



# The searches for neutrinoless double beta decay and other physics with EXO-200

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#### **Double Beta Decay**



Atomic Number Z

- Double beta decay is possible for certain even-even nuclei, like <sup>130</sup>Te (CUORE), <sup>76</sup>Ge (GERDA, MAJORANA), <sup>116</sup>Cd (COBRA), etc.
- <sup>136</sup>Xe, being the heaviest xenon isotope, is relatively easy to enriched
  - Being a noble gas, xenon can be purified in situ
  - Relatively high Q-value of 2458 keV





#### Two Modes of Double Beta Decay



Two neutrino mode:

- •Standard model process
- Second order
- • $\Delta$ L = 0 (lepton number conserved)





Neutrinoless mode:

- •Hypothetical "Beyond the Standard
- Model" process
- •Can happen if:
  - neutrino has nonzero mass
  - neutrino is its own antiparticle

(Majorana neutrino)

•Total lepton number violating ( $\Delta L = 2$ )

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## Mass Measurement with $\beta\beta0\nu$

• Half-life depends on the effective mass,







#### $\beta\beta0\nu$ signature: a peak in the $\beta\beta$ energy spectrum



Summed electron energy in units of the kinematic endpoint (Q)



#### EXO-200 Liquid Xe Time Projection Chamber



~110 kg active mass Xe enriched to 80% in <sup>136</sup>Xe, ultralow background construction Readout plane is made up of LAAPDs (scintillation) + crossed wire grid (ionization) Achieve electron lifetime in liquid xenon  $\tau_e > 3$  ms Began operating with enriched Xe at the Waste Isolation Pilot Plant (WIPP) in May 2011





### Underground Detector Site

Dewey Lake

stler Form

850 ft

WIPP Surface and Underground Facilities

EXO

- Waste Isolation Pilot Plant in New Mexico, USA
- Overburden of 1585 meters water equivalent
- low salt radioactivity





#### **Active and Passive Shielding**







### Xe Purity over Run 2



- Estimation based upon data from <sup>228</sup>Th source runs
- Purity strongly correlated with circulation pump speed
- At  $\tau_e$  = 3 ms: drift time <110  $\mu s,$  loss of charge: 3.6% at full drift length





### **Calibration System**

- Periodic campaigns with
   <sup>228</sup>Th, <sup>60</sup>Co, and
   <sup>137</sup>Cs, <sup>226</sup>Ra
- Main calibration
   is done with 2615
   keV gamma line
   from <sup>228</sup>Th
   source.







## **Event reconstruction**

V-wire signals Charge clusters **1.** Event position U-wire signals 2. Event multiplicity Linked 3. Energy measurement **APD** signals Scintillation clusters Ionization Ability to 100% reconstruct in 3D individual charge Scintillation cluster down to ~200 keV (limited by induction signal)





Allows for background measurement and reduction



0*νββ*: ~90% SS

Wires are ~1 cm apart

#### $\gamma$ -rays: ~30% SS at $0\nu\beta\beta$ Q-value

Total error in fiducial volume from position reconstruction: **1.73%** 





# **Energy measurement**

#### Anti-correlation of charge and light



- Rotation angle determined weekly using <sup>228</sup>Th source data, defined as angle which gives best rotated resolution
- Energy resolution is dominated by APD noise



copper vessel



Most precise measurement of the  $2\nu\beta\beta$  half-life  $T_{1/2}^{2\nu\beta\beta} = 2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{sys}) \times 10^{21} \text{ yr}$ [PRC **89**, 015502 (2014)]





# Majoron Mode 🗿 🗤 **Search Results**

\*Data fit for each Majoron mode separately

\*Background shown is for fit to n=1 mode

> (Arbitrary Units) 0.6

0.4

0.2

0.0<sup>l</sup>

 $0\nu\beta\beta\chi_0$  $0\nu\beta\beta\chi_0$ 

 $0
uetaeta\chi_0\chi_0$ 

 $0
uetaeta\chi_0\chi_0$ 

 $0\nu\beta\beta\chi_0$ 

0.0

Decay mode

0.5

1.0







# **Decay to Excited State**



Knowledge of the decay to excited state rate check the various Nuclear Matrix Element calculations -> smaller uncertainty on the effective neutrino mass

- Search for the  $2\nu\beta\beta$  decay to the excited state,  $0^+_1$ , of  $^{136}$ Ba
- Decays to excited state has been observed in two isotopes: <sup>100</sup>Mo and <sup>150</sup>Nd
- De-excitation from the 0<sup>+</sup><sub>1</sub> state produce two γs – unique signature
- Main background is 2νββ
   to the ground state





# **Decay to Excited State**

- 1<sup>st</sup> EXO-200 physics analysis to use machine learning
- Being healthily skeptical of our robotic friends

- Algorithm and input variables decided prior to the final fit
- Method is chosen based on the toy MC studies of uncertainty (excited state normalization) after the "unskewing" of source agreements applied to PDFs





#### Phys Rev C 93 035501 (2016)



#### **Decay to Excited State**



The "discriminator" variable is produced from a combination of 6 variables as determined by BDT, boosted decision tree, machine learning algorithm,

Y.-R. Yen



**Drexel** 

# **Decay to Excited State**

 Machine learning improves our sensitivity compares to traditional EXO-200 method



Sensitivity of 1.7 x 10<sup>24</sup> yr



#### Phys Rev C 93 035501 (2016) Decay to Excited State





• 2D Maximum Likelihood fit in Energy + Discriminator



#### Phys Rev C 93 035501 (2016) Decay to Excited State



 $T_{1/2} (0^+ - > 0^+_1) > 6.9 \times 10^{23} \text{ yr}$ 

- Best fit value of 43 events
  - 90% CL of 104 events
  - Consistent with null hypothesis at 1.6σ







#### Search for Lorentz- and CPT- violation in double beta decay within the Standard-Model Extension

- First search for Lorentz-violation in double beta decay within the Standard-Model Extension framework
- Able to search for an oscillation-free, momentum independent neutrino coupling parameter,  $\mathring{a}_{of}^{(3)}$
- A perturbation to the standard 2spectrum defines the magnitude of  $a_{of}^{(3)}$  $\Gamma_0 = g_A^4 m_e^2 |M^{2\nu}|^2 G_0^{2\nu}$

$$\Gamma = \Gamma_0 + \delta \Gamma$$



 $\delta G^{2\nu} \propto \mathring{a}_{\rm of}^{(3)}$ 

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 $\delta\Gamma = g_A^4 m_e^2 |M^{2\nu}|^2 \delta G^{2\nu}$ 





# Search for Lorentz- and CPT- violation in double beta decay within the Standard-Model Extension



#### \* First limit on this parameter set by a direct search

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#### **Other Recent EXO-200 Physics Papers**

- Cosmological background: *arXiv:1512.06835, accepted by JCAP* 
  - Improved understanding of Xe-137 (a muon induced background above our  $0\nu\beta\beta$  Q-value)
- Alphaion fraction and mobility: *PRC 92 4 (2015)* Interesting info for potential future Ba tagging work
- Radioactivity-induced background: PRC 92 1 (2015)
  - Understanding of our backgrounds are detailed



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# EXO-200 Status to Jan. 2016

- WIPP incidents (NOT OUR FAULT!):
  - Feb. 5 2014 Fire in WIPP underground
  - Feb. 14 2014 Airborne radiological event
- EXO-200 timeline:
  - In late Feb. 2014, with remote system access, Xe was successfully recovered (as designed), followed by controlled warm up of TPC/Cryostat.

GEOFF BRUMFIEL

MARCH 26, 2015 6:40 PM E

the two-way BREAKING NEWS FROM NPP

Caused By Wrong Cat Litter

Official Report: Nuclear Waste Accide

- In Sept. 2014, lost underground power but regained access.
- Power restored in Feb. 2015
- Sample salt near the experiment show virtually zero contamination from the radiological event
- Ongoing cleanup and equipment repair/replacement
- Cooling and filling LXe TPC in the winter 2015/2016

EXO-200 is nearly 4000 feet from the radiation event









#### EXO-200 is Back!



#### A Bi-Po alpha event right after turning the detector back on





- Hardware upgrade to improve energy resolution at Q-value to 1% via improvement in minimizing APD noise
- Many of the ROI background counts come from radon daughters external to the detector. Deradonator should improve our background in the ROI.



http://nest.physics.ucdavis.edu/







# Sensitivity outlook







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### nEXO ("next EXO")





#### Preliminary design of nEXO



Preliminary design of the nEXO detector in SNOLab's cryopit



#### nEXO R&D

Charge readout tiles



#### **Barium Tagging (O background possible if we can** ID the daughter barium ion)



RF Funnel for Gas (Int. J. of Mass. Spec., V379, 110, 2015)







#### nEXO vs. EXO-200

Parameter	nEXO	EXO-200	Effective Majorana mass vs. $M_{total}$ For the mean values of oscillation parameters (dashed) and for the 3 $\sigma$ errors (full)
Fiducial mass (kg)	4780	98.5	0.1 Final EXO-200 0.1 Final EXO-200 nEXO 5yr 0.01 nEXO 10yr w/ Ba tag 0.00 0.1 M <sub>total</sub> (eV)
Enrichment (%)	80-90	80	
Data taking time (yr)	5	5	
Energy resolution @ Q <sub>ββ</sub> (keV)	58	88 (58)	
Depth (m.w.e)	6010	1500	
Background within FWHM of endpoint (events/yr/mol <sub>136</sub> )	6.1x10 <sup>-4</sup>	0.022 (0.0073)	
Background within FWHM of endpoint inner 3000kg (events/yr/mol <sub>136</sub> )	1.6x10 <sup>-4</sup>		



#### **NOT the Summary**





- EXO-200 has restarted and hope to do great physics in running for 3 more years.
- Effort toward nEXO R&D is just one of the byproduct.

- Problem at WIPP (yes, organic kitty litter - not EXO's fault!)
   caused a two year biatus, but we
- caused a two year hiatus, but we have overcame that.
- (Yes, the organic kitty litter thing is rather memorable if not confusing.)







- 100 kg·yr (736 mol·yr) <sup>136</sup>Xe exposure of EXO-200 data have resulted in
  - Precision  $2\nu\beta\beta$  measurement (*PRC* **89**, 015502 (2014))
  - 0νββ limit (*Nature* **510**, 229 (2014))
  - Majoron mode limits (*PRD* **90,** 092004 (2014))
  - $-2\nu\beta\beta$  to the excited state limit (*PRC* **93**, 035501 (2016))
- After 2 yr hiatus, EXO-200, one of the most sensitive  $0\nu\beta\beta$  experiment currently, has restarted to take more data (Jan. 2016)
- Upgrades (electronics and deradonator) will help with nEXO (5 ton next-gen LXe experiment) R&D currently in progress





# THANK YOU FOR YOUR ATTENTION!

#### EXO-200 $0\nu\beta\beta$ Half-life Sensitivity



 $T_{1/2}^{0\nu\beta\beta} > 1.1 \cdot 10^{25}$ yr (90%CL)  $< m_{v} > < 190 - 450 \text{ meV}$ Median  $T_{1/2}^{0\nu\beta\beta}$  sensitivity:

1.9.10<sup>25</sup> yr

J.B.Albert et al. (EXO-200), Nature (6 June, 2014) A. Gando et al. (KamLAND-ZEN), PRL 110 (2013) 062502 M. Agostini et al. (GERDA), PRL 111 (2013) 122503

With upgraded detector and 2 yrs of live-time, EXO-200  $T_{1/2}^{0\nu\beta\beta}$  median sensitivity will increase by a factor of 3.

One of the most sensitive  $0\nu\beta\beta$  experiments in the next 3 - 5 years.