

*Imitation of the inverse beta-decay reaction  
in the neutrino scintillator detectors  
with the controllable UV/visible LED flasher*

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# Motivation

- 1) Imitation of the inverse beta-decay reaction and other processes in the organic liquid scintillator detectors
- 2) Calibration at high energies (10-50 MeV)

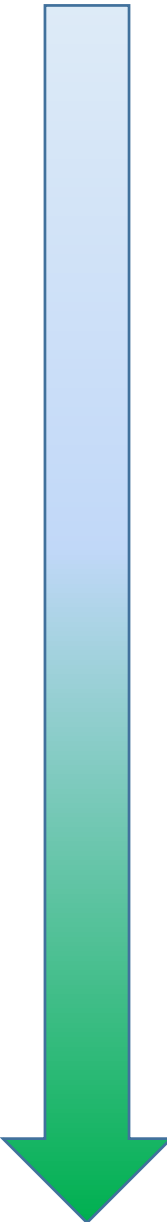
# Two complementary approaches to simulation

- 1) Computer simulation (Monte Carlo and so on)
- 2) Simulation events inside the detector with a UV sources

The second approach necessarily requires three ingredients:

- a) nanosecond LED driver (flasher)
- b) monitoring line to control the energy of the simulated event
- c) special diffuser that increases the isotropy of the outgoing UV radiation

# Historical overview



1985	<p><b>Original idea: the electrical scheme of nanosecond LED flasher</b> <b>Kapustinsky's driver:</b> Kapustinsky J S et al. 1985 <i>Nucl. Instrum. Meth. A</i> 241 612–13</p>
1990s, 2000s	<p><b>Some implementations (not all!):</b> &gt; 1997 the <b>Baikal experiment NT-200</b> Belolaptikov I A et al. 1997 <i>Astropart. Phys.</i> 7 263–82 &gt; 2001 the <b>ANTARES underwater neutrino telescope</b> McMillan J E et al. 2001 <i>Proc. of 27th ICRC (Hamburg)</i> p 1287</p>
2006	<p><b>Upgrade the driver for using with ultraviolet and blue LEDs:</b> Lubsandorzhiev B K and Vyatchin Y E 2006 <i>JINST</i> 1 T06001 (Preprint arXiv:physics/0410281v1)</p>
2015, 2016	<p><b>The first prototypes for online calibration of neutrino liquid scintillator detectors at energies above 10 MeV:</b> Chepurnov A S, Gromov M B and Shamarin A F 2016 <i>J. Phys.: Conf. Series</i> 675 012008 Chepurnov A S, Gromov M B et al. 2017 <i>J. Phys.: Conf. Series</i> 798 012118</p>
2017	<p><b>The UV LED calibration system for Borexino</b> It has not been applied because the calibration campaign is postponed for indefinite period</p>

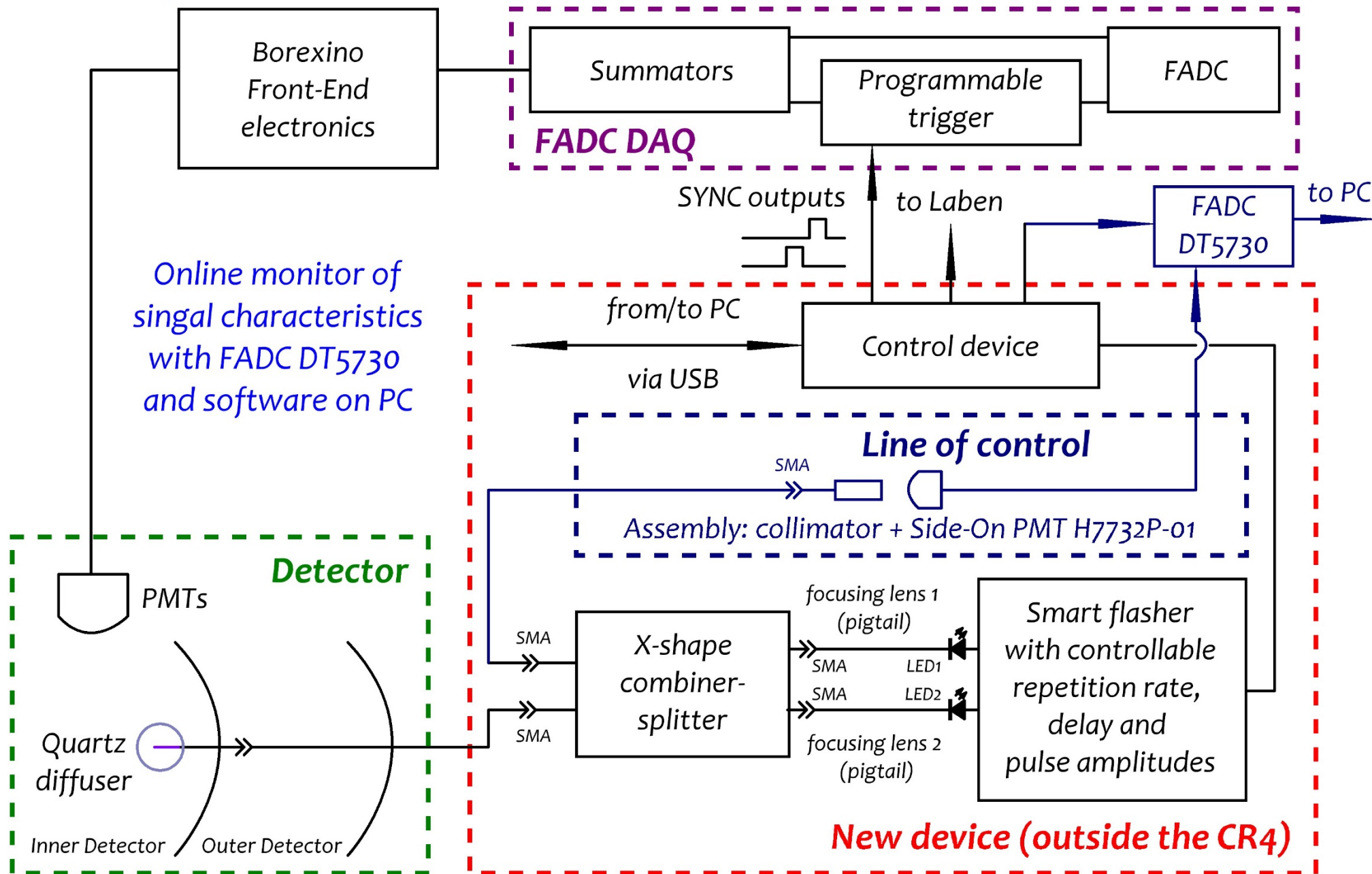
# Main ideas

- Mimic of real events in any neutrino liquid scintillator detector with a controllable UV double-LEDs flasher
- The energy of the imitated signal is directly proportional to the number of emitted photons but depends on the source position

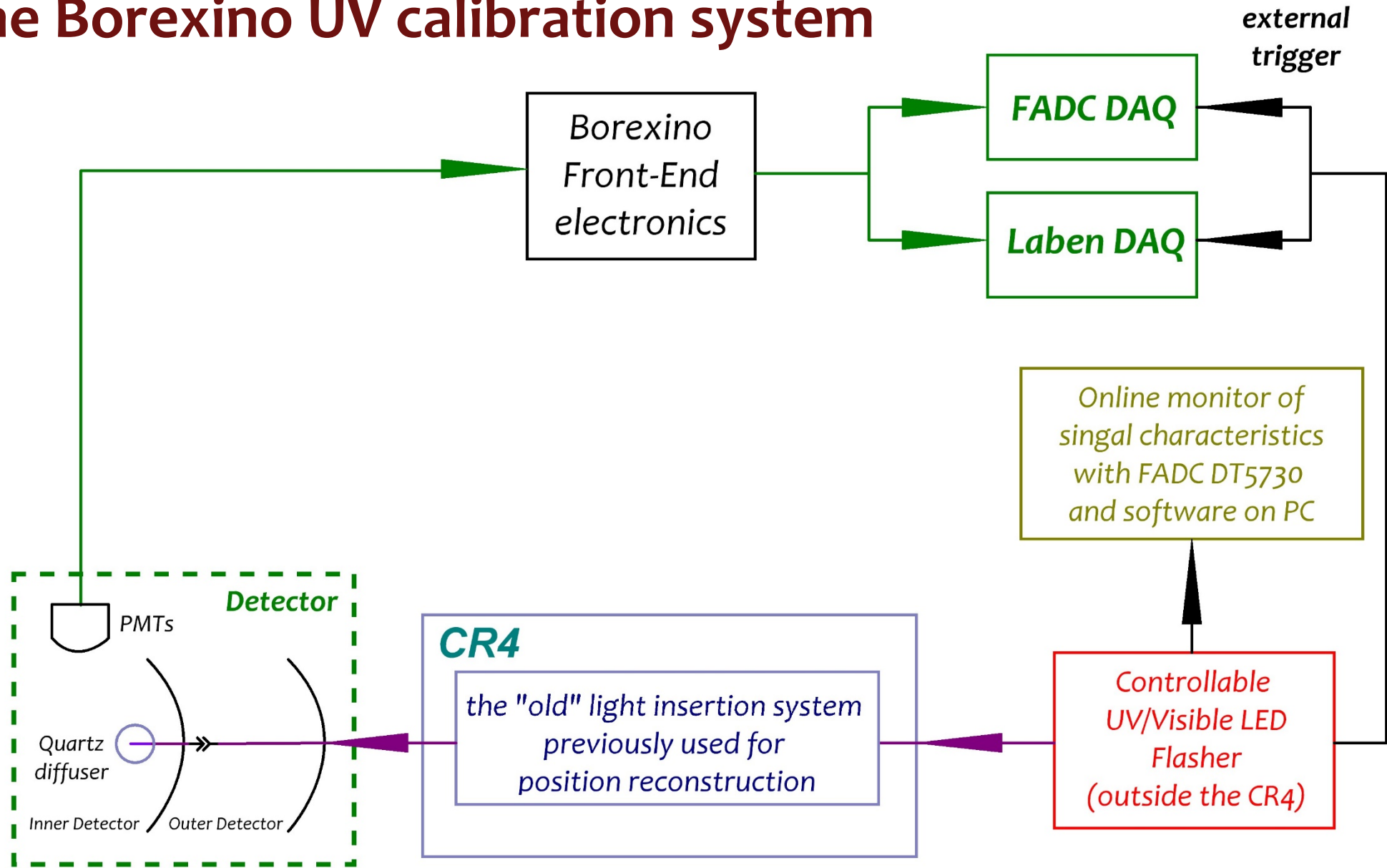
# Foundations of the conception

- 1) Adjustable level of the output power of the LED flashers
- 2) Double-LEDs scheme
- 3) Remote control
- 4) Pulse to pulse monitoring of the flasher output
- 5) All components except the end of the fiber with diffuser can be placed outside a cleanroom
- 6) To increase the isotropy of the UV radiation the diffuser is mounted at the tip of the fiber inside the detector
- 7) Hot swapping
- 8) UV or/and Visible LEDs

## The Calibration System Based On the Controllable UV/Visible LED Flasher simplified scheme

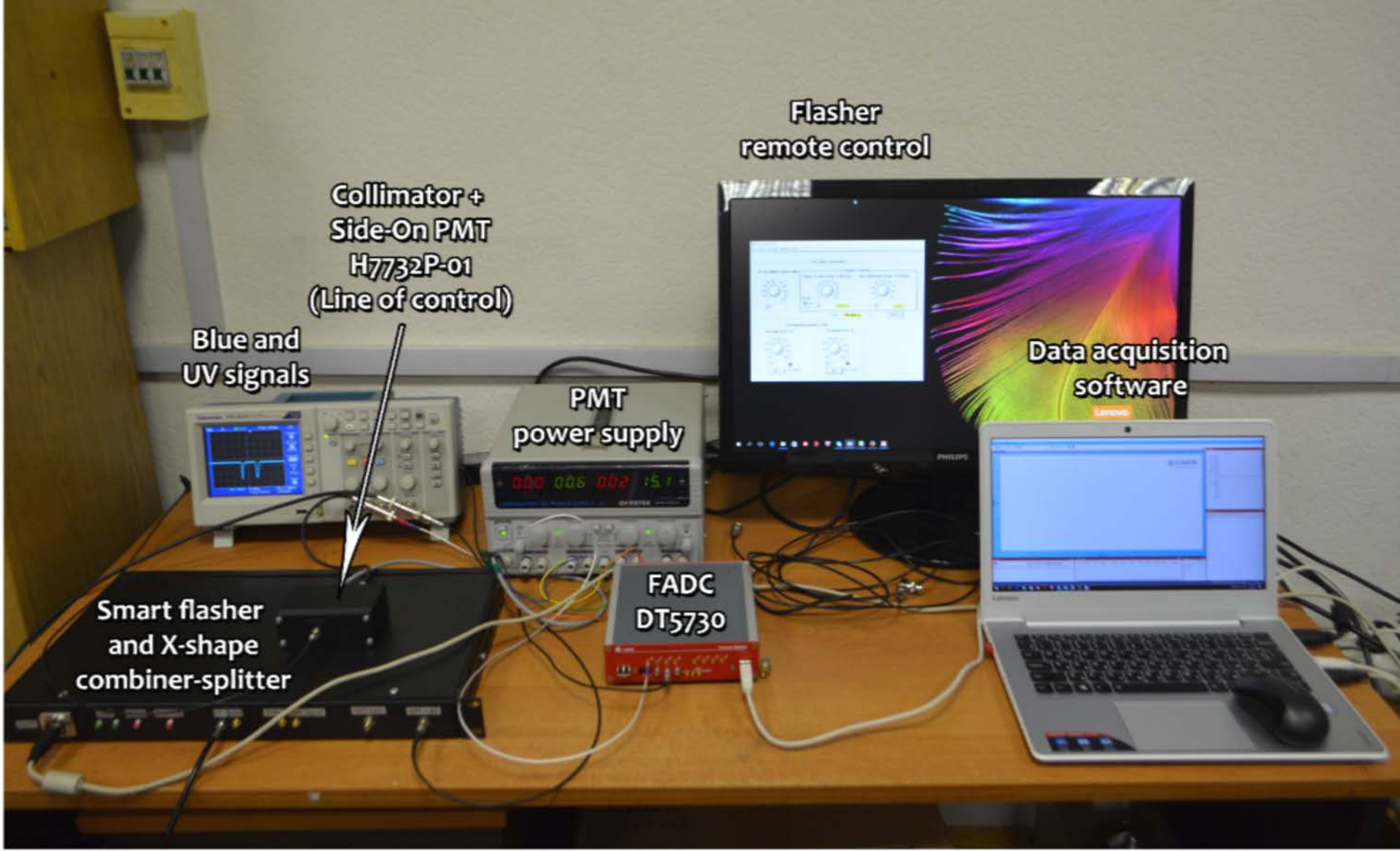


# General overview of the Borexino UV calibration system





# SETUP FOR TESTING @ SINP MSU



Flasher  
remote control

Collimator +  
Side-On PMT  
H7732P-01  
(Line of control)

Blue and  
UV signals

PMT  
power supply

Data acquisition  
software

Smart flasher  
and X-shape  
combiner-splitter

FADC  
DT5730

# UV LED power (for $\tau = 20$ ns, $\lambda = 260$ nm)

The number of UV photons  
in Borexino (MC simulation):

$$m_{\text{bx}} \sim 10^4/\text{MeV}$$

Attenuation:

$$n_{\text{exp}} \sim 10^2 - 10^3 \text{ (based on tests)}$$

The number of photons  
emitted by the UV LED:

$$m \sim 10^7/\text{MeV}$$

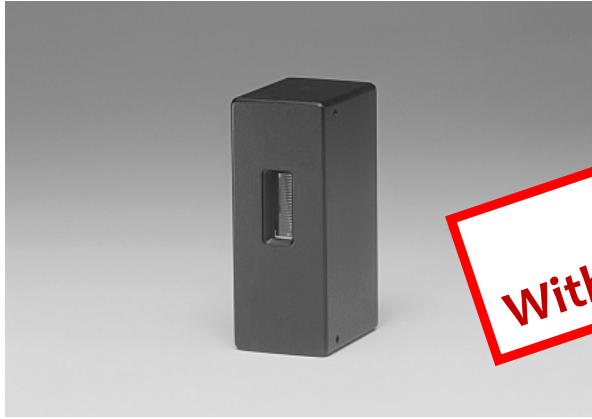
The UV LED power:

$$P \sim \frac{mhc}{\lambda\tau} \sim 4 \text{ mW (for 10 MeV)}$$

The standard UV LED power:

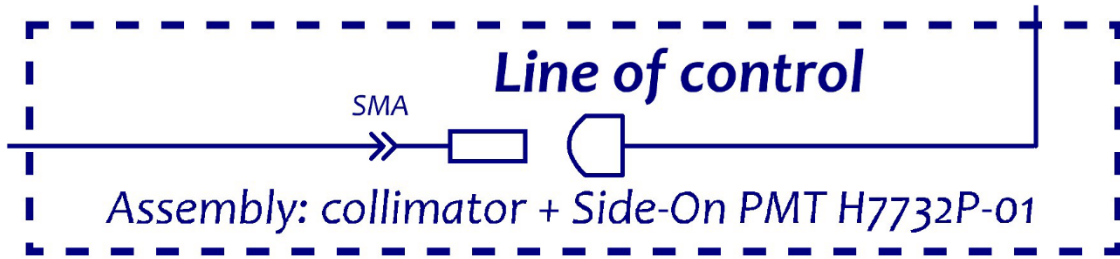
$$P \sim (0.1 - 1.0) \text{ mW}$$

**UV LED in pulse mode  
(high current mode)**



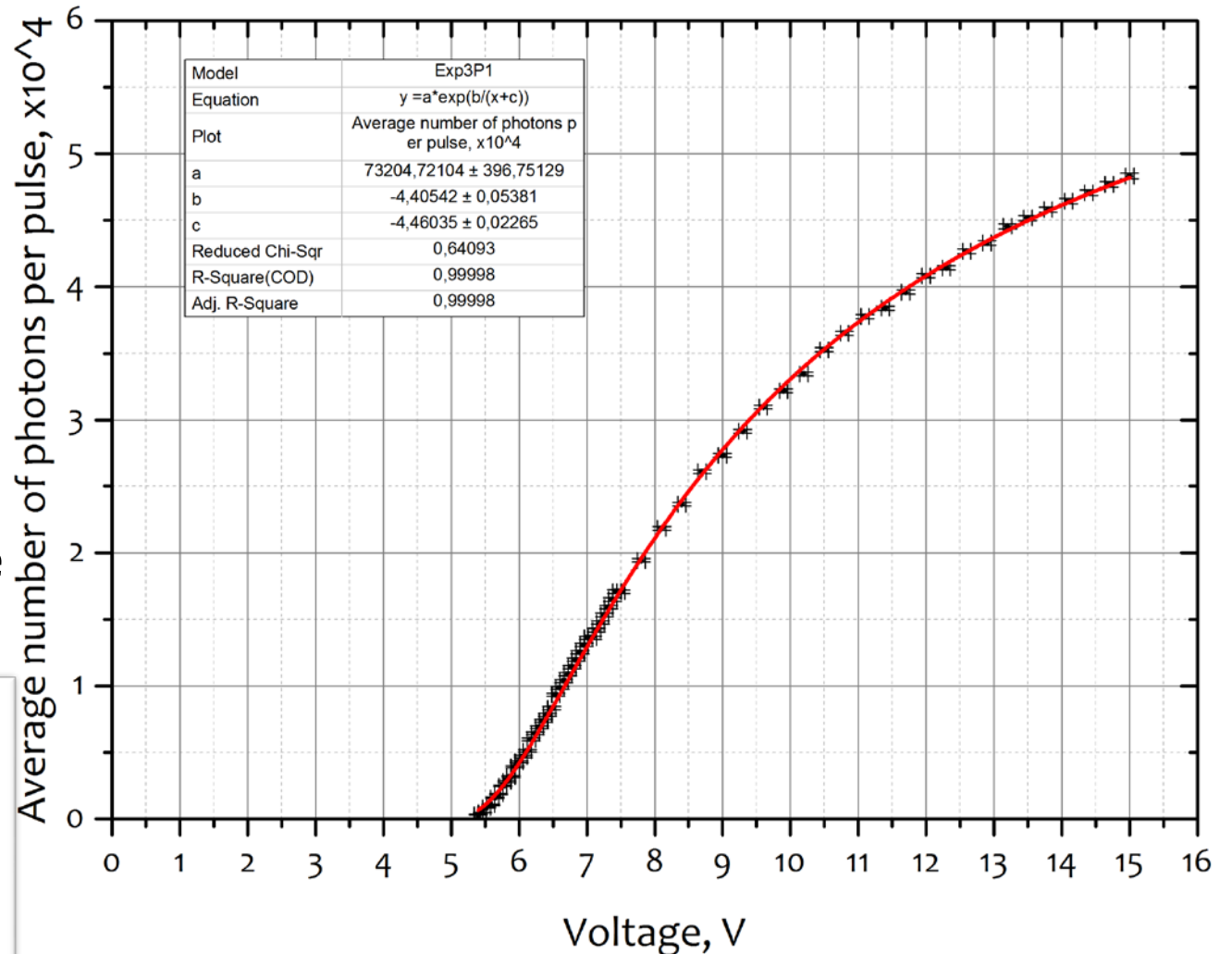
**Side-On PMT  
With adjustable sensitivity**

# Line of control



**Number of photons emitted by the UV LED  
as a function of the voltage applied to the diode  
(empirical dependence)**

$$\langle k_\gamma \rangle = a \exp\left(\frac{b}{V_d + c}\right)$$



# Line of control:

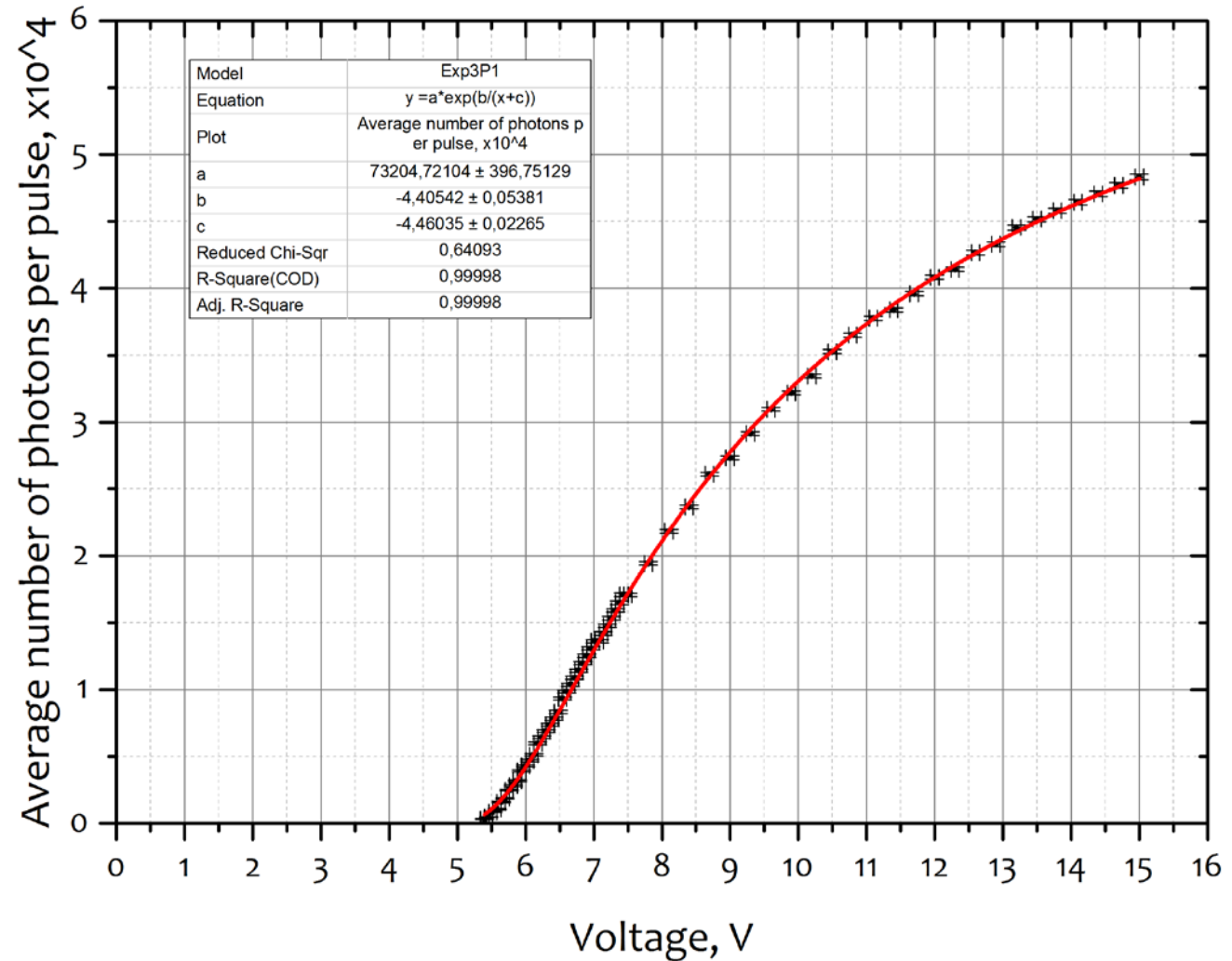
## side-on PMT **VS** photodiode + preamplifier

- The characteristics of both monitoring system are practically the same
- The system based on a photodiode and preamplifier was applied initially
- A good optical system for focusing radiation in case of the control line with a photodiode is difficult to manufacture
- Instead of manufacturing the system can be purchased but it's quite expensive
- It's easier to apply a side-on PMT

# Measurement features

All measurements are relative

So that the calibration with  $\gamma$  source  
( $E \sim$  a few MeV)  
can be required in each position in the detector



# Windows application for control

Adjustable parameters:

- simulated energy
- delay time between two continuous pulses
- repetition rate



Simulation physical event  
in the detector,  
moreover  
physical spectra

Flasher emulator

File Port Terminal About! Exit!

The pulse parameters

The repetition period, mks

Channel 2 delay

Rough scale delay, x 16.67ns

The additional delay, x 0.167ns

Scale

x100

x1

15 250.001 ns

0 0.000 ns

Delay = 250.001 ns ns SET

The brightness of the LEDs

Channel1 LED, %

Channel2 LED, %

30.8 SET On/Off

35.6 SET On/Off

RS232 monitor

Connection Error

From RS232 Clear

P1115

P293

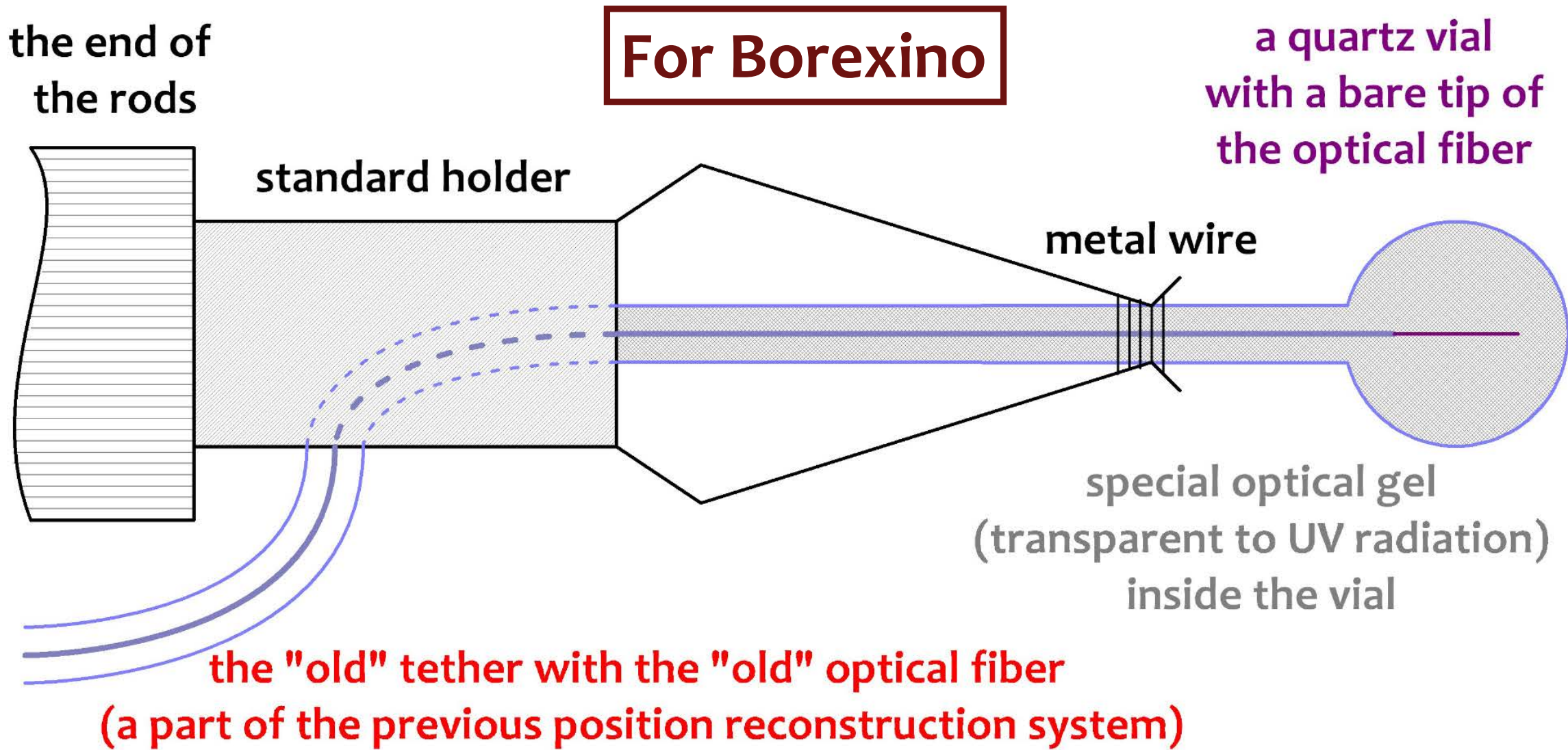
s21

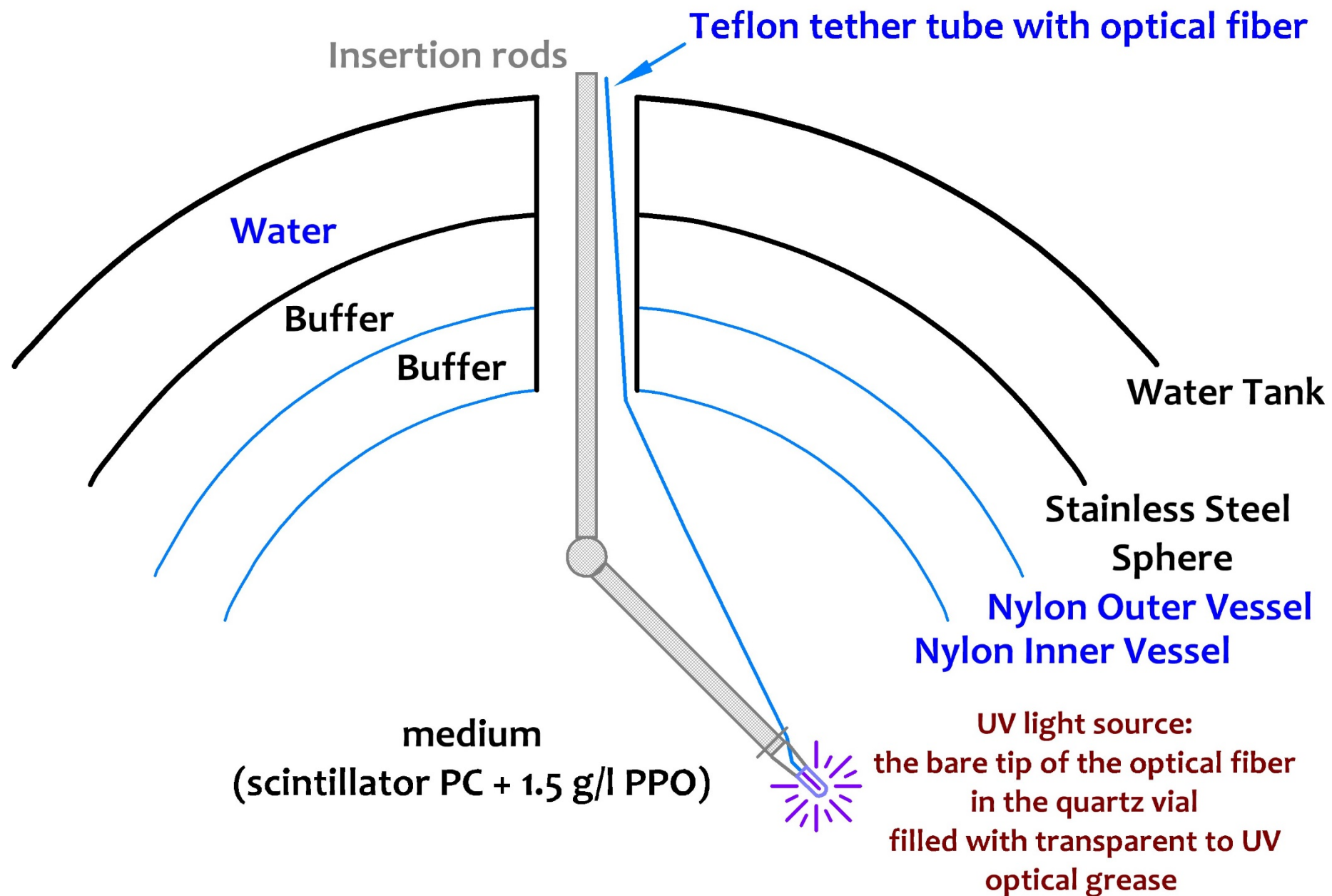
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# Diffuser and its holder

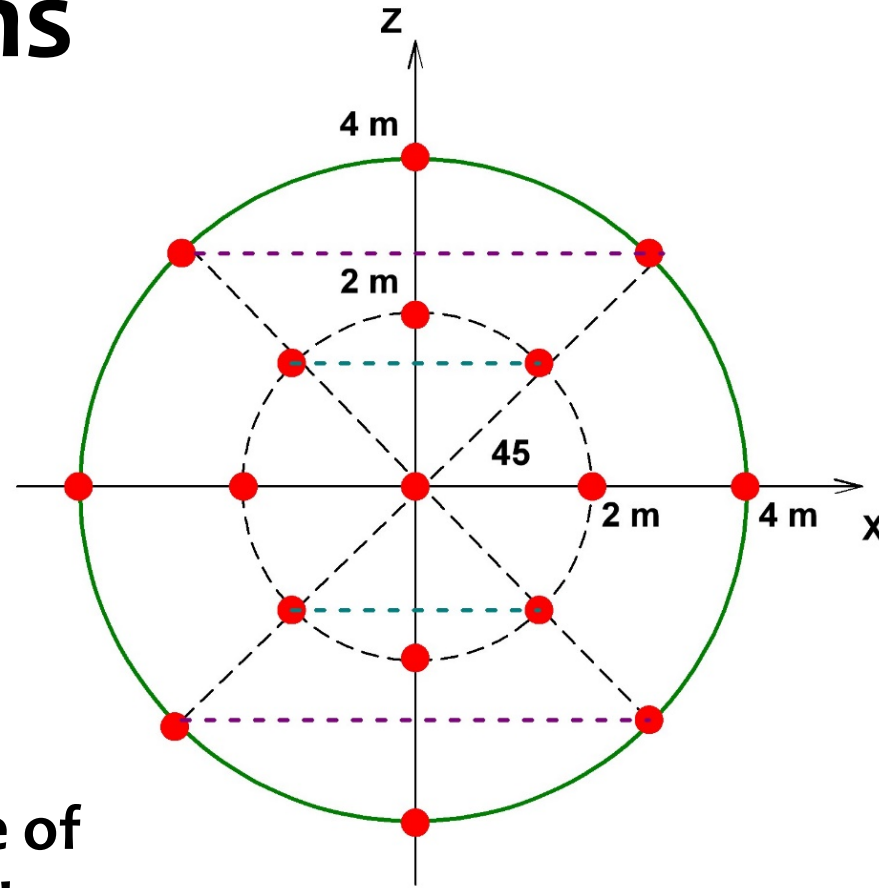




Principle scheme of the Borexino off-axis calibration

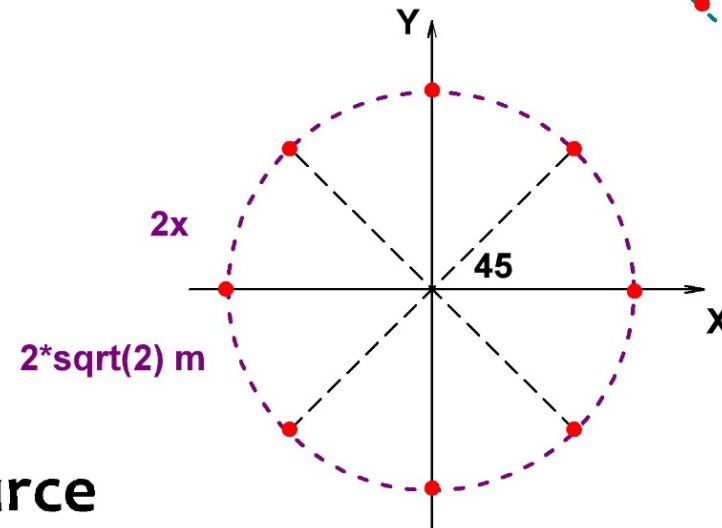
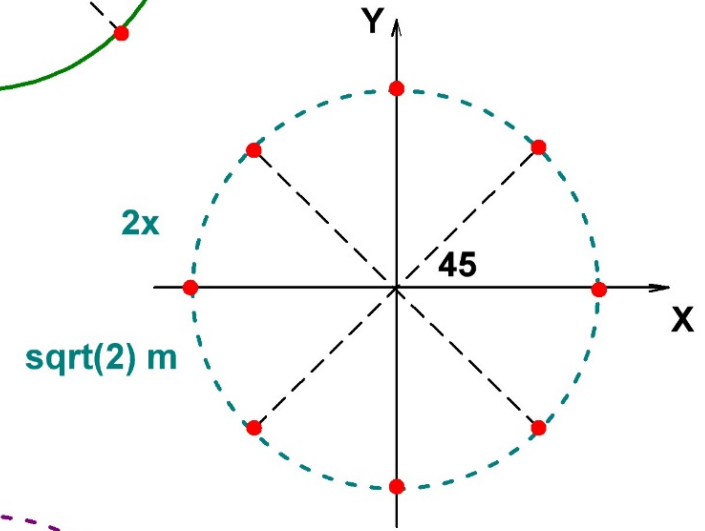
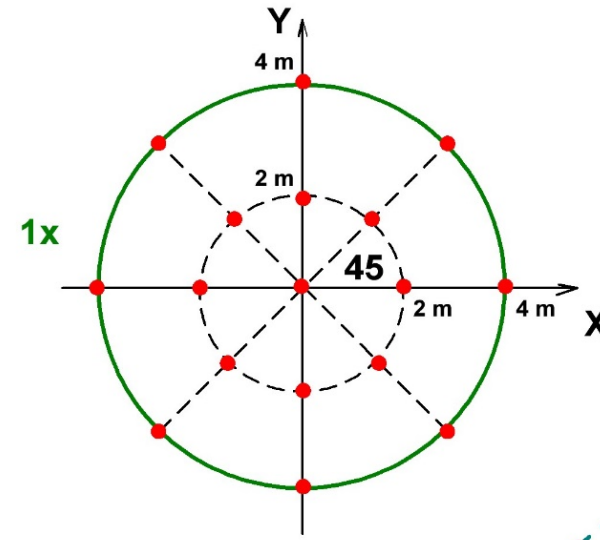


# Positions



example of  
a simple scheme  
for energy calibration of Borexino

53 positions of the UV source



# Advantages

- **A non-radioactive source**
- ✓ **safety for experimenters and technical staff**
- ✓ **safety for low background environment**
- ✓ **no papers**
- ✓ **it calms and save nerves**
- ✓ **it saves time**
- **Wide range of energies (100 keV – 50 MeV)**
- **A cheap calibration source**



# Advantages



- **Easy operation and control**
- **Mounted along with other systems  
(using infrastructure of other calibration systems)**
- **Possible fast calibration campaign:  
simultaneous calibration measurements  
(a few different sources at the end of the insertion rods)**

# **In conclusion: other possible applications**

- 1) Updating and correction of the energy scale  
in the range 1 - 10 MeV**
- 2) Study of the scintillator properties**
- 3) Reference information for different spatial  
reconstruction algorithms**
- 4) Pile-up study**
- 5) Additional detector hardware and  
software checks**

**Thank you for your attention!**