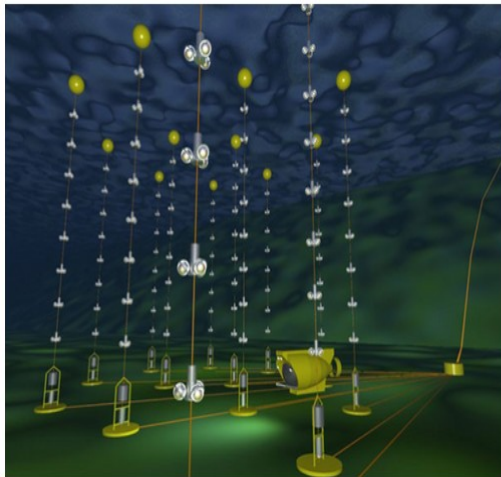
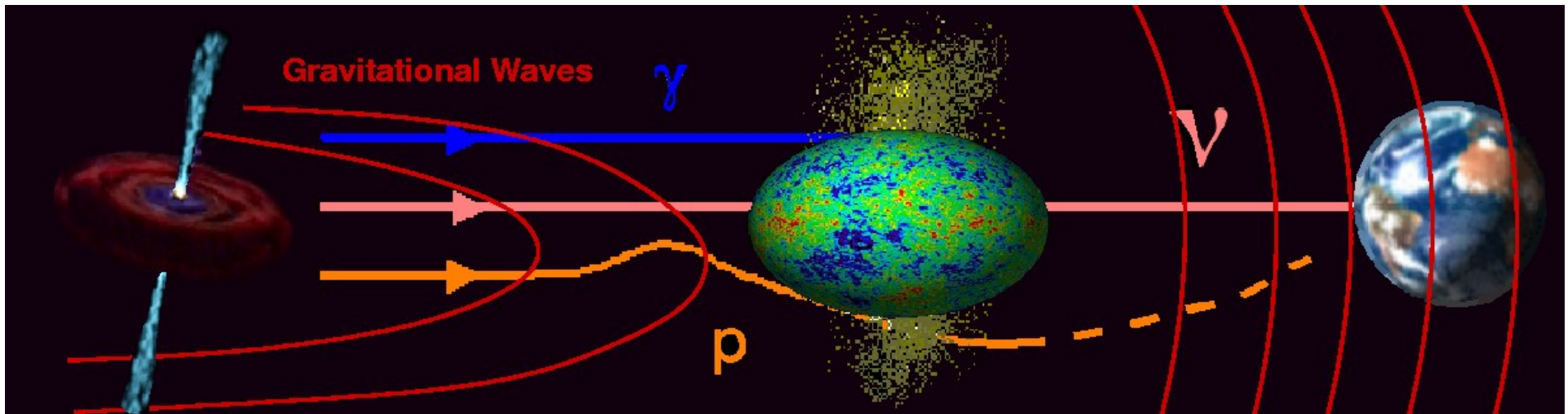


Joint searches of gravitational waves (GW) and high-energy neutrinos (HEN)

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for the LIGO Scientific Collaboration and the Virgo Collaboration



Gravitational waves and High Energy Neutrinos



GW and HEN as cosmic messengers

- *no absorption/diffusion*: travel “cosmological” distances as opposed to photons (dust, gaz, MW or IR background)
- *no deflection* by magnetic fields: trace back (as opposed to charged cosmic rays)
- *weakly interacting*: escape from dense objects

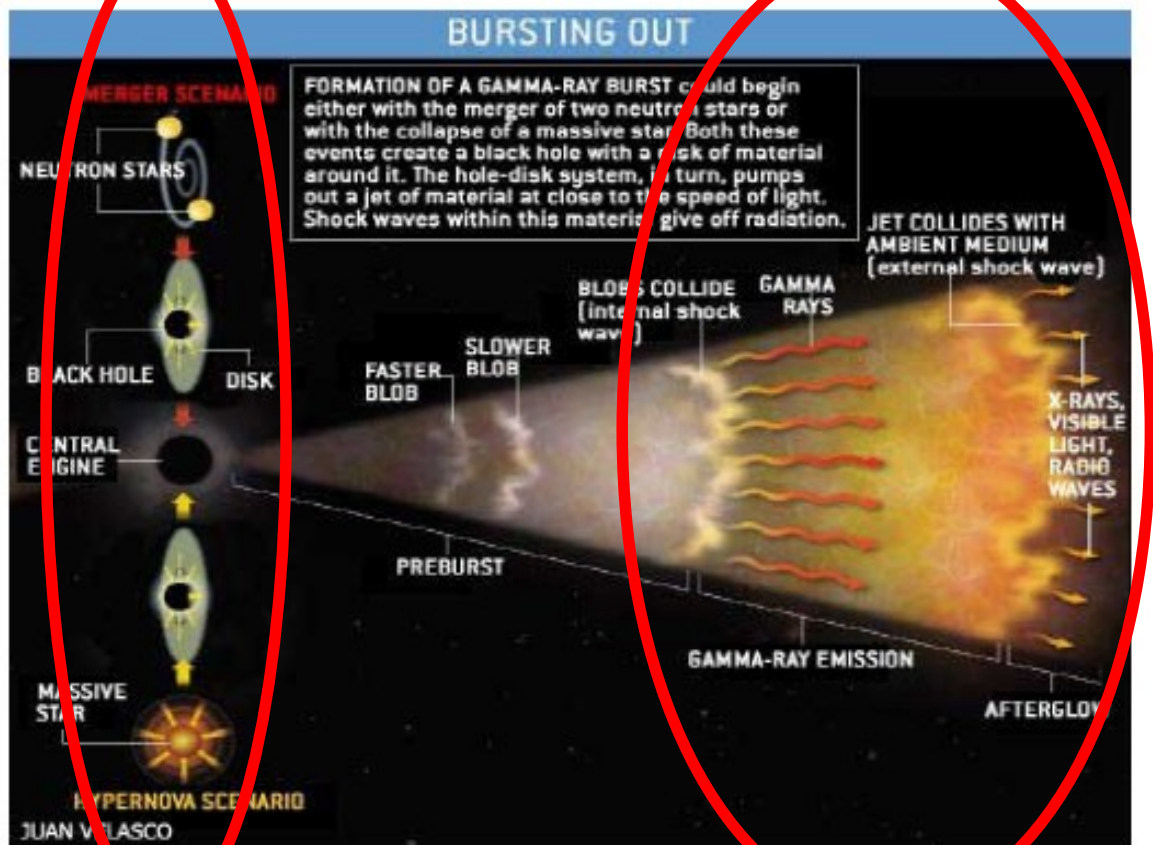
Potential GW+HEN sources

Requirements

- Massive, compact, relativistic (\rightarrow GW)
 - Sudden <1 s (\rightarrow LIGO/Virgo)
 - Baryons (\rightarrow neutrino)
 - Close/frequent enough
- **Galactic**
 - Soft γ repeater
 - Micro quasar
 - **Extra-galactic**
 - *Long GRBs*
 - Short GRBs
 - *Low-lumin. GRBs*

GW+HEN sources (1) : GRBs

short binary mergers



long supernovae hypernovae

GW

high-energy radiation
 $\gamma + \nu$

Fireball model: colliding relativistic shells

accel. electrons produce **gamma rays** by synchrotron

accel. protons interact and produce pions, which decay in **high-energy neutrinos HEN**

caveat: Fermi observations puts the "internal shock model" in troubles. Basic scenario under reconsideration

GW+HEN sources (2) : “failed” GRBs

- **Why GRB jets are relativistic? (compactness pb)**
non-relativistic: optical depth due to absorption $\gamma\gamma \rightarrow e^- e^+ \gg 1$
includ. relativistic effects, optical depth is $\times \Gamma^{-2\alpha}$ (Lorentz fact.)
optically thin if $\Gamma = O(100)$, required to see flash of γ -rays
- **Baryon (heavy) pollution \rightarrow mildly relativistic jet $\Gamma = O(1)$**
optically thick, photon don't escape! No GRB. (“failed”)
more baryons means more neutrinos
- **Events hidden from conventional telescopes**
accessible only to GW+HEN observation
unknown rate, could be large

GW+HEN sources (3): connection between SN and GRB?

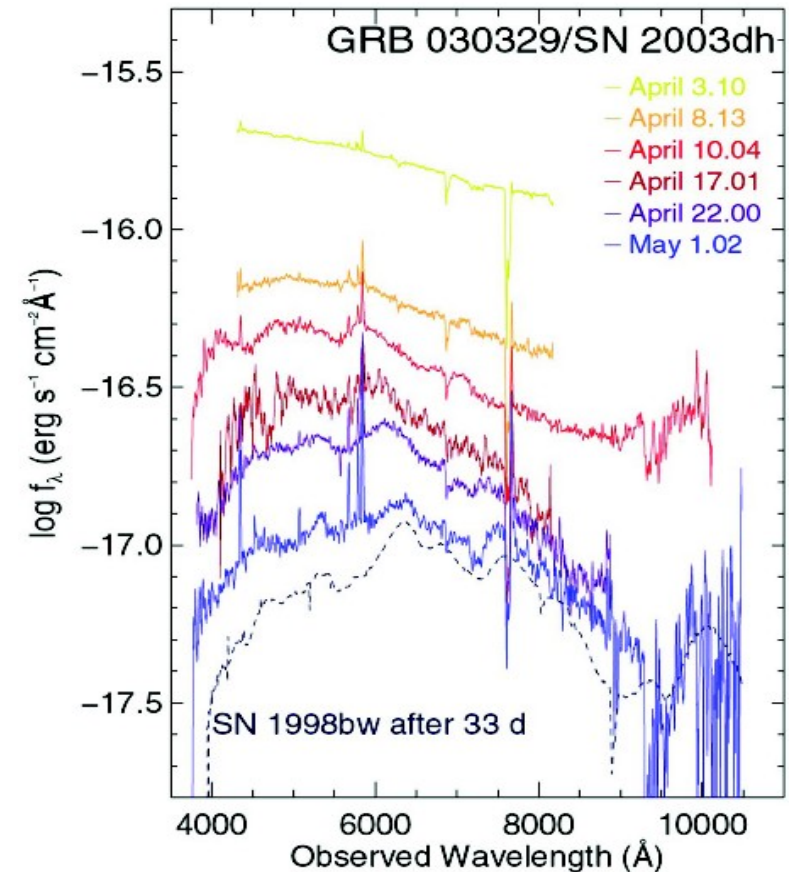
	SN	“Failed” GRB	GRB
Energy	10^{51} erg	10^{51} erg	10^{51} erg
Rate/gal	$\sim 10^{-2}$ yr $^{-1}$	10^{-5} – 10^{-2} yr $^{-1}$	$\sim 10^{-5}$ yr $^{-1}$
Γ	~ 1	~ 3 – 100	~ 100 – 10^3

Barion rich
Nonrelativistic
Frequent

↔
 Similar kinetic energy

Baryon poor
Relativistic jets
Rare

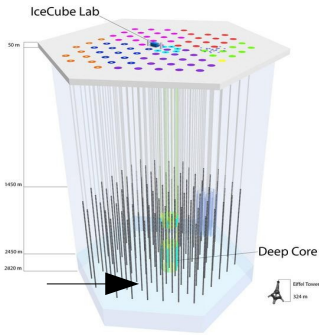
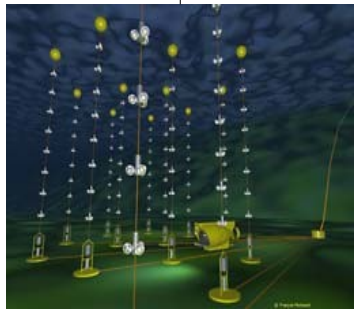
taken from Ando (2009)



Hjorth et al. 2003; Stanek et al. 2003

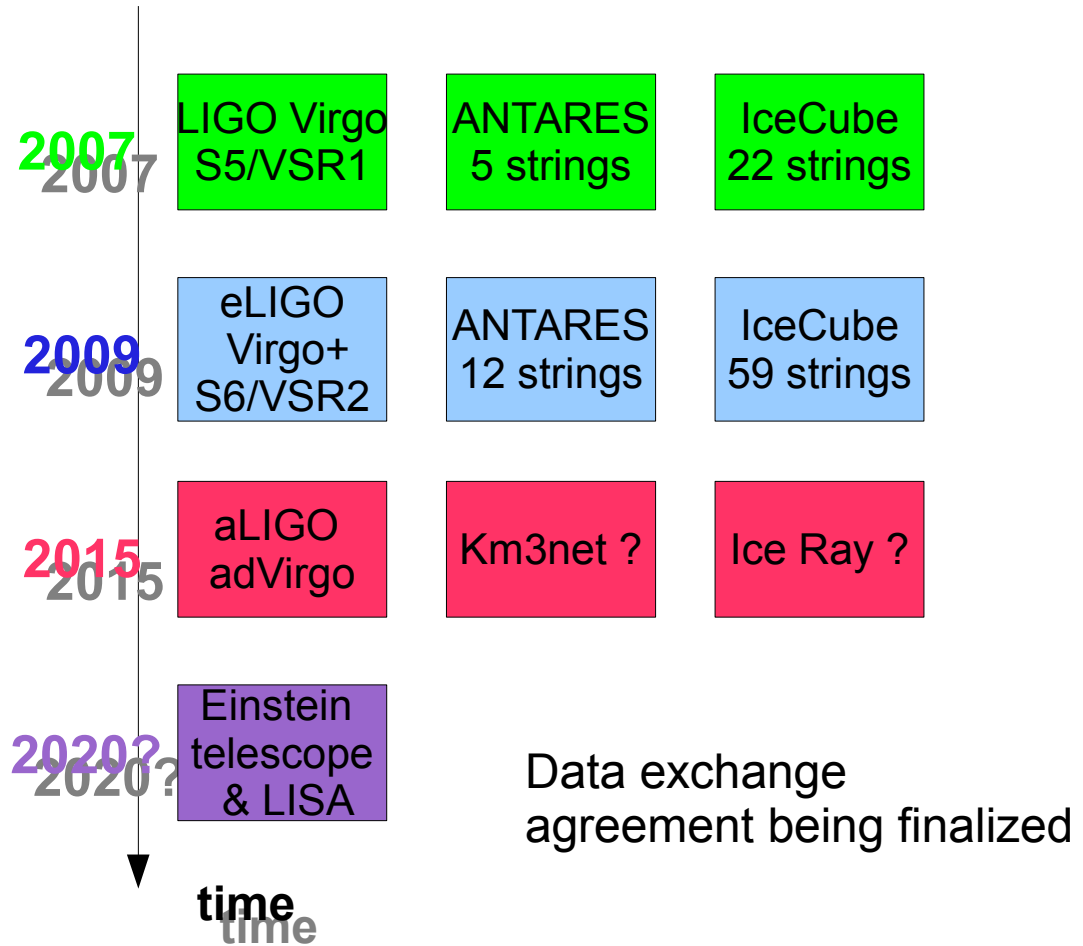
missing link between SN and GRB?

Common data sets with HEN telescopes



ANTARES
(Mediterranean sea)

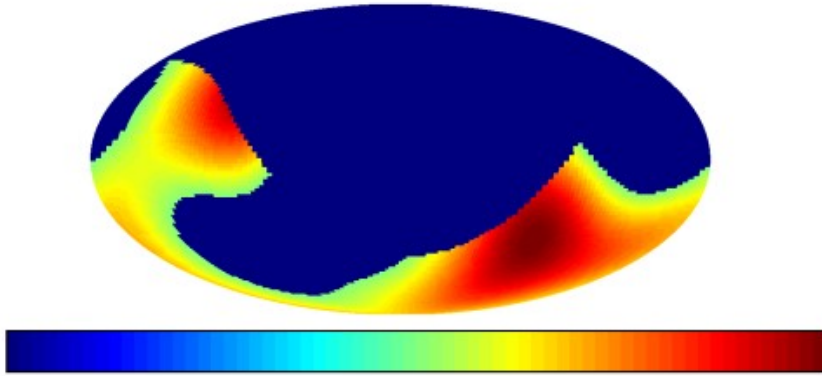
IceCube
(South pole)



Feasibility: basic ingredients

ANTARES & GW det.

common sky GW and ANTARES = 3.98 sr



Sky coverage

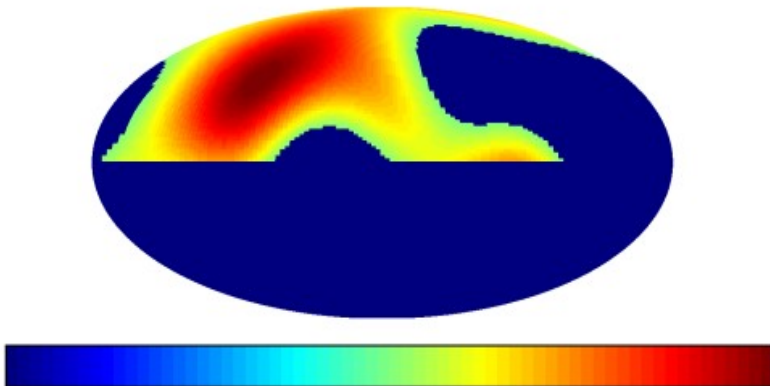
- ANTARES and IceCube sky complementary
- Each have ~30 % common sky with GW det.

Resolution of source localization

- ANTARES has sub-degree error box
- IceCube has ~ degree error box
- GW network has few degree error box

IceCube & GW det.

common sky GW and Ice Cube = 3.94 sr



Exploring possible data analysis strategies

- GW and HEN = same search style
few small signal buried in background noise
- *rationale for a coincidence search* : independent detectors : prob. of accidental coincidence (backgrounds) is **very low** if coinc. observed, high confidence in detection

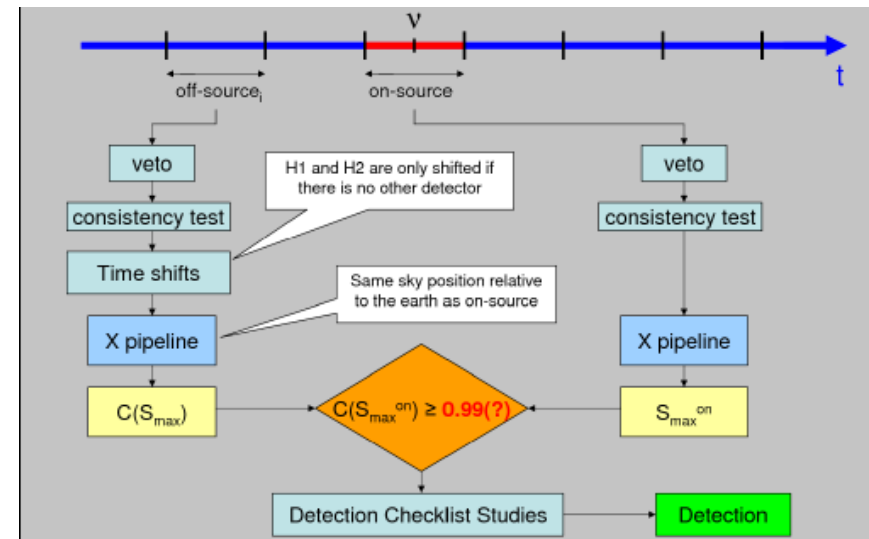
- first studies initiated within LIGO/Virgo and Icecube and independently within ANTARES

detect an excess of time/spatial coincidence

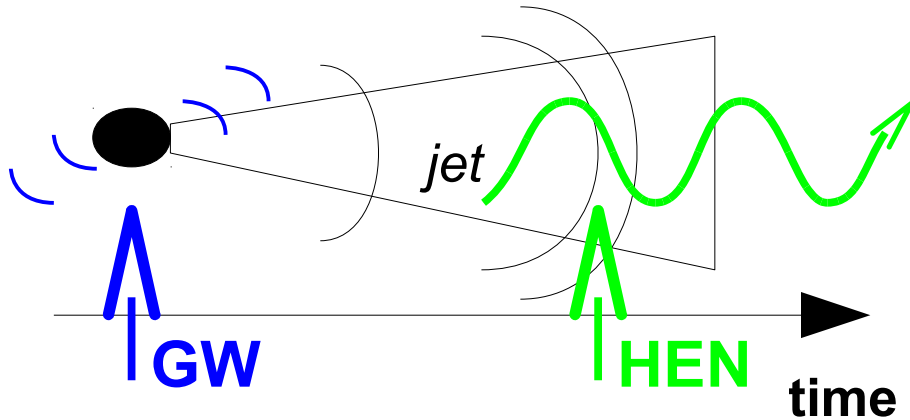
reduce false alarm rate, dig deeper into background



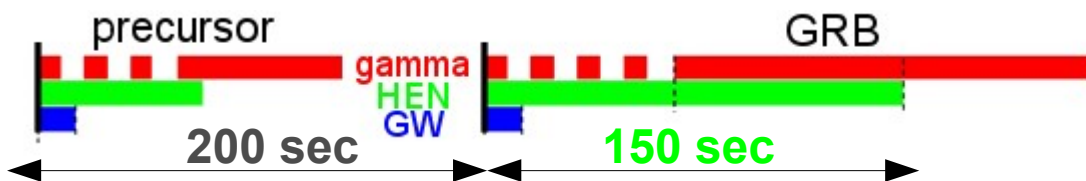
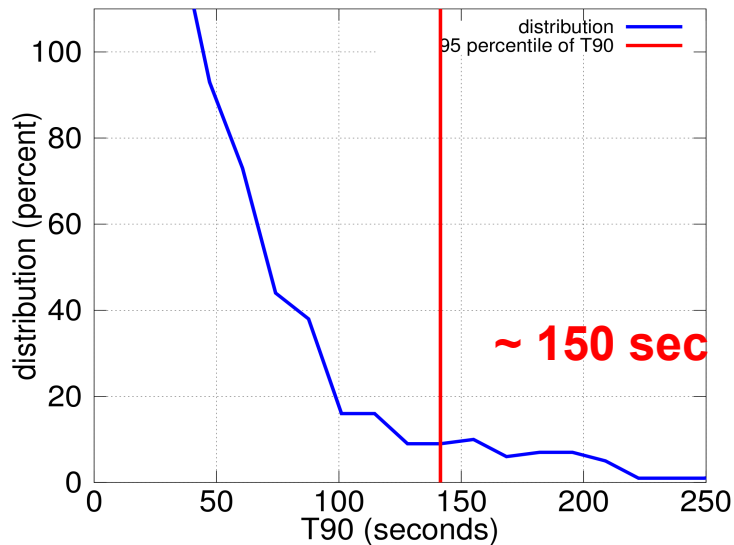
Investigate the use of X pipeline currently used for burst searches in coincidence with GRB



Coincidence time window



GRB durations from 4th BATSE catalog (1234 GRBs)



- Time delay between GW and HEN
- Source/model dependent
- Case study: *long GRBs* (Bartos et al.)

GW emission is prompt

Neutrinos emission simultaneous to γ

GRB duration as indicator for time window: $< \sim 150$ sec (from 4th BATSE catalog)

- GRB may be preceded by precursor
final window is [-350, +200] sec
- Reconsider this window in light of Fermi observations (low statistic for now)

Conclusions

- Working group joining GW and HE neutrinos (IceCube and ANTARES) just formed
- In the process of signing data exchange agreement
- Individuate scenarios for potential joint sources
- Propose procedure for the time/spatial coincidence of GW and HEN events
- Pathfinder effort for advanced detectors