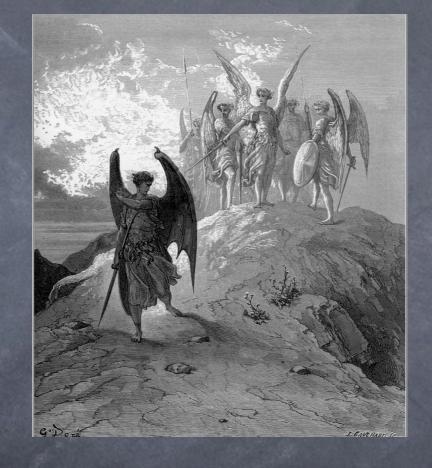
LUCIFER a new Technique for Double Beta Decay

(and an unusual funding source)







Fernando Ferroni Sapienza Universita` di Roma INFN Sezione di Roma European Research Council





XXIV Rencontre de Physique de la Vallee d'Aoste March, 1 - 2010

Outline

Meutrino-less Double Beta Decay

The experimental (standing) problem

Toward O background : LUCIFER

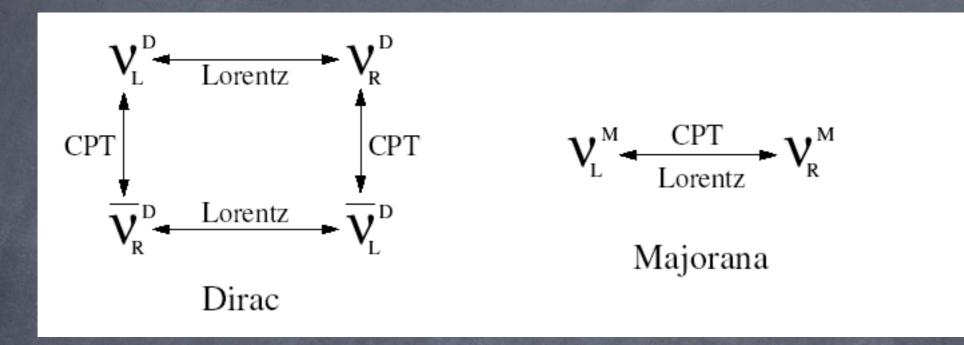


Majorana conjecture

V = V

Practical consequence : Lepton Number Violation Caveat: massless neutrinos do not allow testing of the Majorana nature

Indeed nobody payed much attention to the Furry hypothesis (1939) that a Majorana neutrino could induce Neutrino-less DBD via helicity flip Massive neutrinos makes the story much more attractive



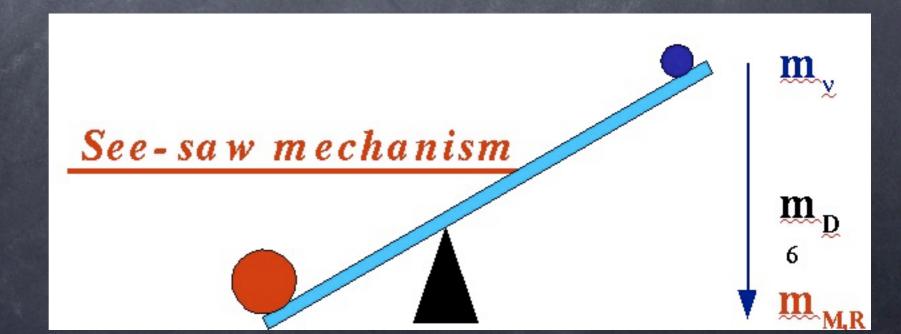
Now helicity flip can happen in both Dirac and Majorana cases. However Dirac forbids the absorption of an anti-neutrino right that was emitted as a neutrino left because the Lepton Number Conservation one elegant explanation (beyond the SM)

Mass Term $\frac{1}{2} \begin{bmatrix} v_L & (v_R)^c \end{bmatrix} C \begin{pmatrix} M_{M,L} & m_D \\ m_D & M_{M,R} \end{pmatrix} \begin{bmatrix} v_L \\ (v_R)^c \end{bmatrix} + h.c.$

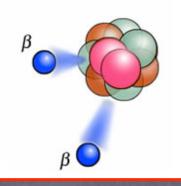
where $M_{M,L} \sim 0$ $M_D ~ M_{EW} \sim 100 \mbox{ GeV} \label{eq:mbox} M_{M,R} \sim Gauge singlet \mbox{ unprotected} \sim M_{GUT}$

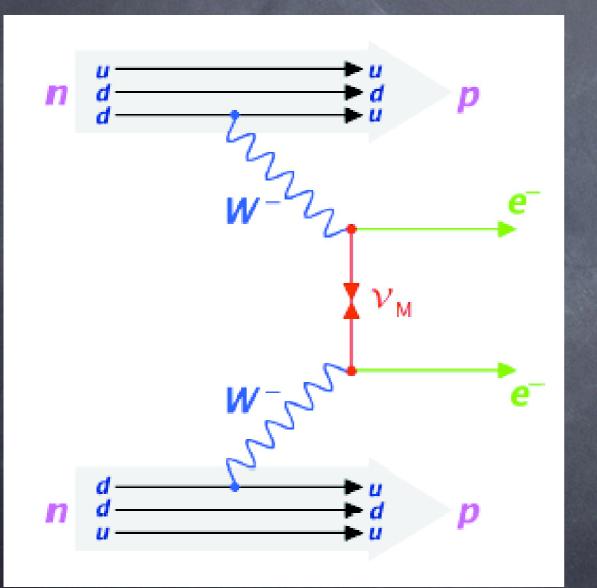
$$m_N \simeq M_{M,R}$$

 $m_\nu \simeq \frac{m_D^2}{M_{M,R}}$



Neutrino-less DBD ($0\nu\beta\beta$)





Only if:

Majorana Neutrinos

Massive Neutrinos

If observed:

Proof of the Majorana nature of Netrino

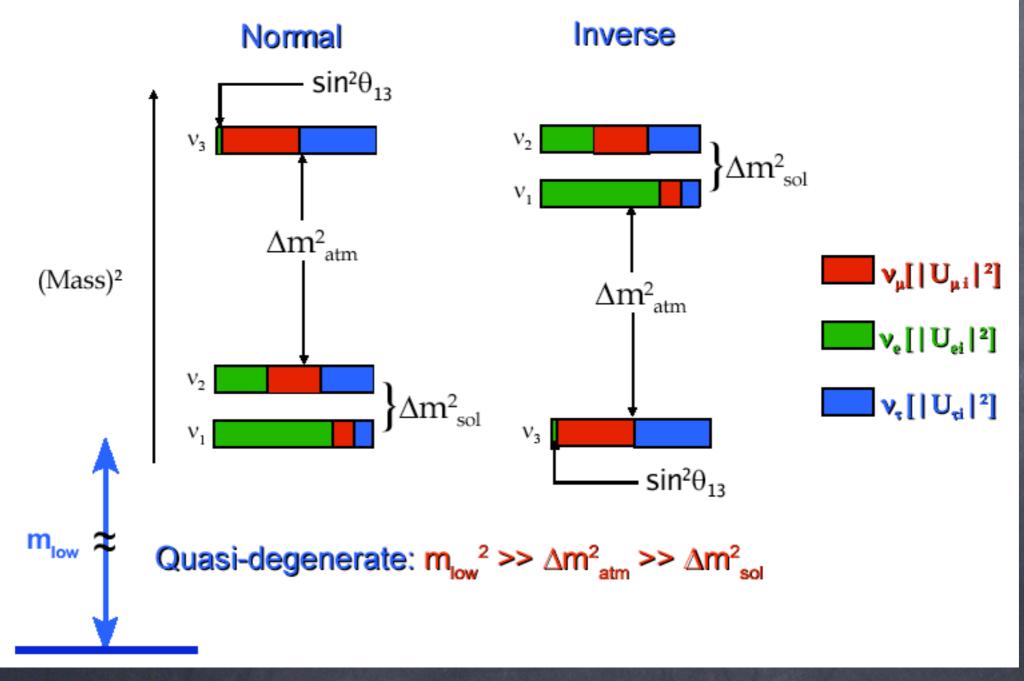
Does it also measure the mass?

$$m_{\beta\beta} = \sum m_{\nu_{k}} U_{ek}^{2} = \cos^{2} \theta_{13} (m_{1} \cos^{2} \theta_{12} + m_{2} e^{2i\alpha} \sin^{2} \theta_{12}) + m_{3} e^{2i\beta} \sin^{2} \theta_{13}$$

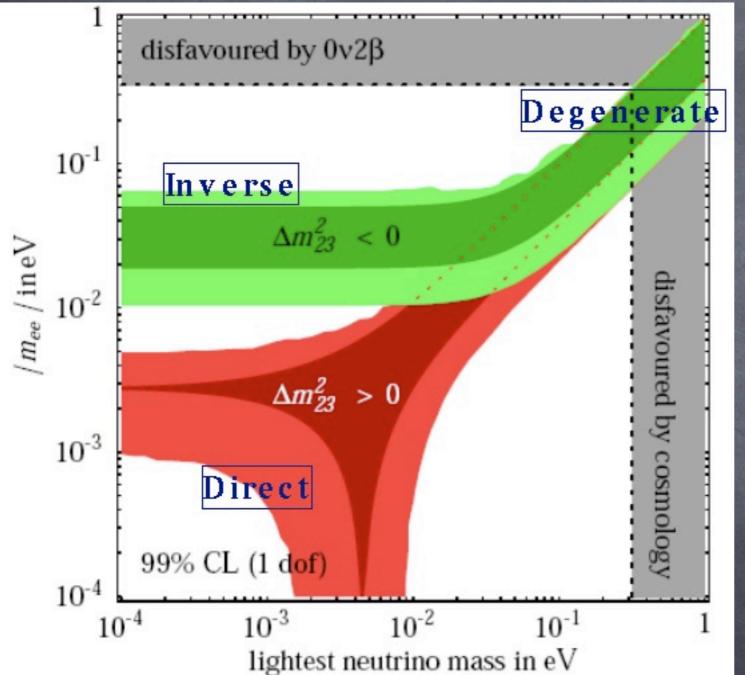
well...not so straight. It comes as a combination of the three neutrino masses, the mixing angles and the Majorana phases.

Exercise: parameterize as a function of the known parameters: $m_{\beta\beta} = f(U_{ek}, m_{lightest}, \delta m_{sol}, \Delta m_{atm})$

Three possibilities:

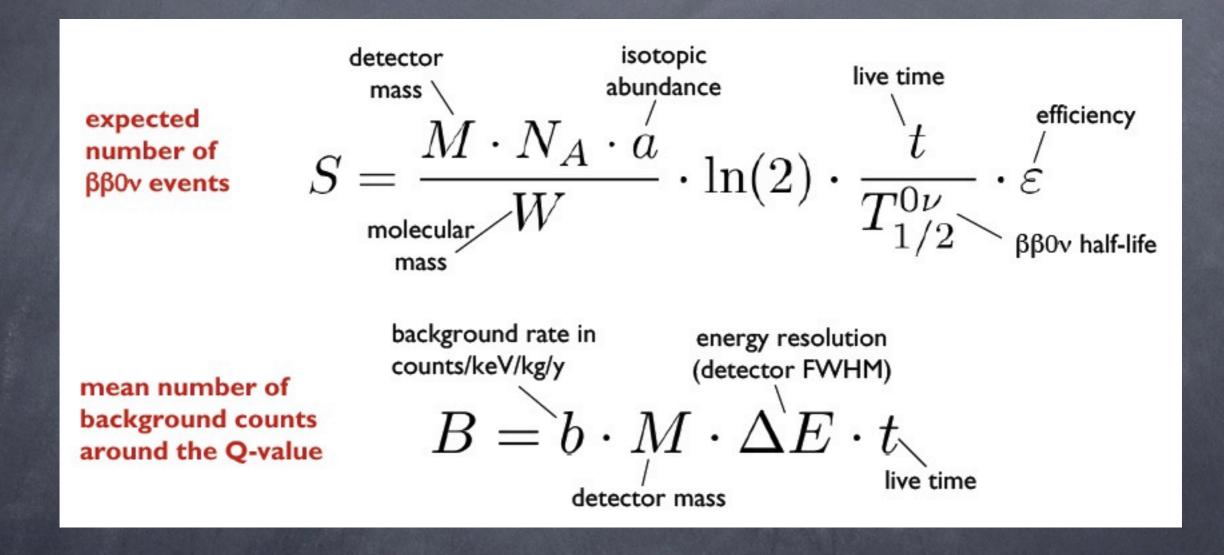


that translates into a nice plot



The question is which, if any, part of this phase space can be attained by a realistic experiment.

The name of the game



Sensitivity and background

Sensitivity $\propto K_{1} \frac{M \cdot t}{B \cdot \Delta E}$ (i.a. • ϵ)

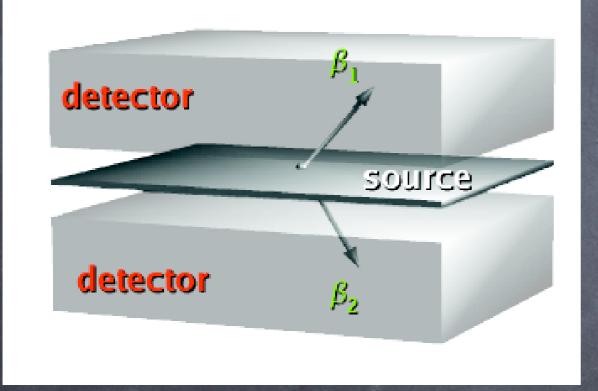
The (limited) isotope choice

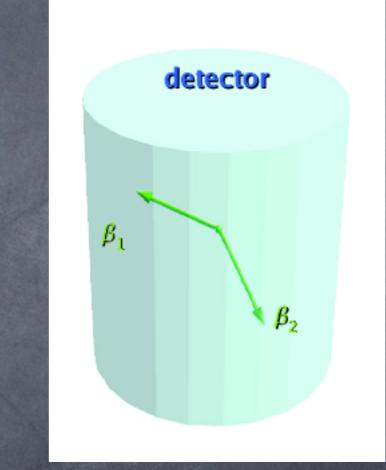
Isotope	$Q_{\beta\beta}$ (MeV)	Isotopic abundance (%)
⁴⁸ Ca	4.271	0.0035
⁷⁶ Ge	2.039	7.8
⁸² Se	2.995	9.2
⁹⁶ Zr	3.350	2.8
100 Mo	3.034	9.6
^{116}Cd	2.802	7.5
¹²⁸ Te	0.868	31.7
¹³⁰ Te	2.533	34.5
¹³⁶ Xe	2.479	8.9
150 Nd	3.367	5.6

Two techniques (and a few variations)

Source *≠* Detector

Source \subseteq Detector





+++ Topology, Background +++ M, ΔE , ϵ --- Μ, ΔΕ, ε

--- Topology, Background

The Problem

TeO₂ case T^{0v} [y] 1σ Sensitivity 0 12222 (CUORE) ηύρηκα baseline x 2.94 740 Kg x 1.96 10²⁶ demonstrated of which 0.001 counts/keV/kg/y 0.01 counts/keV/kg/y 0.04 counts/keV/kg/y 200 Kg 3 10 2 8 9 6 4 Live time [y] of ¹³⁰Te ton•y

Where is the Physics ?

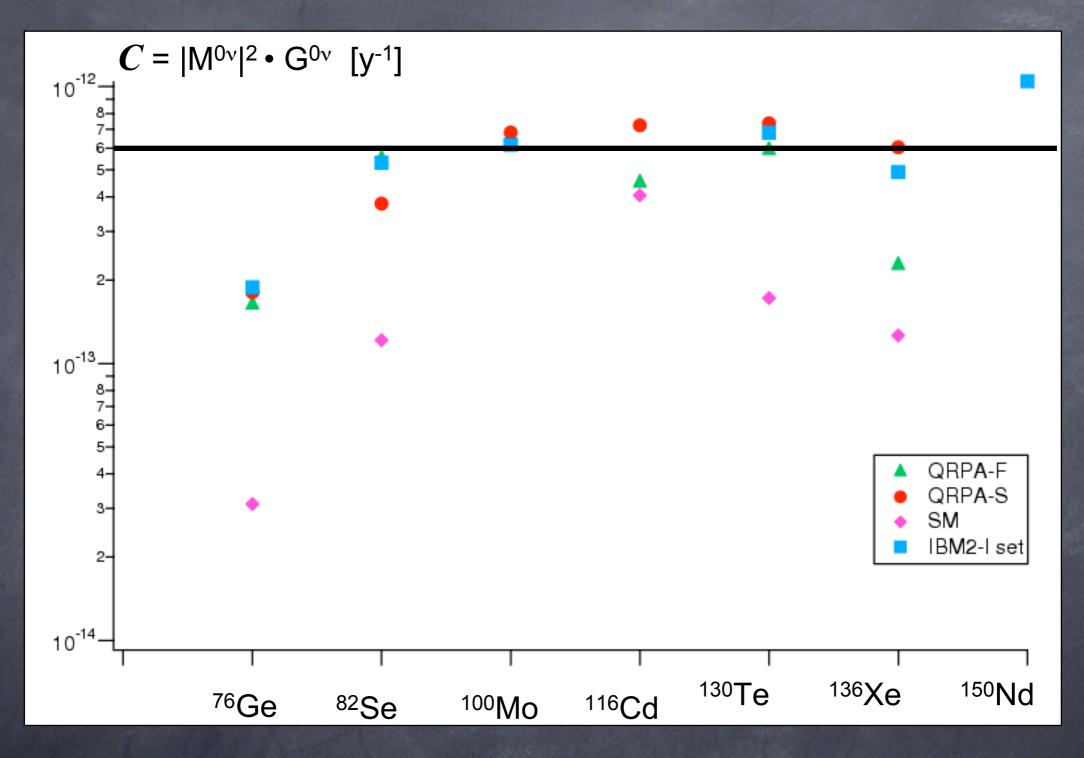
parameter containing the physics

$1/\tau = G(Q,Z) |M_{nucl}|^2 \langle M_{\beta\beta} \rangle^2$

what the **experimentalists** try to measure

what the **nuclear theorists** try to calculate

The role of NME



Sensitivity and background

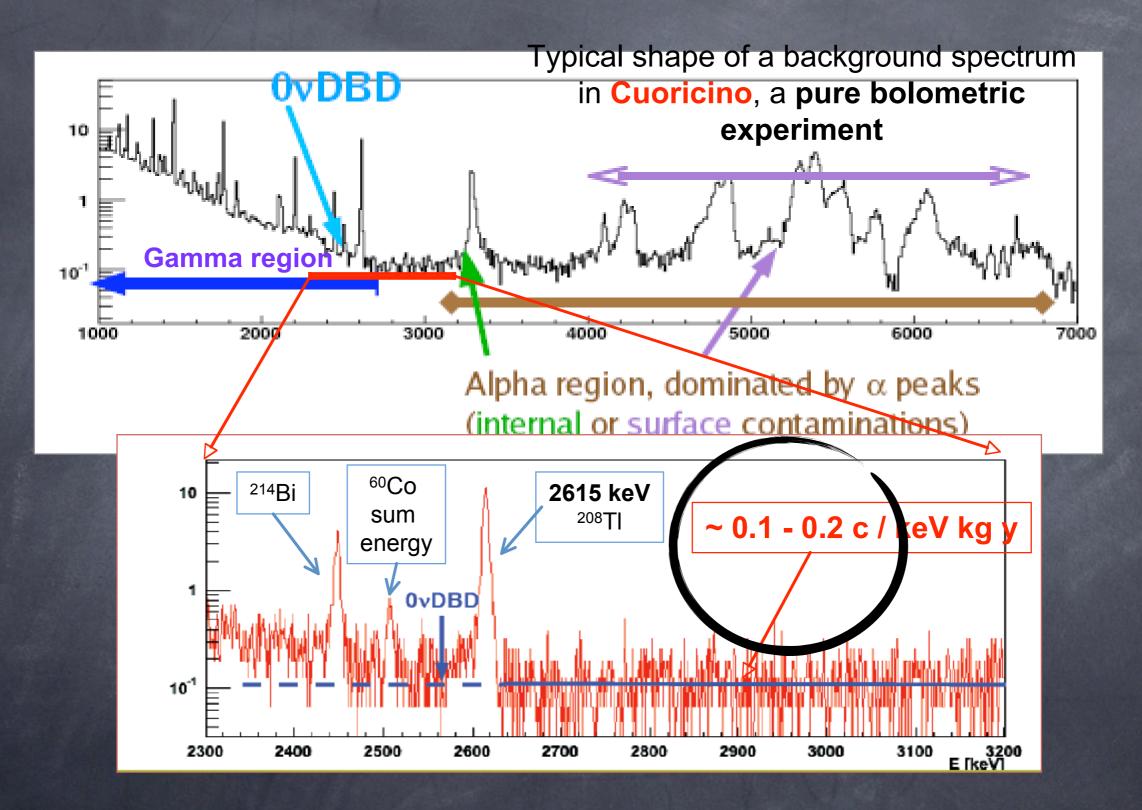
Sensitivity $\propto K_{1} \frac{M \cdot t}{B \cdot \Delta E}$ (i.a. • ϵ)

$m_{\beta\beta} \propto \sqrt{(1/\tau)}$ To get a factor 10 in $m_{\beta\beta}$ you have a choice :

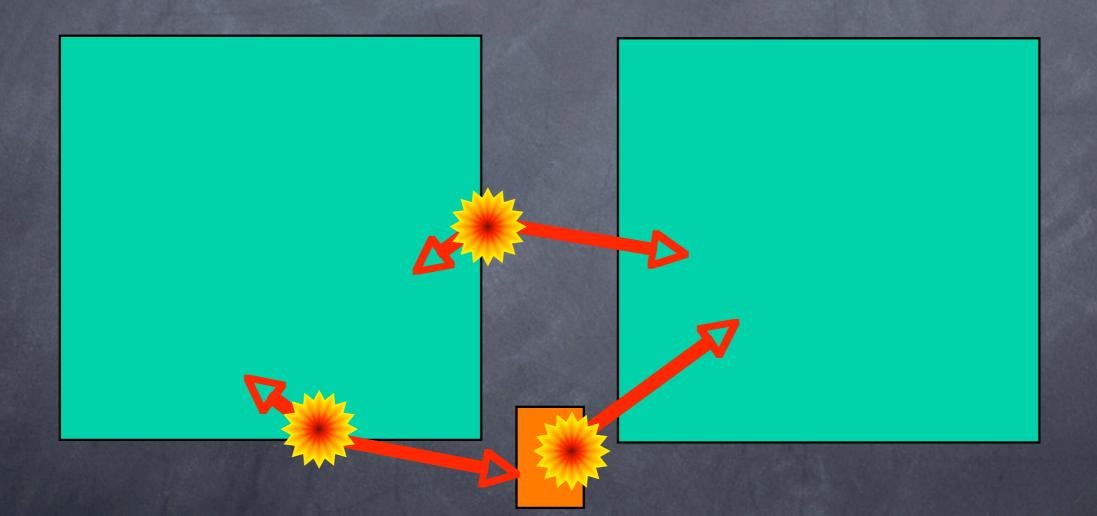
100 Ton instead of 1 Ton M + 500 y instead of 5 y 50 eV instead of 5 keV ΔE

0.001 instead of 0.1 B

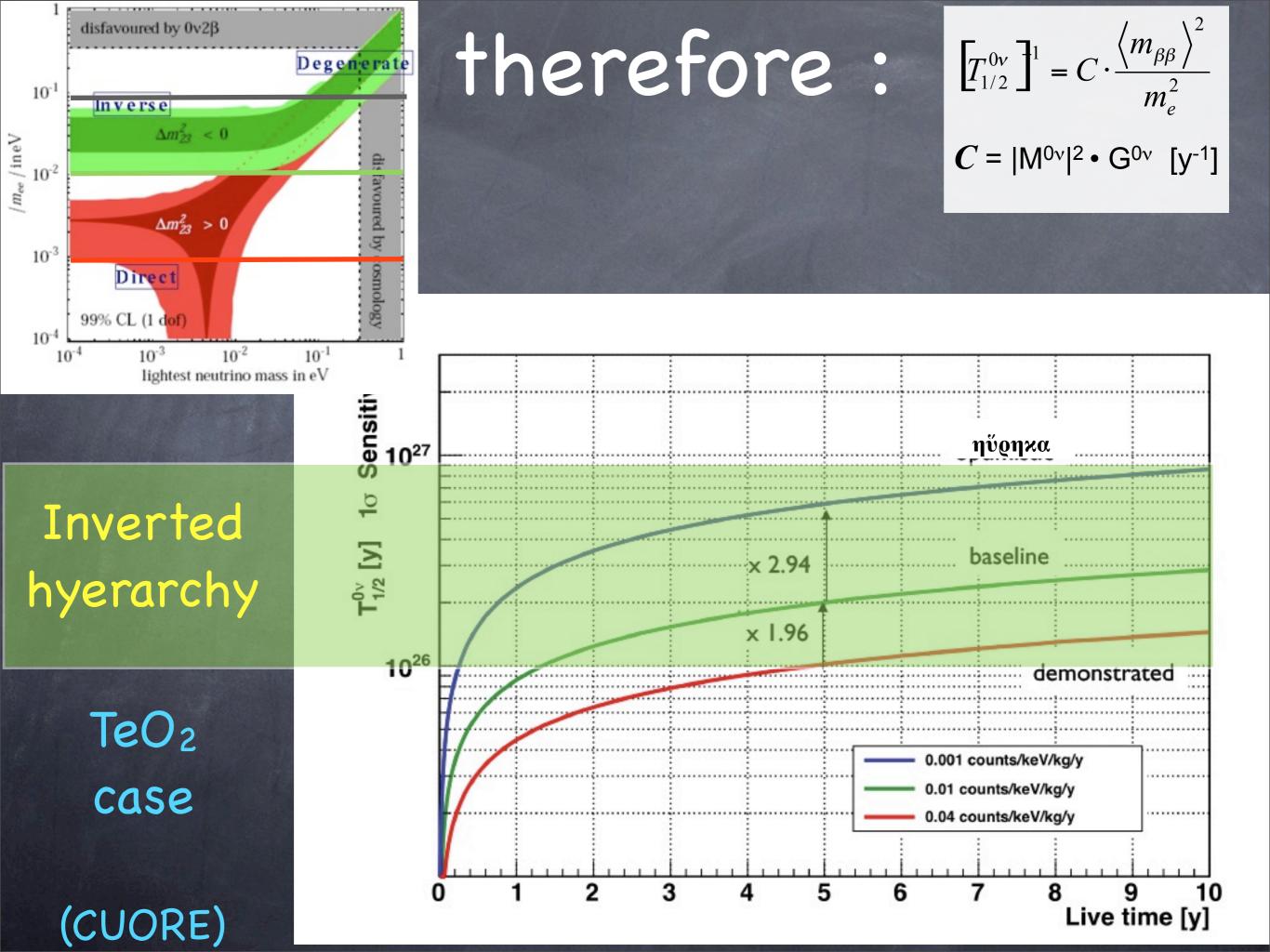
Where we start from



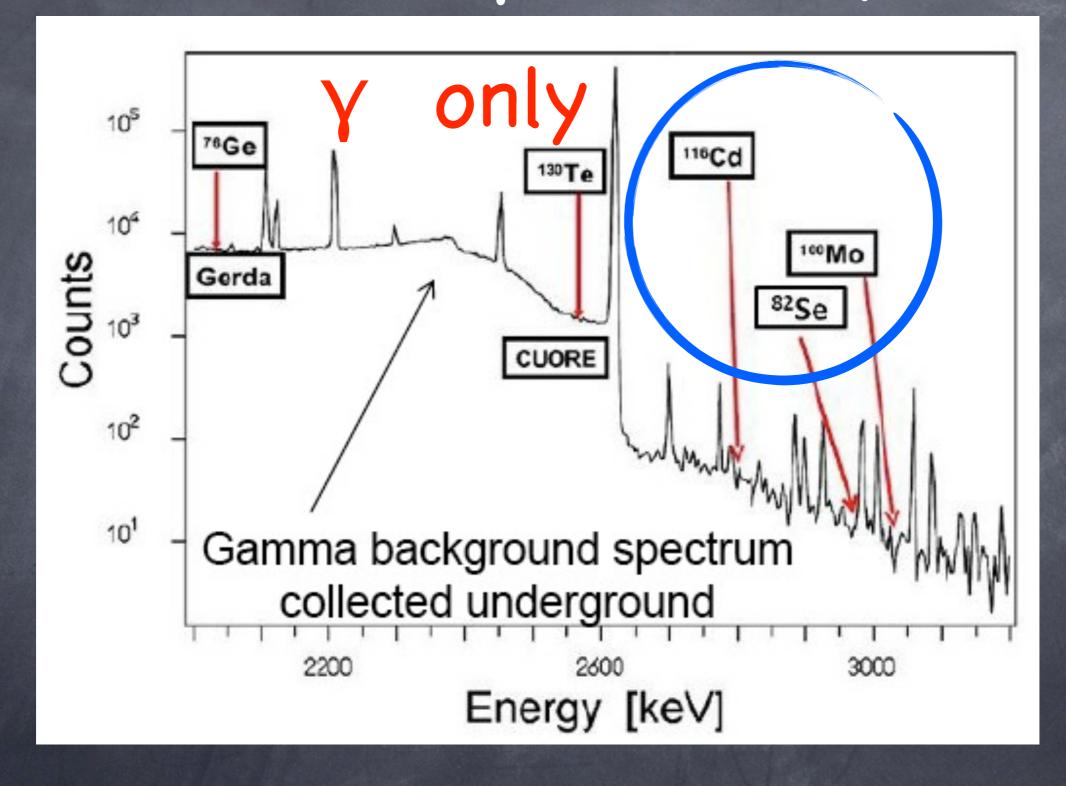
The standard interpretation of the background



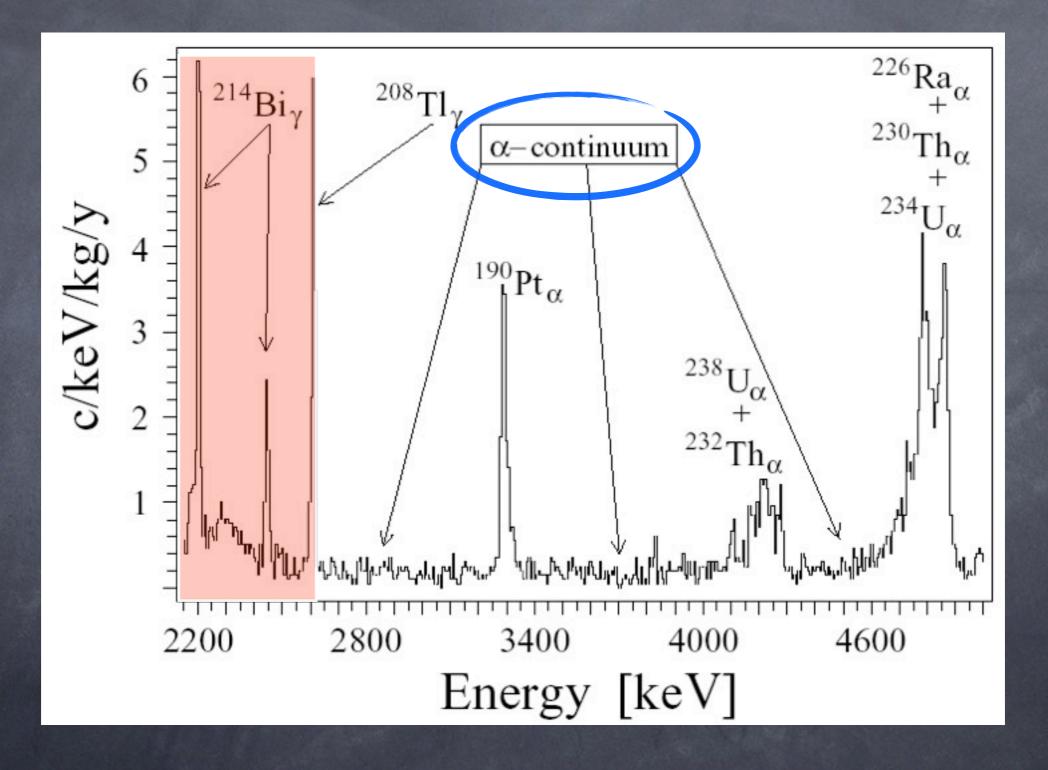
Degraded a's



The y step is 'easy'



The *c* step is 'difficult'



(very) Low Temperature Calorimeter

A True Calorimeter

heat sink (T₀) (thermal conductance G) (C) ββ atom x-tal

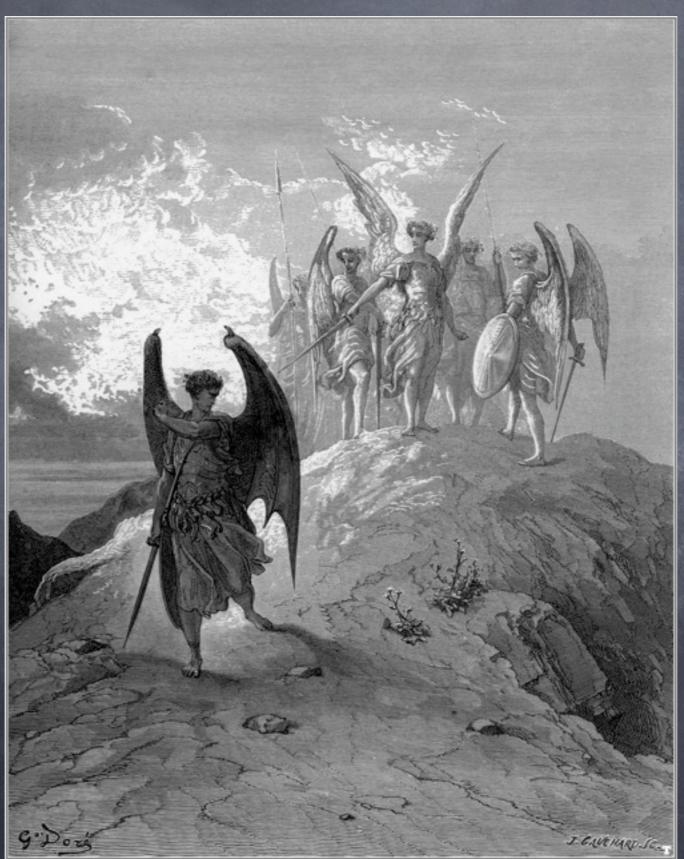
Basic Physics: $\Delta T = E/C$ (Energy release/ Thermal capacity) Implication: Low $C \Rightarrow$ Low T Bonus: (almost) No limit to ΔE (k_BT^2C) Not for all : $T = C/G \sim 1s$

$$C(T) = \beta \frac{m}{M} \left(\frac{T}{\Theta_D}\right)^3$$

$$\Delta T(t) = \frac{\Delta E}{C} \exp\left(-\frac{t}{\tau}\right)$$

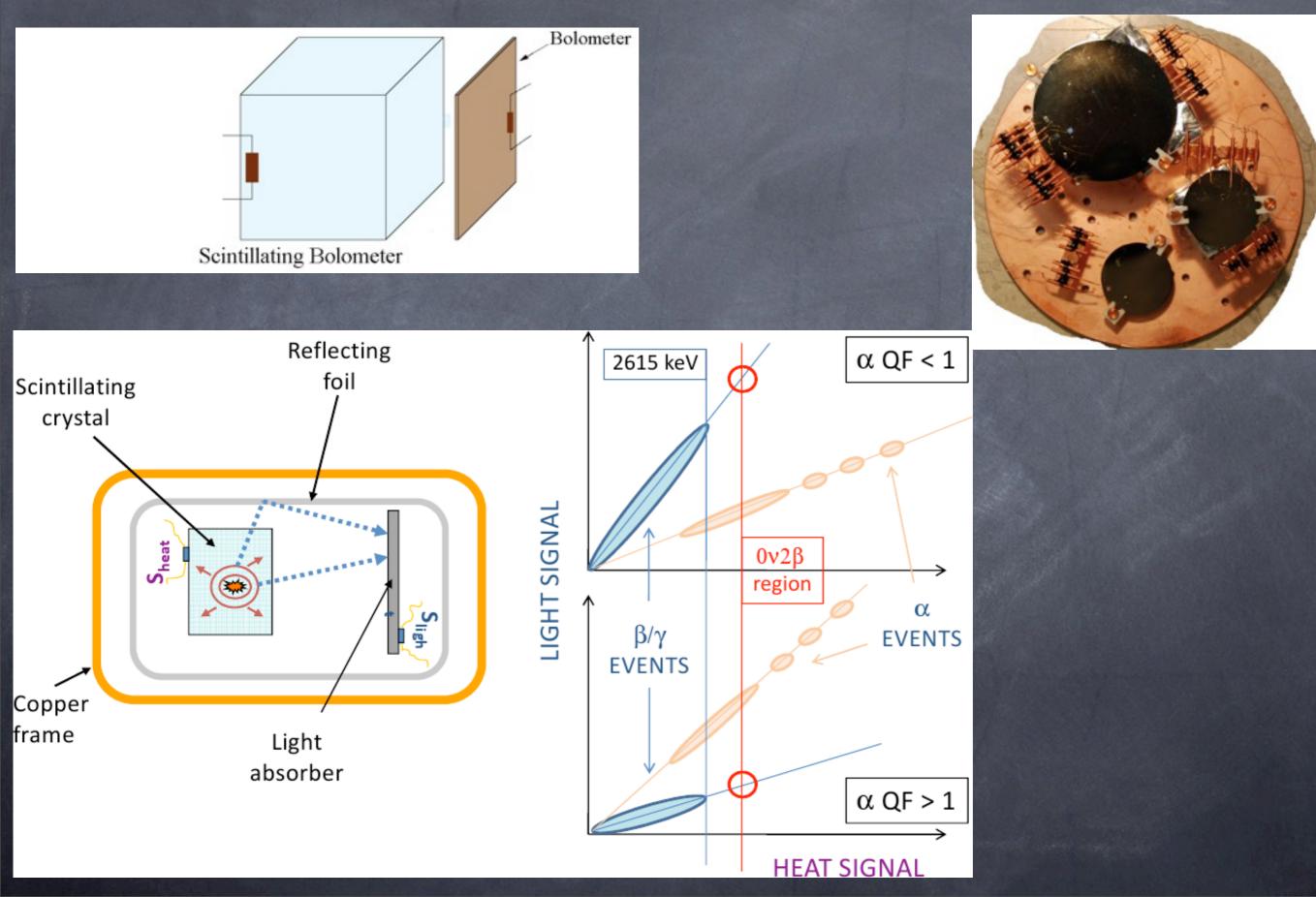
LUCIFER concept

Lucifer is a Latin word (from the words *lucem ferre*), literally meaning "light-bearer", which in that language is used as a name for the dawn appearance of the planet Venus, heralding daylight.

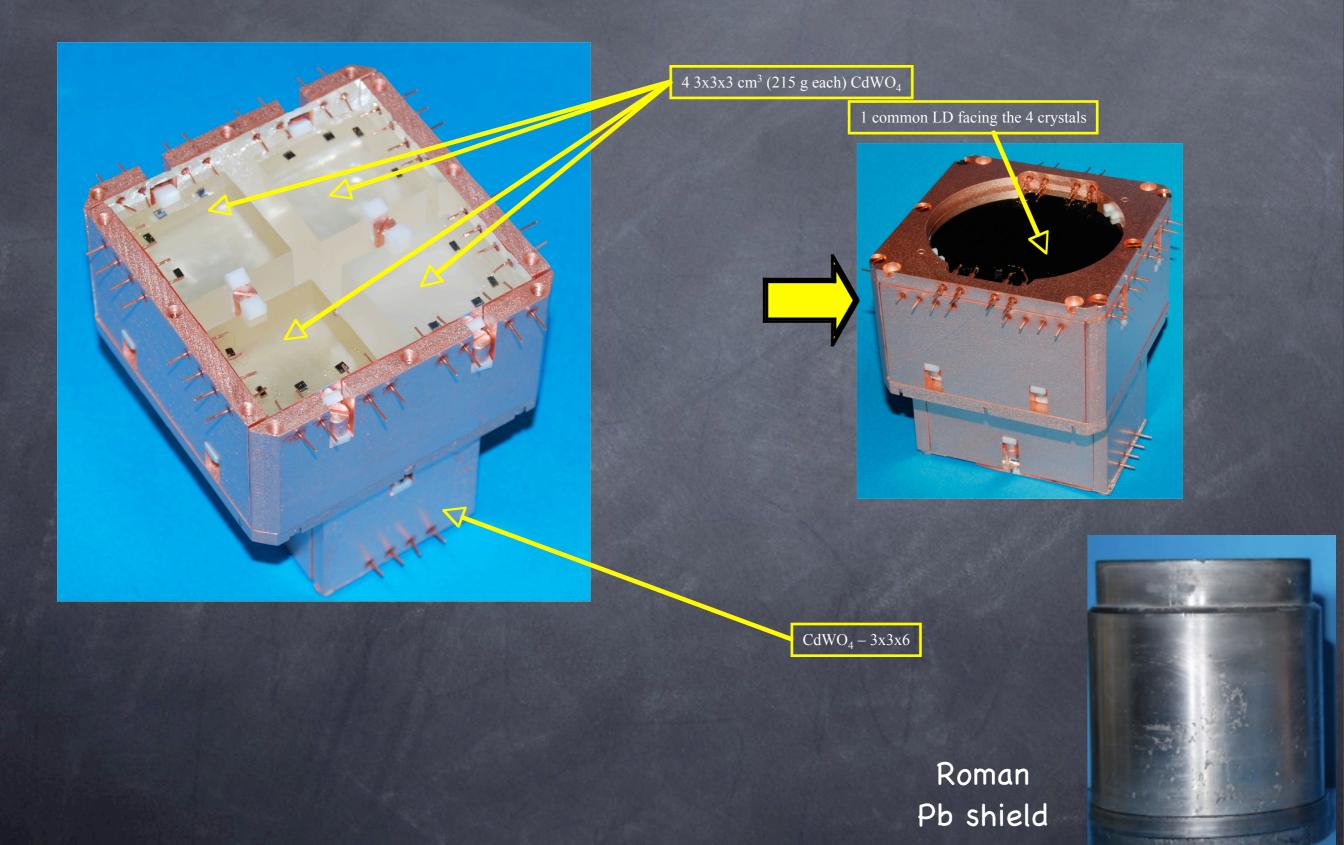


Bringing light underground

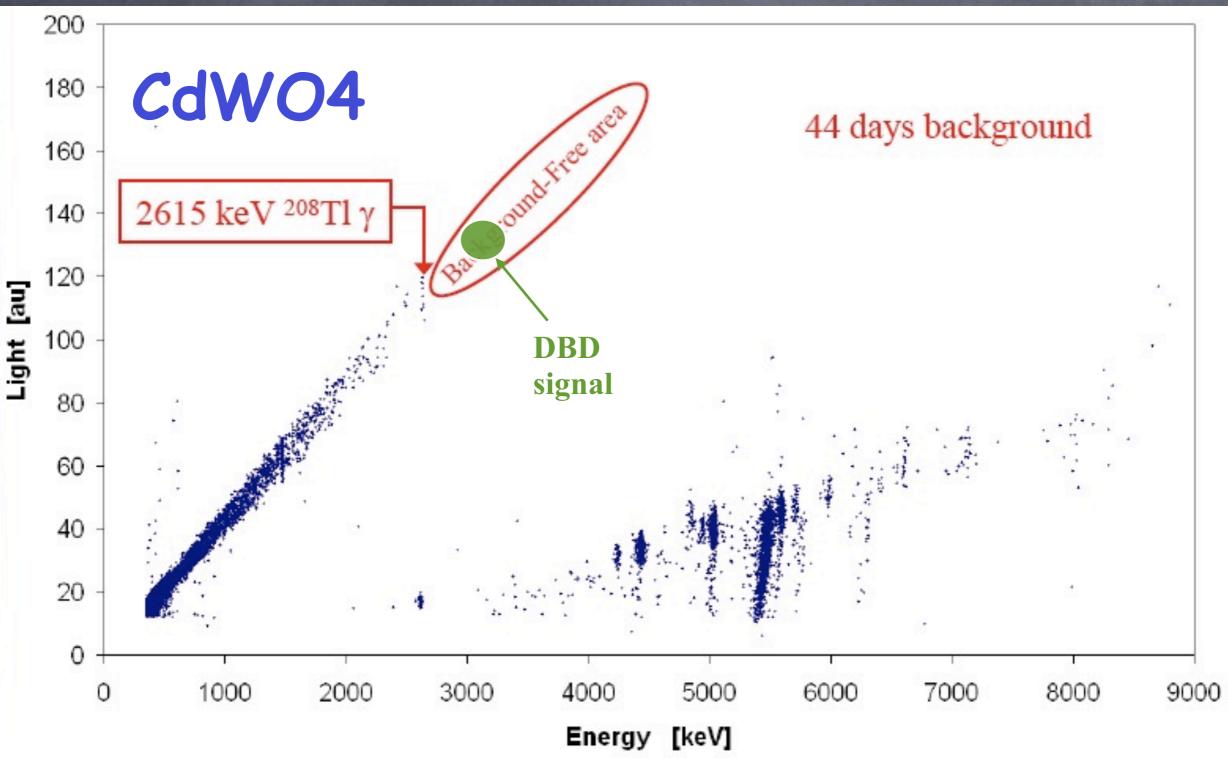
Double read-out



The best so-far

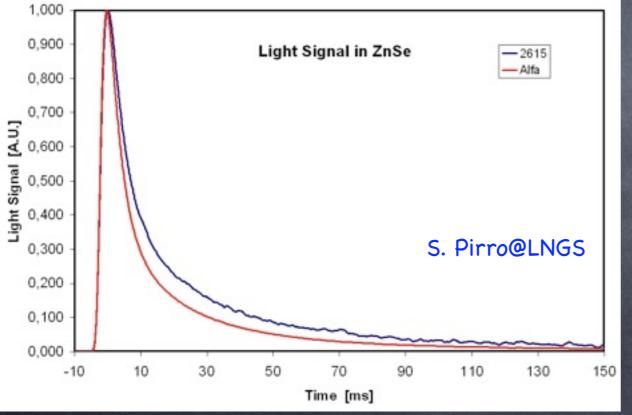


just to make the case clear

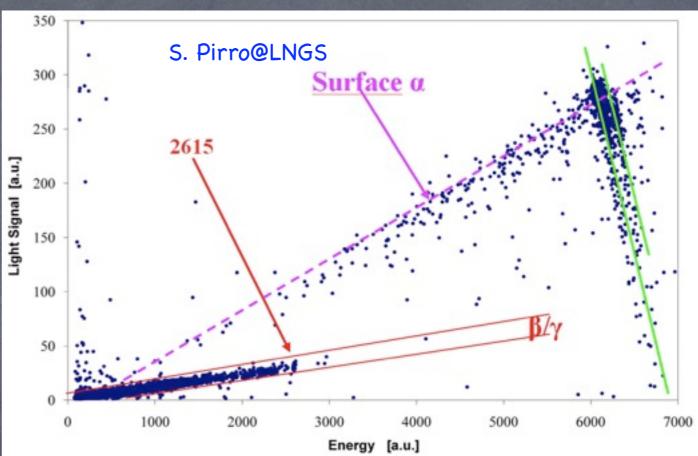


The best compromise (possibly)





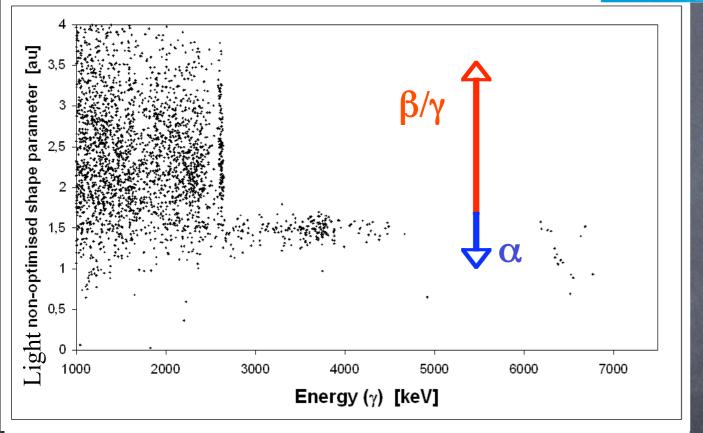
ZnSe



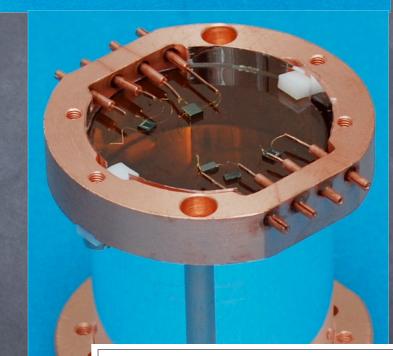
Goal: b < 0.001 c/Kg/KeV/y

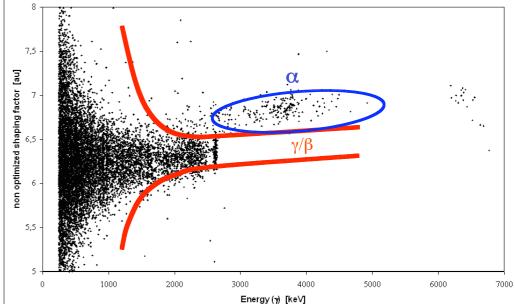
Pulse Shape Analysis

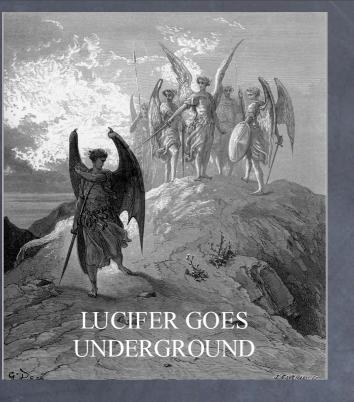
337 g "new" ZnSe Crystal



Calibration with ^{232}Th and a smeared α source



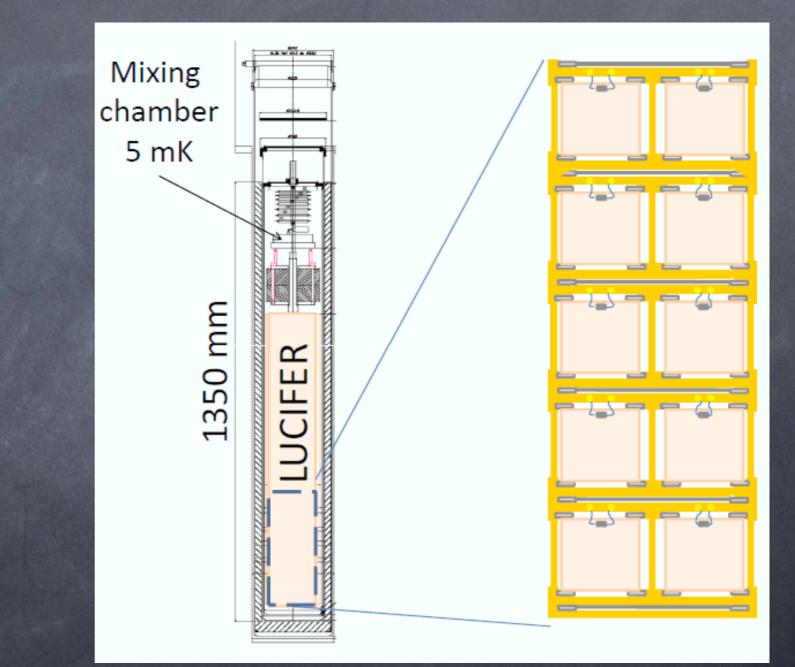


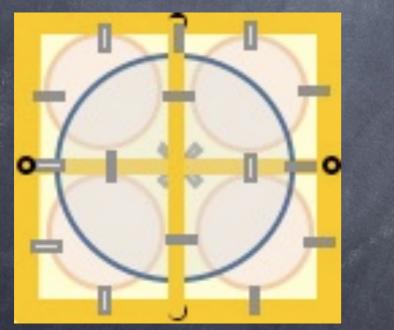


LUCIFER

Low-background Underground Cryogenic Installation For Elusive Rates

ERC-2009-AdG 247115





The challenge

- negotiate a good contract for ⁸²Se separation
- øget chemically pure and radiopure isotope after enrichment
- ø efficient cristallization

Got Physics ?

In the second second

- \odot T_{1/2} in excess of 10²⁶ y are expected
- The reach depends critically on cost, purification and efficient cristallization (high risk- high reward !)
- It would be great already to open the way to 3rd generation experiment

Conclusions

Neutrino Physics is one of the leading field in HEP today

Dirac or Majorana nature of neutrino mass is a fundamental question that needs to be answered at (almost) all cost(s)

Neutrino-less DBD might possibly be the sole chance to give a measure of neutrino mass

The second generation experiments might not be enough to win.

We have to prepare for third generation.
 Toward O background.