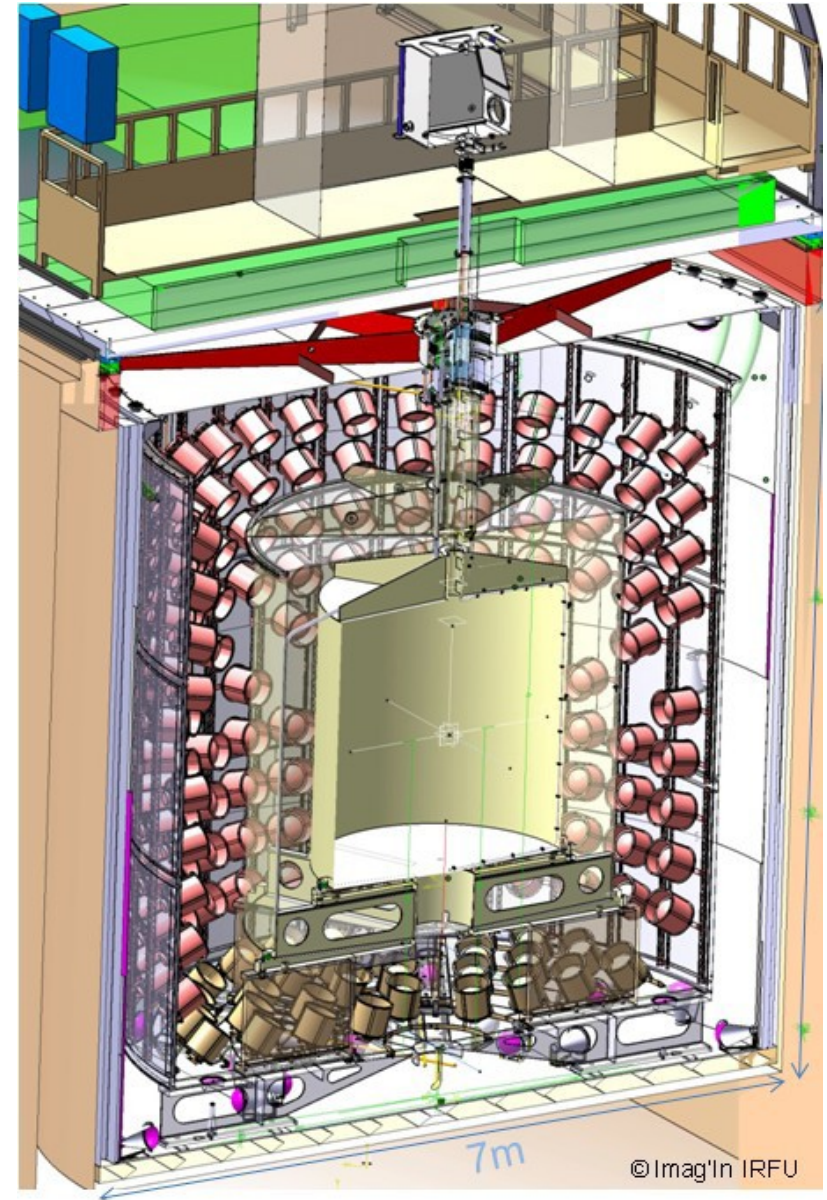
- 115 mwe overburden
- Neutrino rate $\sim 500/\text{day}$
- Muon rate $\sim 250 \text{ Hz (IV)}$

Far site:

- 300 mwe (old CHOOZ lab)
- Neutrino rate $\sim 50/\text{day}$
- Muon rate $\sim 20 \text{ Hz (IV)}$

Detector volumes and instrumentation:

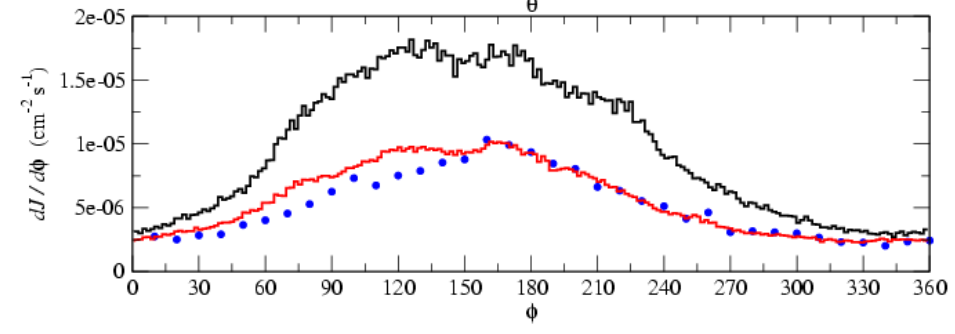
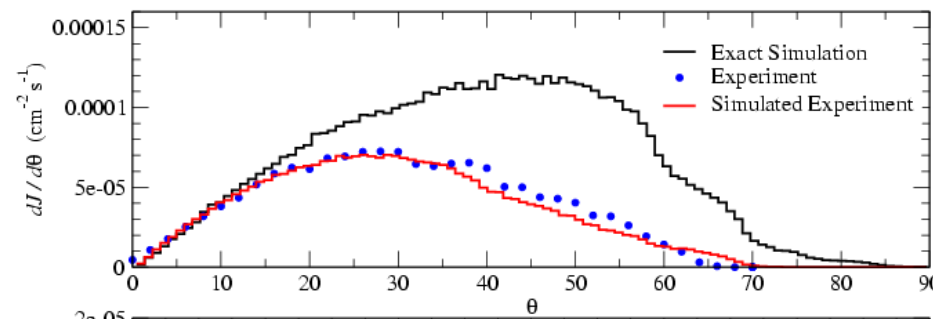
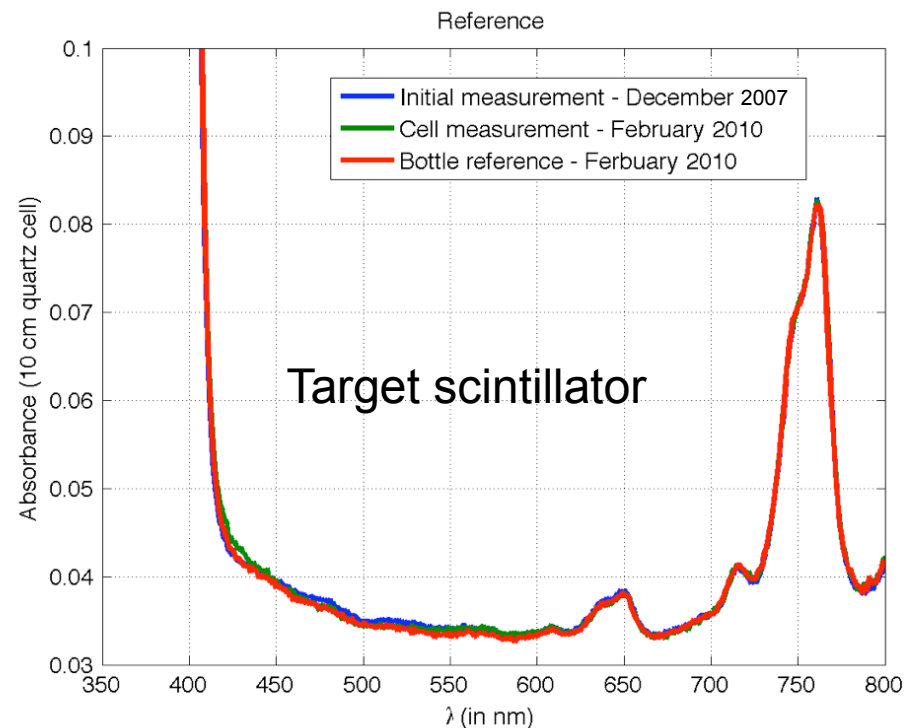
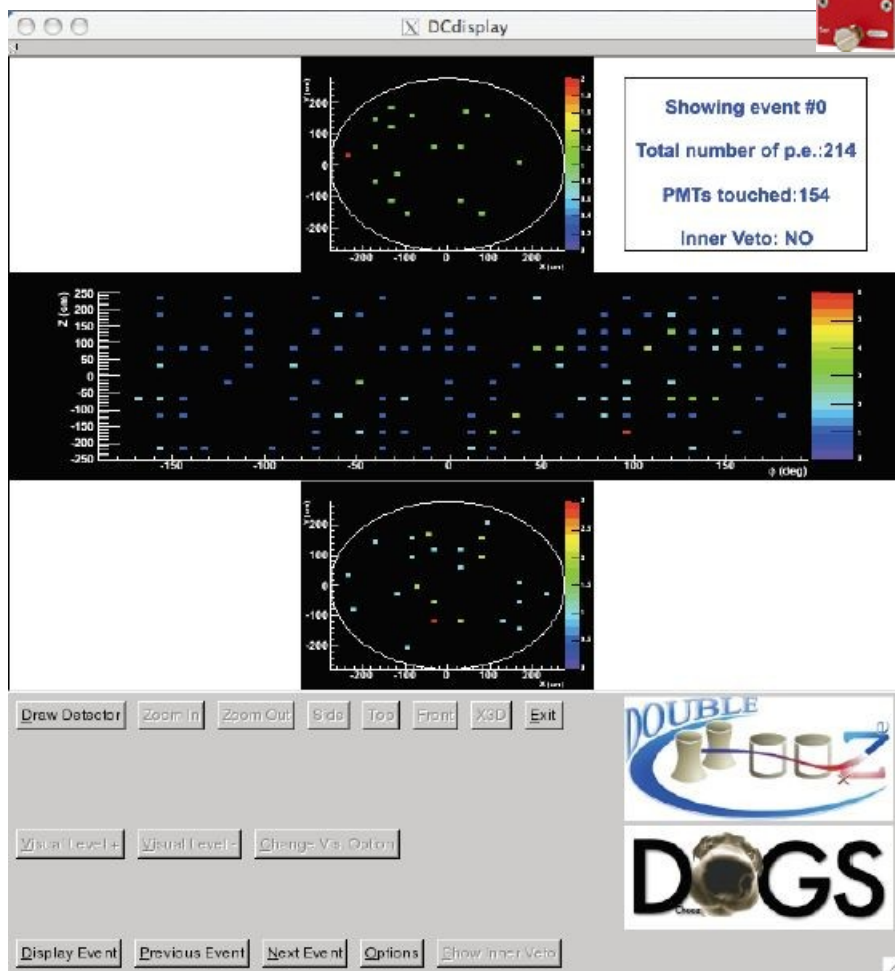
- **Outer veto:** four layers of plastic scintillator panels covering the top of the detector and the glove box
- **Shield:** 17 cm steel
- **Inner veto:** 50 cm of LAB based scintillator, 78 8" PMTs
LED calibration system
- **Buffer:** 105 cm of mineral oil, 390 10" PMTs, calibration source guide tube, LED system
- **γ -catcher:** acrylics vessel, 55 cm of PXE based scintillator
calibration source guide tube
- **Target:** acrylics vessel, ~8t of Gd-doped PXE based scintillator, fish line and articulated arm calibration



© Imag'In IRFU

Substantial R&D effort

- Scintillator development
- Detector Monte Carlo
- Reactor spectrum



Double Chooz: Backgrounds

• Accidental background

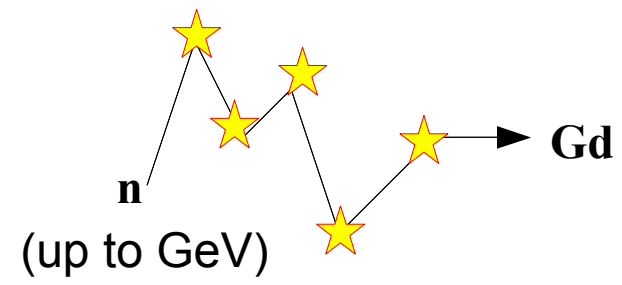
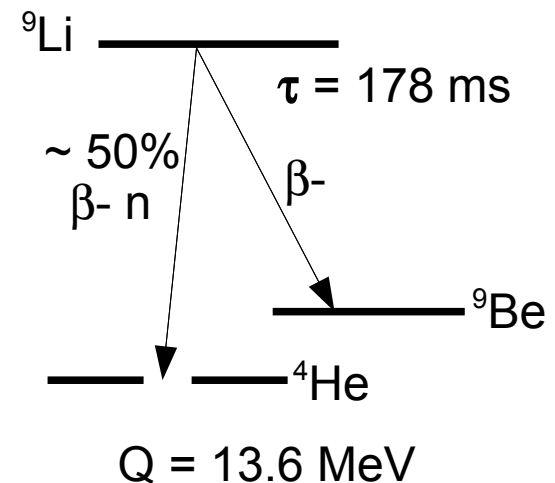
“2 independent processes happen to look like ν event”

- Natural radioactivity of components or rock
- Secondaries of μ interaction
- High purity materials, shielding & μ veto
- Expected rates: ~ 12 / day (near detector)
 ~ 2 / day (far detector)

• Correlated background

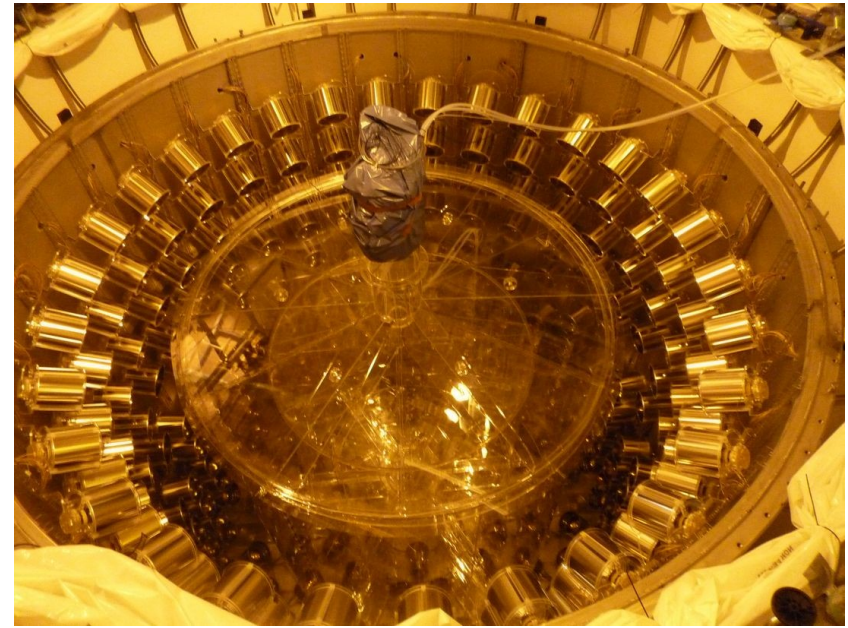
“single process mimicking ν signature”

- Spallation product β -n cascades
- Fast n created by μ outside of detector
- Deeper site, μ monitoring, modelling
- Expected rates: ~ 8 / day (near detector)
 ~ 2 / day (far detector)

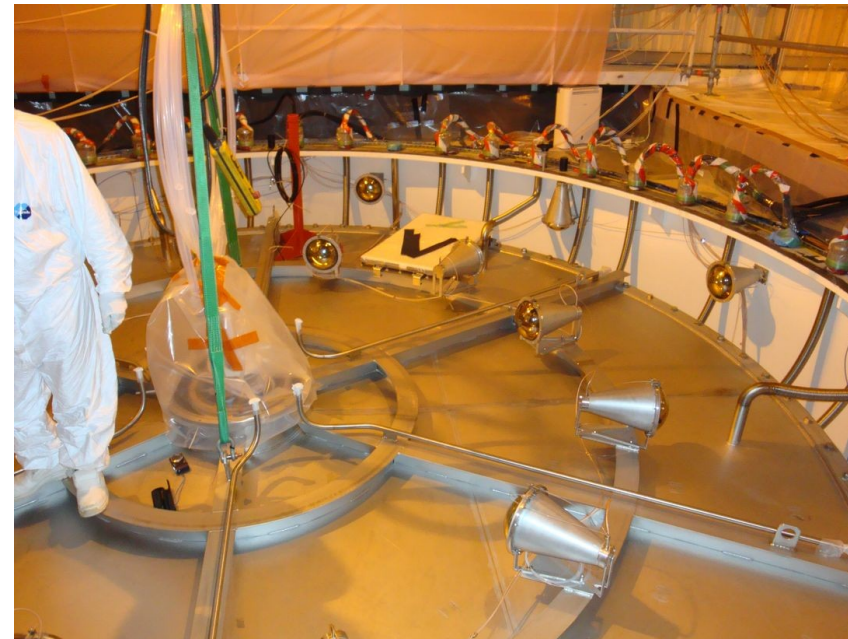


Double Chooz: Status & Time schedule

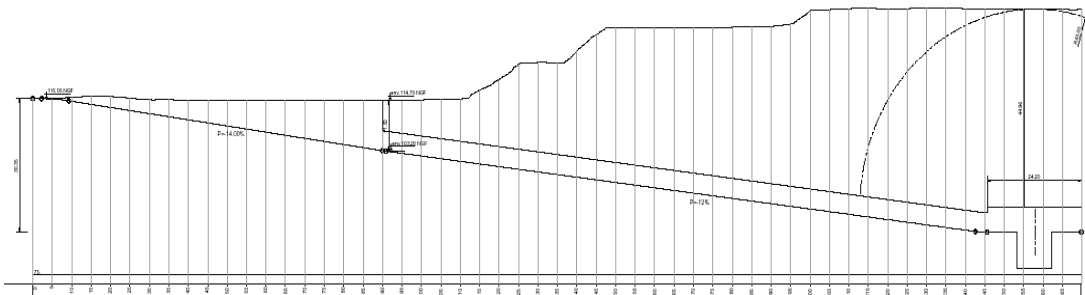
- Far detector closing imminent
 - Electronics installation ongoing
 - Awaiting fluid delivery
 - Commissioning run ~ May
 - In parallel: closing of steel shield and outer veto installation
- Near laboratory construction starting
 - Fully funded
 - Design approved by EDF
 - To be finished till end 2010



Acrylics integration, 10/2009



Top veto PMT installation, 1/2010

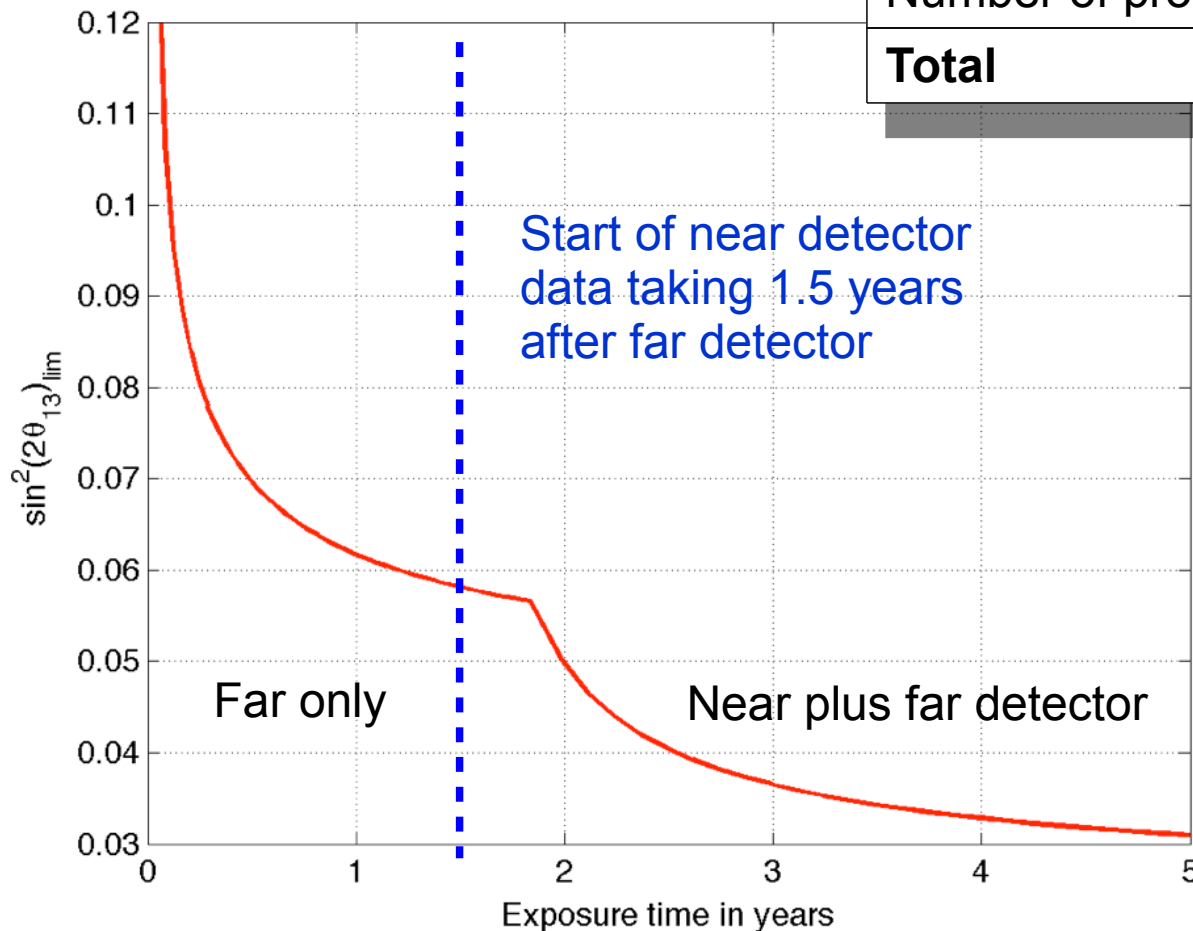


Side view of near lab planning sketch

Double Chooz: Sensitivity

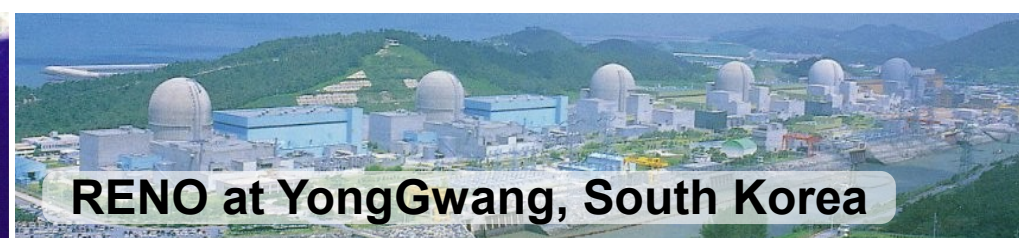
- Two distinct phases:
 - Far detector only
 - Near plus far detector

| Systematic errors | Absolute | Relative |
|---------------------|--------------|--------------|
| Production σ | 1.9 % | - |
| Reactor power | 0.7 % | - |
| Energy per fission | 0.6 % | - |
| Detector efficiency | 1.5 % | 0.5 % |
| Number of protons | 0.8 % | 0.2 % |
| Total | 2.7 % | 0.6 % |



- 1.5 years far only data:
 $\sin^2 2 \theta_{13} < 0.06$
- 4 years of near + far data
 $\sin^2 2 \theta_{13} < 0.03$

Daya Bay & Reno


RENO at YongGwang, South Korea

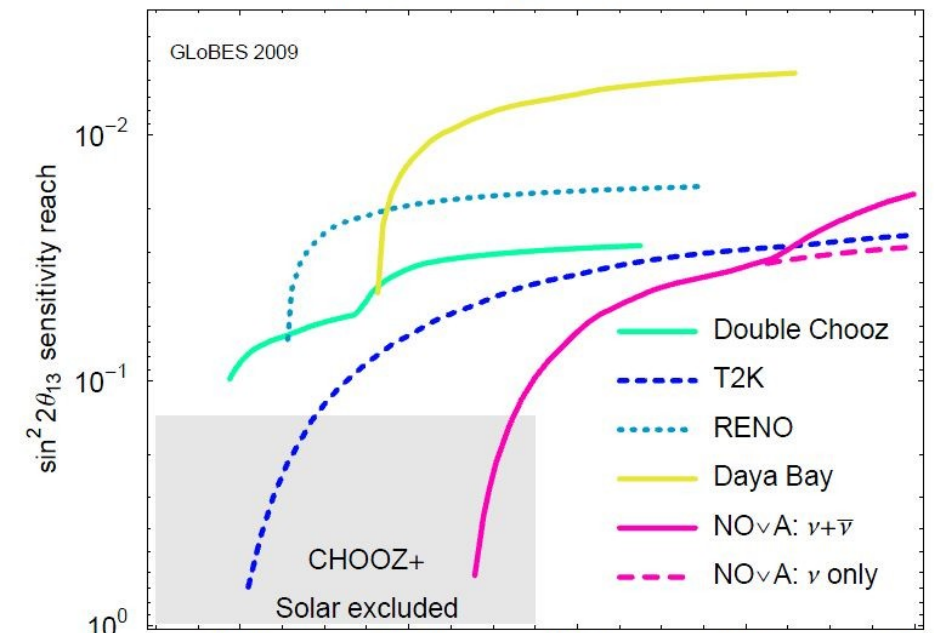
Reno expects to start mid-2010

Daya Bay power plant complex, China

DB near detectors to start mid-2010, far mid-2011

| | Power [GW _{th}] | L _{near} [m] | L _{Far} [m] | M _{target} [t] | S _{stat} [%] | S _{sys} [%] | sin ² 2θ ₁₃ > (90 % CL) |
|----------|------------------------------|--------------------------|-------------------------|----------------------------|--------------------------|-------------------------|--|
| DC | 8.6 | 400 | 1050 | 8.3 | 0.5 | 0.6 | 0.03 |
| RENO | 17.3 | 290 | 1380 | 16 | 0.3 | 0.5 | 0.02 |
| Daya Bay | 17.4 | 360 (500) | 1990 (1620) | 80 | 0.2 | 0.4 | 0.01 |

- **Daya Bay:** 2 near sites, two 20t detectors each and one far site with four. Exchange detectors between sites for cross calibration.
- **RENO** consists of one near and one far detector, similar to DC
- Both have comparable detector design, but use water Cerenkov veto instead of LS
- Both sites need to correct for differing ν flux contributions @ near and far site(s). No reactor-off data for both sites available – and total shutdown of complex unlikely

 sin²2θ₁₃ sensitivity limit (NH, 90% CL)


Summary

- Double Chooz aims at **measuring θ_{13}** , the last undetermined neutrino mixing angle.
- Fixing θ_{13} allows to solve degeneracies of beam experiments \rightarrow **access to δ_{CP} , mass hierarchy.**
- DC far detector will **start data taking ~ June** and reach sensitivity of **$\sin^2 2\theta_{13} < 0.06$ (90 % CL)** after 1.5 years, improving the current limit by a factor of two.
- With four years of both near and far detector data the limit will be **$\sin^2 2\theta_{13} < 0.03$ (90 % CL)** in 2015.
- Very active field – several experiments start data taking in the near future.
- Expect to learn more about neutrino nature soon!



CBPF, UNICAMP



CIEMAT Madrid



**Hiroshima Inst. Tech.,
Kobe Univ.,
Miyagi Univ.,
Niigata Univ.,
Tohoku Univ.,
Tohoku Gakuin Univ.,
Tokyo Metro. Univ.,
Tokyo Inst. Tech.**



**INR-RAS, IPC-RAS,
RRC Kurchatov**



**RWTH Aachen,
Univ. Hamburg,
MPIK Heidelberg,
TU München,
Univ. Tübingen**

Fin

Univ of Sussex



**APC Univ. of Paris,
SUBATECH (Nantes),
DAPNIA CEA/Saclay,
IPHC Strasbourg**



**Univ. of Alabama, ANL,
Univ. of Chicago, Columbia,
U.C. Davis, Drexel Univ.,
Kansas State, Illinois Inst. Tech.,
LLNL, Notre Dame,
SNL, Univ. of Tennessee**



Oscillation probabilities

$$P_{reactor} \simeq \sin^2 2\theta_{13} \sin^2 \Delta + \alpha^2 \Delta^2 \cos^4 \theta_{13} \sin^2 2\theta_{12},$$

$$\begin{aligned} P_{long-baseline} &\simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \Delta \\ &\mp \alpha \sin 2\theta_{13} \sin \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin^3 \Delta \\ &+ \alpha \sin 2\theta_{13} \cos \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \sin^2 \Delta \\ &+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \sin^2 \Delta \end{aligned}$$

with $\alpha \equiv \Delta m_{21}^2 / \Delta m_{23}^2$ and $\Delta \equiv \Delta m_{31}^2 L / (4E_\nu)$.

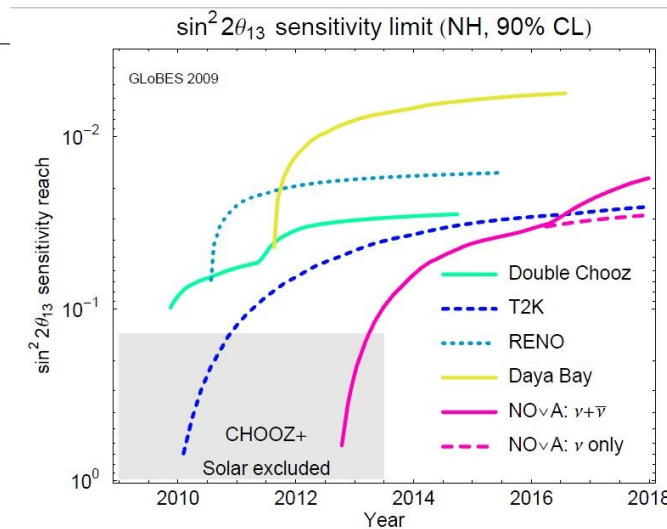
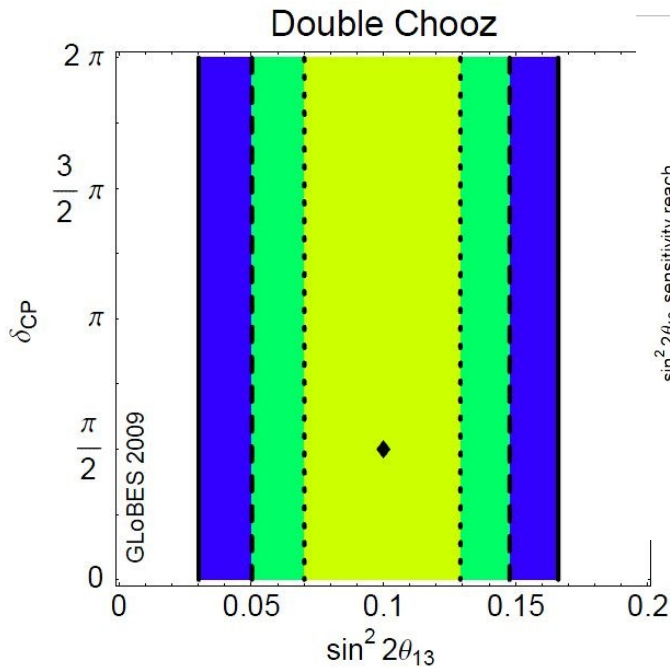
Mahn et al
[hep-ex/0409028v4](http://arxiv.org/abs/hep-ex/0409028v4)

Reactor experiments

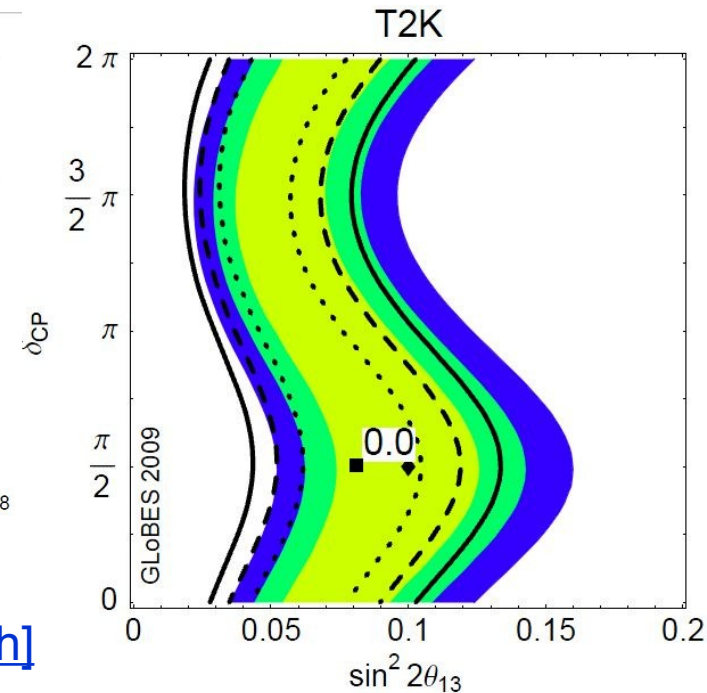
- Disappearance: $\bar{\nu}_e \rightarrow \bar{\nu}_x$
- Clean Θ_{13} measurement
- No information on δ_{CP}

Beam experiments

- Appearance $\nu_\mu \rightarrow \nu_e$
- Θ_{13} coupled to δ_{CP} and Θ_{23}
- δ_{CP} if Θ_{13} not too small



[Huber et al., arXiv:0907.1896v1 \[hep-ph\]](#)



→ Need information from both to determine ν parameters



- Shape vs rate information

