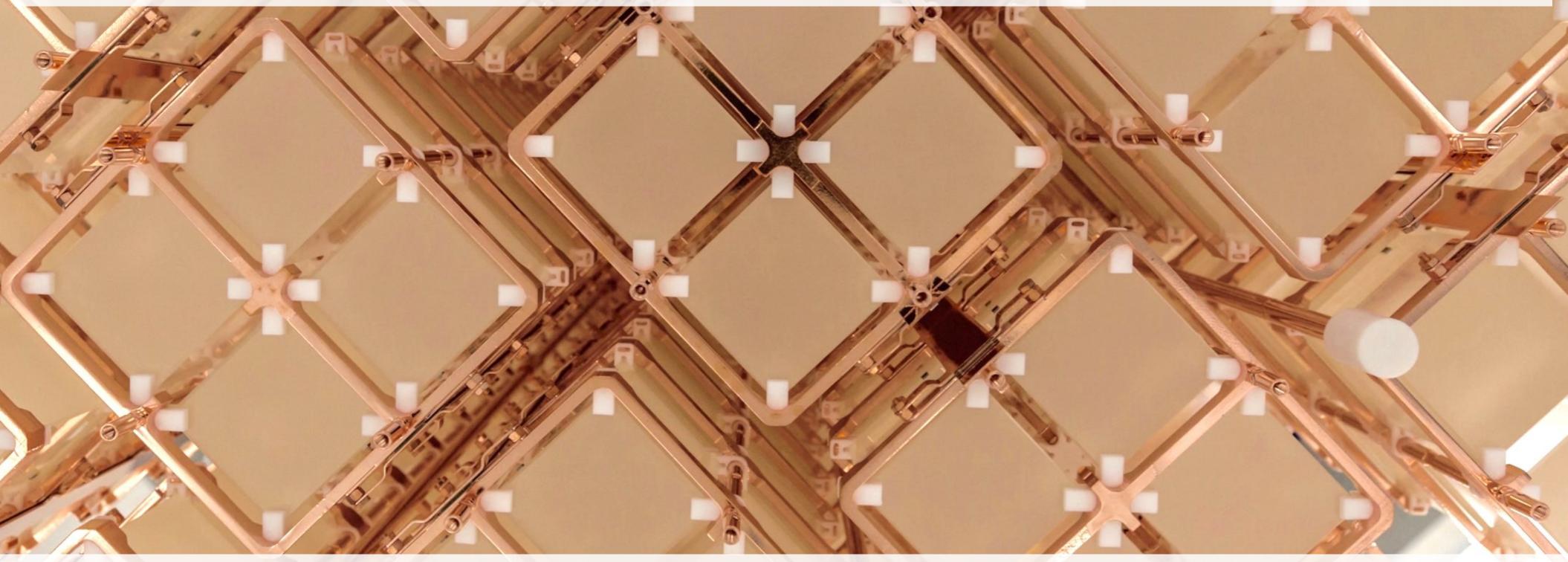
First Results from CUORE: A Search for Lepton Number Violation via 0vßß Decay of 130Te

Matteo Biassoni on behalf of the **CUORE** Collaboration INFN - Sez. Milano Bicocca



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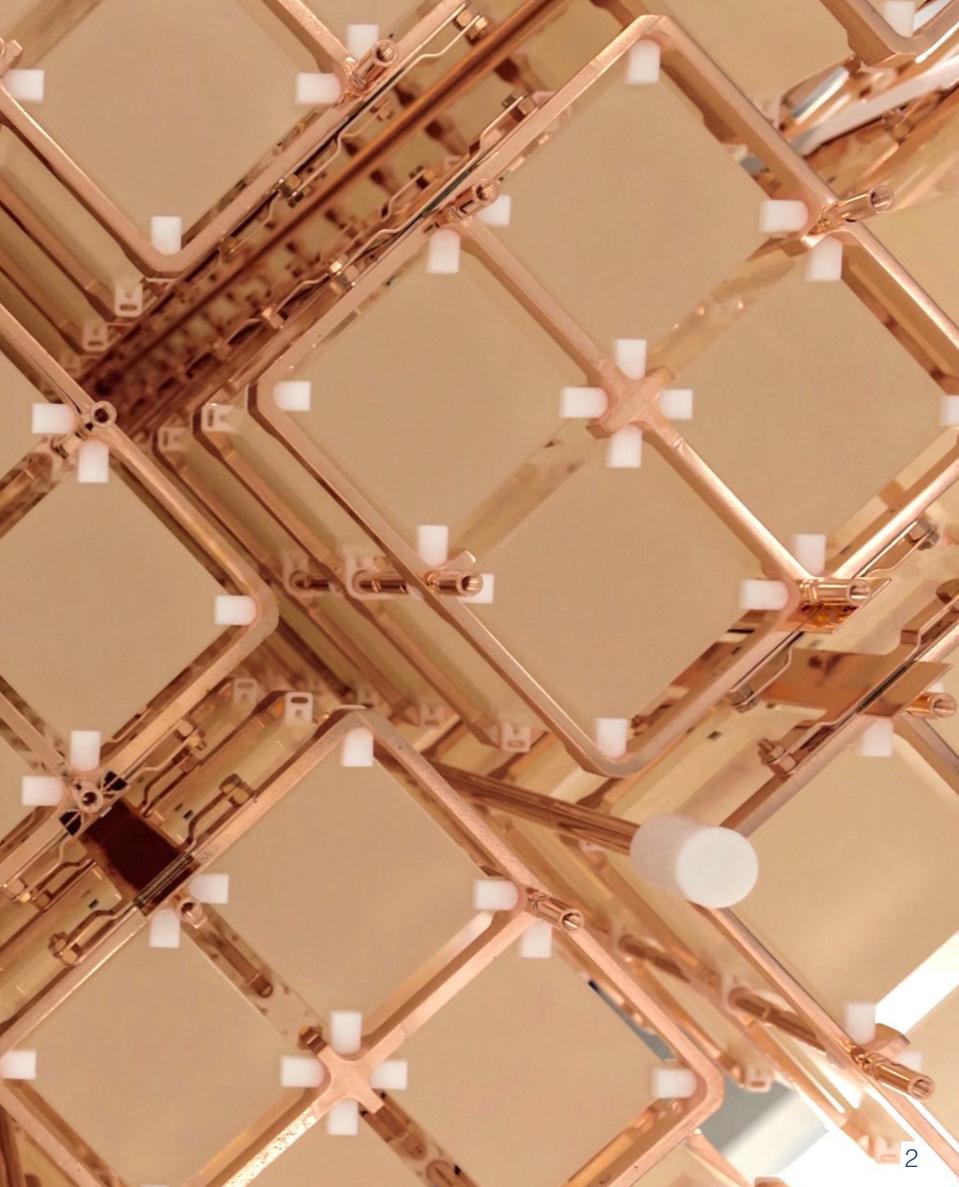






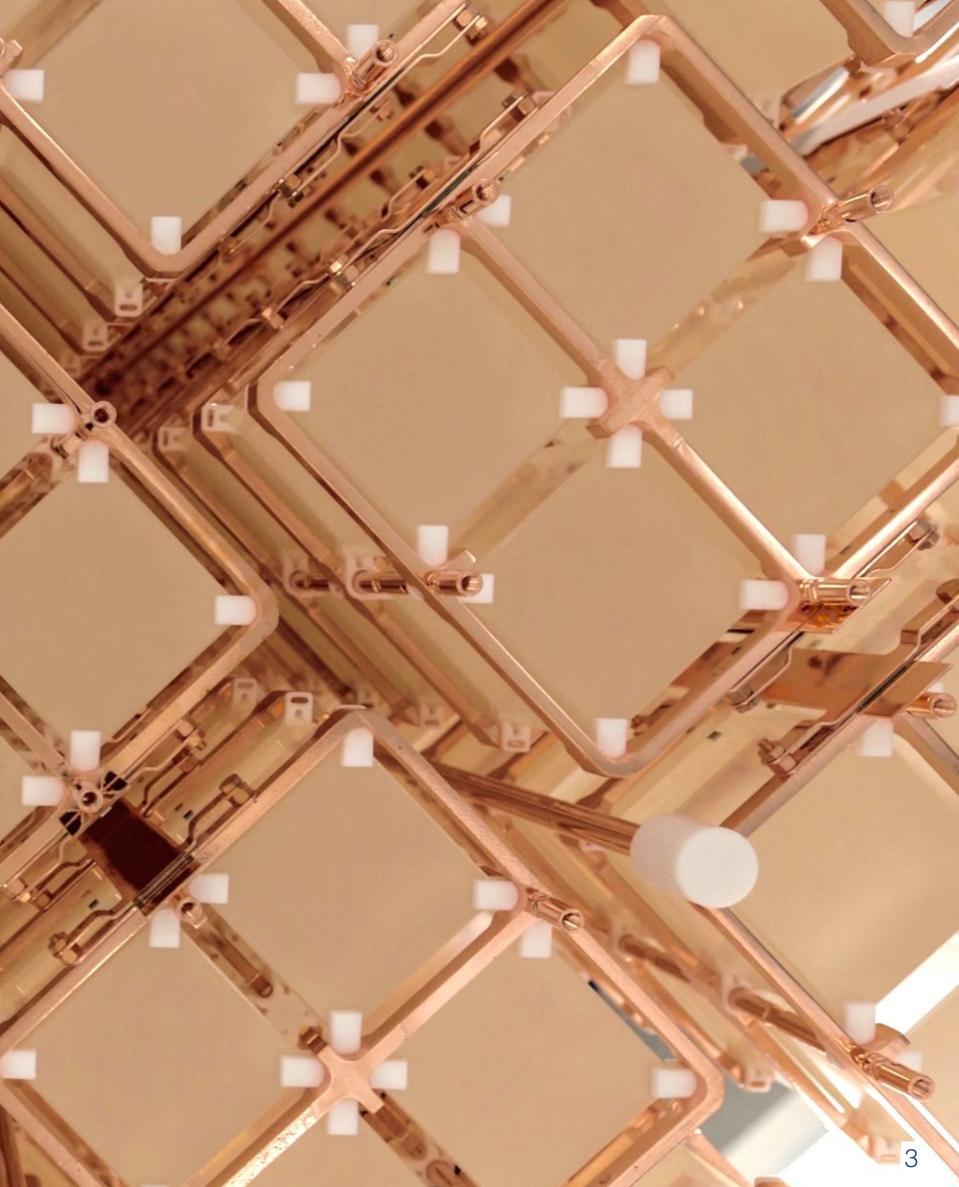
Outline

- TeO₂ and thermal detectors for neutrino-less DBD
- CUORE setup
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 - ► Fit
 - Systematics
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- Conclusions and outlook



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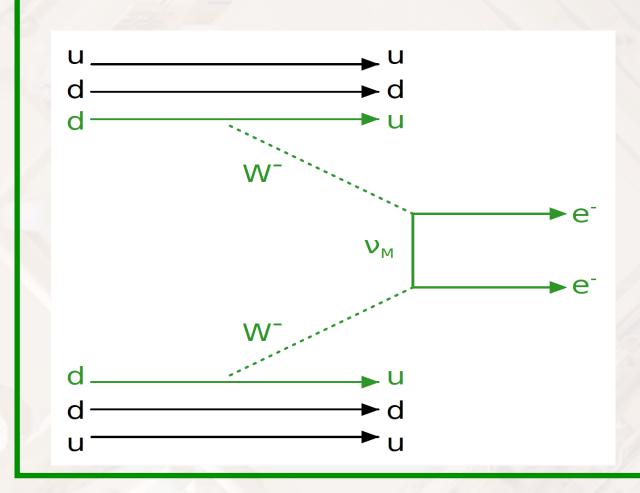


TeO₂ and thermal detectors for 0vDBD

Second order nuclear process, alternative to beta decay forbidden by mass difference for some even-even nuclei

 $(A, Z) \to (A, Z+2) + 2e^- + 2\bar{\nu}_e$ 2nd order SM process, $T_{1/2} \sim 10^{18 \sim 24}$ years

$(A,Z) \to (A,Z+2) + 2e^{-1}$



- SM forbidden, $\Delta L = 2$
- if observed, then neutrino is a Majorana particle
- underlying mechanism can give insight into beyond SM physics
 - light neutrino mass scale and hierarchy
 - heavy neutrino



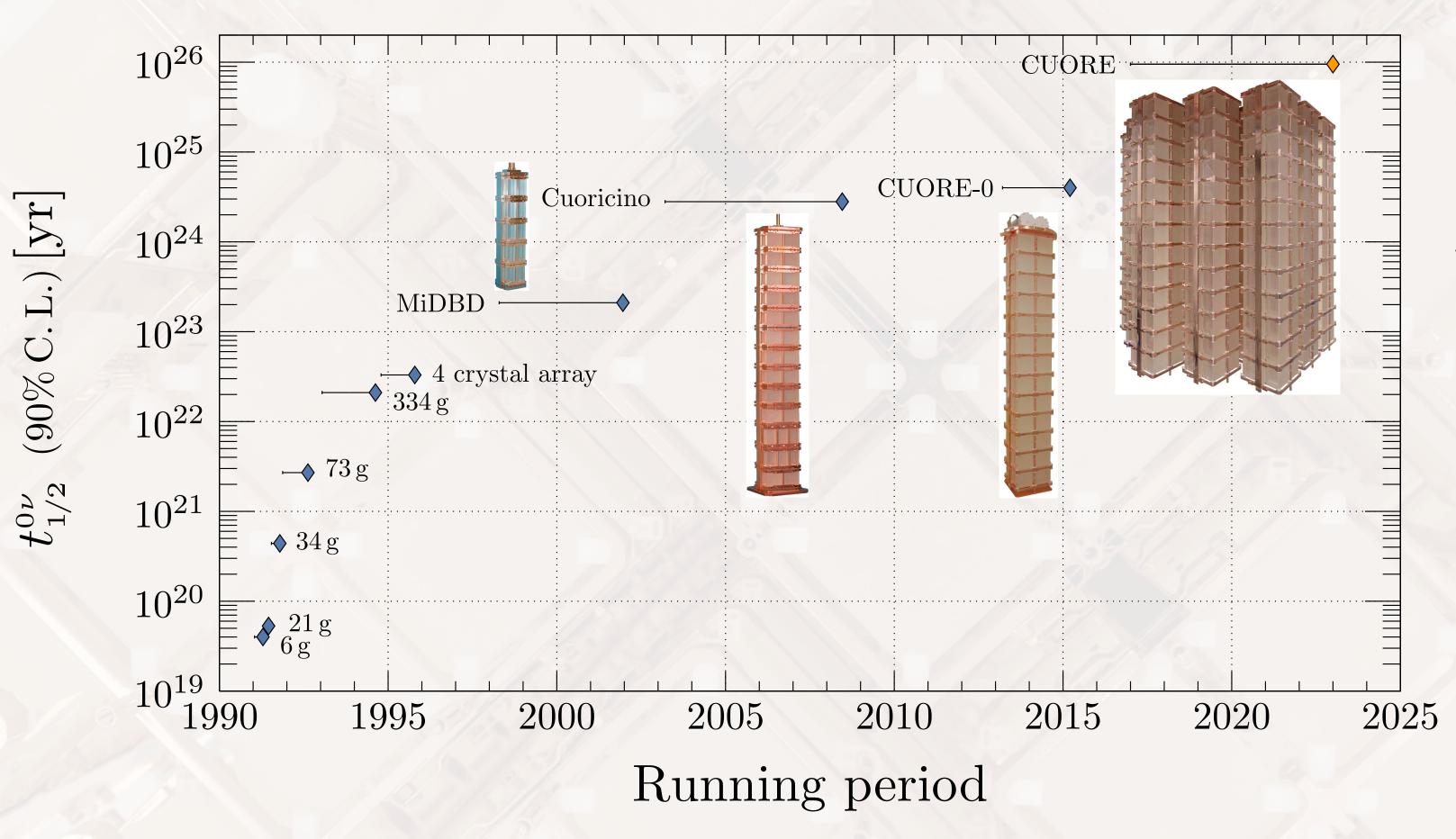
¹³⁰Te is a good candidate source for 0vDBD search:

- high natural isotopic abundance (~34%)
- NME and phase space on average
- Q-value (2528 keV) above most of the natural radioactivity
- easy to mix in convenient chemical compounds (TeO₂)

Thermal detectors are a good choice for 0vDBD search:

- excellent energy resolution
- large active mass and efficiency/unit cost
- fully active source and sensitive volume, no dead-layer

TeO₂ arrays: state of the art



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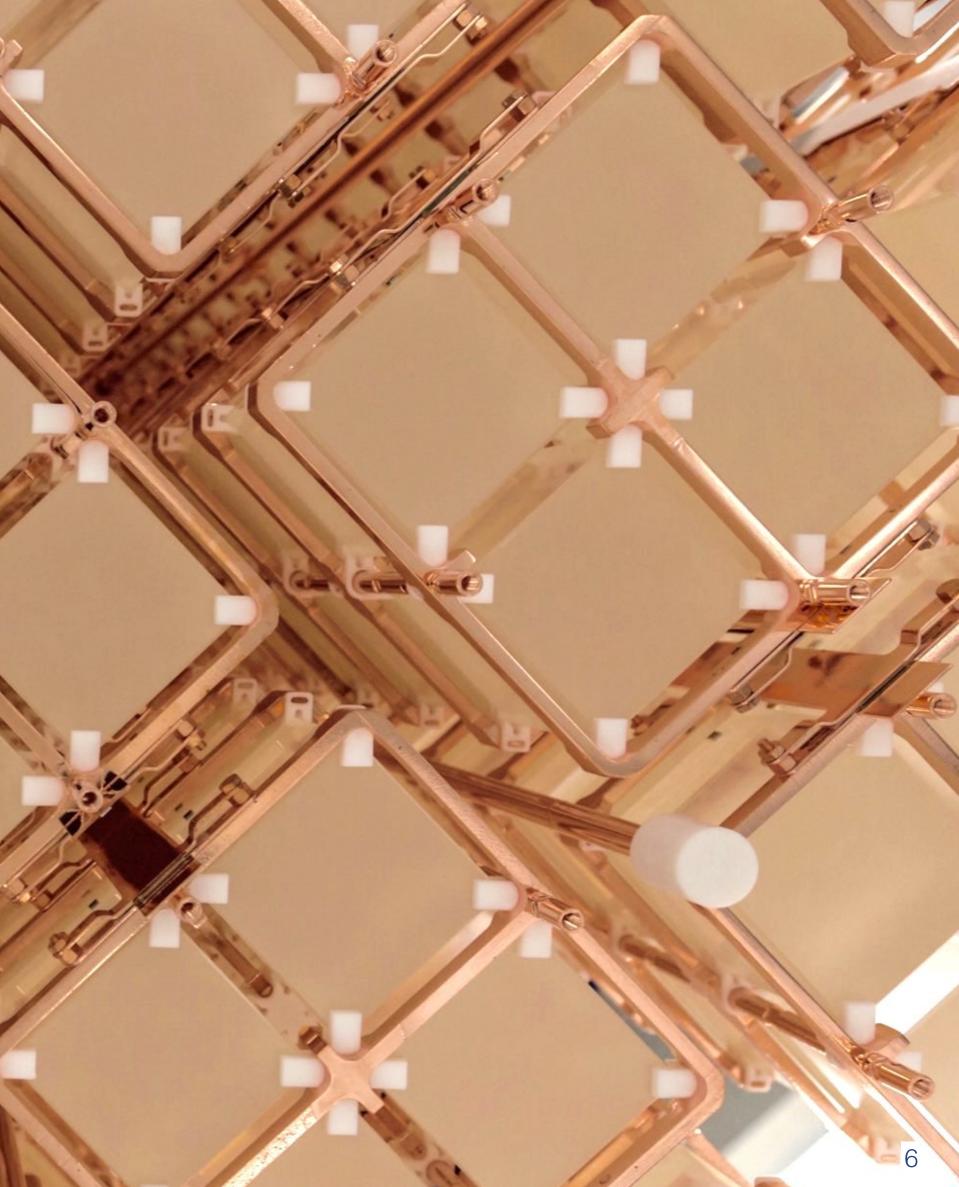


CUORE is the latest in a long progression of TeO₂ detectors which included two large demonstrators:

- Cuoricino (2.8x10²⁴ y)
- CUORE-0 (4.0x10²⁴ y combined)

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CUORE

Cryogenic Underground Observatory for Rare Events **Primary goal:** search for 0vββ decay in ¹³⁰Te

Detector design:

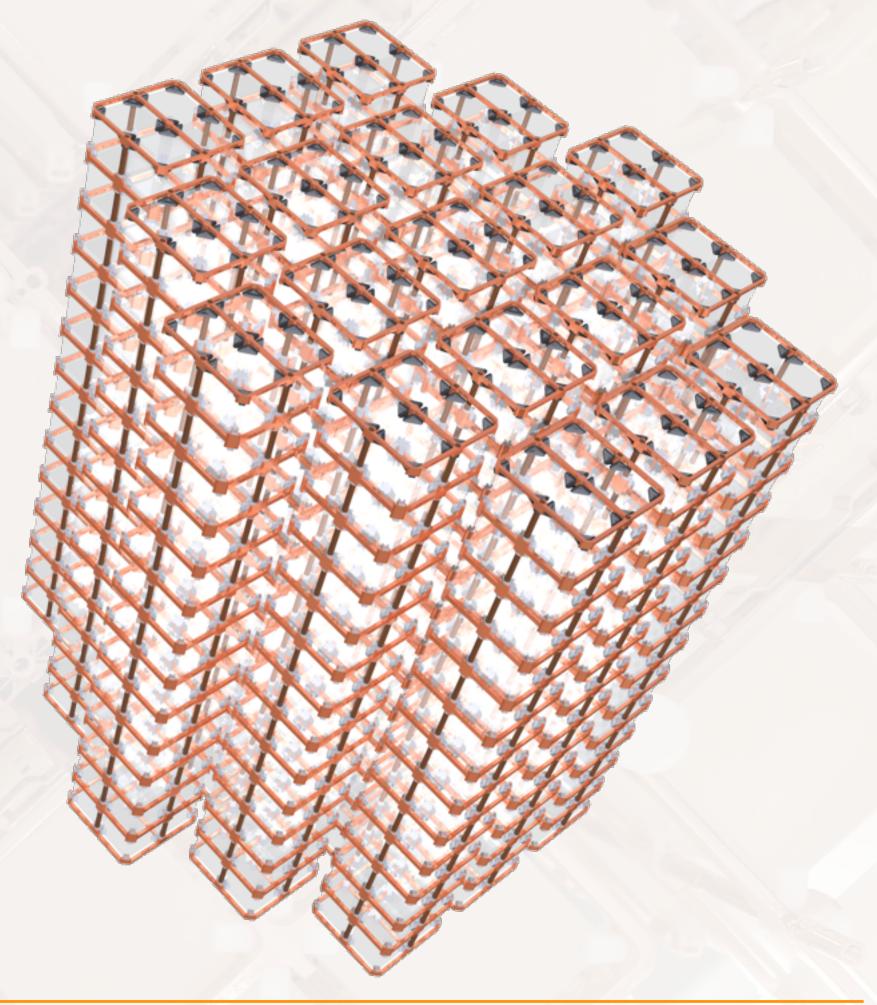
closely packed array of 988 TeO₂ crystals arranged in 19 towers

Design parameters:

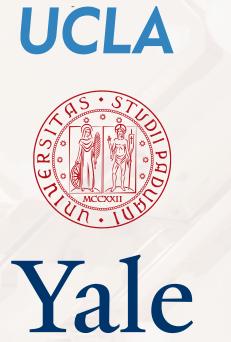
- mass of TeO2: 742 kg (206 kg of ¹³⁰Te)
- low background aim: 10-2 c/(keV·kg·yr)
- target energy resolution: 5 keV FWHM in the Region Of Interest (ROI)
- high granularity
- deep underground location
- strict radio-purity controls on materials and assembly

CUORE projected sensitivity (5 years, 90% C.L.): $T_{1/2} > 9 \times 10^{25}$ yr





The CUORE Collaboration

















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CUORE @ LNGS



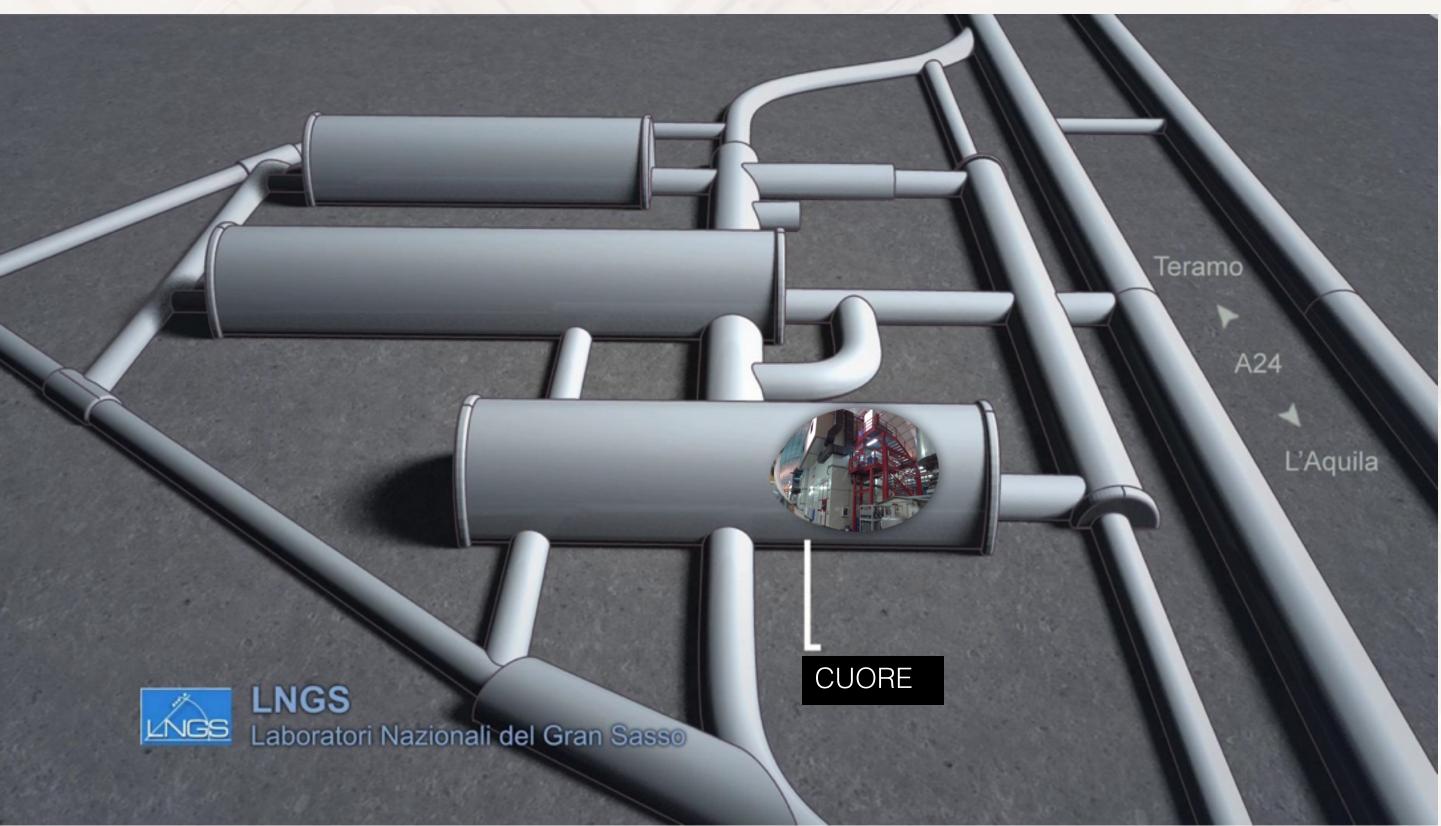


1400 m of rock (~3600 m.w.e.) deep • μ 's: $\sim 3 \times 10^{-8} / (s \cdot cm^2)$ • γ's: ~0.73 / (s·cm²) • neutrons: 4×10^{-6} n/(s·cm²) below 10 MeV



CUORE @ LNGS







Underground Laboratory

- Three-story building
- Hosting the cryostat supporting structure



Main Support Plate

Y beam

Cryostat

H₃BO₃ panels

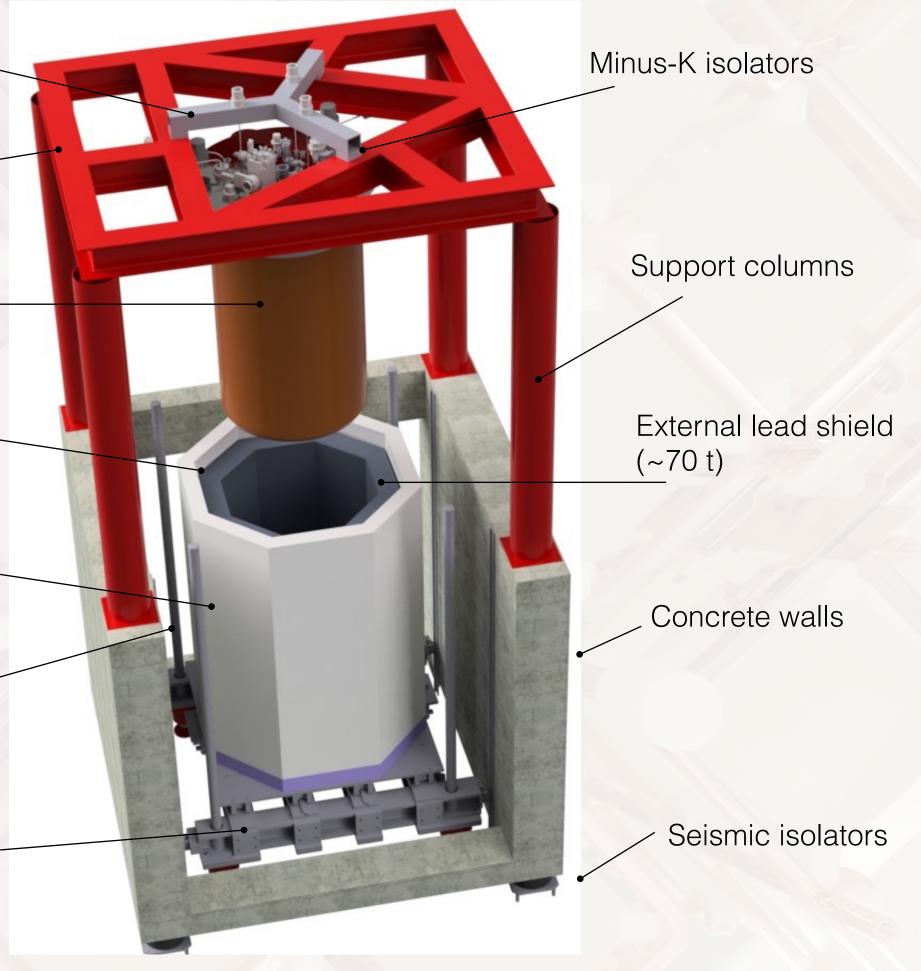
Polyethylene

Screw jacks

Movable platform

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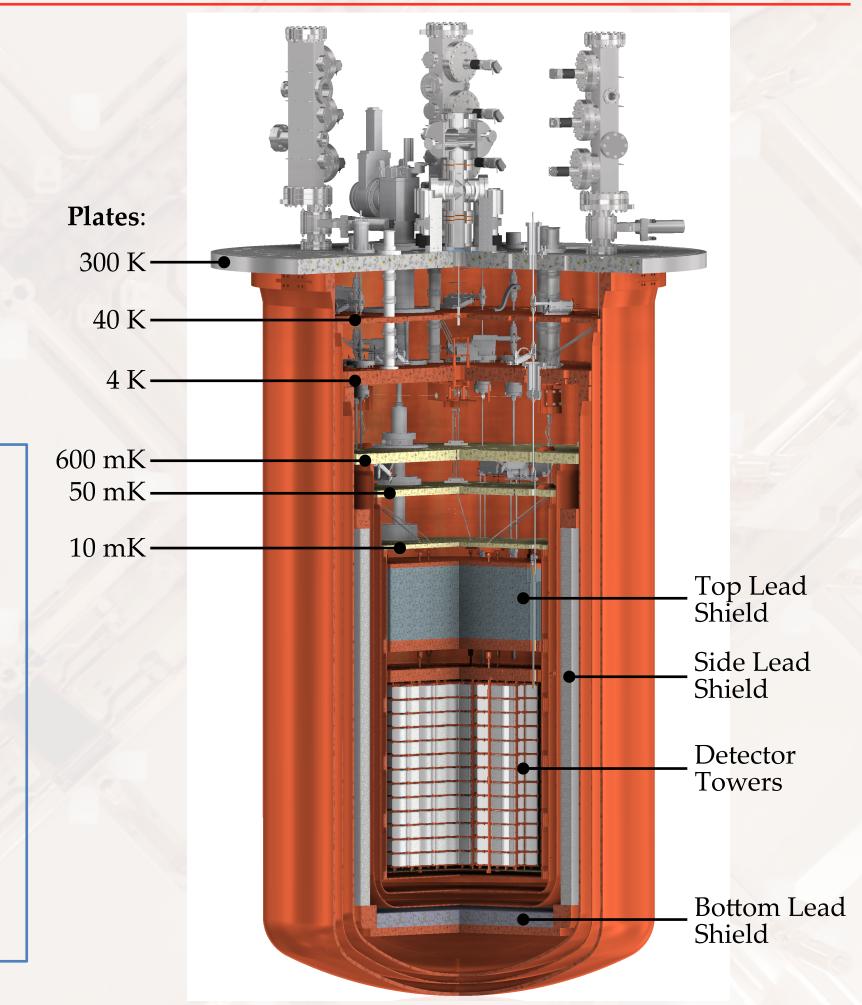
The CUORE cryostat

Challenges:

- Cool down ~1 ton detector to ~10 mK
- Mechanically decoupled for extremely low vibrations
- Low background environment

- Cryogen-free cryostat
- Fast Cooling System (⁴He gas) down to ~50K
- 5 pulse tubes cryocooler down to ~4K
- Dilution refrigerator down to operating temperature ~10 mK
- Nominal cooling power: 3 µW @ 10mK
- Cryostat total mass ~30 tons
- Mass to be cooled < 4K: ~15 tons
- Mass to be cooled < 50 mK: ~3 tons (Pb, Cu and TeO₂)

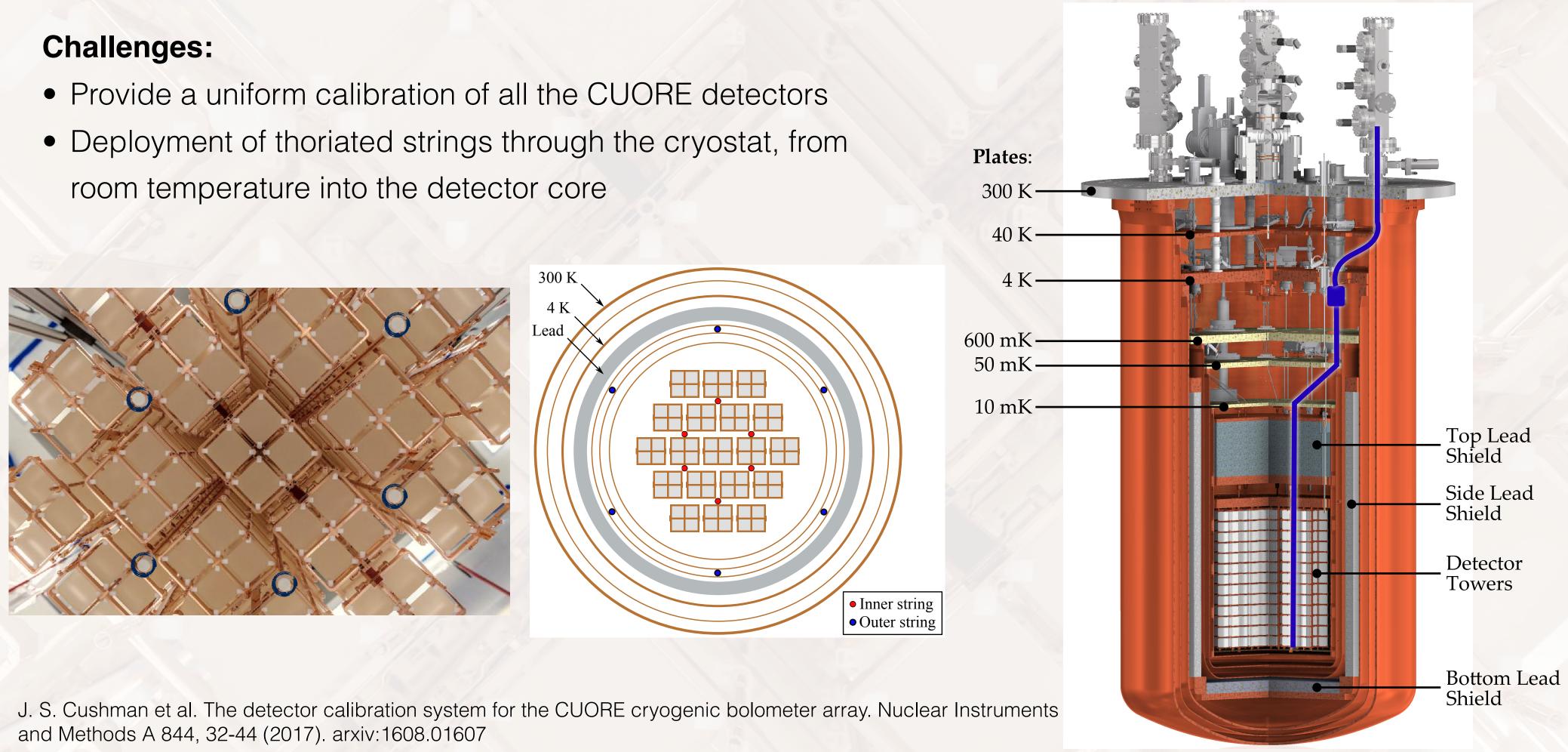




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Detector calibration system

- room temperature into the detector core



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Passive shielding

Challenges:

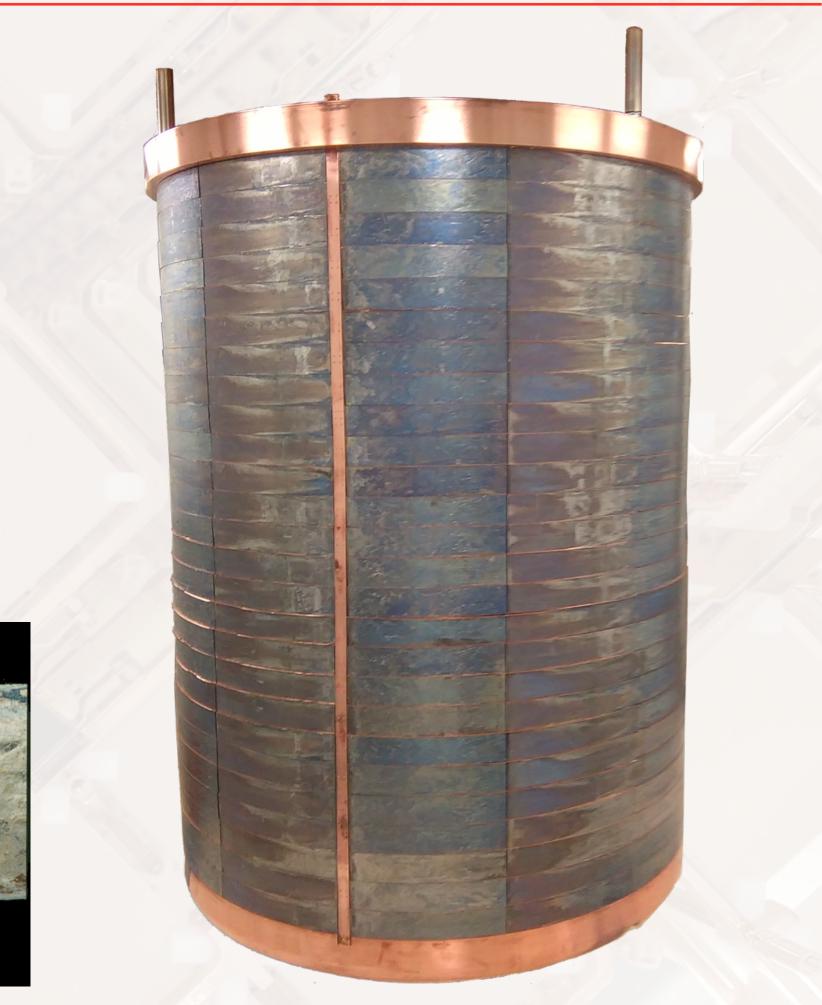
- Protect the detectors with a heavy shield against gamma and neutron activity from external sources (~70 tonnes lead + H₃BO₃)
- Select materials that don't contribute themselves to the background level (ancient roman lead and selected NOSV copper)
- Cool down inner layers of the shielding to the correct temperature (2.5 tonnes @ 50mK + 5.5 tonnes @ 4K)





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Cryogenic system commissioning

In February 2016 we completed the last test cool-down at full load:

- everything but the CUORE detector
- small test detector ("MiniTower")

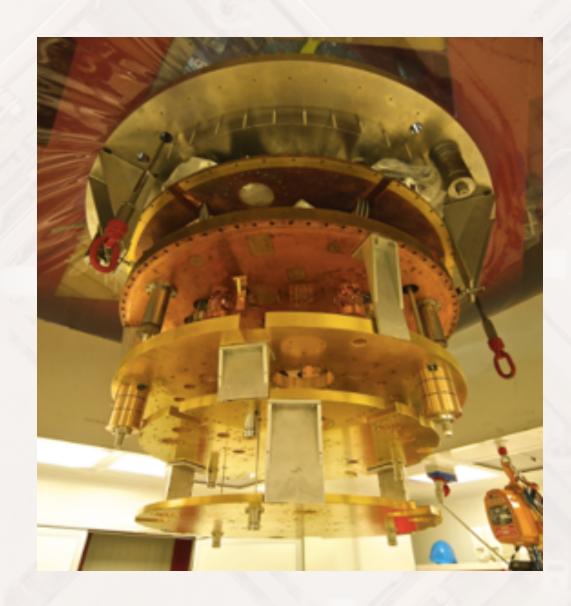
Excellent performance of the cryogenic system:

- base temperature below 7 mK
- stable operation

Important information on the noise sources and abatement

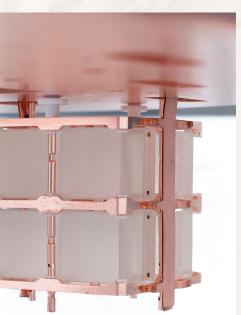
Successful deployment of the calibration sources at base temperature

Ready for the detector installation











Detector installation

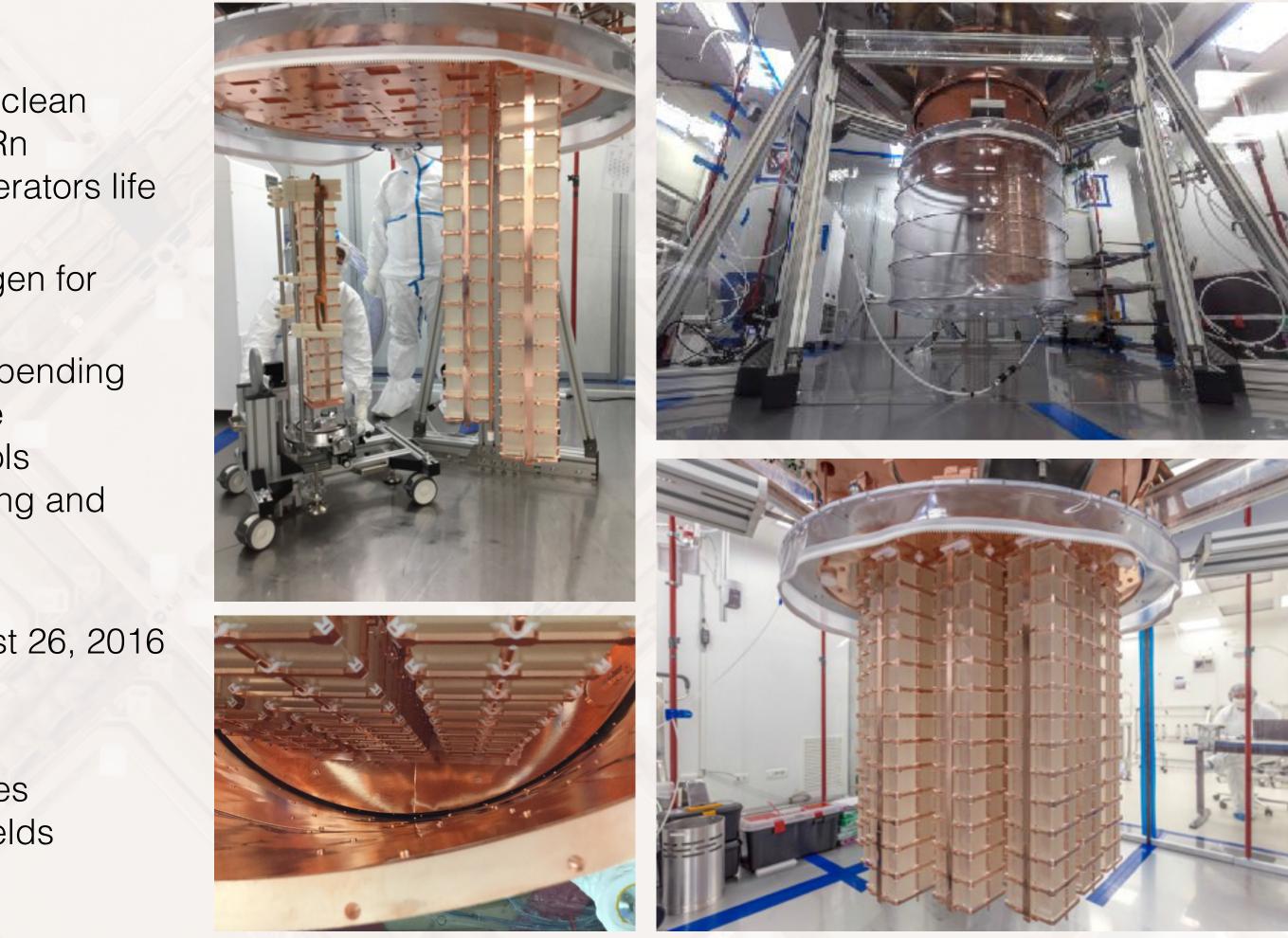
Performed in a radon-free environment:

- protected area inside the CUORE clean room flushed with radon-free air (Rn concentration < 0.1 Bq/m³) for operators life support
- protective bags flushed with nitrogen for overnight and emergency storage
- teams composed of 3 operators spending the minimum amount of time in the cleanroom, following strict protocols developed during months of training and test with mockup components

Towers installation completed on August 26, 2016

September-October 2016:

- installation of the cryostat interfaces (protective tiles) and radiation shields
- read-out tests



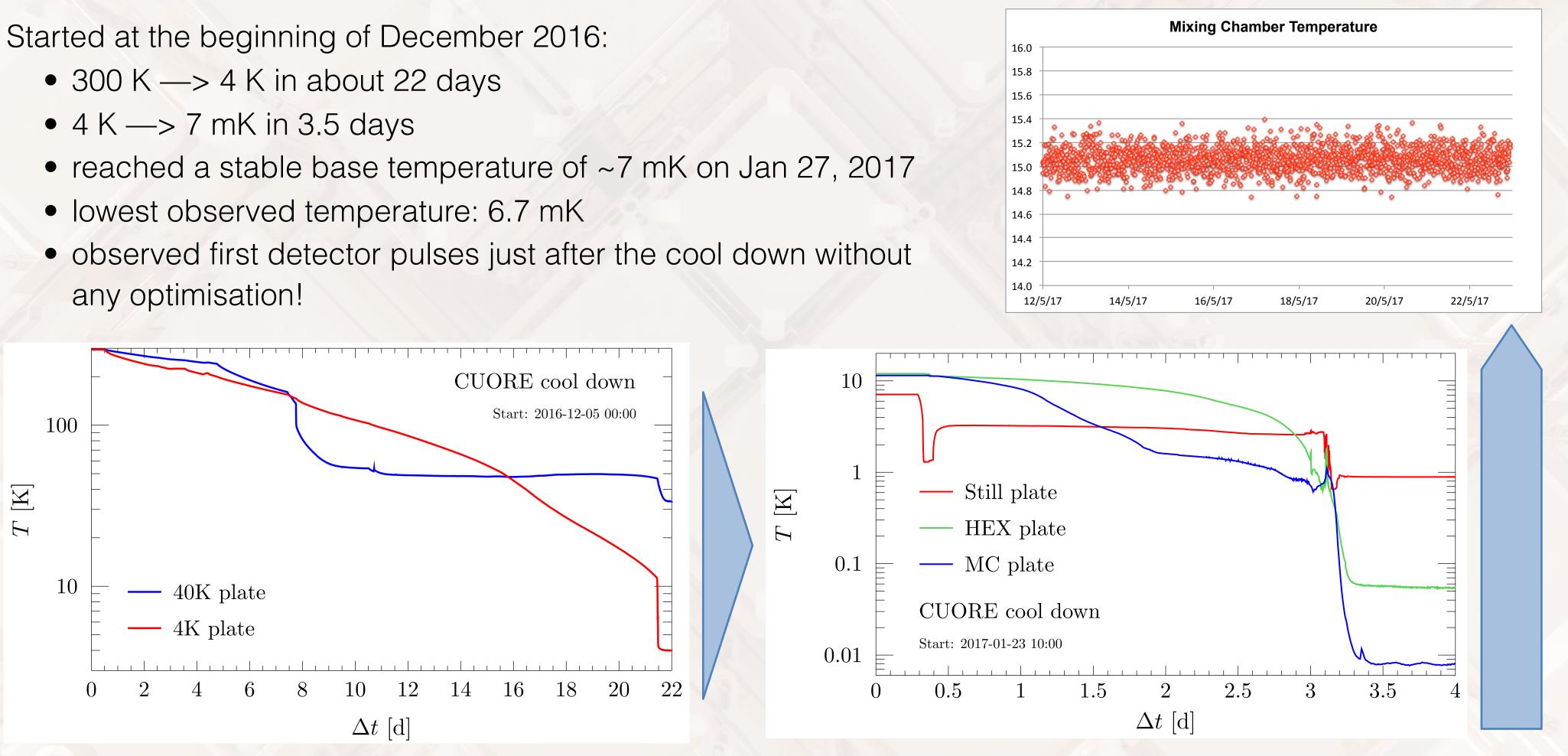




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Detector cool down

- any optimisation!



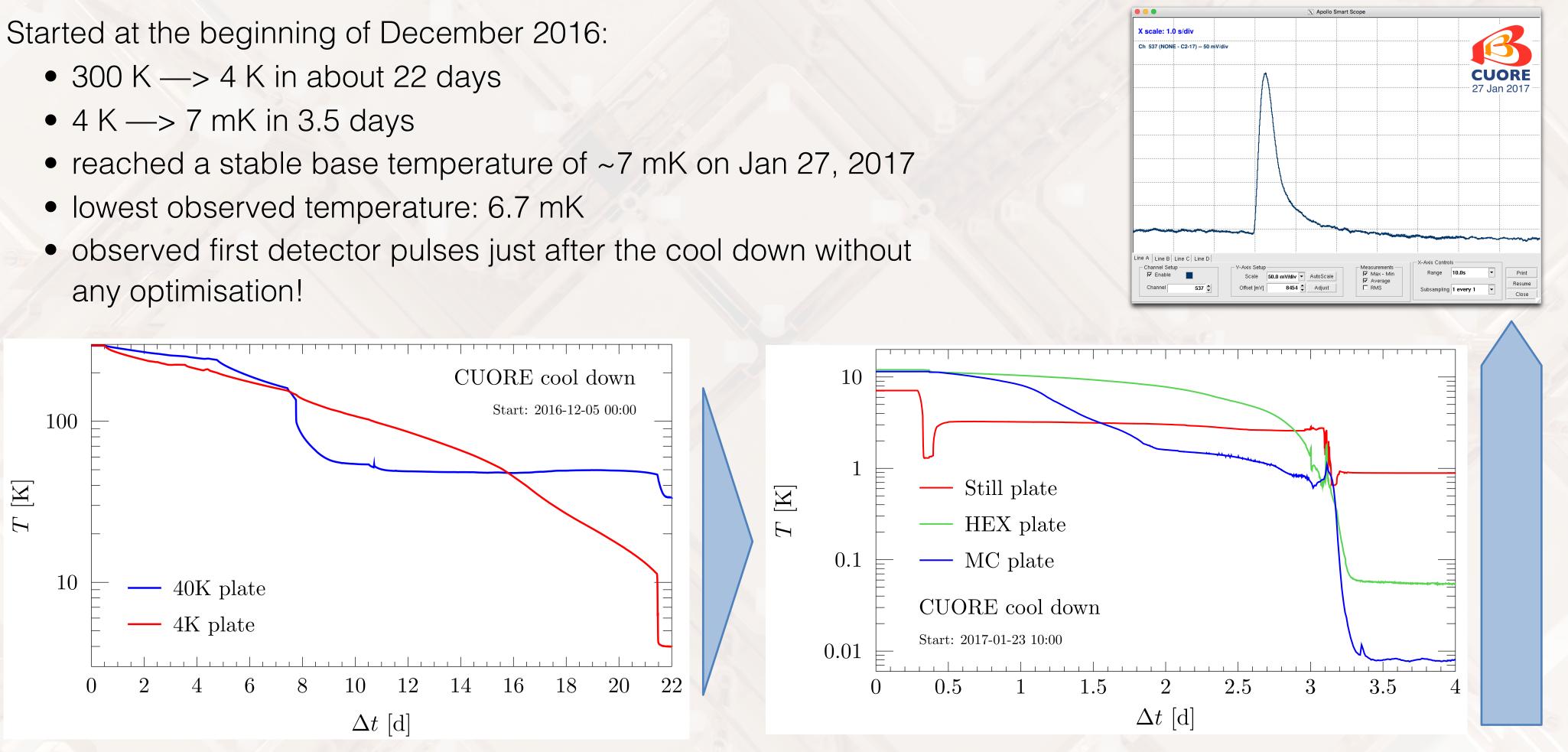
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Detector cool down

- any optimisation!



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Detector pre-operation

After the successful cool-down we faced the challenge to operate a thousand bolometers in a completely new system.

A long list of tests and activities

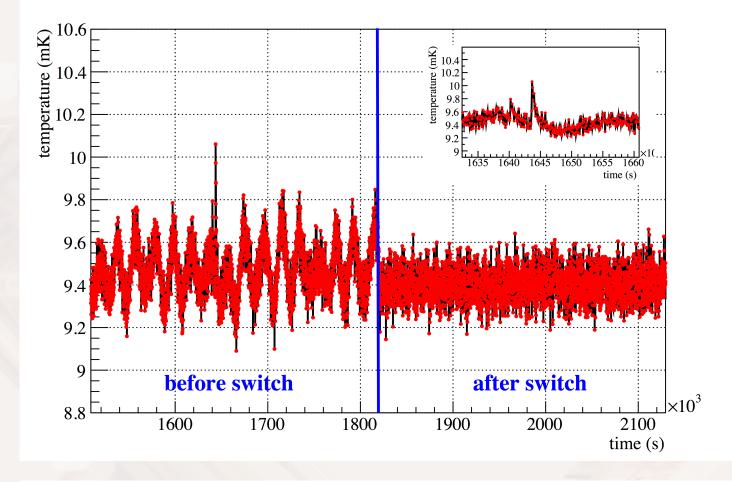
- DAQ and front-end electronics optimization
- Detector working points
 - Select representative subset
 - Load curves (to select optimal working points)
 - Temperature scan for the best operating conditions
- Noise reduction
- Linear drives to control the pulse tube (PT) motor-heads
- Monitor and control the relative phase shifts between different PT's using pressure sensors installed on the PT lines
 - Impressive results both in terms of temperature stabilisation and noise abatement

End of March 2017:

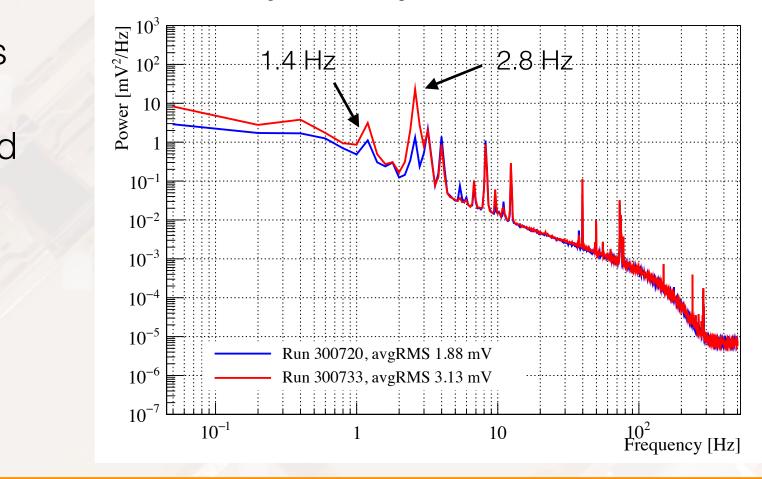
- Closed first optimisation phase
- Ready to start calibrations and science runs
- Selected working temperature: 15 mK

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Average Noise Power Spectrum: ch. 142 runs 300720, 300733



Science runs

Science operations started on April 14, 2017

- Very short commissioning run (identified issue with the thermistor bias on about 1/3 of the channels)
- First optimisation of the detector working point
- **Dataset 1**: 3 weeks of physics data bracketed by 2 calibration periods (May - June 2017)
- Second optimisation campaign
- Dataset 2: August September 2017

Operational performance:

- 984/988 operational channels
- Excellent data-taking efficiency when in operations
- Much improved detector stability, compared to Cuoricino/CUORE-0
- Calibrations/physics data ratio still to be optimised to maximise 0vßß sensitivity

May

June

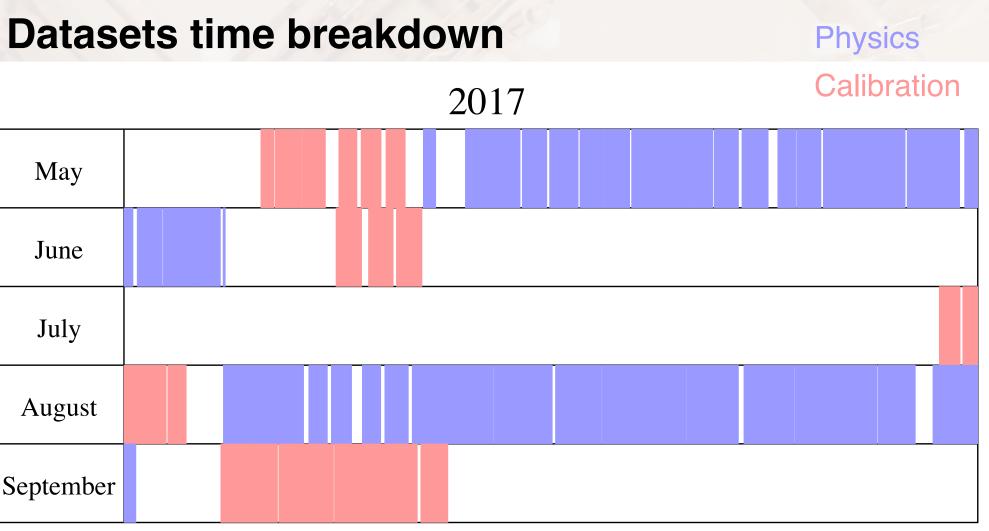
July

August

September

Acquired statistics used for this 0vDBD decay search (Dataset 1 + Dataset 2):

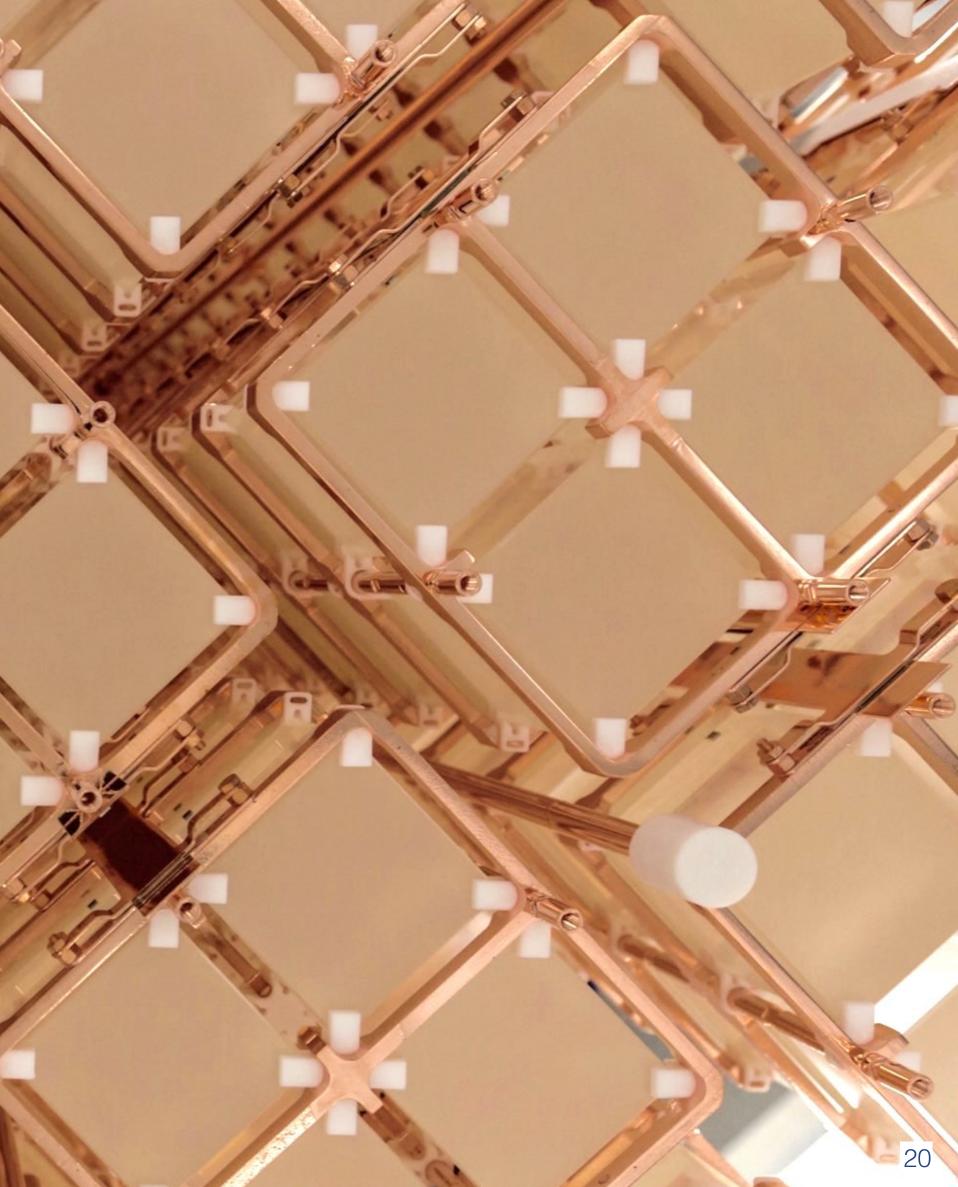




- natTeO₂ exposure: **86.3 kg yr** (37.6 + 48.7)
- ¹³⁰Te exposure: **24.0 kg yr**

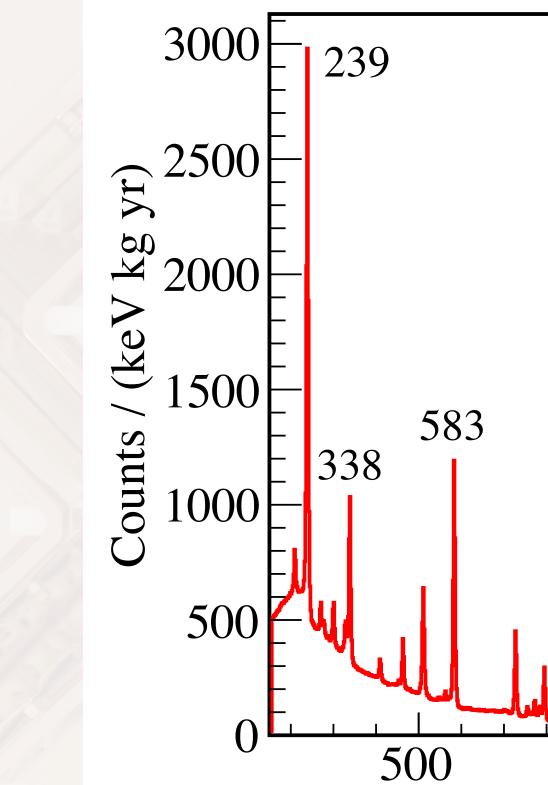
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Calibration spectrum

- Calibration strings deployed inside the CUORE detector
- Summed energy spectrum of all the CUORE detectorsdatasets
- Calibration data used for:
 - energy scale calibration
 - thermal gain stabilisation
 - detector response (line shape) study





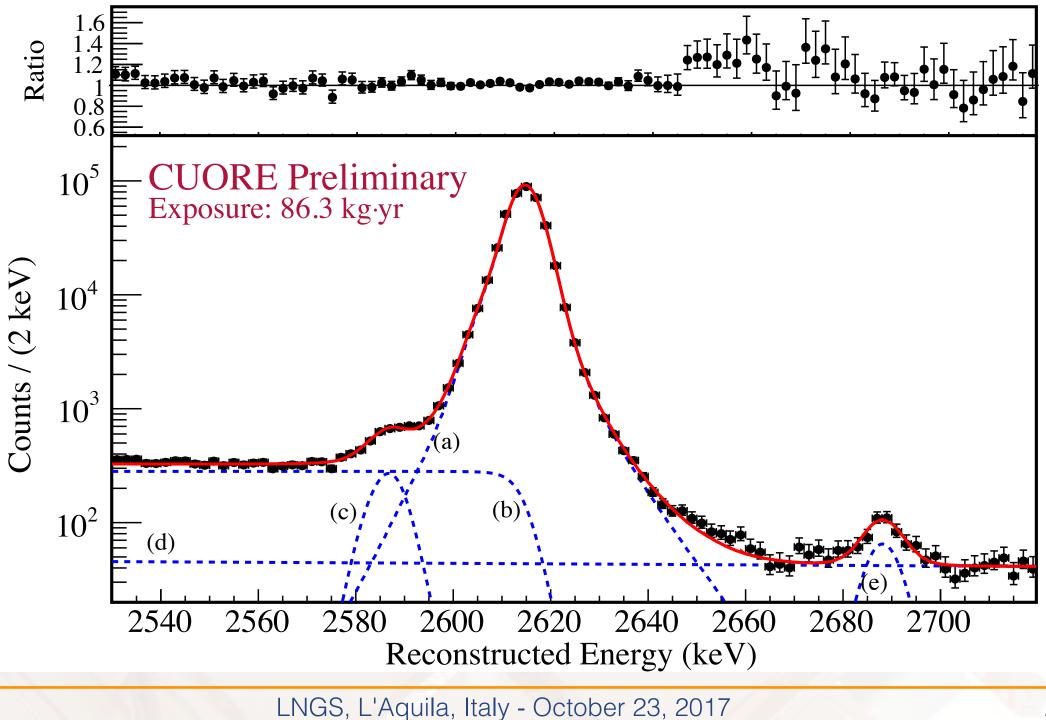
239 keV - ²¹²Pb 338, 911, 969 keV - ²²⁸Ac 583, 2615 keV - ²⁰⁸TI **CUORE** Preliminary Exposure: 86.3 kg·yr 911 969 2615 2500 1500 1000 2000 Reconstructed Energy (keV)

Detector performance: line shape

- The prominent ²⁰⁸TI line in the calibration spectrum was used to model the detector response to a monochromatic electron-like energy deposition
- 19 simultaneous fit on the data from all the channels of each tower, with some parameters (backgrounds) common to the whole tower and some defined channel-by-channel (resolutions, normalisations)
- Eventually only channel-dependent parameters (signal) will be used to build the PDF for the ROI fit
- Fit components:

(a) triple gaussian for the photopeak

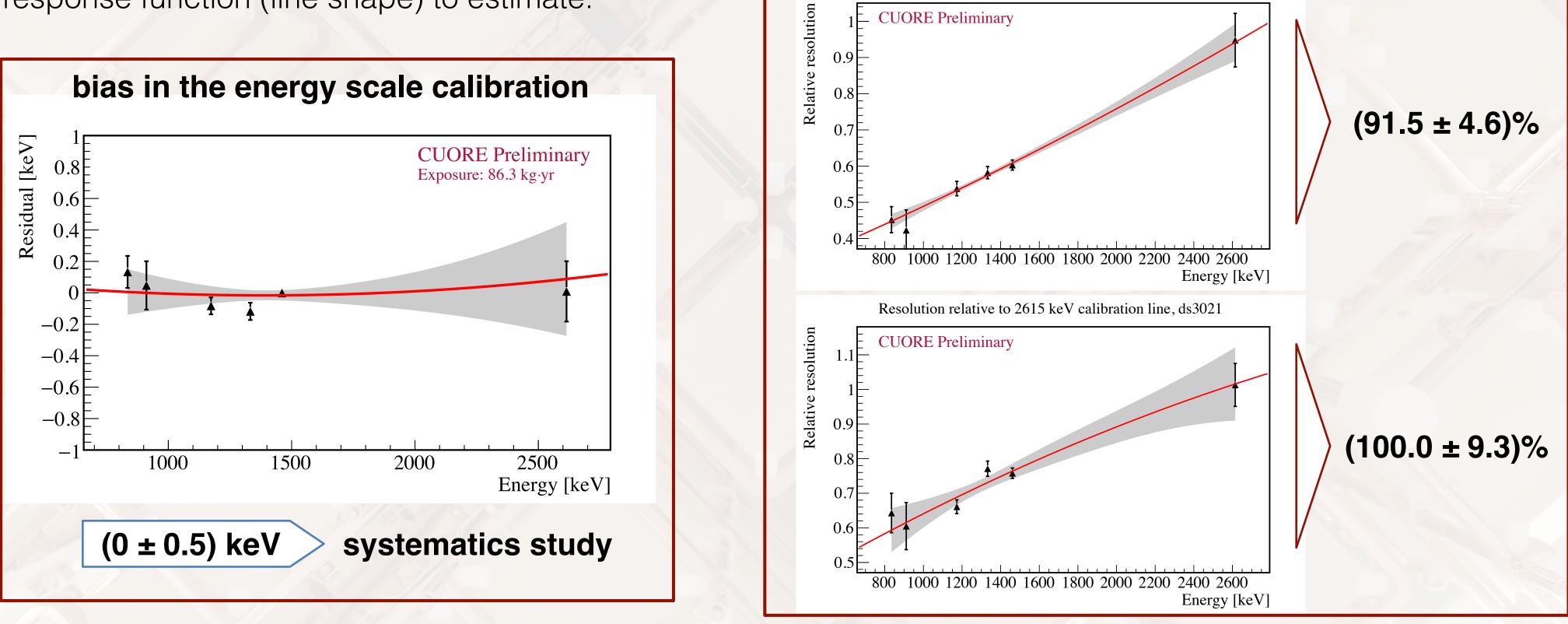
- (b) step-wise smeared multi-compton background
- (c) combination of gaussian X-rays escape lines
- (d) linear background
- (e) single gaussian line for the coincident absorption of 2615-keV and 583-keV followed by a single escape process





Energy scale and resolution scaling

The **gamma lines in the background spectrum** have been fitted with the complete detector response function (line shape) to estimate:



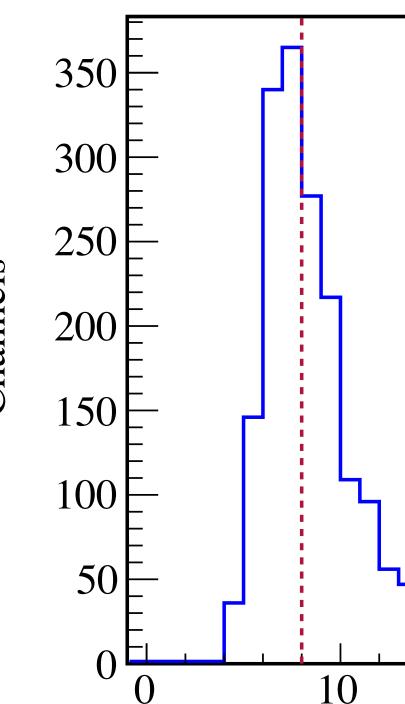


dependence of resolution from energy to define a scaling factor @ Q-value Resolution relative to 2615 keV calibration line, ds3018

Energy resolution - Calibration runs

A total of **1811** (92% of live channels) channels-dataset couples were used in this analysis; discarded channels had poor line or pulse shapes, or the energy couldn't be reconstructed accurately

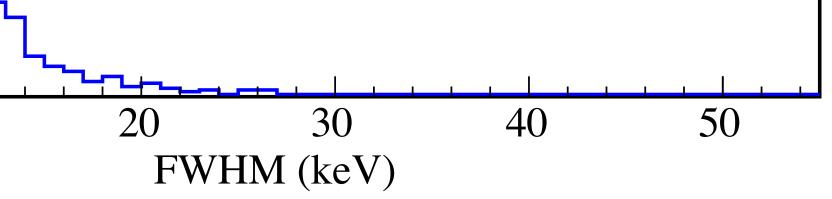
Channels @ 2615 keV ds3018: 9.0 keV FWHM ds3021: 7.4 keV FWHM effective (exposure-weighted): 8.0 keV FWHM





Calibration resolution at 2615 keV

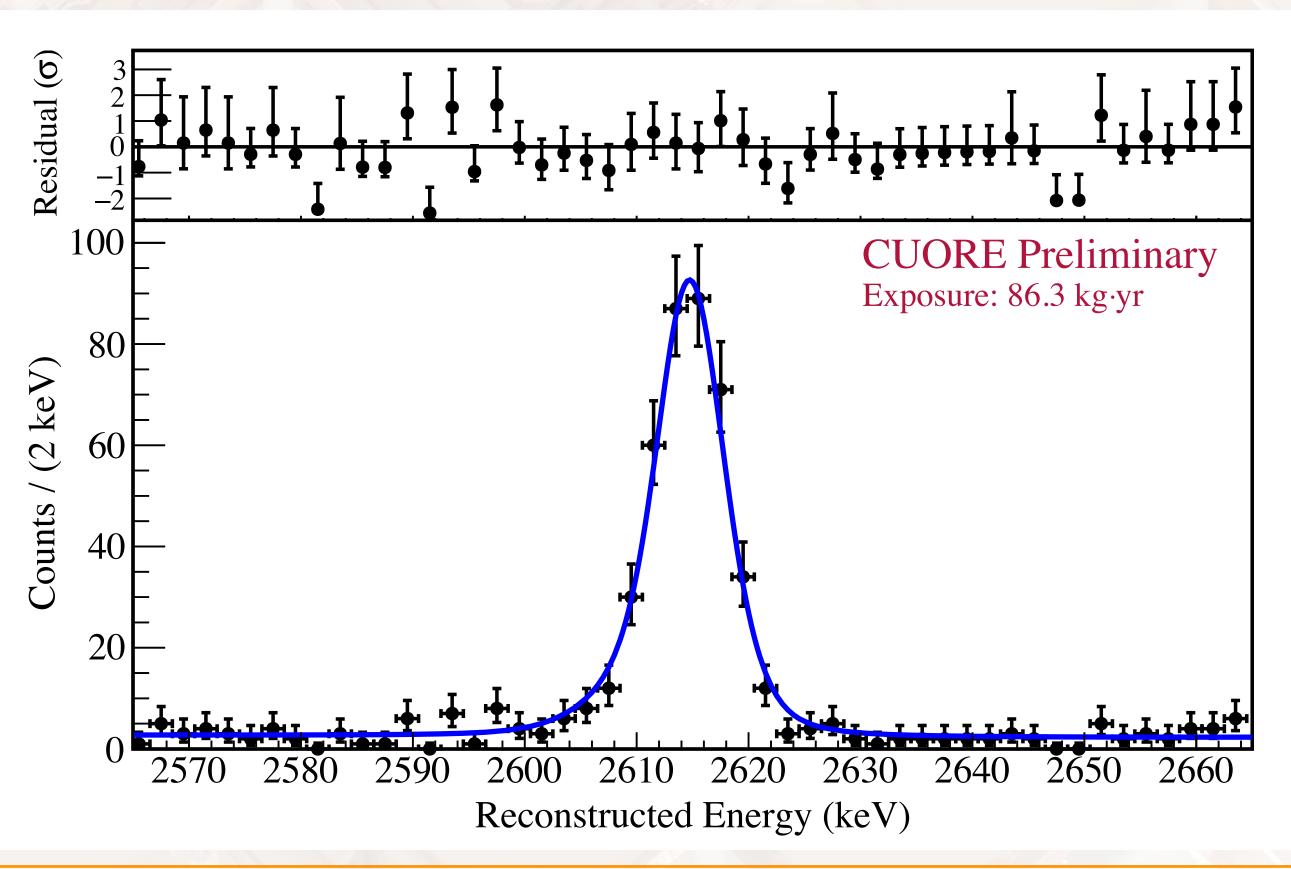
CUORE Preliminary Exposure: 86.3 kg·yr



Energy resolution - Physics runs

A total of **1811** (92% of live channels) channels-dataset couples were used in this analysis; discarded channels had poor line or pulse shapes, or the energy couldn't be reconstructed accurately

@ Q-value ds3018: (8.3 ± 0.4) keV FWHM ds3021: (7.4 ± 0.7) keV FWHM effective (exposure-weighted): (7.7 ± 0.5) keV FWHM







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Analysis procedure

- Acquisition of continuous waveforms
- Triggering
- Data preprocessing: estimation of raw parameters
- Pulse filtering with Optimum Filter
- Thermal Gain Stabilization (TGS): calibration and heater-based
- Energy calibration and best energy estimator selection
- Particle event selection Pulse Shape Analysis
- Coincidence analysis w/ detector response synchronisation and software threshold @ 150 keV (to prevent any spectral shape distortion due to threshold effects in the ROI)
- Energy spectrum

Very similar to what was developed and used for CUORE-0 (Phys. Rev. C 93, 045503 (2016))



Event selection: efficiencies

Event selection occurs after periods of low-quality data (~1% of the total live time) are removed.

- Trigger: compare the number of pulses generated by injection of power through a heater (pulser events) that were triggered as normal events
- Energy reconstruction: number of pulser-generated events reconstructed within 3 sigma from the average of their energy distribution
- Base cuts: computed on pulser events
- Anti-coincidence: calculated on ⁴⁰K line @ 1460 keV (no physical coincidences expected) with side-bands subtraction
- Pulse shape analysis: calculated on ²⁰⁸TI line @ 2615 keV (not used for optimisation) with side-bands subtraction
- Containment efficiency: calculated with Monte Carlo simulation of the detector geometry



Event selection: efficiencies

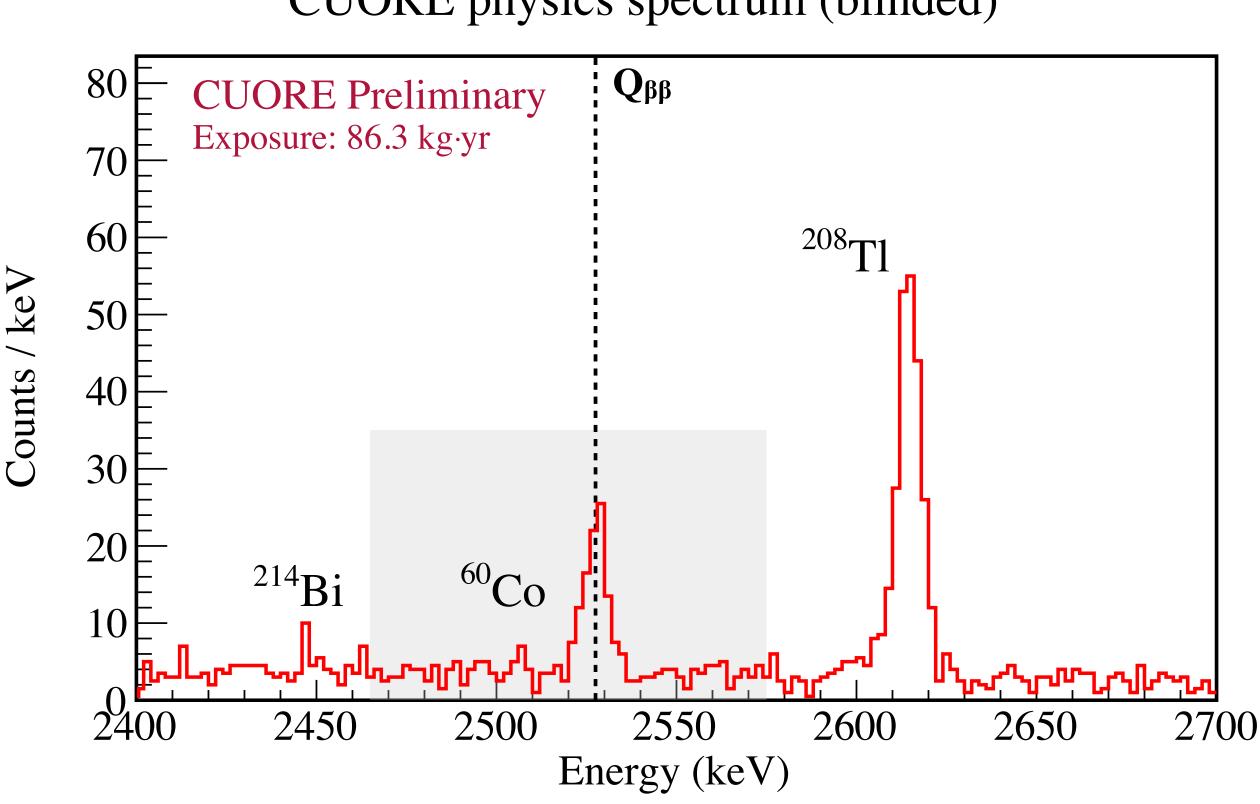
Event selection occurs after periods of low-quality data (~1% of the total live time) are removed.

Dataset 1	Dataset 2
(99.766 ± 0.003) %	(99.735 ± 0.004) %
(99.168 ± 0.006) %	(99.218 ± 0.006) %
(95.63 ± 0.01) %	(96.69 ± 0.01) %
(99.4 ± 0.5) %	(100.0 ± 0.4) %
(91.1 ± 3.6) %	(98.2 ± 3.0) %
(85.7 ± 3.4) %	(94.0 ± 2.9) %
(88.35 ± 0.09) %	
(75.7 ± 3.0) %	(83.0 ± 2.6) %
	$(99.766 \pm 0.003) \%$ $(99.168 \pm 0.006) \%$ $(95.63 \pm 0.01) \%$ $(99.4 \pm 0.5) \%$ $(91.1 \pm 3.6) \%$ $(85.7 \pm 3.4) \%$ $(88.35 \pm 0.01) \%$



Blinding procedure

- To blind our data we randomly move a fraction of events from +/- 20 keV of 2615 keV to the Q-value and vice versa
- The blinding algorithm produces an artificial peak around the 0vDBD Qvalue hiding the real 0vDBD rate of ¹³⁰Te
- This method of blinding the data preserves the integrity of the possible **OvDBD** events while maintaining the spectral characteristics with measured energy resolution and introducing no discontinuities in the spectrum
- When all data analysis procedures are fixed the data are eventually unblinded

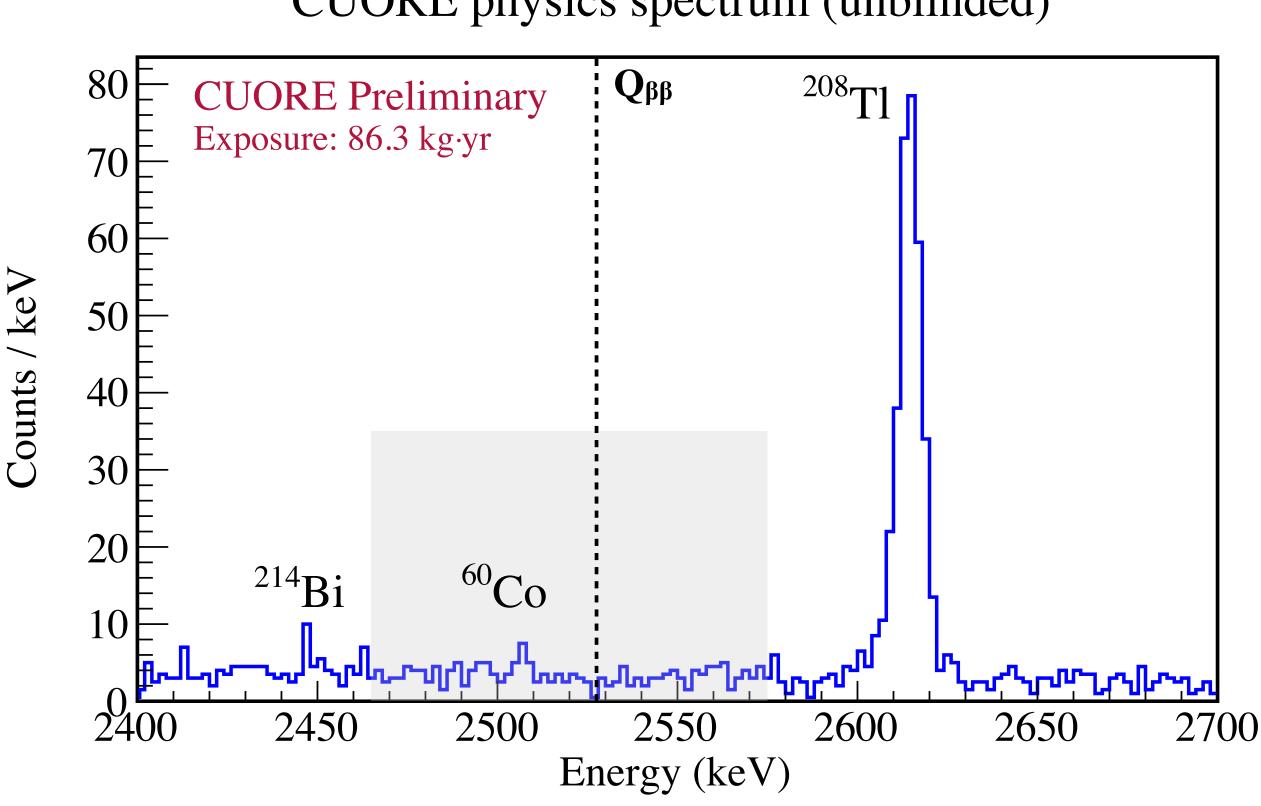




CUORE physics spectrum (blinded)

Blinding procedure

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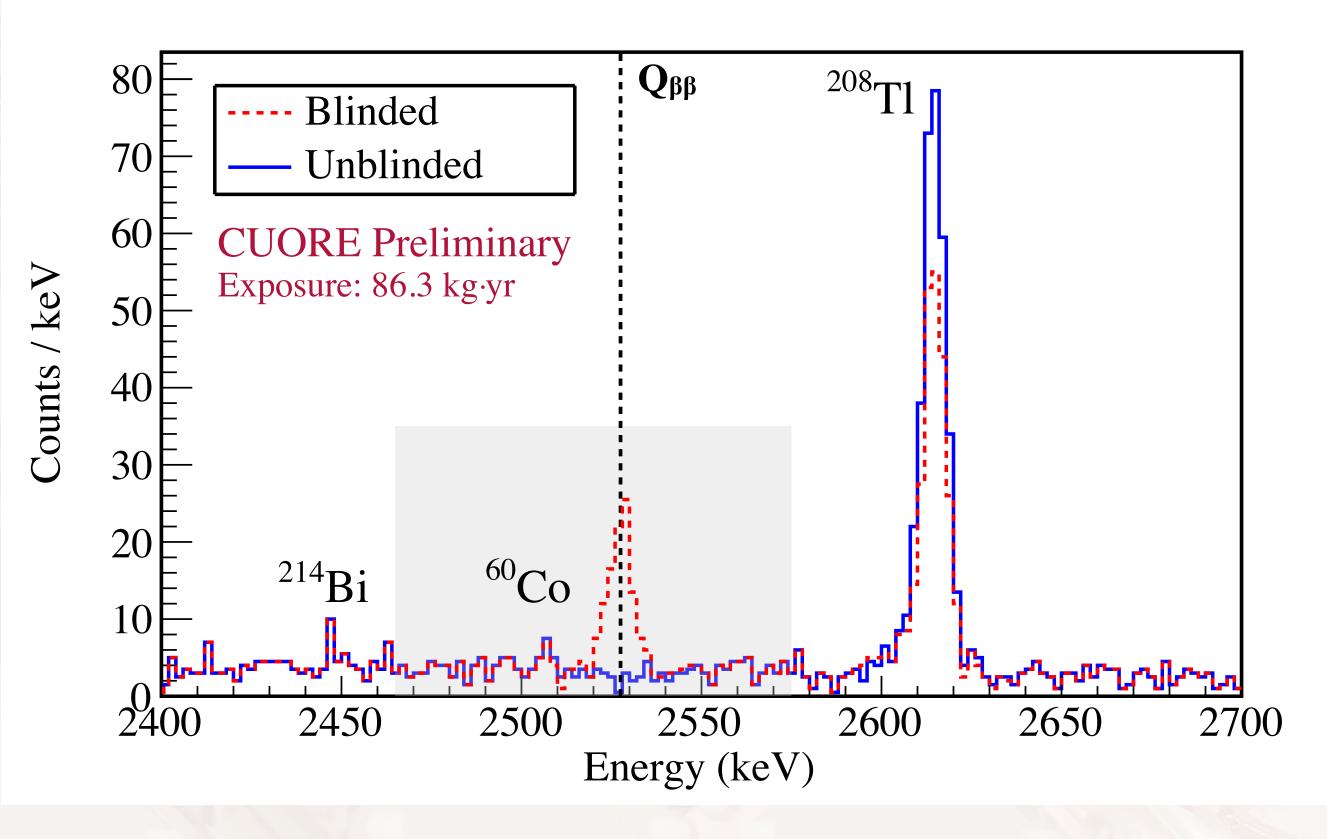




CUORE physics spectrum (unblinded)

Blinding procedure

- To blind our data we randomly move a fraction of events from +/- 20 keV of 2615 keV to the Q-value and vice versa
- The blinding algorithm produces an artificial peak around the 0vDBD Q-value hiding the real 0vDBD rate of ¹³⁰Te
- This method of blinding the data preserves the integrity of the possible 0vDBD events while maintaining the spectral characteristics with measured energy resolution and introducing no discontinuities in the spectrum
- When all data analysis procedures are fixed the data are eventually unblinded





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Event selection

-2000 --3000 --34000 --34000 --34000 --40000 --40000 --40000 --40000 -Dataset: 3021 Channel: 860 Tower: 17 Energy: 2471.7 Dataset: 3021 Channel: 544 Tower: 11 Energy: 2516.2 Dataset: 3021 Channel: 193 Tower: 4 Energy: 2469.9 Dataset: 3021 Channel: 548 Tower: 11 Energy: 2508.7 Dataset: 3018 Channel: 283 Tower: 6 Energy: 2505.1 -45000 -55000 -65000 -65000 15000 -4000 --4000 --4000 --4000 --4000 --43300 Dataset: 3021 Channel: 4 Tower: 1 Energy: 2481.9 Dataset: 3018 Channel: 319 Tower: 7 Energy: 2488.6 Dataset: 3021 Channel: 126 Tower: 3 Energy: 2475.5 -2000 -2000 -2000 -34000 -34000 -34000 -34000 -34000 -40000 -40000 Dataset: 3018 Channel: 441 Tower: 9 Energy: 2505.5 Dataset: 3021 Channel: 904 Tower: 18 Energy: 2565.1 -10000 -15000 -25000 -30000 Dataset: 3021 Channel: 50 Tower: 1 Energy: 2520.0 Dataset: 3021 Channel: 729 Tower: 15 Energy: 2571.0 Dataset: 3018 Channel: 748 Tower: 15 Energy: 2493.5 Dataset: 3021 Channel: 98 Tower: 2 Energy: 2531.5 Dataset: 3021 Channel: 463 Tower: 9 Energy: 2544.6 -49000 -49000 -50000 -89000 -49000 -25000 -25000 -30000 -40000 -45000 -60000 -4000 -4000 -5000 -5000 -6000 Detaset: 3021 Channel: 153 Tower: 3 Energy: 2493.6 Dataset: 3021 Channel: 468 Tower: 9 Energy: 2550.2 Dataset: 3021 Channel: 635 Tower: 13 Energy: 2531.6 Dataset: 3018 Channel: 527 Tower: 11 Energy: 2506.7 Dataset: 3021 Channel: 674 Tower: 13 Energy: 2552.2 Dataset: 3021 Channel: 412 Tower: 8 Energy: 2492.2 Dataset: 3018 Channel: 59 Tower: 2 Energy: 2588.4 Dataset: 3021 Channel: 968 Tower: 19 Energy: 2513.2 Dataset: 3021 Channel: 243 Tower: 5 Energy: 2555.9 Detaset: 3021 Channel: 149 Tower: 3 Energy: 2503.9 -60000 -45000 -30000 --35000 --40000 --50000 --50000 -Dataset: 3021 Channet: 26 Tower: 1 Energy: 2472.2 Dataset: 3021 Channel: 547 Tower: 11 Energy: 2516.6 Dataset: 3018 Channel: 251 Tower: 5 Energy: 2521.0 Dataset: 3018 Channel: 640 Tower: 13 Energy: 2489.7 Dataset: 3018 Channel: 586 Tower: 12 Energy: 2502.5 -29000 -39000 -49000 -49000 -59000 -50000 -55000 -66000 -70000 -75000 -40000 -35900 -40900 -45900 -59900 -6000 Dataset: 3021 Channel: 361 Tower: 7 Energy: 2517.2 Dataset: 3018 Channel: 521 Tower: 11 Energy: 2483.1 Dataset: 3021 Channel: 100 Tower: 2 Energy: 2493.6 Dataset: 3021 Channel: 138 Tower: 3 Energy: 2496.8 -60000 Dataset: 3021 Channel: 891 Tower: 18 Energy: 2479.0 -40000 -50000 -55000 -60000 -29000 -29000 -29000 Dataset: 3018 Channel: 907 Tower: 18 Energy: 2567.3 Detaset: 3018 Channel: 95 Tower: 2 Energy: 2507.2 -5000 --4000 --4000 --6000 --6000 --6000 --6000 -Dataset: 3018 Channel: 607 Tower: 12 Energy: 2534.9 Dataset: 3021 Channet: 62 Tower: 2 Energy: 2537.6 Dataset: 3018 Channel: 811 Tower: 16 Energy: 2477.6 - 10000 - 15000 - 29000 -15000 -20000 -25000 -30000 Dataset: 3021 Channel: 981 Tower: 19 Energy: 2474. Dataset: 3021 Channel: 805 Tower: 16 Energy: 2538.3 Dataset: 3018 Channel: 768 Tower: 15 Energy: 2476.1 Dataset: 3018 Channel: 942 Tower: 19 Energy: 2506.6 Dataset: 3018 Channel: 775 Tower: 15 Energy: 2466.5 25000 20000 15000 10000 5000 -60000 -10000 -15000 -20000 -25000 Dataset: 3018 Channel: 892 Tower: 18 Energy: 2491.6 Detaset: 3021 Channel: 684 Tower: 14 Energy: 2553.5 Dataset: 3018 Channel: 50 Tower: 1 Energy: 2572.4 Dataset: 3018 Channel: 191 Tower: 4 Energy: 2477.1 Dataset: 3018 Channel: 703 Tower: 14 Energy: 2507.6 -10000 -15000 -20000 Dataset: 3018 Channel: 62 Tower: 2 Energy: 2561.5 Dataset: 3018 Channel: 409 Tower: 8 Energy: 2544.0 Dataset: 3021 Channel: 717 Tower: 14 Energy: 2496.1 Dataset: 3021 Channel: 479 Tower: 10 Energy: 2505. Dataset: 3021 Channet: 322 Tower: 7 Energy: 2568.3 -14000 -18000 -2000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -20000 -2000 -20000 -200 -19300 -10000 -25000 -40000 -40000 -50000 -60000 -60000 -60000 Dataset: 3021 Channel: 614 Tower: 12 Energy: 2480. Dataset: 3021 Channet: 863 Tower: 17 Energy: 2543.8 Detaset: 3021 Channel: 134 Tower: 3 Energy: 2518.7 Dataset: 3021 Channel: 338 Tower: 7 Energy: 2560.3 Dataset: 3021 Channel: 608 Tower: 12 Energy: 2544.8 -50000-

Eventually 155 (65+90) events survive the selection cuts in the unblinded region of interest

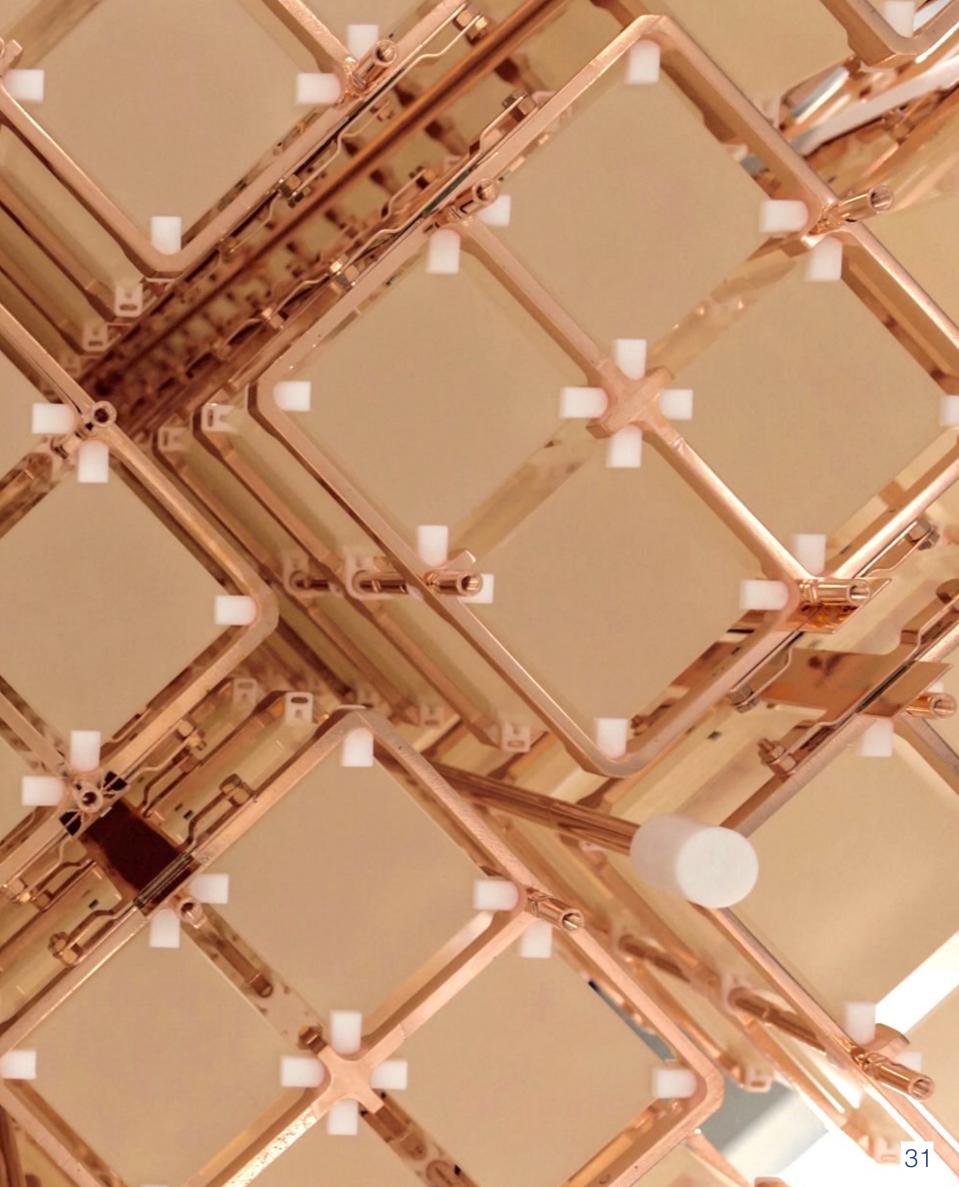
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Fit in the ROI

- Simultaneous UEML (Unbinned Extended Maximum Likelihood) fit in the energy region 2465-2575 keV
- The fit has 3 components:
 - 1. posited peak at the Q-value of 130Te:
 - energy scale defined relative to the ²⁰⁸TI line in calibration data to account for residual mis-calibration between channels
 - signal normalisation common to all detectors-datasets (1 free parameter)

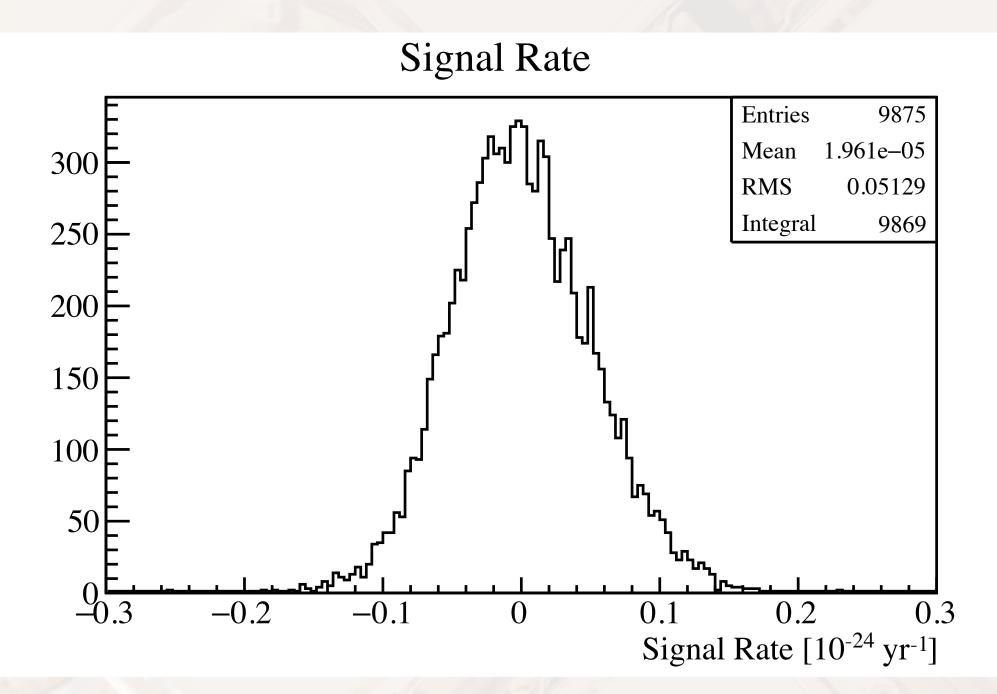
2. floating peak to account for the 60Co sum gamma line (2505 keV):

- energy scale defined relative to the ²⁰⁸TI line in calibration data to account for residual mis-calibration between channels
- rate common to all detectors-dataset, with a correction accounting for the time elapsed between the two datasets (1 free parameter)
- 3. flat background, attributed to multi scatter Compton events from ²⁰⁸Tl and surface alpha events:
 - common to all detectors in a single dataset, two independent parameters for the two datasets to account for differences in the background rejection efficiency (2 free parameters)
- The peaks in each channel-dataset are fitted with its own line shape (fixed from calibration data)



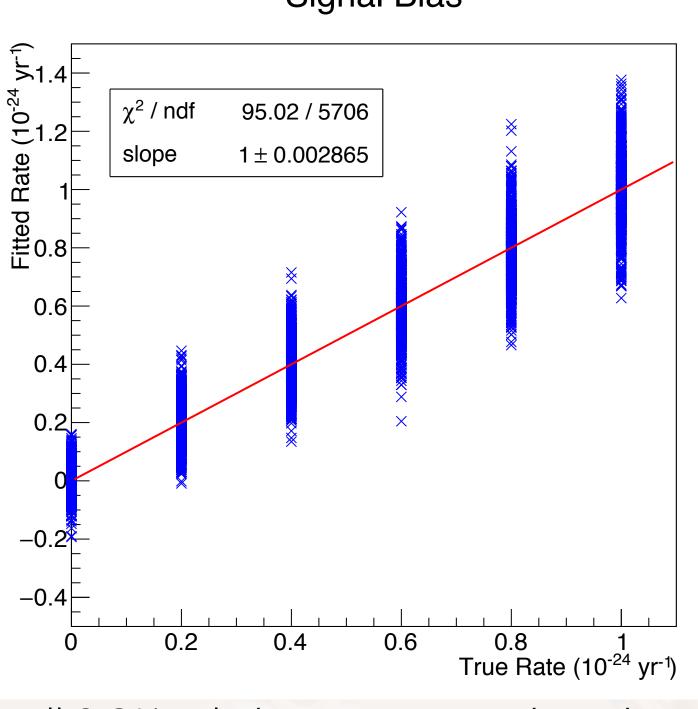
Bias and systematic errors

Any bias in the fitting procedure is searched for. A large number of pseudo-experiments is generated starting from a model where the signal rate Γ^{0v} is changed from 0 to a large value. Each pseudo-experiment is fitted with the full model (H₁), and the distribution of the fitted signal rates are examined.



NO EVIDENCE OF ABSOLUTE BIAS is found, and a small 0.3% relative component is estimated





Signal Bias

LNGS, L'Aquila, Italy - October 23, 2017

Bias and systematic errors

Uncertain parameters of the model that can affect the result are considered as potential source of systematic errors, and studied as nuisance parameters.

For each parameter θ_i in the model H₁:

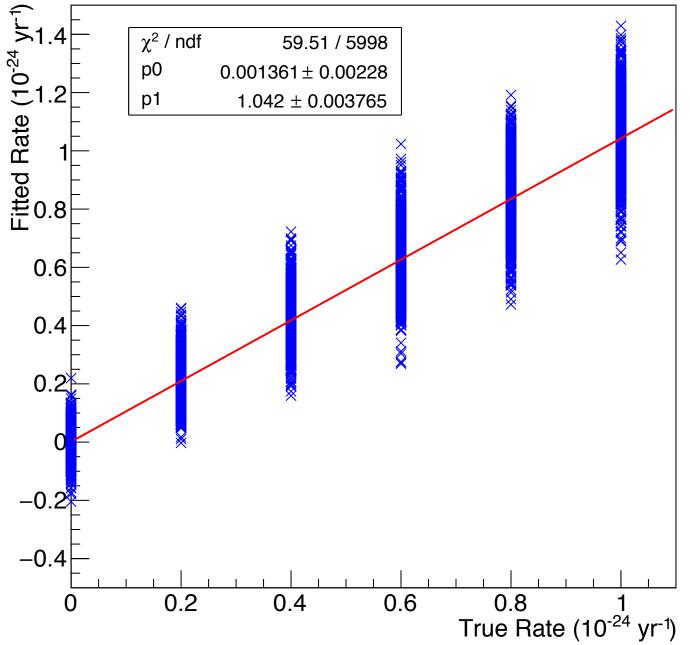
- take H₁ best fit and change by $+(-)1\sigma$ (worst case for the line shape) the value of θi
- generate a large number of pseudoexperiments from the obtained PDF
- fit each pseudo-experiment with H₁ with the original value of θ_i and extract Γ^{0v}_{fit}
- repeat the procedure with different values of the signal Γ^{Ov}_{true} , from 0 to a large value
- plot Γ^{0v}_{fit} vs Γ^{0v}_{true}
- parametrise the bias as p0 (additive) +p1. Γ^{0v} (scaling)

Rate Fitted F 0.6 0.4 0.2 -0.2-0.4

Only deviations from 0 in p0 and from 1 in p1 that are statistically significant are considered as systematic errors and added in quadrature and propagated to the limit







Systematic errors

The following nuisance parameters are considered:

- energy resolution (higher and lower by 1σ)
- Q-value (higher and lower by 0.5 keV from energy scale uncertainty)
- no sub-peak in the detector response (simple gaussian lineshape)
- linear background (higher and lower by 1σ)

The systematic error associated to efficiency is computed directly from the statistical uncertainty on the efficiency

Systematic	Absolute uncertainty [10-24 yr]	Relative uncertainty
Resolution	_	1.5%
Q-value location	_	0.2%
No subpeaks	0.002	2.4%
Efficiency	_	2.4%
Linear fit	0.005	0.8%

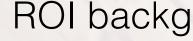


Region of interest: 2465 to 2575 keV

Overall signal efficiency: (75.7 ± 3.0)% - ds3018

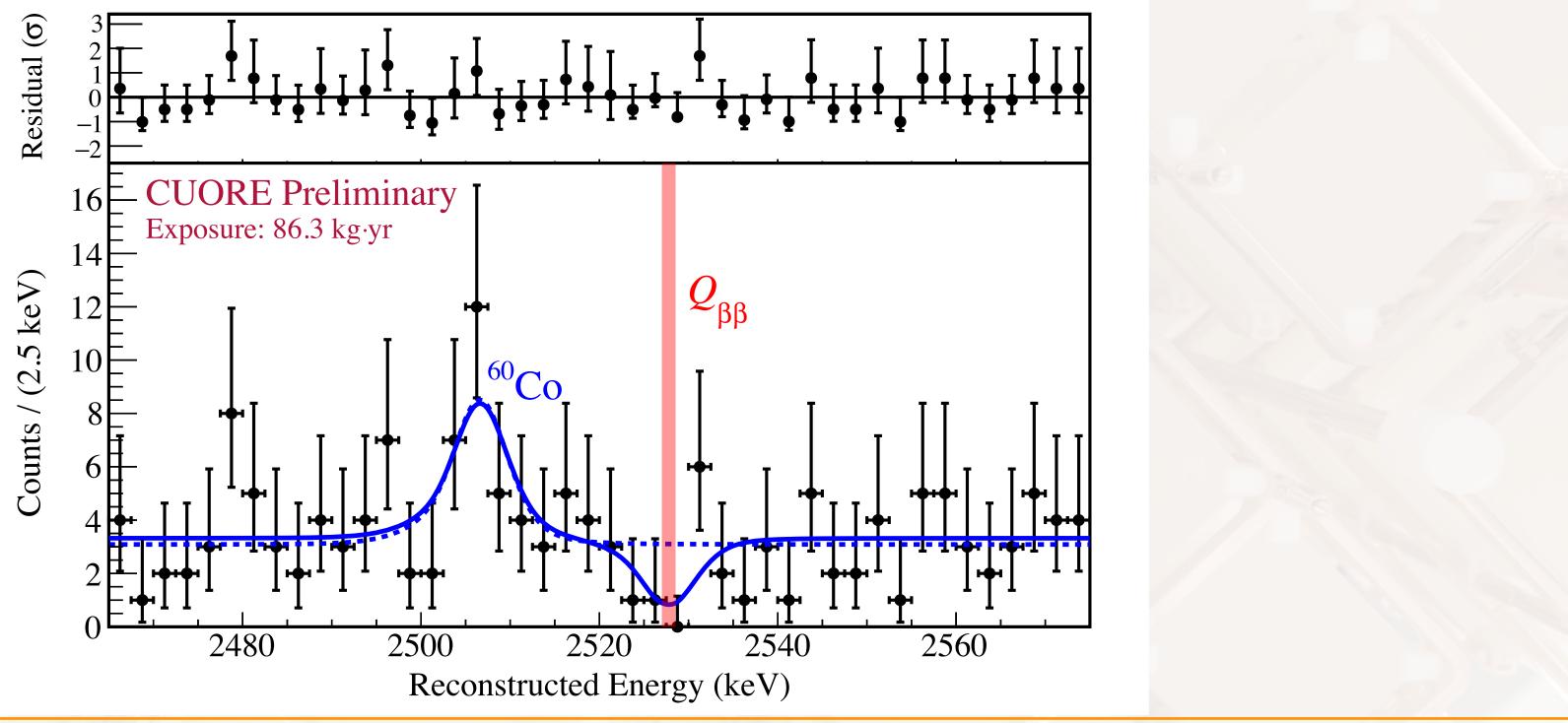
 $(83.0 \pm 2.6)\% - ds3021$

Events in the region of interest: **155**



Best fit for 60Co mean: (2506.4 ± 1.2) keV

Best fit decay rate: (-1.0_{-0.3}+0.4 (stat.) ± 0.1 (syst.))×10⁻²⁵ / yr





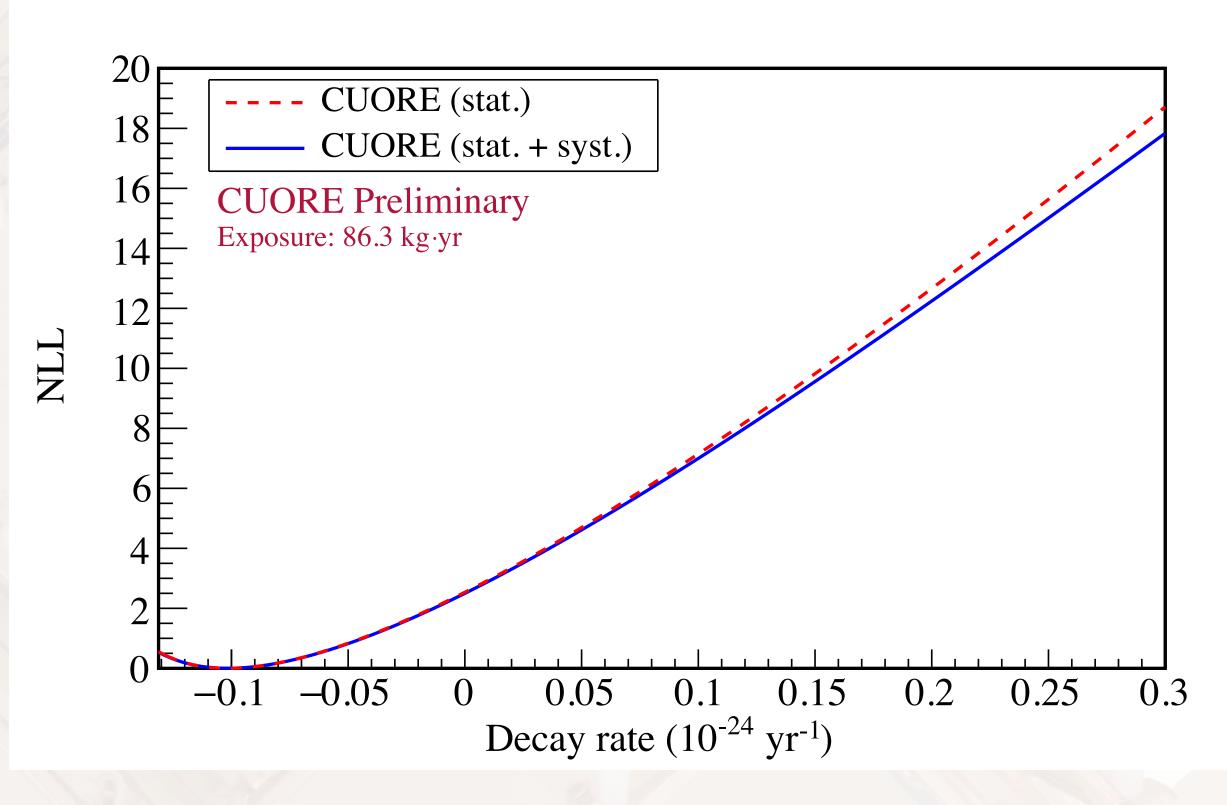
ROI background index: (1.49-0.17+0.18) × 10-2 c/(keV·kg·yr) $(1.35_{-0.18}^{+0.20}) \times 10^{-2} \text{ c/(keV \cdot kg \cdot yr)}$

Best fit decay rate:

$(-1.0_{-0.3}^{+0.4} \text{ (stat.)} \pm 0.1 \text{ (syst.)})\times 10^{-25} / \text{ yr}$

No evidence of signal

Limit calculation Profile likelihood integrated on the physical region ($\Gamma^{0v} > 0$)





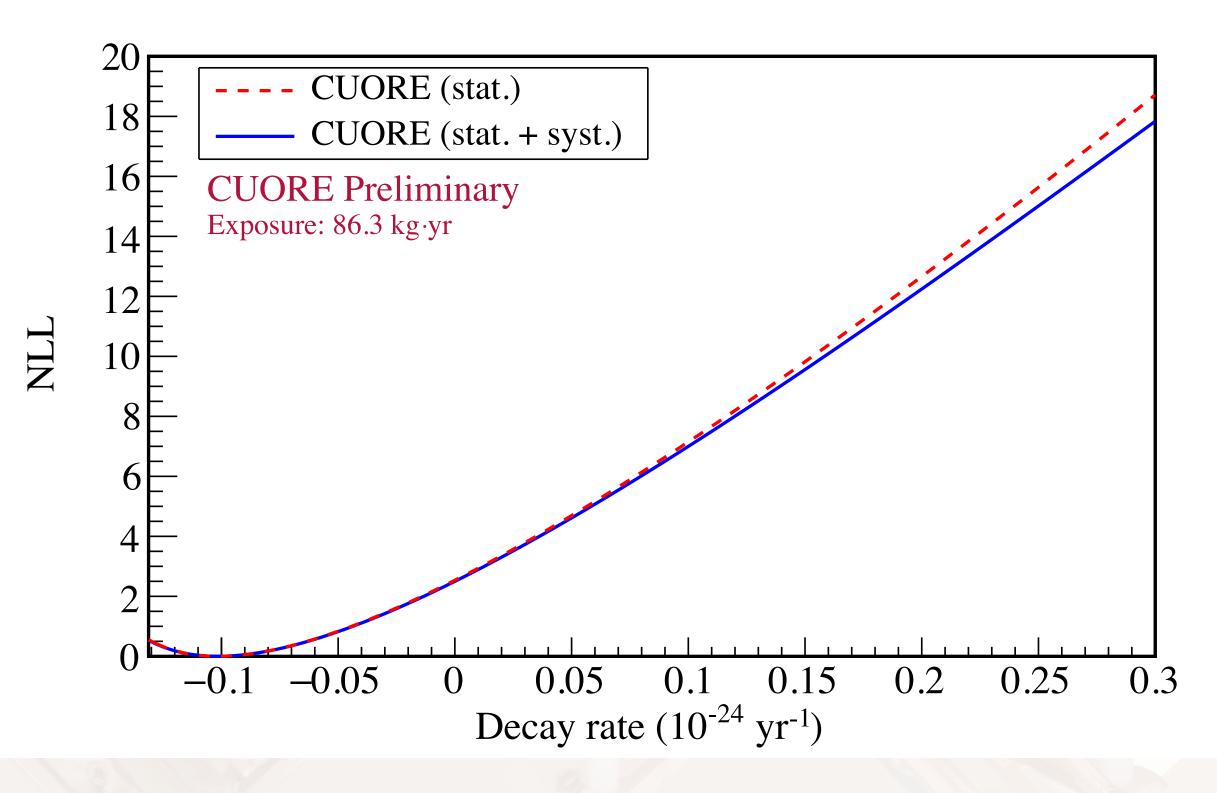
Best fit decay rate:

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No evidence of signal

Limit calculation Profile likelihood integrated on the physical region ($\Gamma^{0v} > 0$)

Decay rate limit (90% CL, including systematics): 0.51 × 10⁻²⁵ / yr Half-life limit (90% CL, including systematics): 1.3 × 10²⁵ yr Median expected sensitivity: 7.0 × 10²⁴ yr





We also perform an independent, fully-bayesian analysis based on the MCMC Gibbs sampler implemented in BAT.

We put a flat positive prior on the signal rate and compute the limit on the signal rate by integrating the marginalised posterior.

The result is compatible with the one obtained by integration of the profile likelihood.



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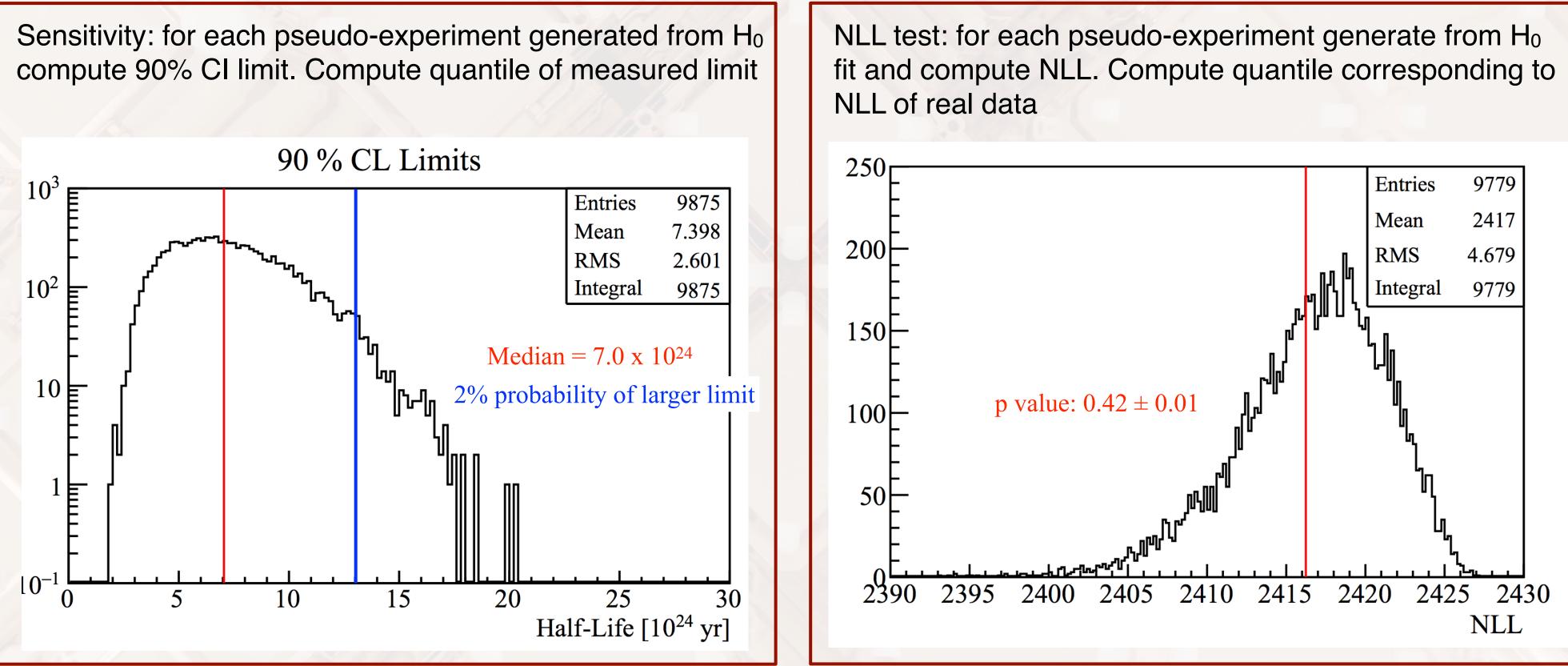
551, 493-503 (2005)":

We have also evaluated frequentist limits according to "W. Rolke et al., Nucl. Instrum. Meth. A - Decay rate limit (90% CL, including systematics): 0.33×10⁻²⁵ / yr - Half-life limit (90% CL, including systematics): 2.1×10²⁵ yr



Statistical considerations

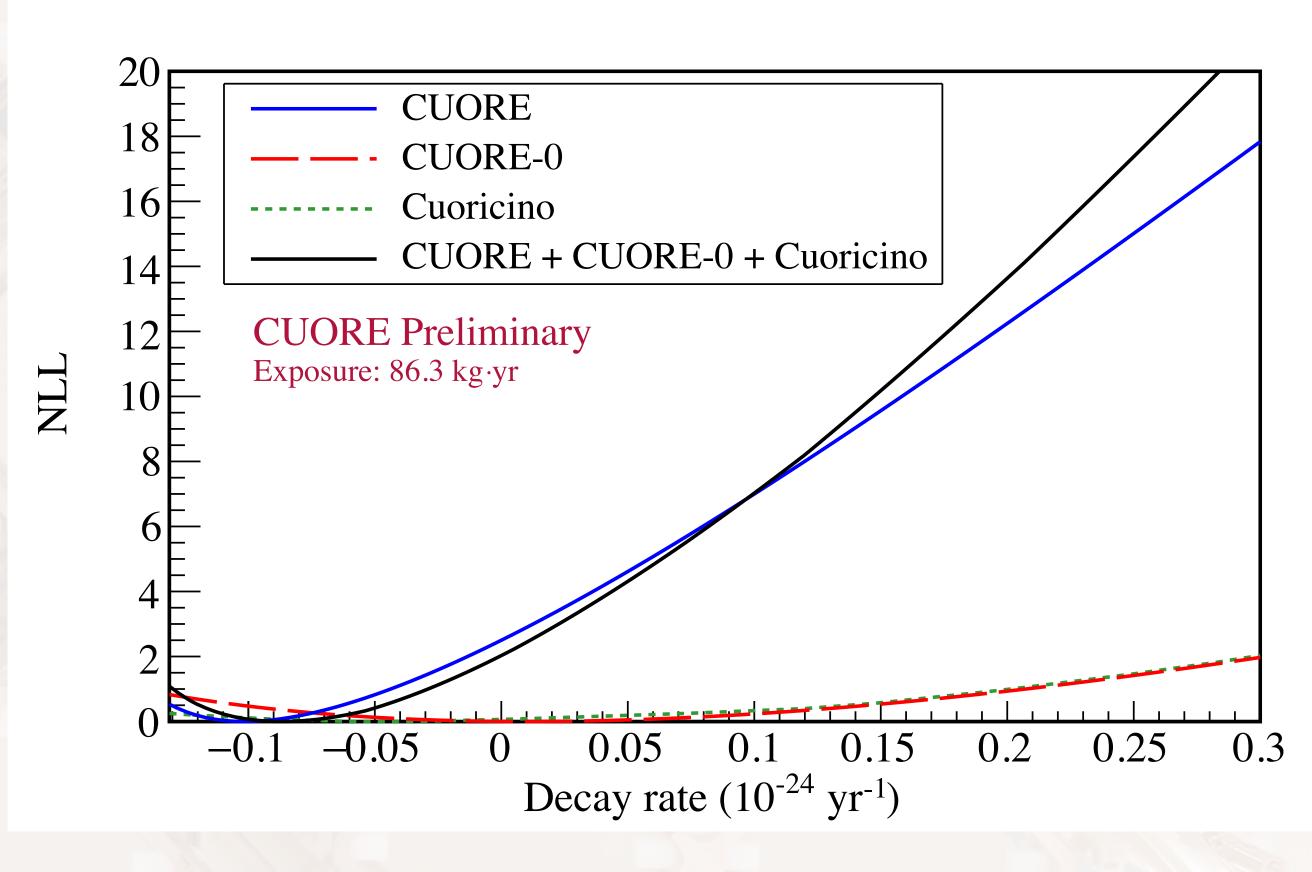
We perform a number of tests to establish the statistical significance of the result and the goodness of fit.





Combination with previous results

- We combined the CUORE result with the existing ¹³⁰Te
 - 19.75 kg·yr of Cuoricino
 - 9.8 kg·yr of CUORE-0
- The combined 90% C.L. limit is $T_{0v} > 1.5 \times 10^{25} \text{ yr}$



Combined "Rolke" limit: 2.2×10²⁵ yr

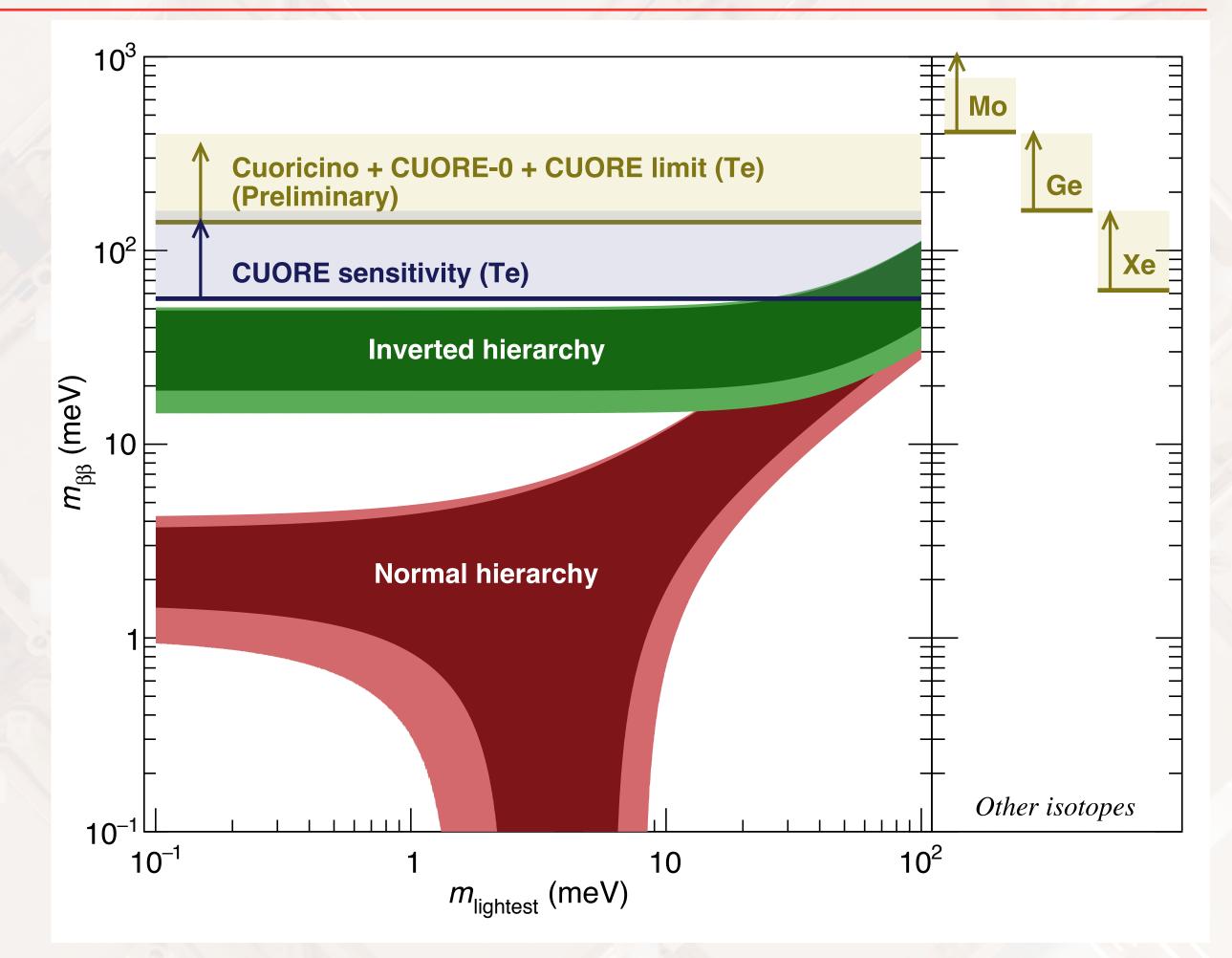
M. Biassoni - CUORE Inauguration Public Seminar



Combination with previous results

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 - 19.75 kg·yr of Cuoricino
 - 9.8 kg·yr of CUORE-0

• The combined 90% C.L. limit is $T_{0v} > 1.5 \times 10^{25} \text{ yr}$ $m_{\beta\beta} < 140-400 \text{ meV}$



NME:

Phys. Rev. C 91, 034304 (2015) Phys. Rev. C 87, 045501 (2013) Phys. Rev. C 91, 024613 (2015) Nucl. Phys. A 818, 139 (2009) Phys. Rev. Lett. 105, 252503 (2010)

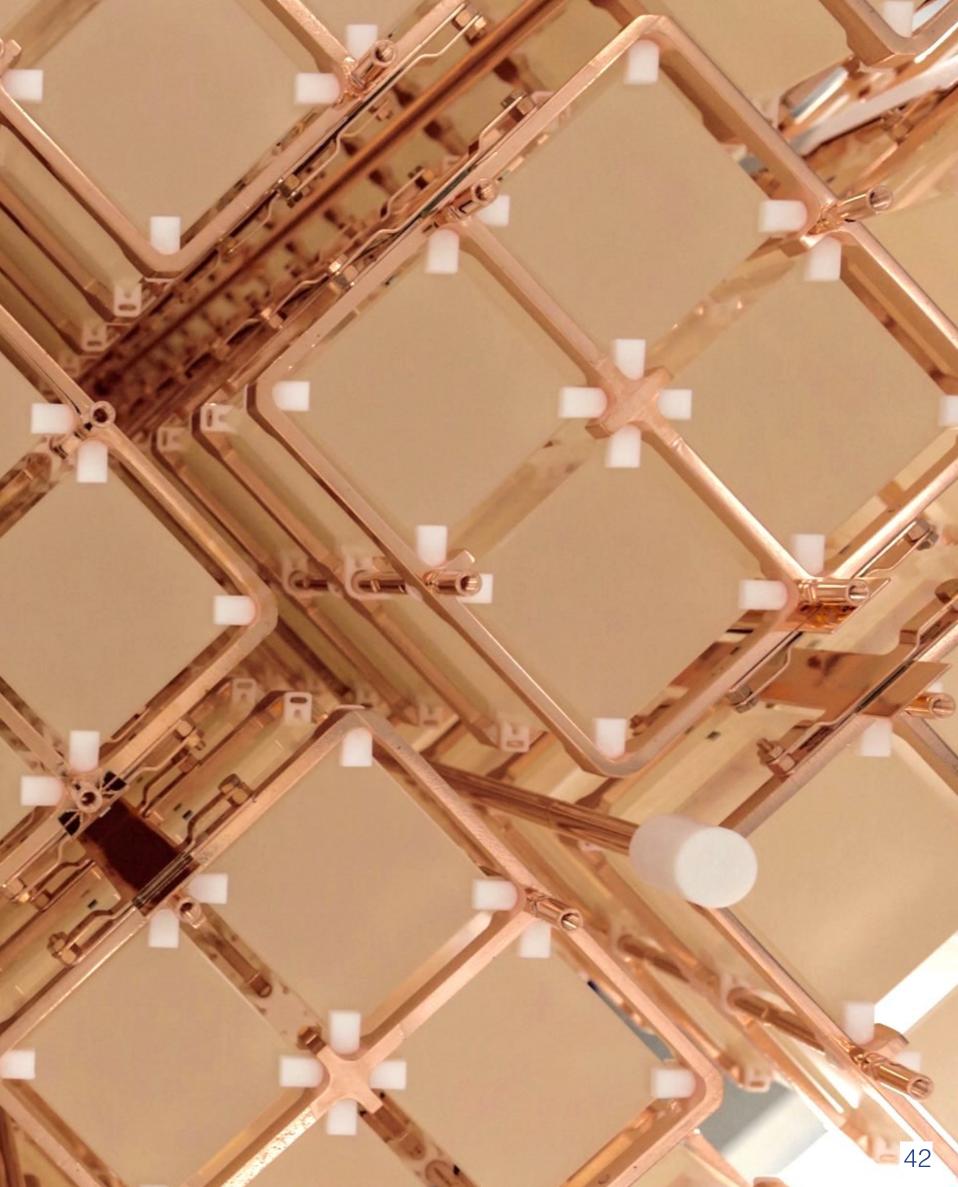
Experiments:

¹³⁰Te: 1.5×10^{25} yr from this analysis ⁷⁶Ge: 5.3×10^{25} yr from Nature 544, 47–52 (2017) ¹³⁶Xe: 1.1×10^{26} yr from Phys. Rev. Lett. 117, 082503 (2016) ¹⁰⁰Mo: 1.1×10^{24} yr from Phys. Rev. D 89, 111101 (2014) CUORE sensitivity: 9.0×10^{25} yr



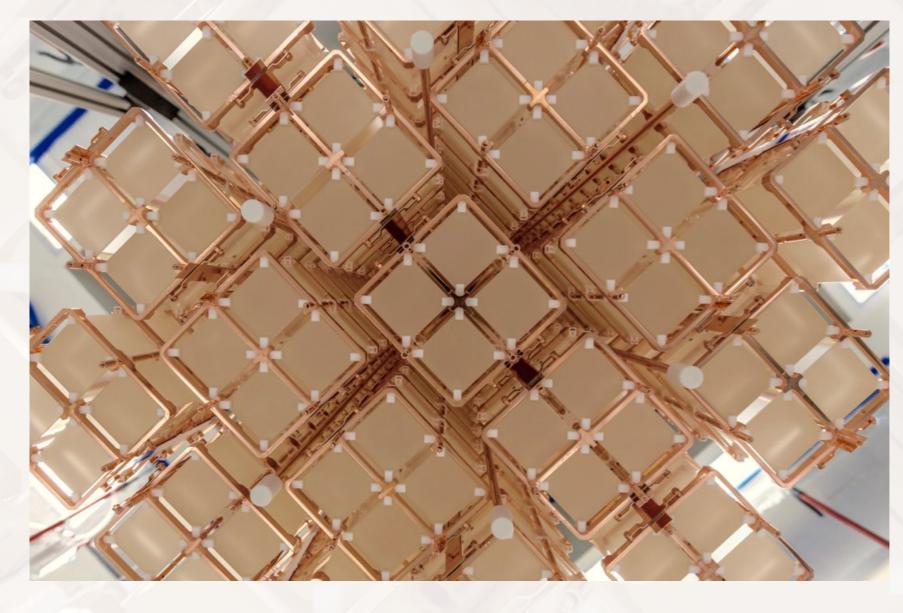
Outline

- TeO₂ and thermal detectors for neutrino-less DBD
- CUORE setup
 - Cryogenics
 - Installation
 - Pre-operation
 - Science data taking
- Analysis procedures
 - Calibration and detector response
 - Event selection
 - Blinding
- Physics results
 - Fit
 - Systematics
 - Combination with other experiments
- Conclusions and outlook



Conclusions

- With the first two datasets CUORE have:
 - accumulated a total exposure of almost 100 kg \cdot y
 - Invaluable operational experience
 - collected important information on detector performance, noise, resolutions, background levels
 - pushed for the first time the limit on neutrino-less double beta decay half life of ¹³⁰Te beyond 10²⁵ years



- The largest and most complex cryogenic experiment is taking physics data
- The first analysis efforts were focused on the neutrino-less double beta decay
- Physics results on more processes are on their way
- With an unprecedented amount of data, CUORE is the best tool to study and model the backgrounds for the next generation experiments
- Paper will appear on arXiv tomorrow, to be submitted to PRL



CUORE cryogenic system is working spectacularly well